

Carroll gravity from the conformal approach

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Introduction

- Carroll field theory = ultralocal ($c \rightarrow 0$) limit of relativistic field theory. Invariant under Carroll algebra = $c \rightarrow 0$ contraction of Poincaré algebra. (Lévy-Leblond, Gupta)
- Relevant for e.g.:
 - Flat space/celestial holography (Duval, Gibbons, Horvathy; Bagchi, Grumiller et al.; Ciambelli, Marteau, Petkou, Petropoulos, Siampos; Donnay, Fiorucci, Herfray, Ruzziconi;...)
 - Physics on null hypersurfaces, black hole horizon (Donnay, Marteau; Bagchi, Banerjee, Hartong, Have, Kolekar, Mandlik)
 - Tensionless limit of strings (Bagchi et al.), new decoupling limits of string theory (Bidussi, Harmark, Hartong, Obers, Oling; Blair, Lahnsteiner, Obers, Yan)
 - Dark energy and inflation (de Boer, Hartong, Obers, Sybesma, Vandoren), fluid dynamics (de Boer, Hartong, Obers, Sybesma, Vandoren), ...
- This talk: Carroll gravity. At two- ∂ level, usually studied by carefully defining $c \rightarrow 0$ limit of General Relativity. (Henneaux; Dautcourt; Bergshoeff, Gomis, Rollier, JR, ter Veldhuis; Campoleoni, Henneaux, Pekar, Perez, Salgado-Rebolledo)
- Can be subtle. More complete understanding and generalizations require intrinsic constructions.
- Here: conformal approach.

Outline

- Einstein gravity from the conformal approach
- Carroll gravity from the $c \rightarrow 0$ limit of General Relativity
- Carroll gravity from the conformal approach
- Conclusions

Einstein gravity from the conformal approach

- Einstein gravity is not invariant under local dilatations. Can be reintroduced using a compensating scalar ϕ :

$$E_\mu^{\hat{A}} \rightarrow \phi^{1/w} E_\mu^{\hat{A}}.$$

Invariant under

$$\delta\phi = w\Lambda_D(x)\phi, \quad \delta E_\mu^{\hat{A}} = -\Lambda_D(x)E_\mu^{\hat{A}}.$$

- This replacement gives the Lagrangian of a conformally coupled scalar:

$$ER \rightarrow E \left(\frac{(D-1)(D-2)}{w^2} \partial_{\hat{A}}\phi \partial^{\hat{A}}\phi + R\phi^2 \right).$$

$\phi = 1$ dilatation gauge fixing gives Einstein-Hilbert.

- Strategy: use this to construct gravity theories by reversing the logic.
 - Construct action of conformally coupled scalar intrinsically, by coupling ϕ to appropriate gauge fields of the conformal algebra.
 - Retrieve gravity by gauge-fixing superfluous conformal symmetries.

Einstein gravity from the conformal approach

- Step 1: obtain “minimal” representation of gauge fields of relativistic conformal algebra.
- Introduce gauge fields of relativistic conformal algebra

$$P_{\hat{A}} \rightarrow E_{\mu}^{\hat{A}}, \quad M_{\hat{A}\hat{B}} \rightarrow \Omega_{\mu}^{\hat{A}\hat{B}}, \quad K_{\hat{A}} \rightarrow F_{\mu}^{\hat{A}}, \quad D \rightarrow B_{\mu},$$

with corresponding field strengths $R_{\mu\nu}(P^{\hat{A}})$, $R_{\mu\nu}(M^{\hat{A}\hat{B}})$, $R_{\mu\nu}(K^{\hat{A}})$, $R_{\mu\nu}(D)$.

- “Reducible” representation, realized by too many *independent* gauge fields.
To get a minimal representation, impose so-called conventional constraints:

$$R_{\mu\nu}(P^{\hat{A}}) \equiv 2 \partial_{[\mu} E_{\nu]}^{\hat{A}} + 2 \Omega_{[\mu}^{\hat{A}\hat{B}} E_{\nu]\hat{B}} + 2 B_{[\mu} E_{\nu]}^{\hat{A}} = 0,$$

$$E_{\hat{B}}^{\nu} R_{\mu\nu}(M^{\hat{A}\hat{B}}) \equiv E_{\hat{B}}^{\nu} \left(2 \partial_{[\mu} \Omega_{\nu]}^{\hat{A}\hat{B}} + 2 \Omega_{[\mu}^{\hat{A}|\hat{C}|} \Omega_{\nu]}^{\hat{C}\hat{B}} + 8 F_{[\mu}^{\hat{A}} E_{\nu]}^{\hat{B}} \right) = 0.$$

Allows to solve $\Omega_{\mu}^{\hat{A}\hat{B}}$ and $F_{\mu}^{\hat{A}}$ in terms of $E_{\mu}^{\hat{A}}$ and B_{μ} . E.g.:

$$\Omega_{\mu}^{\hat{A}\hat{B}}(E, B) = \Omega_{\mu}^{\hat{A}\hat{B}}(E) + 2 E_{\mu}^{[\hat{A}} E^{\hat{B}]\nu} B_{\nu}, \quad F_{\hat{A}}^{\hat{A}} = -\frac{1}{4(D-1)} R'_{\hat{A}\hat{B}}(M^{\hat{A}\hat{B}}).$$

where $\Omega_{\mu}^{\hat{A}\hat{B}}(E)$ is the usual Levi-Civita spin-connection and

$$R'_{\mu\nu}(M^{\hat{A}\hat{B}}) = 2 \partial_{[\mu} \Omega_{\nu]}^{\hat{A}\hat{B}}(E, B) + 2 \Omega_{[\mu}^{\hat{A}|\hat{C}|}(E, B) \Omega_{\nu]}^{\hat{C}\hat{B}}(E, B).$$

Einstein gravity from the conformal approach

- Step 2: Construct action of a conformally coupled compensating scalar field ϕ with dilatation weight w

$$\delta\phi = w \Lambda_D \phi \quad \Rightarrow \quad D_{\hat{A}}\phi \equiv E_{\hat{A}}{}^\mu (\partial_\mu - w B_\mu) \phi.$$

- Under $K_{\hat{A}}$, $B_{\hat{A}}$ transforms with a shift:

$$\delta B_{\hat{A}} = 2 \Lambda_{K\hat{A}} \quad \Rightarrow \quad \delta D_{\hat{A}}\phi = -2w \Lambda_{K\hat{A}}\phi.$$

The proper definition of a covariant conformal box operator on ϕ is then

$$\square^C \phi \equiv E^{\hat{A}\mu} \left[\partial_\mu (D_{\hat{A}}\phi) + \Omega_{\mu\hat{A}}{}^{\hat{B}}(E, B) D_{\hat{B}}\phi - (w + 1) B_\mu D_{\hat{A}}\phi + 2w F_\mu{}^{\hat{A}}(E, B) \phi \right]$$

- For $w = (D - 2)/2$, the following action is conformally invariant:

$$\mathcal{L}_{\text{conf}} = -\frac{1}{2} E \phi \square^C \phi,$$

B_μ does not appear!

$\Rightarrow \mathcal{L}_{\text{conf}}$ describes the conformally coupled scalar and is thus gauge-equivalent to Einstein-Hilbert.

Carroll gravity

- Carroll gravity = $c \rightarrow 0$ limit of General Relativity. Taken by reintroducing c as

$$\hat{A} \rightarrow \{0, A = 1, \dots, D-1\}$$

$$E_\mu^0 = c\tau_\mu, \quad E_\mu^A = e_\mu^A, \quad E_0^\mu = c^{-1}\tau^\mu, \quad E_A^\mu = e_A^\mu.$$

Redefining also the local Lorentz transformation parameters $\Lambda^{\hat{A}\hat{B}}$ as

$$\Lambda^{0A} = c\lambda^{0A}, \quad \Lambda^{AB} = \lambda^{AB},$$

one finds

$$\delta E_\mu^{\hat{A}} = -\Lambda^{\hat{A}\hat{B}} E_\mu^{\hat{B}} \xrightarrow{c \rightarrow 0} \begin{cases} \delta\tau_\mu = -\lambda^{0A} e_{\mu A}, \\ \delta e_\mu^A = -\lambda^A_{\ B} e_\mu^B \end{cases}.$$

- τ_μ, e_μ^A are Vielbeine for a Carrollian geometry. Give along with their dual “inverse” Vielbeine τ^μ, e_A^μ a degenerate Carroll metric structure

$$\tau^{\mu\nu} = \tau^\mu \tau^\nu, \quad h_{\mu\nu} \equiv e_\mu^A \delta_{AB} e_\nu^B.$$

τ^μ, e_A^μ used to turn curved into flat indices:

$$X_0 \equiv \tau^\mu X_\mu, \quad X_A \equiv e_A^\mu X_\mu, \quad X_{0A} \equiv \tau^\mu e_A^\nu X_{\mu\nu}, \quad X_{AB} \equiv e_A^\mu e_B^\nu X_{\mu\nu}.$$

Carroll gravity

- Einstein-Hilbert action can be expanded as ($e = \det(\tau_\mu, e_\mu^a)$)

$$\mathcal{L}_{\text{EH}} = c^{-2} \mathcal{L}^{(0)} + \mathcal{L}^{(1)} + c^2 \mathcal{L}^{(2)} + \mathcal{O}(c^4).$$

$c \rightarrow 0$ limit can then be taken in two ways.

- First way: retain leading order term $\mathcal{L}^{(0)}$ = “electric Carroll gravity”:

$$\mathcal{L}_{\text{el}} = \mathcal{L}^{(0)} \propto e \left(T'_{0(A,B)} T_0'^{(A,B)} - T_{0A}'^A T_{0B}'^B \right),$$

with “intrinsic torsion” $T'_{0(A,B)} \equiv 2\tau^\mu e_{(A}{}^\nu \partial_{[\mu} e_{\nu]B)}$.

- Second way: first rewrite the expansion of \mathcal{L}_{EH} in the classically equivalent form

$$\mathcal{L}_{\text{EH}} \sim \mathcal{L}^{(0)} + e \left(\lambda^{AB} T'_{0(AB)} - \lambda T_{0A}'^A \right) + c^2 \left(\mathcal{L}^{(2)} - \frac{1}{4} e \lambda^{AB} \lambda_{AB} + \frac{1}{4} e \lambda^2 \right) + \mathcal{O}(c^4).$$

Retaining the leading order term gives “magnetic Carroll gravity”:

$$\mathcal{L}_{\text{magn}} \propto \mathcal{L}^{(0)} + e \left(\lambda^{AB} T'_{0(AB)} - \lambda T_{0A}'^A \right).$$

λ^{AB} , λ are Lagrange multipliers for geometric constraints.

Carroll gravity from the conformal approach

- Can the two Carroll gravity theories be obtained via a conformal approach?
- Starting point: conformal Carroll algebra = $c \rightarrow 0$ contraction of relativistic conformal algebra ($A = 1, \dots, D - 1$)

$$\begin{aligned} H &\rightarrow \tau_\mu, & P_A &\rightarrow e_\mu^A, & J_{AB} &\rightarrow \omega_\mu^{AB}, & J_{0A} &\rightarrow \omega_\mu^{0A}, \\ K &\rightarrow f_\mu, & K_A &\rightarrow g_\mu^A, & D &\rightarrow b_\mu, \end{aligned}$$

Field strengths: $R_{\mu\nu}(H), R_{\mu\nu}(P_A), R_{\mu\nu}(J_{AB}), R_{\mu\nu}(J_{0A}), R_{\mu\nu}(K_A), R_{\mu\nu}(K), R_{\mu\nu}(D)$.

- Reducible representation. Impose conventional constraints:

$$R_{\mu\nu}(H) = R_{BC}(P^A) = R_{\mu B}(J^{AB}) = R_{\mu A}(J^{0A}) = 0.$$

- Can be used to express

$$\omega_\mu^{AB}, \quad \omega_0^{0A}, \quad \omega^{[A|,0|B]}, \quad b_0, \quad g_\mu^A, \quad f_\mu,$$

in terms of the remaining independent gauge fields τ_μ, e_μ^A, b_A and $\omega^{(A|,0|B)}$.

Carroll gravity from the conformal approach

- There exist two kinds of free massless Carroll scalars. Start from the Lagrangian of a massless relativistic scalar in Hamiltonian form:

$$\mathcal{L} = \Pi_{\Phi} \partial_t \Phi - \frac{c^2}{2} \Pi_{\Phi}^2 - \frac{1}{2} \partial^A \Phi \partial_A \Phi.$$

$c \rightarrow 0$ limit can be taken in two ways.

- First way: rescale

$$\Pi_{\Phi} = c^{-1} \pi, \quad \Phi = c \phi,$$

and then take $c \rightarrow 0$. This gives an electric Carroll scalar:

$$\mathcal{L} = \pi \partial_t \phi - \frac{1}{2} \pi^2 \quad \Leftrightarrow \quad \mathcal{L} = \frac{1}{2} \partial_t \phi \partial_t \phi.$$

- Second way: rename

$$\Pi_{\Phi} = \pi, \quad \Phi = \phi,$$

and then take $c \rightarrow 0$. This gives a magnetic Carroll scalar:

$$\mathcal{L} = \pi \partial_t \phi - \frac{1}{2} \partial^A \phi \partial_A \phi.$$

Carroll gravity from the conformal approach

- First case: construct conformally invariant generalization of electric Carroll scalar

$$\mathcal{L} = -\frac{1}{2}\phi\partial_t\partial_t\phi.$$

- Take ϕ with dilatation weight w . The correct generalizations of $\partial_t\phi$ and $\partial_t\partial_t\phi$ and \mathcal{L} are

$$D_0\phi \equiv \tau^\mu (\partial_\mu - wb_\mu)\phi, \quad D_0D_0\phi \equiv \tau^\mu [\partial_\mu (D_0\phi) - (w+1)b_\mu D_0\phi],$$

$$\mathcal{L} = -\frac{1}{2}e\phi D_0D_0\phi.$$

Only dilatations act non-trivially. Conformally invariant for $w = (D-2)/2$.

- Gauge-fixing dilatations by setting $\phi = 1$ gives ‘electric Carroll gravity’:

$$\mathcal{L} = \frac{w^2}{2}eb_0^2 = \frac{w^2}{2(D-1)^2}eT'_{0A}{}^AT'_{0B}{}^B \quad \text{with } T'_{0A}{}^A = 2\tau^\mu e_A{}^\nu \partial_{[\mu}e_{\nu]}{}^A.$$

- Not the most general electric Carroll gravity possible. Other possibility:

$$\mathcal{L}' \propto eT'_{0\{A,B\}}T'_{0\{A,B\}} \quad \text{with } T'_{0\{A,B\}} = 2\tau^\mu e_{\{A}{}^\nu \partial_{[\mu}e_{\nu]B\}}.$$

Not reproduced from a dynamic scalar Lagrangian, since it transforms homogeneously under dilatations.

Carroll gravity from the conformal approach

- Second case: construct conformally invariant generalization of electric Carroll scalar

$$\mathcal{L} = \pi \partial_t \phi - \frac{1}{2} \partial^A \phi \partial_A \phi.$$

- First attempt to make this conformally invariant: assume ϕ has dilatation weight w and generalize \mathcal{L} to

$$\mathcal{L}_0 = \pi D_0 \phi - \frac{1}{2} D^A \phi D_A \phi,$$

$$\text{with } D_0 \phi = \tau^\mu (\partial_\mu - w b_\mu) \phi, \quad D_A \phi = e_A^\mu (\partial_\mu - w b_\mu) \phi.$$

Invariant under dilatations and boosts for $w = (D - 2)/2$ and

$$\delta \pi = \lambda^{0A} D_A \phi + \frac{1}{2} D \lambda_D \pi.$$

- Not invariant under K_A however:

$$\delta \mathcal{L}_0 = 2w e \lambda_K^A \phi D_A \phi.$$

Carroll gravity from the conformal approach

- This can be cancelled by adding a boost invariant combination of $g_A^A \phi^2$ and $f_0 \phi^2$:

$$\mathcal{L} = \pi D_0 \phi - \frac{1}{2} D^A \phi D_A \phi + we g_A^A \phi^2 + we f_0 \phi^2 .$$

Invariance under K requires assigning a K -transformation to π :

$$\delta_K \pi = (D - 2) \lambda_K \phi .$$

- Fix dilatations, K_A and K transformations by:

$$\begin{aligned} \phi &= 1 , & b_A &= 0 , & \pi &= 0 , \\ \mathcal{L}_{\text{fixed}} &= we \left(g_A^A + f_0 \right) . \end{aligned}$$

- $g_A^A + f_0$ depends on τ_μ, e_μ^a , but also the independent spin connection components $\omega^{(A|,0|B)}$. Can be identified with Lagrange multipliers for the constraint

$$T'_{0(A,B)} \equiv 2\tau^\mu e_{(A}{}^\nu \partial_{[\mu} e_{\nu]B)} = 0 .$$

$\mathcal{L}_{\text{fixed}}$ then agrees with magnetic Carroll gravity Lagrangian.

Conclusions

- Carroll gravity comes in 2 guises: electric and magnetic. Both can be constructed from the conformal approach.
- Starts from a single conformal Carroll algebra, but two different “electric” and “magnetic” massless compensating scalar field theories.
- Difference with relativistic conformal approach: some invariants involving intrinsic torsion can not be reproduced by coupling to dynamic matter.
- Extra difference with relativistic conformal approach: extra independent spin connection components. Match with Lagrange multipliers in magnetic Carroll gravity.
- Outlook:
 - Including matter couplings.
 - Carroll supergravity.
 - Aristotelian gravity, coupled to fracton matter?