

Star-products for Lie-algebraic noncommutative Minkowski space-times

MARIS Valentine
Cost Action CaLISTA General Meeting 2025

ENS de Lyon (LPENSL)/IJCLab
(based on 2503.07176 with F.Pozar and J-C.Wallet)

20/09/2025

Motivations

- ➔ Quantization of space time

Motivations

- ⇒ Quantization of space time
- ⇒ Gelfand Naimark theorem:
Manifold equivalent to commutative C^* algebra

Motivations

- Quantization of space time
- **Geland Naimark theorem:**
Manifold equivalent to commutative C^* algebra
- Quantum version: Noncommutative C^* algebra

Definitions

⇒ \star Algebra :

An algebra $(\mathcal{A}, +, \star)$ equipped with an involutive antihomomorphism $\dagger : \mathcal{A} \rightarrow \mathcal{A}$

⇒ Banach \star algebra:

Banach space $(\mathcal{A}, \|\cdot\|)$ such that

- ✓ It is an \star -algebra for a given product \star and involution \dagger
- ✓ Its norm satisfies $\|fg\| \leq \|f\| \|g\|, \quad \forall f, g \in \mathcal{B}$
- ✓ \dagger is an isometry for the norm: $\|f^\dagger\| = \|f\|$

A Banach \star -algebra is called a $C\star$ -algebra if

$$\|ff^\dagger\| = \|f\|^2, \quad \forall f \in \mathcal{A}$$

Context

- ⇒ Deformed Minkowski space-time and Poincaré symmetry Hopf algebra
- ⇒ Well known examples: κ/ρ -Minkowski [Lukierski, Ruegg, Nowicki, Tolstoy,1991],[Majid, Ruegg,1994][Dimitrijevic, Konjik,Samsarov,2018]

$$[x_0, x_j] = \frac{i}{\kappa} x_j, \quad [x_0, x_1] = \rho x_2, \quad [x_0, x_2] = -\rho x_1$$

- ⇒ Classification of deformed Poisson structures of the Poincaré group [Zakrzewski, 2001]
- ⇒ Deformed Minkowski space-time [Lukierski,Woronowicz, 1995][Mercati, 2001]
- ⇒ Lie algebraic deformations:

$$[x^\mu, x^\nu] = i f_\rho^{\mu\nu} x^\rho \quad (1)$$

- ⇒ Convolution algebra and Weyl quantization map [von Neumann, 1929]

Context

- ⇒ Deformed Minkowski space-time and Poincaré symmetry Hopf algebra
- ⇒ Well known examples: κ/ρ -Minkowski [Lukierski, Ruegg, Nowicki, Tolstoy,1991],[Majid, Ruegg,1994][Dimitrijevic, Konjik,Samsarov,2018]

$$[x_0, x_j] = \frac{i}{\kappa} x_j, \quad [x_0, x_1] = \rho x_2, \quad [x_0, x_2] = -\rho x_1$$

- ⇒ Classification of deformed Poisson structures of the Poincaré group [Zakrzewski,1997]
- ⇒ Deformed Minkowski space-time [Lukierski,Woronowicz,2005][Mercati,2024],
- ⇒ Lie algebraic deformations:

$$[x^a, x^b] = if_{\rho}^{ab} x^c \quad (1)$$

- ⇒ Convolution algebra and Weyl quantization map von Neumann

Context

- ⇒ Deformed Minkowski space-time and Poincaré symmetry Hopf algebra
- ⇒ Well known examples: κ/ρ -Minkowski [Lukierski, Ruegg, Nowicki, Tolstoy,1991],[Majid, Ruegg,1994][Dimitrijevic, Konjik,Samsarov,2018]

$$[x_0, x_j] = \frac{i}{\kappa} x_j, \quad [x_0, x_1] = \rho x_2, \quad [x_0, x_2] = -\rho x_1$$

- ⇒ Classification of deformed Poisson structures of the Poincaré group [Zakrzewski,1997]
- ⇒ Deformed Minkowski space-time [Lukierski,Woronowicz,2005][Mercati,2024].
- ⇒ Lie algebraic deformations:

$$[x^\mu, x^\nu] = i f_\rho^{\mu\nu} x^\rho \tag{1}$$

- ⇒ Convolution algebra and Weyl quantization map [von Neumann,1931]

Construction of $*$ -algebra

- ⇒ Start from Lie algebra \mathfrak{g} with locally compact Le group \mathcal{G} .
- ⇒ Derive its (right and left) Haar measures and its convolution algebra $\mathbb{C}(\mathcal{G}) = (L^1(\mathcal{G}), \circ)$
- ⇒ Study and characterize representations of the convolution algebra $\pi : \mathbb{C}(\mathcal{G}) \rightarrow \mathcal{B}(\mathcal{H})$
- ⇒ Define quantization map, $Q := \pi \circ \mathcal{F}$, where \mathcal{F} is the Fourier transform of the convolution algebra.
- ⇒ Define star product \star as $Q(f \star g) = Q(f)Q(g)$ and involution $Q(f)^\star = Q(f^\star)$, where \star is involution of $\pi(\mathcal{B}(\mathcal{H}))$.

Construction of $*$ -algebra

- ⇒ Start from Lie algebra \mathfrak{g} with locally compact Lie group \mathcal{G} .
- ⇒ Derive its (right and left) Haar measures and its convolution algebra $\mathbb{C}(\mathcal{G}) = (L^1(\mathcal{G}), \circ)$
- ⇒ Study and characterize representations of the convolution algebra $\pi : \mathbb{C}(\mathcal{G}) \rightarrow B(\mathcal{H})$
- ⇒ Define quantization map, $Q := \pi \circ \mathcal{F}$, where \mathcal{F} is the Fourier transform of the convolution algebra.
- ⇒ Define star product $*$ as $Q(f * g) = Q(f)Q(g)$ and involution $Q(f)^* = Q(f^*)$, where $*$ is involution of $\pi(B(\mathcal{H}))$.

Construction of $*$ -algebra

- ⇒ Start from Lie algebra \mathfrak{g} with locally compact Lie group \mathcal{G} .
- ⇒ Derive its (right and left) Haar measures and its convolution algebra $\mathbb{C}(\mathcal{G}) = (L^1(\mathcal{G}), \circ)$
- ⇒ Study and characterize representations of the convolution algebra $\pi : \mathbb{C}(\mathcal{G}) \rightarrow \mathcal{B}(\mathcal{H})$
- ⇒ Define quantization map, $Q := \pi \circ \mathcal{F}$, where \mathcal{F} is the Fourier transform of the convolution algebra.
- ⇒ Define star product \star as $Q(f \star g) = Q(f)Q(g)$ and involution $Q(f)^\star = Q(f^\dagger)$, where \dagger is involution of $\pi(\mathcal{B}(\mathcal{H}))$

Construction of $*$ -algebra

- ⇒ Start from Lie algebra \mathfrak{g} with locally compact Lie group \mathcal{G} .
- ⇒ Derive its (right and left) Haar measures and its convolution algebra $\mathbb{C}(\mathcal{G}) = (L^1(\mathcal{G}), \circ)$
- ⇒ Study and characterize representations of the convolution algebra $\pi : \mathbb{C}(\mathcal{G}) \rightarrow \mathcal{B}(\mathcal{H})$
- ⇒ Define quantization map, $Q := \pi \circ \mathcal{F}$, where \mathcal{F} is the Fourier transform of the convolution algebra.
- ⇒ Define star product \star as $Q(f \star g) = Q(f)Q(g)$ and involution $Q(f)^\star = Q(f^\star)$, where \star is involution of $\pi(\mathcal{B}(\mathcal{H}))$

Construction of \ast -algebra

- ⇒ Start from Lie algebra \mathfrak{g} with locally compact Lie group \mathcal{G} .
- ⇒ Derive its (right and left) Haar measures and its convolution algebra $\mathbb{C}(\mathcal{G}) = (L^1(\mathcal{G}), \circ)$
- ⇒ Study and characterize representations of the convolution algebra $\pi : \mathbb{C}(\mathcal{G}) \rightarrow \mathcal{B}(\mathcal{H})$
- ⇒ Define quantization map, $Q := \pi \circ \mathcal{F}$, where \mathcal{F} is the Fourier transform of the convolution algebra.
- ⇒ Define star product \star as $Q(f \star g) = Q(f)Q(g)$ and involution $Q(f)^\ast = Q(f^\dagger)$, where \ast is involution of $\pi(\mathcal{B}(\mathcal{H}))$

Convolution algebra and Weyl quantization map (I)

↪ General structure of Lie groups :

$$\mathcal{G} = H \ltimes \mathbb{R}^3, \quad (2)$$

where $H \subset GL(n, \mathbb{R})$ one parameter abelian subgroup.

$$\Leftrightarrow (a(p^M), \vec{p}) \in \mathcal{G} = H \ltimes \mathbb{R}^3 \implies (p^M, \vec{p}) \in \mathbb{R} \times \mathbb{R}^3$$

The group laws of \mathcal{G} are given by:

$$(a_1, \vec{p}_1)(a_2, \vec{p}_2) = (a_1 a_2, \vec{p}_1 + a_1 \vec{p}_2), \quad (3)$$

$$(a, \vec{p})^{-1} = (a^{-1}, -a^{-1} \vec{p}), \quad \mathbb{I}_{\mathcal{G}} = (\mathbb{I}_H, 0) \quad (4)$$

Convolution algebra and Weyl quantization map (I)

↪ General structure of Lie groups :

$$\mathcal{G} = H \ltimes \mathbb{R}^3, \quad (2)$$

where $H \subset GL(n, \mathbb{R})$ one parameter abelian subgroup.

$$\Leftrightarrow (a(p^M), \vec{p}) \in \mathcal{G} = H \ltimes \mathbb{R}^3 \implies (p^M, \vec{p}) \in \mathbb{R} \times \mathbb{R}^3$$

The group laws of \mathcal{G} are given by:

$$(a_1, \vec{p}_1)(a_2, \vec{p}_2) = (a_1 a_2, \vec{p}_1 + a_1 \vec{p}_2), \quad (3)$$

$$(a, \vec{p})^{-1} = (a^{-1}, -a^{-1} \vec{p}), \quad \mathbb{I}_{\mathcal{G}} = (\mathbb{I}_H, 0) \quad (4)$$

Convolution algebra and Weyl quantization map (II)

Right and left invariante Haar measures and modular function for $(a, \vec{p}) \in \mathcal{G}$:

$$d\nu_{\mathcal{G}}((a, \vec{p})) = \Delta_{\mathcal{G}}((a, \vec{p})^{-1}) d\mu_{\mathcal{G}}((a, \vec{p})) \quad \left\{ \begin{array}{l} d\mu_{\mathcal{G}}((a, \vec{p})) = d^3\vec{p} |\det(a)|^{-1} d\mu_H(a) \\ \Delta_{\mathcal{G}}((a, \vec{p})) = |\det(a)|^{-1} \underbrace{\Delta_H(a)}_{=1, H \text{ abelian}} \end{array} \right.$$

where $\Delta_{\mathcal{G}}$ is a group homomorphism from \mathcal{G} in \mathbb{R}^+

Convolution algebra and Weyl quantization map (II)

Right and left Haar measures and modular function for $(a, \vec{p}) \in \mathcal{G}$:

$$d\nu_{\mathcal{G}}((a, \vec{p})) = \Delta_{\mathcal{G}}((a, \vec{p})^{-1})d\mu_{\mathcal{G}}((a, \vec{p})) \quad \left\{ \begin{array}{l} d\mu_{\mathcal{G}}((a, \vec{p})) = d^3\vec{p} |\det(a)|^{-1} d\mu_H(a) \\ \Delta_{\mathcal{G}}((a, \vec{p})) = |\det(a)|^{-1} \underbrace{\Delta_H(a)}_{=1, H \text{ abelian}} \end{array} \right.$$

where $\Delta_{\mathcal{G}}$ is a group homomorphism from \mathcal{G} in \mathbb{R}^+

⇒ Expression of the convolution product and involution on $(L^1(\mathcal{G}, \circ, *))$:

$$(F \circ G)(s) = \int_{\mathcal{G}} d\nu_{\mathcal{G}}(t) F(st^{-1})G(t), \quad t, s \in \mathcal{G}$$
$$F^*(s) = \overline{F}(s^{-1})\Delta_{\mathcal{G}}(s)$$

This convolution algebra can be completed to a C^* algebra, with C^* norm $\|\cdot\|_{C^*}$

Convolution algebra and Weyl quantization map (II)

Right and left Haar measures and modular function for $(a, \vec{p}) \in \mathcal{G}$:

$$d\nu_{\mathcal{G}}((a, \vec{p})) = \Delta_{\mathcal{G}}((a, \vec{p})^{-1})d\mu_{\mathcal{G}}((a, \vec{p})) \quad \left\{ \begin{array}{l} d\mu_{\mathcal{G}}((a, \vec{p})) = d^3\vec{p} |\det(a)|^{-1} d\mu_H(a) \\ \Delta_{\mathcal{G}}((a, \vec{p})) = |\det(a)|^{-1} \underbrace{\Delta_H(a)}_{=1, H \text{ abelian}} \end{array} \right.$$

where $\Delta_{\mathcal{G}}$ is a group homomorphism from \mathcal{G} in \mathbb{R}^+

⇒ Derivation of \star -product and involution :

$$(F \circ G)(s) = \int_{\mathcal{G}} d\nu_{\mathcal{G}}(t) F(st^{-1})G(t) \xrightleftharpoons[\text{Inverse}]{\text{Fourier}} f \star g = \mathcal{F}^{-1}(\mathcal{F}f \circ \mathcal{F}g)$$

$$F^*(s) = \overline{F}(s^{-1})\Delta_{\mathcal{G}}(s) \xrightleftharpoons[\text{Inverse}]{\text{Fourier}} f^\dagger = \mathcal{F}^{-1}(\mathcal{F}f^*)$$

Explicit formulation of star product and involution

$$\begin{aligned}(f \star g)(x^M, \vec{x}) &= \int \frac{dp^M dy^M}{(2\pi)} e^{-ip^M y^M} f(x^M + y^M, \vec{x}) g(x^M, a(p^M)x) \\ f^\dagger(x^M, \vec{x}) &= \int \frac{dp^M dy^M}{(2\pi)} e^{-ip^M y^M} \overline{f(x^M + y^M, a(p^M)\vec{x})}\end{aligned}\tag{5}$$

where $x = (x^M, \vec{x}) \in \mathbb{R} \times \mathbb{R}^3$

→ Construction of 10 such \star product \star -algebras \mathcal{M}_λ

Explicit formulation of star product and involution

$$\begin{aligned}(f \star g)(x^M, \vec{x}) &= \int \frac{dp^M dy^M}{(2\pi)} e^{-ip^M y^M} f(x^M + y^M, \vec{x}) g(x^M, a(p^M)x) \\ f^\dagger(x^M, \vec{x}) &= \int \frac{dp^M dy^M}{(2\pi)} e^{-ip^M y^M} \overline{f(x^M + y^M, a(p^M)\vec{x})}\end{aligned}\tag{5}$$

where $x = (x^M, \vec{x}) \in \mathbb{R} \times \mathbb{R}^3$

⇒ Construction of 10 such \star product \star -algebras \mathcal{M}_λ

Commutators

Commutators of lie algebra of coordinates \mathfrak{g} reads:

$$[x^M, x^\mu]_\star = x^M \star x^\mu - x^\mu \star x^M = -i \left[\partial_{p^M} a(p^M) \Big|_{p^M=0} \right]^\mu{}_\sigma x^\sigma, \quad (6)$$

x^M can be a:

- ⇒ timelike coordinate
- ⇒ spacelike coordinate
- ⇒ lightlike coordinate

Rescaling $p^M \rightarrow \lambda p^M$ and thus (6) become

$$[x^M, x^\mu] = -i\lambda \left[\partial_{p^M} a(\lambda p^M) \Big|_{p^M=0} \right]^\mu{}_\sigma x^\sigma$$

Commutators

Commutators of lie algebra of coordinates \mathfrak{g} reads:

$$[x^M, x^\mu]_\star = x^M \star x^\mu - x^\mu \star x^M = -i [\partial_{p^M} a(p^M)|_{p^M=0}]^\mu \sigma x^\sigma, \quad (6)$$

x^M can be a:

- ⇒ timelike coordinate
- ⇒ spacelike coordinate
- ⇒ lightlike coordinate

Rescaling $p^M \rightarrow \lambda p^M$ and thus (6) become

$$[x^M, x^\mu] = -i\lambda [\partial_{p^M} a(\lambda p^M)|_{p^M=0}]^\mu \sigma x^\sigma$$

Nonunimodular group and twisted trace

$$\mathcal{G} \text{ nonunimodular} \implies \left\{ \begin{array}{l} \Delta_{\mathcal{G}}(p^M) = e^{-n\lambda p^M} \quad \Delta_{\mathcal{G}}, \text{ modular function} \\ x^M = x^0 \quad x^M \text{ is timelike} \\ \text{tr}(f \star g) = \text{tr}((\sigma \triangleright g) \star f), \quad \text{tr} := \int dx^\mu \quad \text{twisted trace} \end{array} \right.$$

where the twist $\sigma \in \text{Aut}(\mathcal{M}_\lambda)$ reads:

$$(\sigma \triangleright g)(x^0, \vec{x}) = (e^{in\lambda\partial_0} \triangleright g)(x^0, \vec{x}) = g(x^0 + in\lambda, \vec{x}) \quad (7)$$

note that such twist already appear in κ Minkowski [Poulain,Wallet 1994]

Nonunimodular group and twisted trace

$$\mathcal{G} \text{ nonunimodular} \implies \left\{ \begin{array}{l} \Delta_{\mathcal{G}}(p^M) = e^{-n\lambda p^M} \quad \Delta_{\mathcal{G}}, \text{ modular function} \\ x^M = x^0 \quad x^M \text{ is timelike} \\ \text{tr}(f \star g) = \text{tr}((\sigma \triangleright g) \star f), \quad \text{tr} := \int dx^\mu \quad \text{twisted trace} \end{array} \right.$$

where the twist $\sigma \in \text{Aut}(\mathcal{M}_\lambda)$ reads:

$$(\sigma \triangleright g)(x^0, \vec{x}) = (e^{in\lambda\partial_0} \triangleright g)(x^0, \vec{x}) = g(x^0 + in\lambda, \vec{x}) \quad (7)$$

note that such twist already appear in κ Minkowski [\[Poulain,Wallet,1994\]](#)

KMS weight

With (7) we can define a one-parameter group of \star -automorphisms of \mathcal{M}_λ :

$$\{\sigma_t := e^{itn\lambda\partial_0}\}_{t \in \mathbb{R}}, \quad (8)$$

→ $\{\sigma_t\}_{t \in \mathbb{R}}$ **modular group** for the KMS weight given by the twisted trace

A KMS weight ω on a \star algebra \mathcal{A} for a modular group of \star -automorphism $\{\sigma_t\}_{t \in \mathbb{R}}$

- A weight $\omega : \mathcal{A} \rightarrow \mathbb{R}$ densely define and lower semi-continuous.
- $\{\sigma_t\}_{t \in \mathbb{R}}$ admit analytic extension $\{\sigma_z\}_{z \in \mathbb{C}}$
- It exist a function $F : \mathbb{C} \rightarrow \mathbb{C}$ holomorphic on the strip $\{0 \leq \text{Im}(z) \leq \beta\}_{z \in \mathbb{C}}$, $\beta > 0$ which satisfy:

$$F(t) = \omega(f\sigma_t(g)), \quad F(t + i\beta) = \omega(\sigma_t(g)f), \quad \forall t \in \mathbb{R}$$

KMS weight

With (7) we can define a one-parameter group of \star -automorphisms of \mathcal{M}_λ :

$$\{\sigma_t := e^{itn\lambda\partial_0}\}_{t \in \mathbb{R}}, \quad (8)$$

$\Rightarrow \{\sigma_t\}_{t \in \mathbb{R}}$ **modular group** for the **KMS weight** given by the twisted trace

A KMS weight ω on a \star algebra \mathcal{A} for a modular group of \star -automorphism $\{\sigma_t\}_{t \in \mathbb{R}}$

- \Rightarrow A weight $\omega : \mathcal{A} \rightarrow \mathbb{R}$ densely define and lower semi-continuous.
- $\Rightarrow \{\sigma_t\}_{t \in \mathbb{R}}$ admit analytic extension $\{\sigma_z\}_{z \in \mathbb{C}}$
- \Rightarrow It exist a function $F : \mathbb{C} \rightarrow \mathbb{C}$ holomorphic on the strip $\{0 \leq \text{Im}(z) \leq \beta\}_{z \in \mathbb{C}}$, $\beta > 0$ which satisfy:

$$F(t) = \omega(f\sigma_t(g)), \quad F(t + i\beta) = \omega(\sigma_t(g)f), \quad \forall t \in \mathbb{R}$$

KMS weight

With (7) we can define a one-parameter group of \star -automorphisms of \mathcal{M}_λ :

$$\{\sigma_t := e^{itn\lambda\partial_0}\}_{t \in \mathbb{R}}, \quad (8)$$

$\Rightarrow \{\sigma_t\}_{t \in \mathbb{R}}$ **modular group** for the **KMS weight** given by the twisted trace

A KMS weight ω on a $*$ algebra \mathcal{A} for a modular group of \star -automorphism $\{\sigma_t\}_{t \in \mathbb{R}}$

- \Rightarrow A weight $\omega : \mathcal{A} \rightarrow \mathbb{R}$ densely define and lower semi-continuous.
- $\Rightarrow \{\sigma_t\}_{t \in \mathbb{R}}$ admit analytic extension $\{\sigma_z\}_{z \in \mathbb{C}}$
- \Rightarrow It exist a function $F : \mathbb{C} \rightarrow \mathbb{C}$ holomorphic on the strip $\{0 \leq \mathcal{I}(z) \leq \beta\}_{z \in \mathbb{C}}$, $\beta > 0$ which satisfy:

$$F(t) = \omega(f\sigma_t(g)), \quad F(t + i\beta) = \omega(\sigma_t(g)f), \quad \forall t \in \mathbb{R}$$

KMS weight

⇒ Why KMS weight ω ?

- Equilibrium in thermodynamic \implies time flow on the algebra of observable $\alpha_t(A) = e^{iHt} A e^{-iHt}$ where A is an observable.
- C^* algebra: Tomita/modular flow α_t given by the Tomita-Takesaki theorem and acting on the GNS representation π_ω of the $*$ -algebra

$$\alpha_s(\pi_\omega(f)) = \Delta_T^{-is} \pi_\omega(f) \Delta_T^{is}$$

where Δ_T Tomita operator.

$\alpha_s = -\sigma_{it}$ and $\omega = \text{tr}$

Thermal times hypothesis: Connes, Rovelli

The physical time depends on the state. When the system is in a state ω , the physical time is given by the modular group σ_t of ω .

Time translation $(\sigma_t \triangleright f)(x^0, \vec{x}) = f(x^0 + it\lambda t, \vec{x}), \quad f \in \mathcal{M}_\lambda$

KMS weight

- ⇒ Why KMS weight ω ?
- ⇒ Equilibrium in thermodynamic \implies time flow on the algebra of observable $\alpha_t(A) = e^{iHt} A e^{-iHt}$ where A is an observable.
- ⇒ C^* algebra: Tomita/modular flow α_t given by the Tomita-Takesaki theorem and acting on the GNS representation π_ω of the $*$ -algebra

$$\alpha_s(\pi_\omega(f)) = \Delta_T^{-is} \pi_\omega(f) \Delta_T^{is}$$

where Δ_T Tomita operator.

- ⇒ $\alpha_s = -\sigma_{\beta t}$ and $\omega = \text{tr}$

Thermal times hypothesis: Connes, Rovelli

The physical time depends on the state. When the system is in a state ω , the physical time is given by the modular group σ_t of ω .

Time translation $(\sigma_t \triangleright f)(x^0, \vec{x}) = f(x^0 + it(\lambda), \vec{x})$, $f \in \mathcal{M}_\lambda$

KMS weight

- ⇒ Why KMS weight ω ?
- ⇒ Equilibrium in thermodynamic \implies time flow on the algebra of observable $\alpha_t(A) = e^{iHt} A e^{-iHt}$ where A is an observable.
- ⇒ C^* algebra: Tomita/modular flow α_t given by the Tomita-Takesaki theorem and acting on the GNS representation π_ω of the \star -algebra

$$\alpha_s(\pi_\omega(f)) = \Delta_T^{-is} \pi_\omega(f) \Delta_T^{is}$$

where Δ_T Tomita operator.

- ⇒ $\alpha_s = -\sigma_{it}$ and $\omega = \text{tr}$
- ⇒ Thermal times hypothesis: Connes, Rovelli (1994)

The physical time depends on the state. When the system is in a state ω , the physical time is given by the modular group σ_t of ω .

Time translation $(\sigma_t \triangleright f)(x^\mu, \vec{x}) = f(x^\mu + it(\lambda), \vec{x}), \quad f \in \mathcal{M}_\lambda$

KMS weight

- ⇒ Why KMS weight ω ?
- ⇒ Equilibrium in thermodynamic \implies time flow on the algebra of observable $\alpha_t(A) = e^{iHt} A e^{-Ht}$ where A is an observable.
- ⇒ $\mathbb{C}\star$ algebra: Tomita/modular flow α_t given by the Tomita-Takesaki theorem and acting on the GNS representation π_ω of the \star -algebra

$$\alpha_s(\pi_\omega(f)) = \Delta_T^{-is} \pi_\omega(f) \Delta_T^{is}$$

where Δ_T Tomita operator.

- ⇒ $\alpha_s = -\sigma_{\beta t}$ and $\omega = \text{tr}$
- ⇒ Thermal times hypothesis: Connes, Rovelli (1994)
 - ⇒ *The physical time depends on the state. When the system is in a state ω , the physical time is given by the modular group σ_t of ω .*
- ⇒ Time translation $(\sigma_t \triangleright f)(x^0, \vec{x}) = f(x^0 + it\lambda t, \vec{x}), \quad f \in \mathcal{M}_\lambda$

KMS weight

- ⇒ Why KMS weight ω ?
- ⇒ Equilibrium in thermodynamic \implies time flow on the algebra of observable $\alpha_t(A) = e^{iHt} A e^{-iHt}$ where A is an observable.
- ⇒ C^* algebra: Tomita/modular flow α_t given by the Tomita-Takesaki theorem and acting on the GNS representation π_ω of the \star -algebra

$$\alpha_s(\pi_\omega(f)) = \Delta_T^{-is} \pi_\omega(f) \Delta_T^{is}$$

where Δ_T Tomita operator.

- ⇒ $\alpha_s = -\sigma_{\beta t}$ and $\omega = \text{tr}$
- ⇒ Thermal times hypothesis: [Connes, Rovelli, 1994]

The physical time depends on the state. When the system is in a state ω , the physical time is given by the modular group σ_t of ω .

- ⇒ Time translation $(\sigma_t \triangleright f)(x^0, \vec{x}) = f(x^0 + it\lambda t, \vec{x})$, $f \in \mathcal{M}_\lambda$

KMS weight

- ⇒ Why KMS weight ω ?
- ⇒ Equilibrium in thermodynamic \implies time flow on the algebra of observable $\alpha_t(A) = e^{iHt} A e^{-Ht}$ where A is an observable.
- ⇒ $\mathbb{C}\star$ algebra: Tomita/modular flow α_t given by the Tomita-Takesaki theorem and acting on the GNS representation π_ω of the \star -algebra

$$\alpha_s(\pi_\omega(f)) = \Delta_T^{-is} \pi_\omega(f) \Delta_T^{is}$$

where Δ_T Tomita operator.

- ⇒ $\alpha_s = -\sigma_{\beta t}$ and $\omega = \text{tr}$
- ⇒ Thermal times hypothesis: [Connes,Rovelli,1994]

The physical time depends on the state. When the system is in a state ω , the physical time is given by the modular group σ_t of ω .

- ⇒ Time translation $(\sigma_t \triangleright f)(x^0, \vec{x}) = f(x^0 + it\lambda t, \vec{x})$, $f \in \mathcal{M}_\lambda$

Conclusion and Outlook

⇒ What we did:

- ✓ Construction of several C^* -algebras
- ✓ Investigation of nonunimodular groups and KMS weights. Derivation of a natural time flow.

⇒ What we want to do:

- ✍ Construction of associated Poincaré symmetry Hopf algebra (already done for half of them)
- ✍ Implementation of gauge theory

Thank you for your attention

Conclusion and Outlook

⇒ What we did:

- ✓ Construction of several C^* -algebras
- ✓ Investigation of nonunimodular groups and KMS weights. Derivation of a natural time flow.

⇒ What we want to do:

- ✍ Construction of associated Poincaré symmetry Hopf algebra (already done for half of them)
- ✍ Implementation of gauge theory

Thank you for your attention