

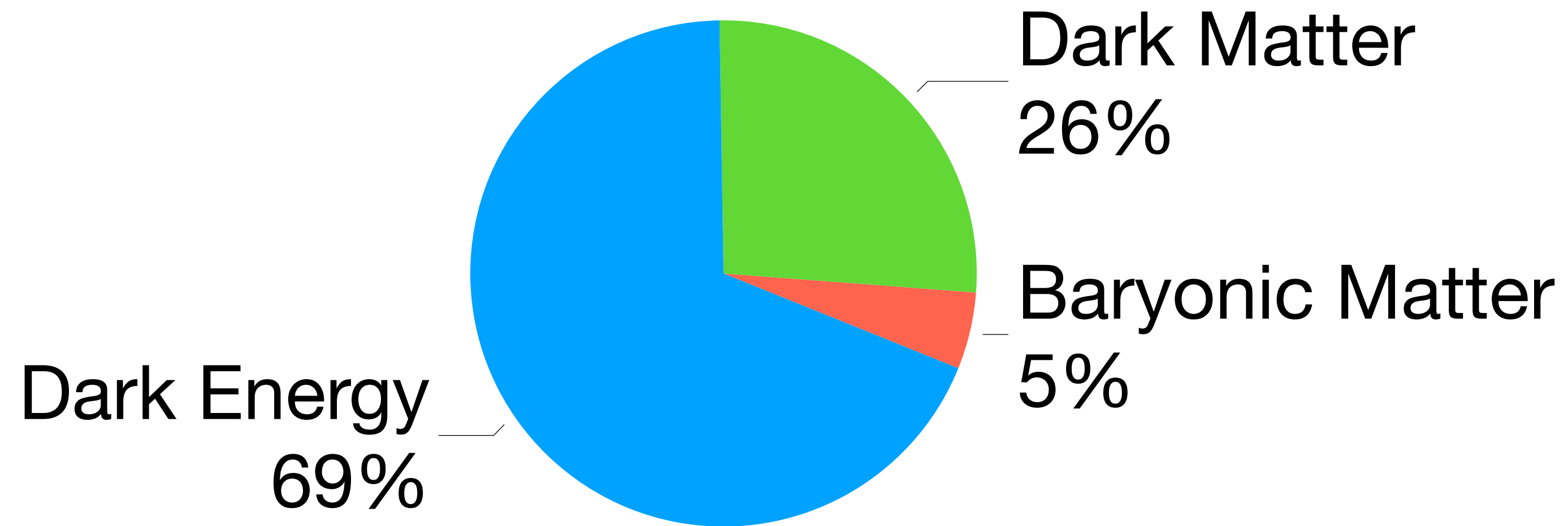
# Enhancing Dark Matter searches at the LHC with Graph Neural Networks

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in collaboration with K. Sakurai and M. Nojiri

DSU2024 12-09-2024

# Dark Matter

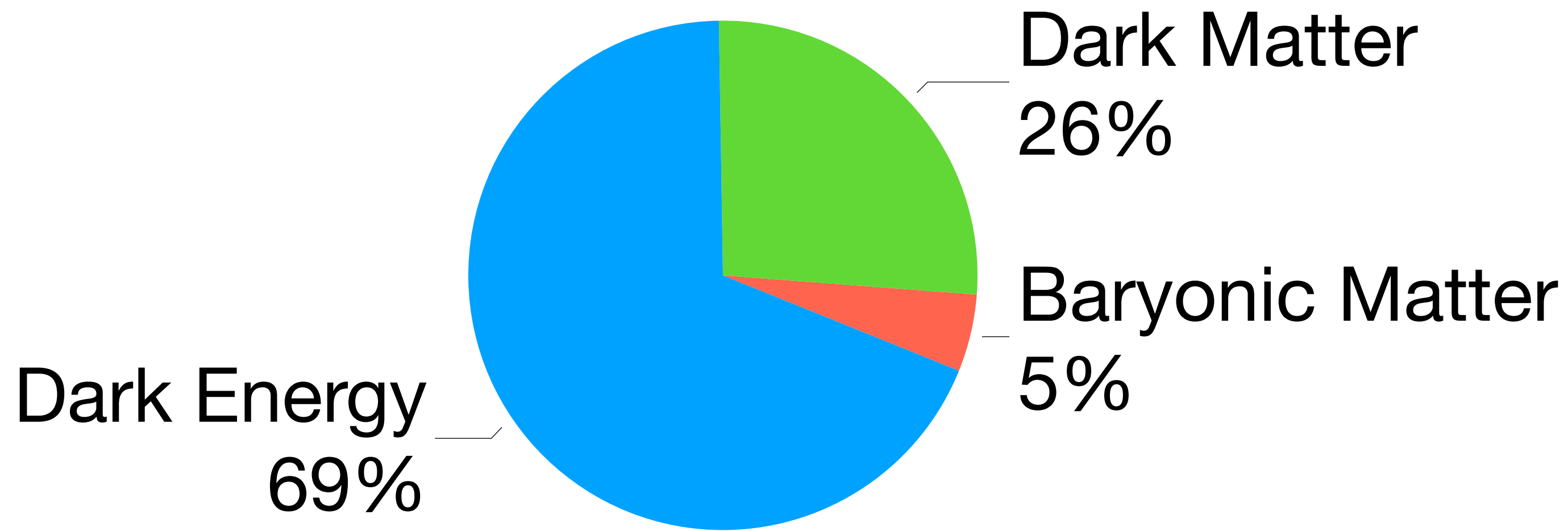


⦿ long-lived over the age of the Universe

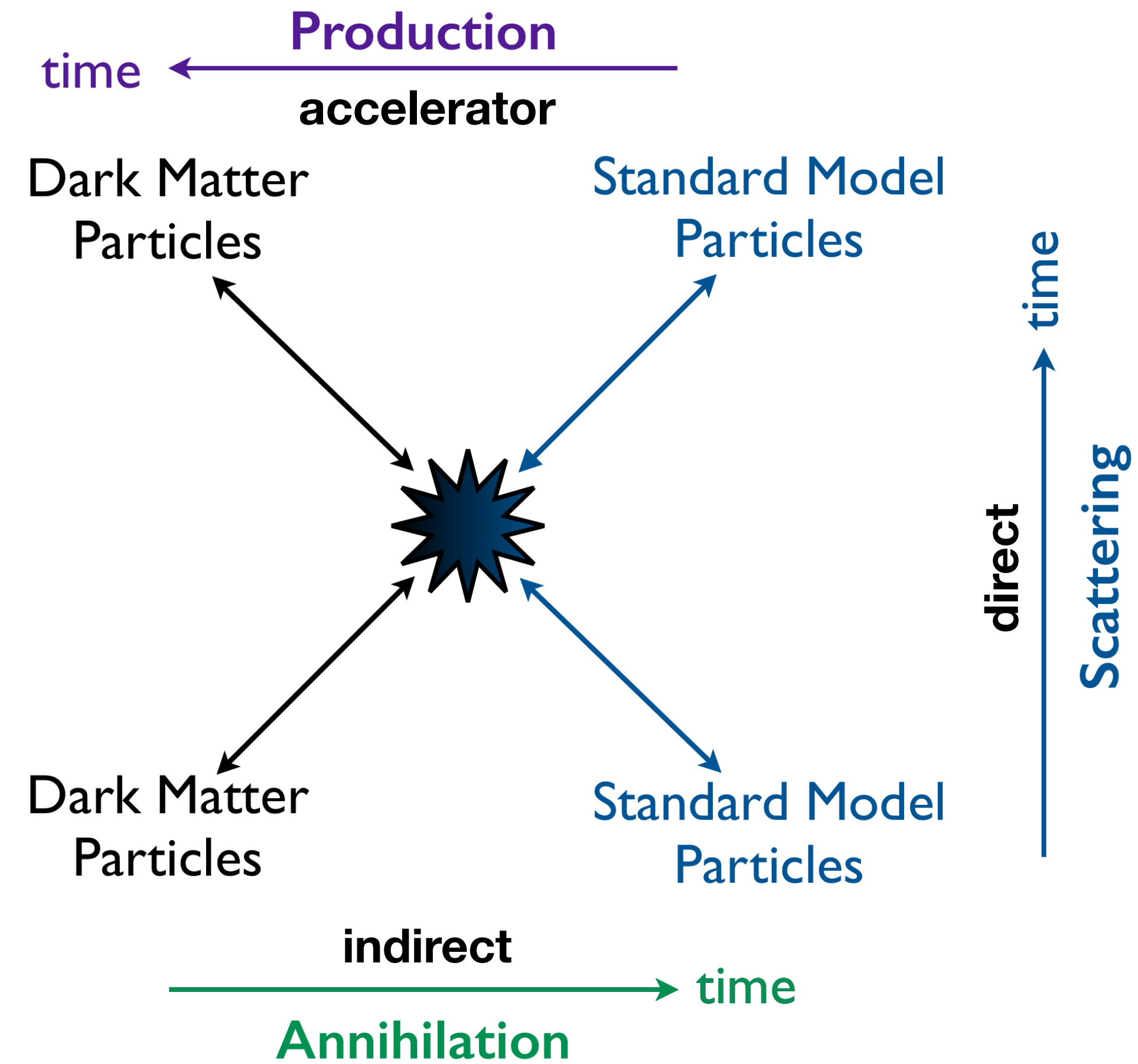
⦿ feebly-interacting with photons and baryons

⦿ not too hot

# Dark Matter



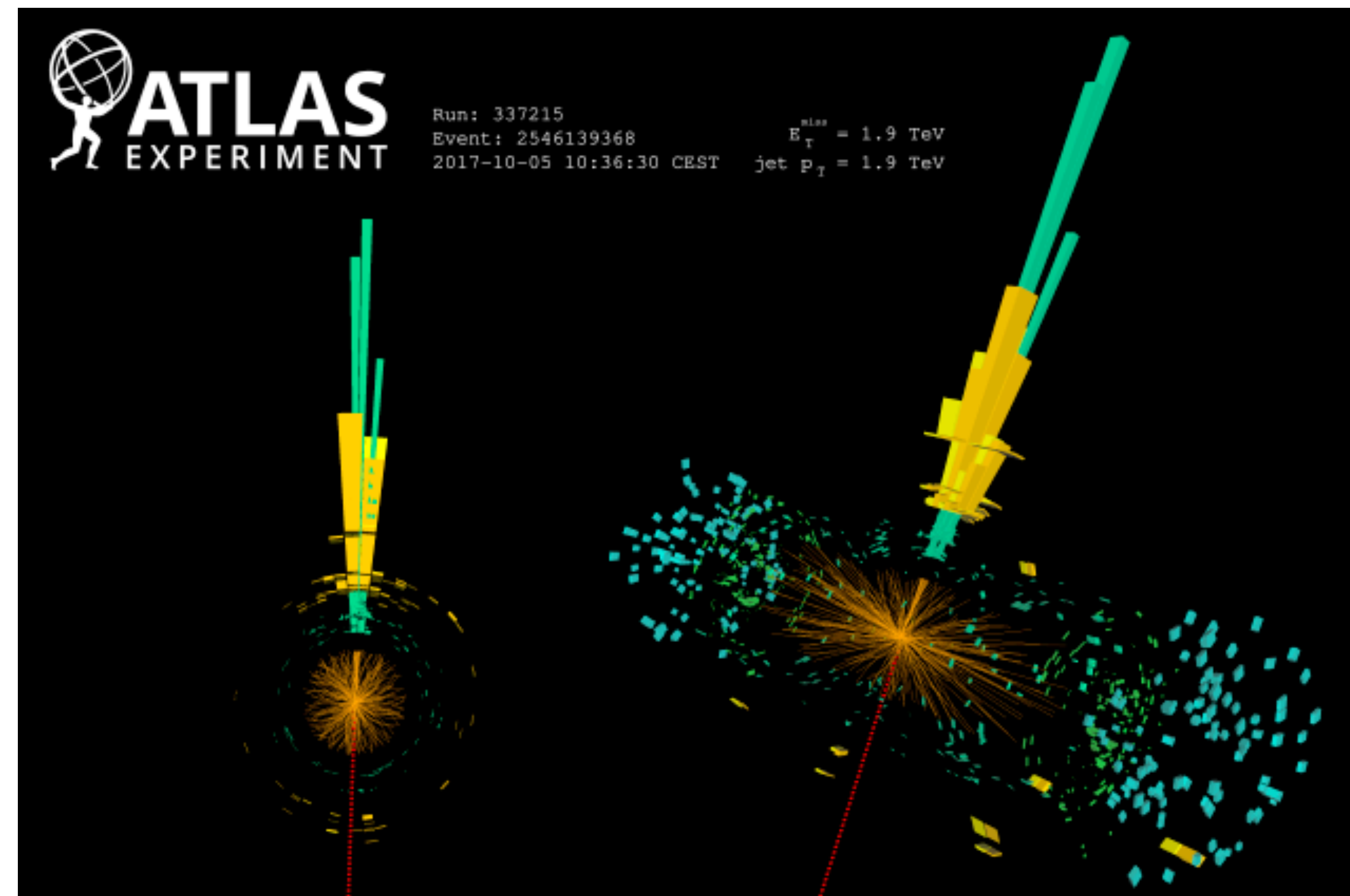
- ⊛ long-lived over the age of the Universe
- ⊛ feebly-interacting with photons and baryons
- ⊛ not too hot



# DM searches @ LHC — Monojet

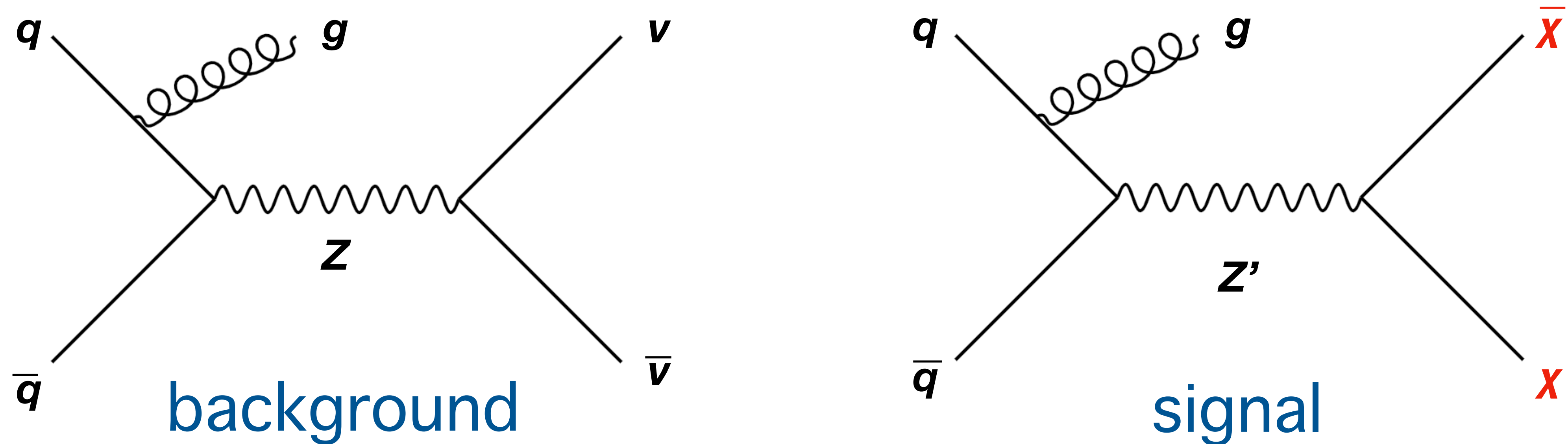
Monojet channel = 1 or more hard jets recoiling against a missing transverse momentum and no isolated leptons

img source: <https://cds.cern.ch/record/2725235>



# DM searches @ LHC — Monojet

Monojet channel = 1 or more hard jets recoiling against a missing transverse momentum and no isolated leptons

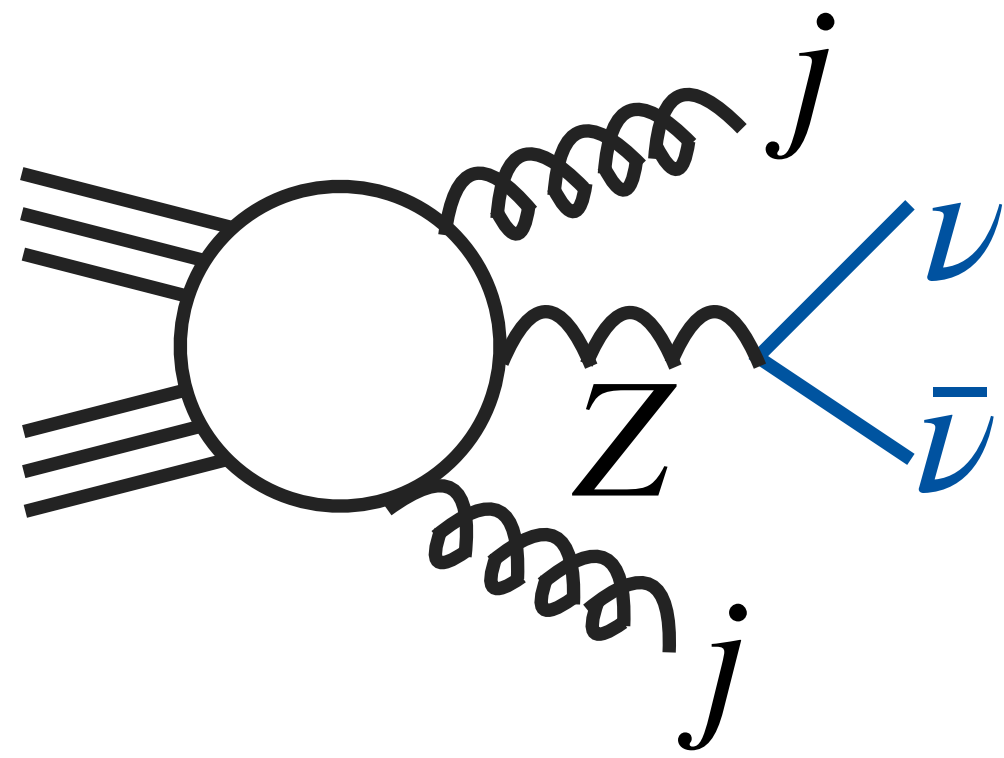


# The idea

- ⊛ One of the challenges for the Monojet searches is that we observe very similar jets for both signal and background
- ⊛ Analysis of jet substructure is needed
- ⊛ With Machine Learning we can effectively analyse particle-level data
- ⊛ ML can learn both local and global correlations
- ⊛ **GOAL: Design new analysis for the monojet using ML**
- ⊛ We want to combine low-level and high-level variables
- ⊛ We want to use information from the soft activity

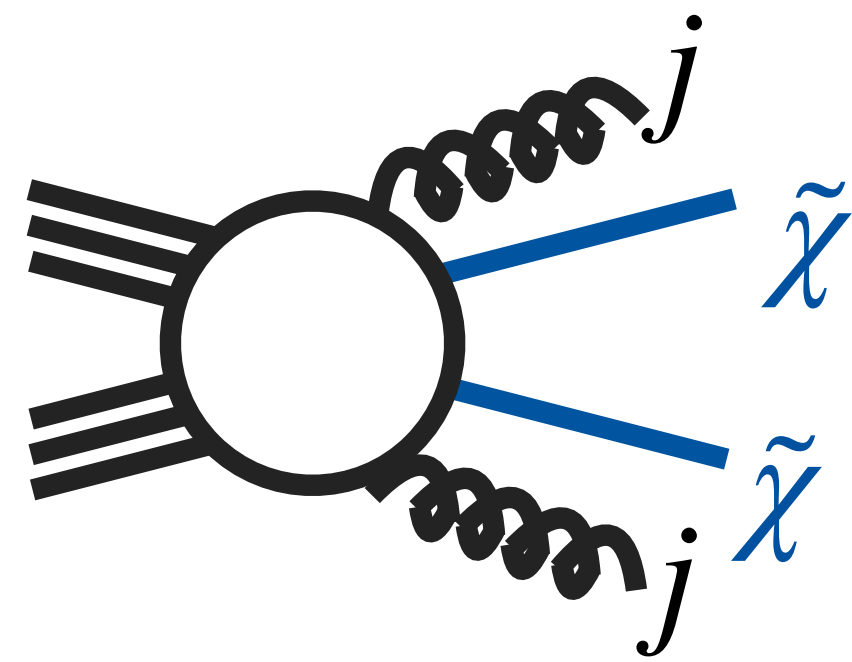
# Benchmark model

SM background



$(Z \rightarrow \nu\bar{\nu}) + \text{jets}$

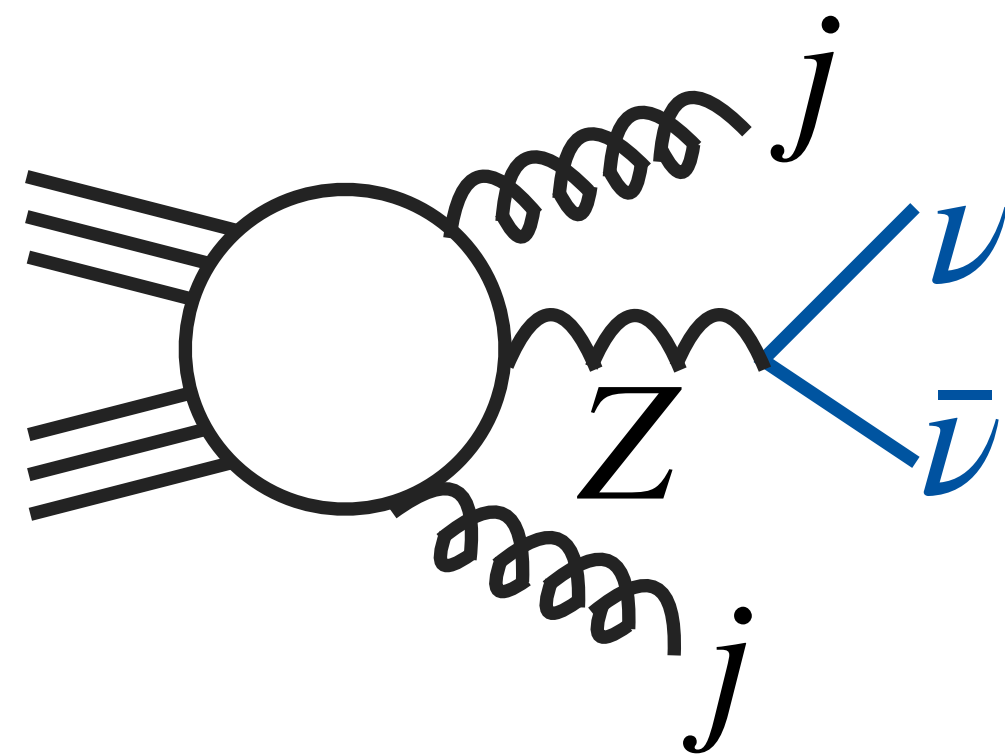
Contributing signal process



EWKino-EWKwino

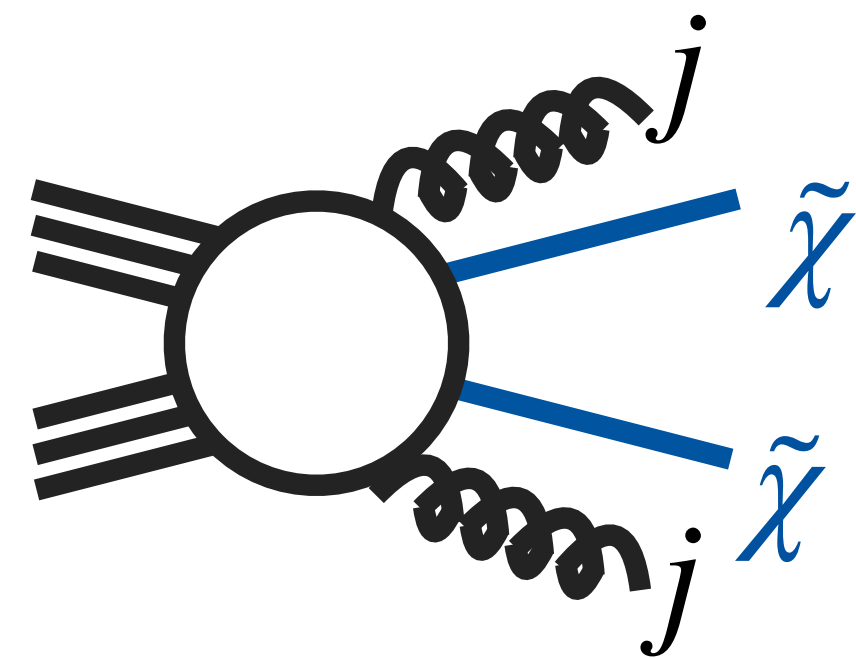
# Benchmark model

SM background

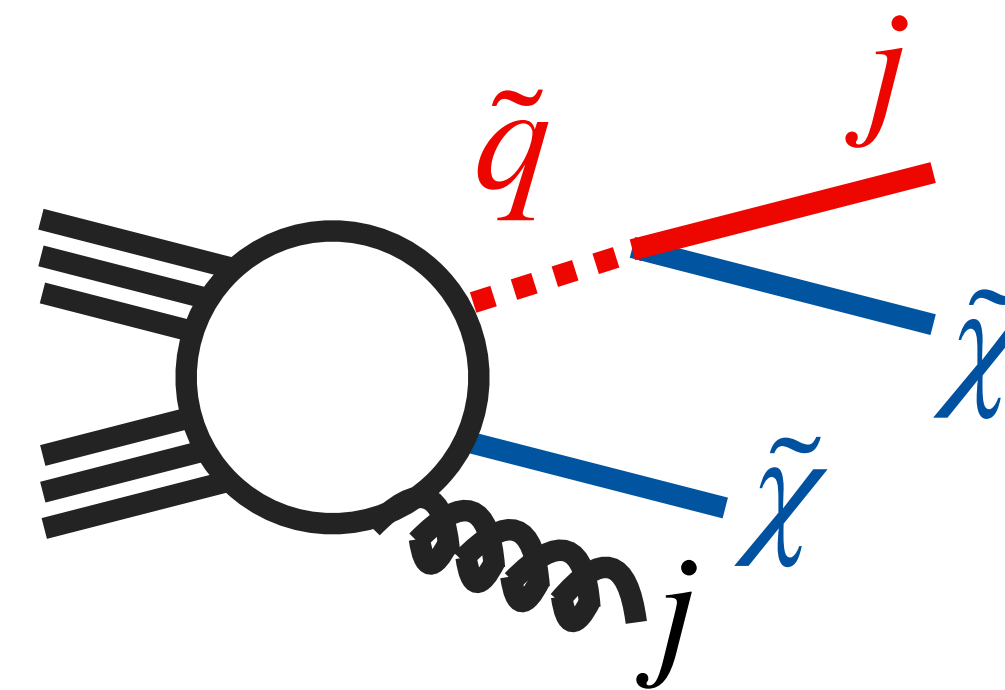


$(Z \rightarrow \nu\bar{\nu}) + \text{jets}$

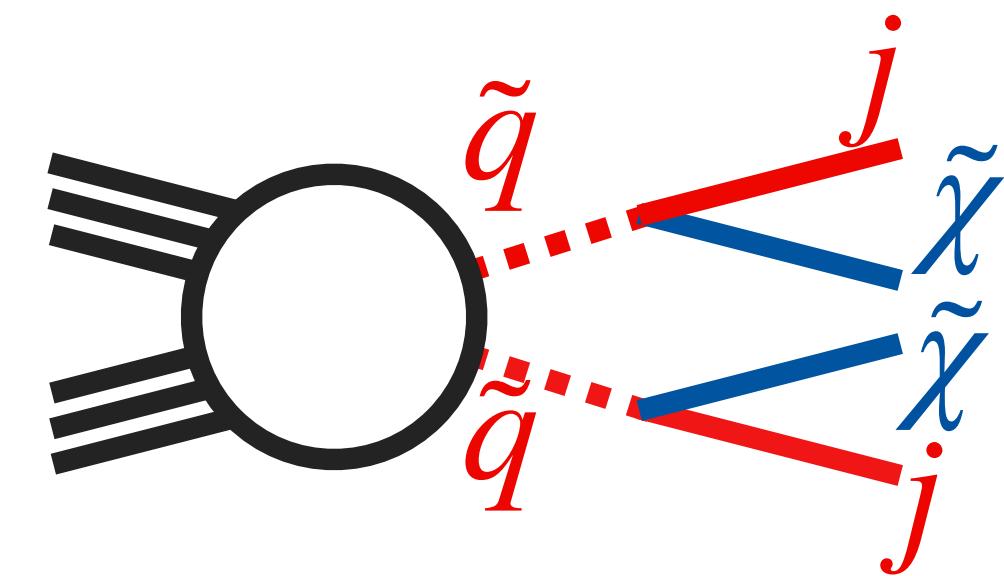
Contributing signal processes



EWKino-EWKwino



EWKino-squark



squark-squark

$$m_{\tilde{\chi}} \in \{200, 300, 500, 700, 900\} \text{ GeV}$$

$$m_{\tilde{q}} \in \{2.0, 2.2, 2.3, 2.7, 3.0\} \text{ TeV}$$

$$m_{\tilde{g}} = 10 \text{ TeV}$$

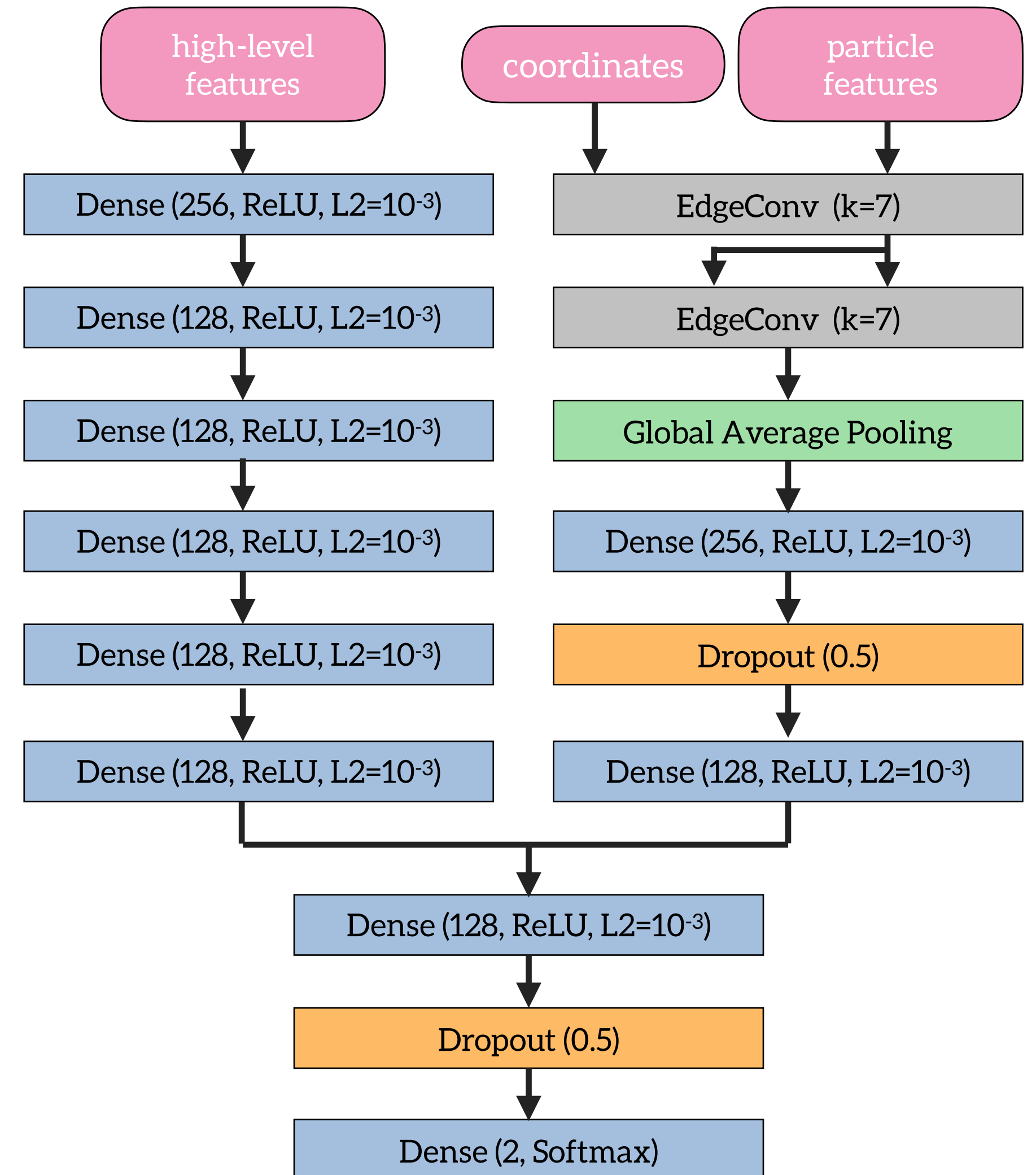


# Analysis

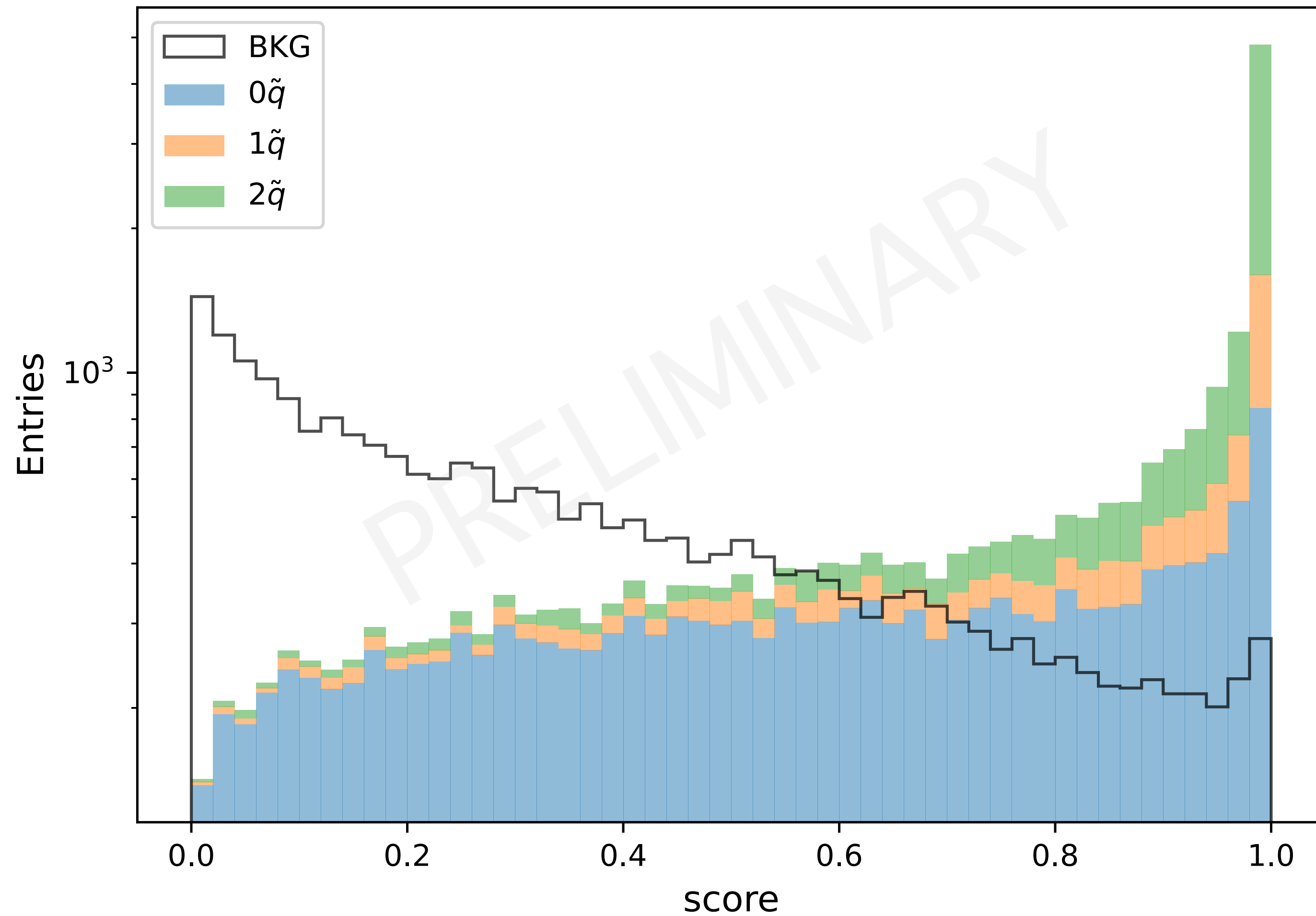
## Preanalysis

- $4 \geq \#jets \geq 2$
- $p_T^{1j} > 520 \text{ GeV}$
- $\eta^{1j} < 2.0$
- $p_T^{2j} > 320 \text{ GeV}$
- $\eta^{2j} < 2.0$
- $MET > 820 \text{ GeV}$
- lepton veto
- $\Delta\phi(j^{1,2,3}, p_T^{\text{miss}}) > 0.8$
- $\Delta\phi(j^4, p_T^{\text{miss}}) > 0.4$
- $MET/\sqrt{H_T} > 16 \sqrt{\text{GeV}}$
- $M_{\text{eff}} > 1600 \text{ GeV}$
- **particle  $p_T > 1 \text{ GeV}$**

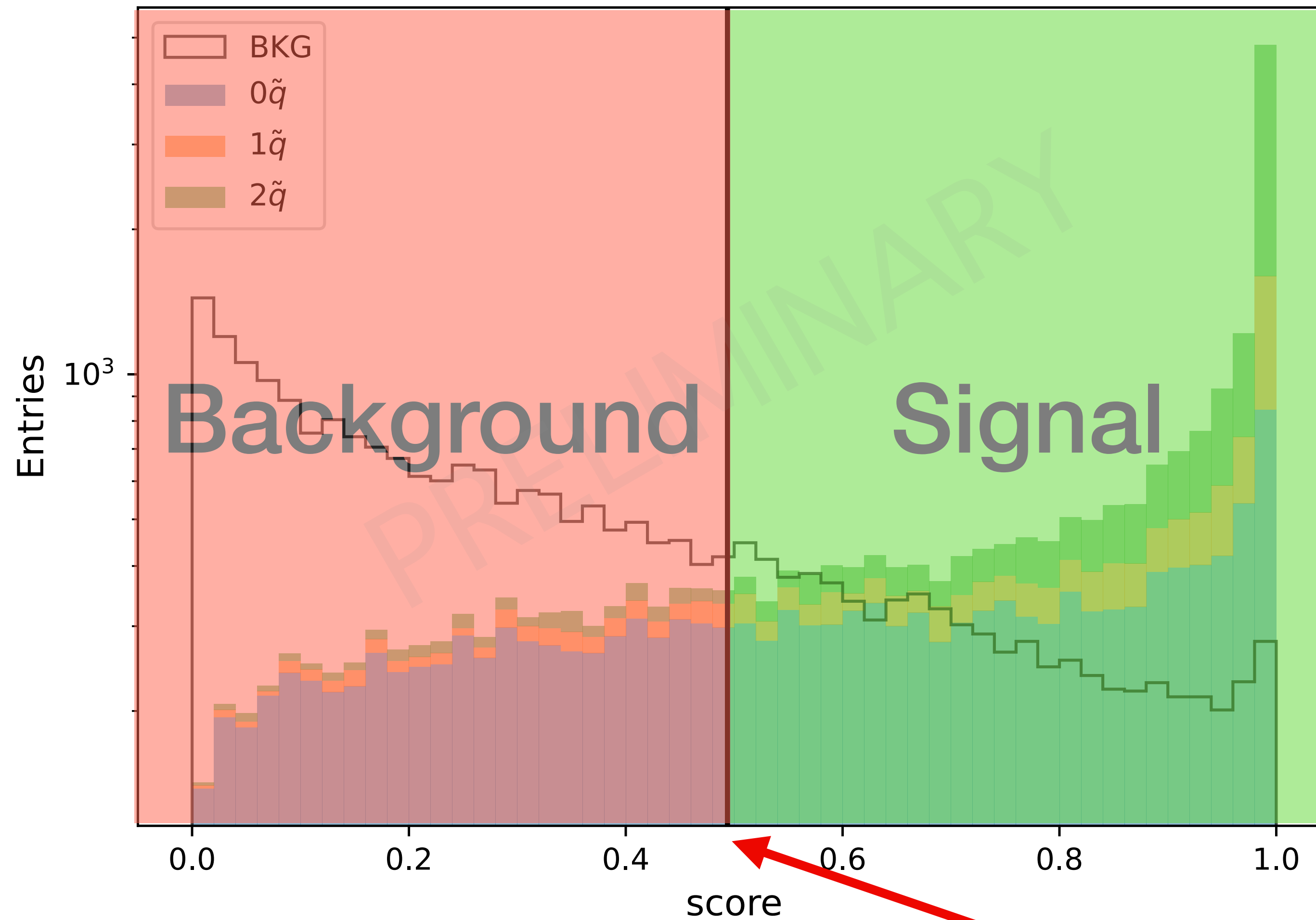
## GNN model



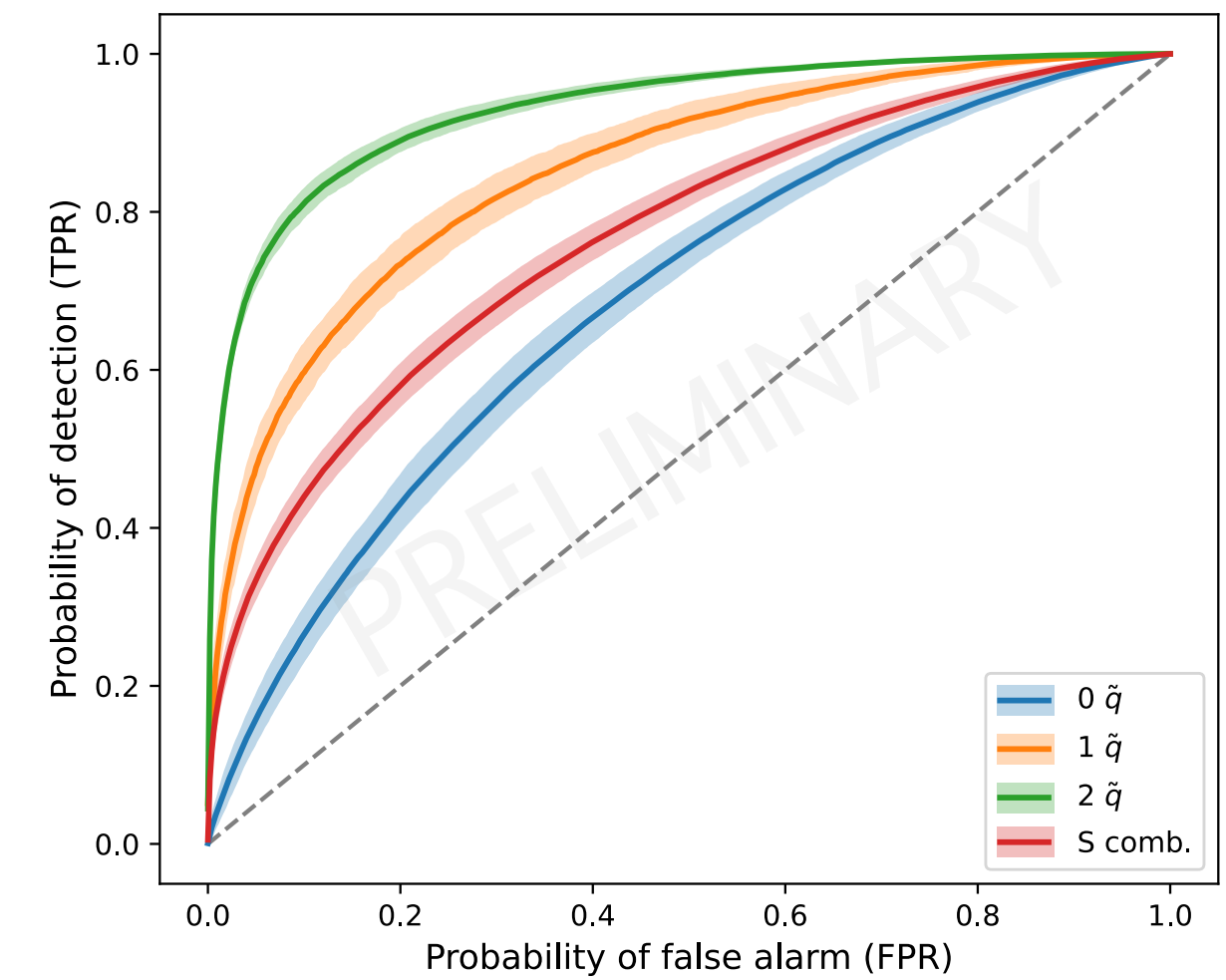
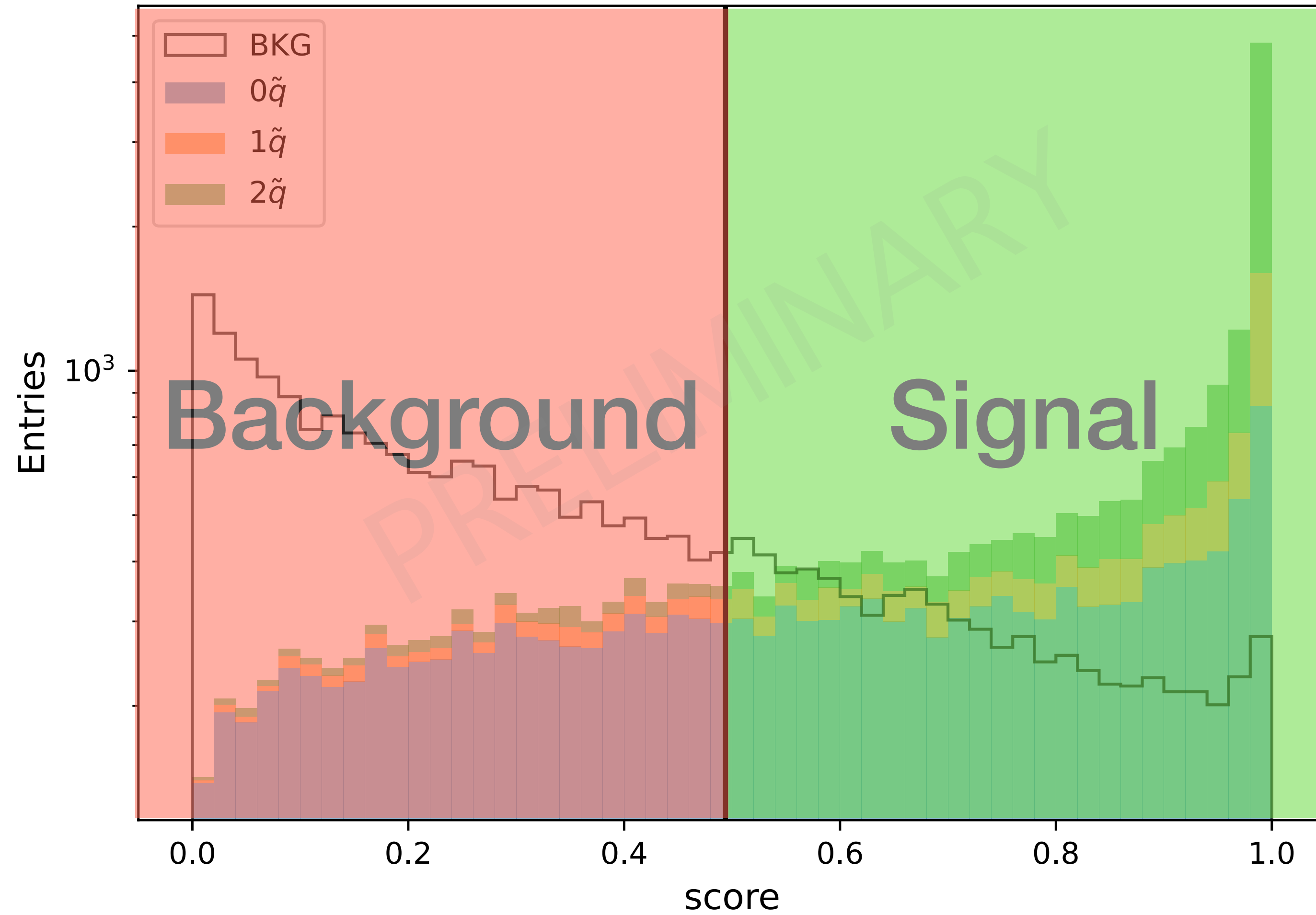
# GNN output (score)



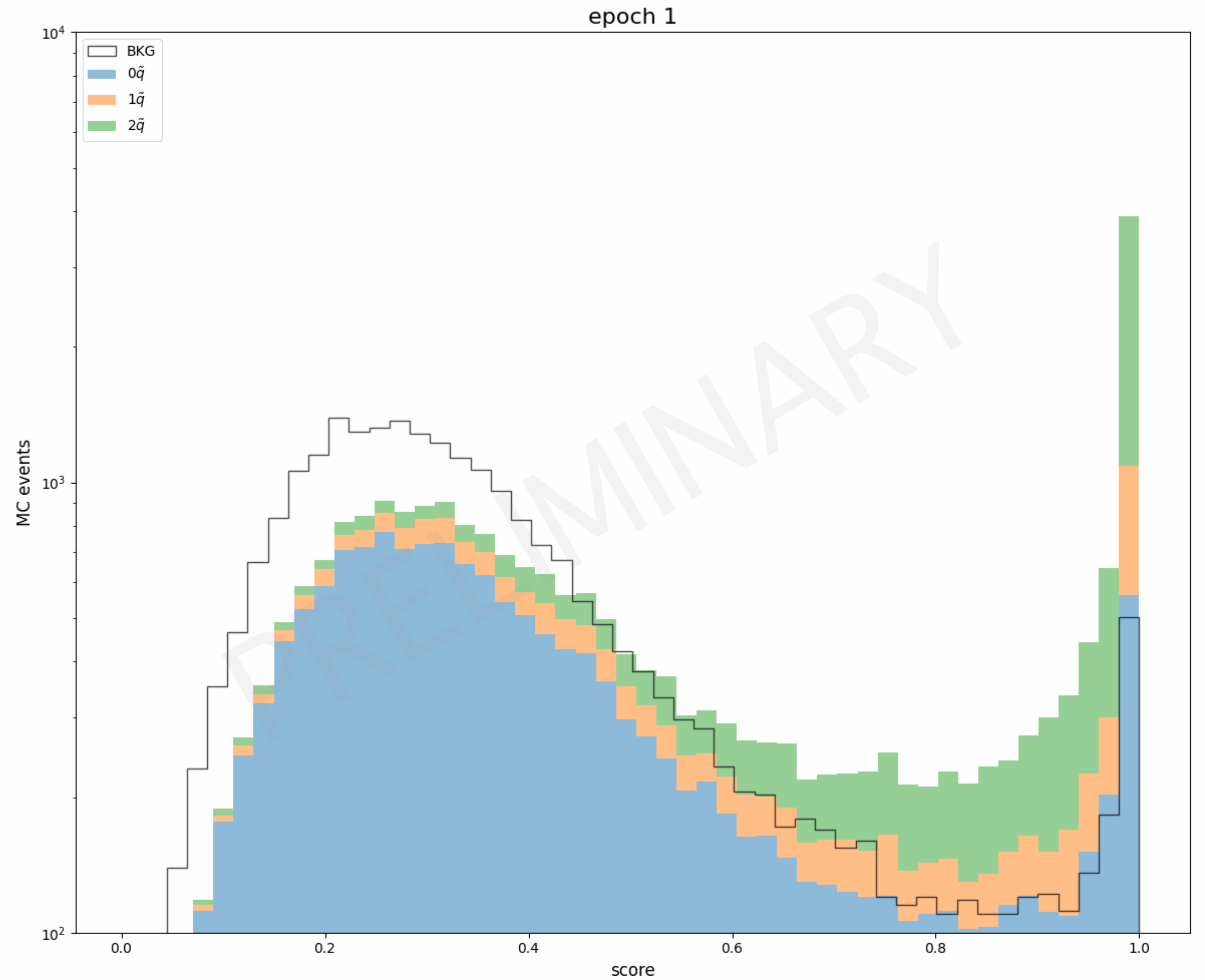
# GNN output (score)



# GNN output (score)



# Training progress (animation)

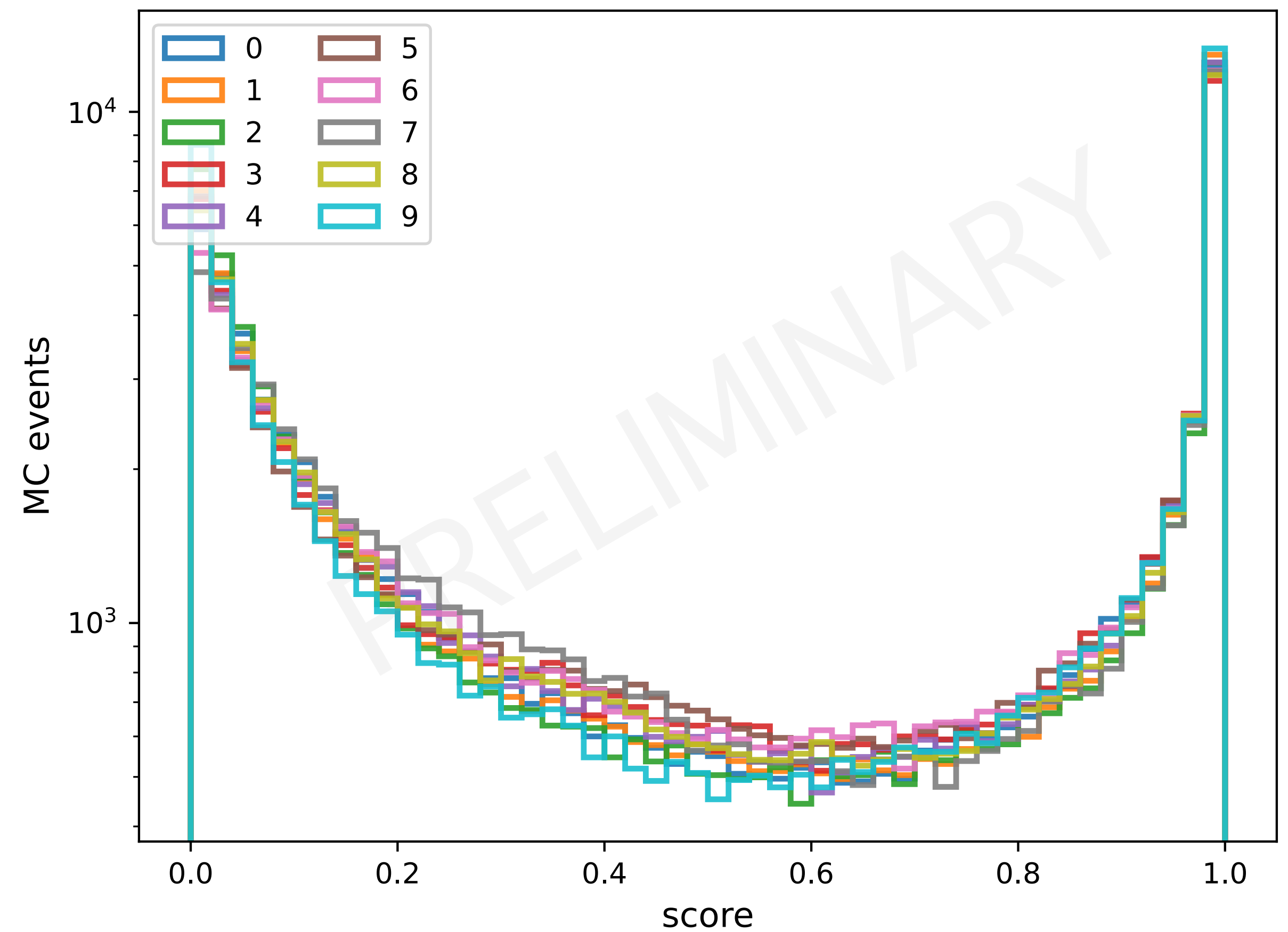
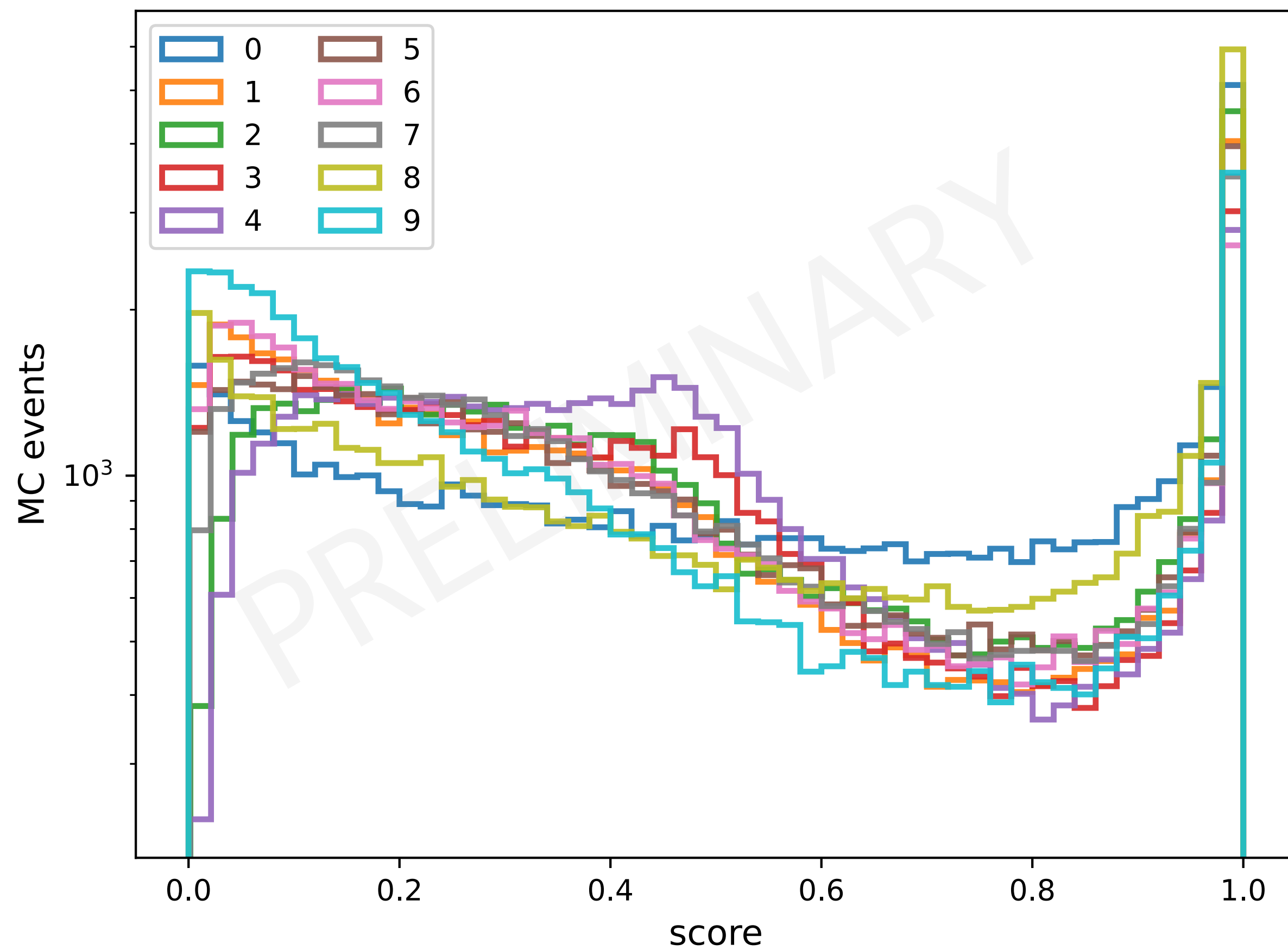


# Ensemble training

**winos**

$$m_{\tilde{\chi}} = 300 \text{ GeV}, m_{\tilde{q}} = 2.2 \text{ TeV}$$

**higgsinos**

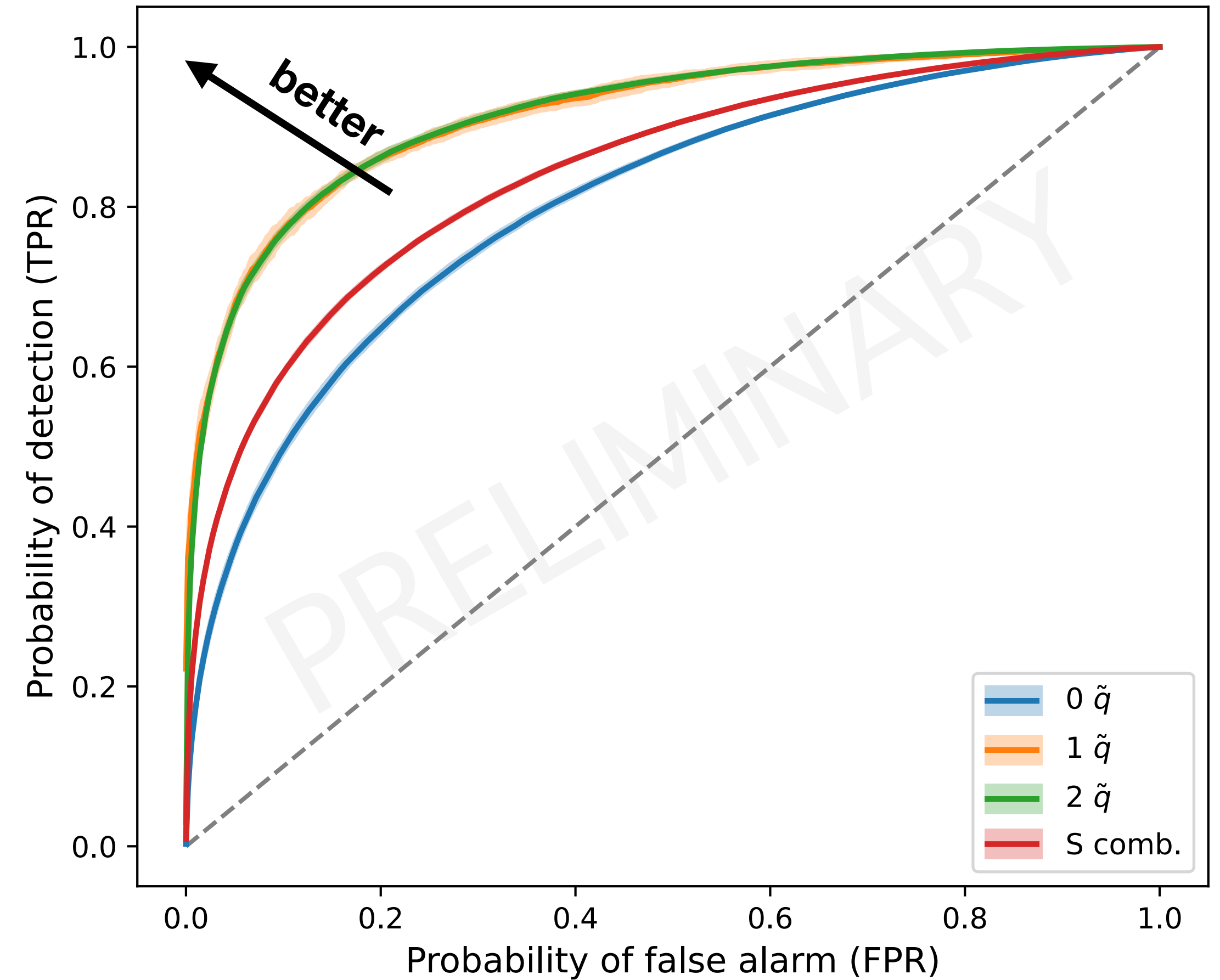
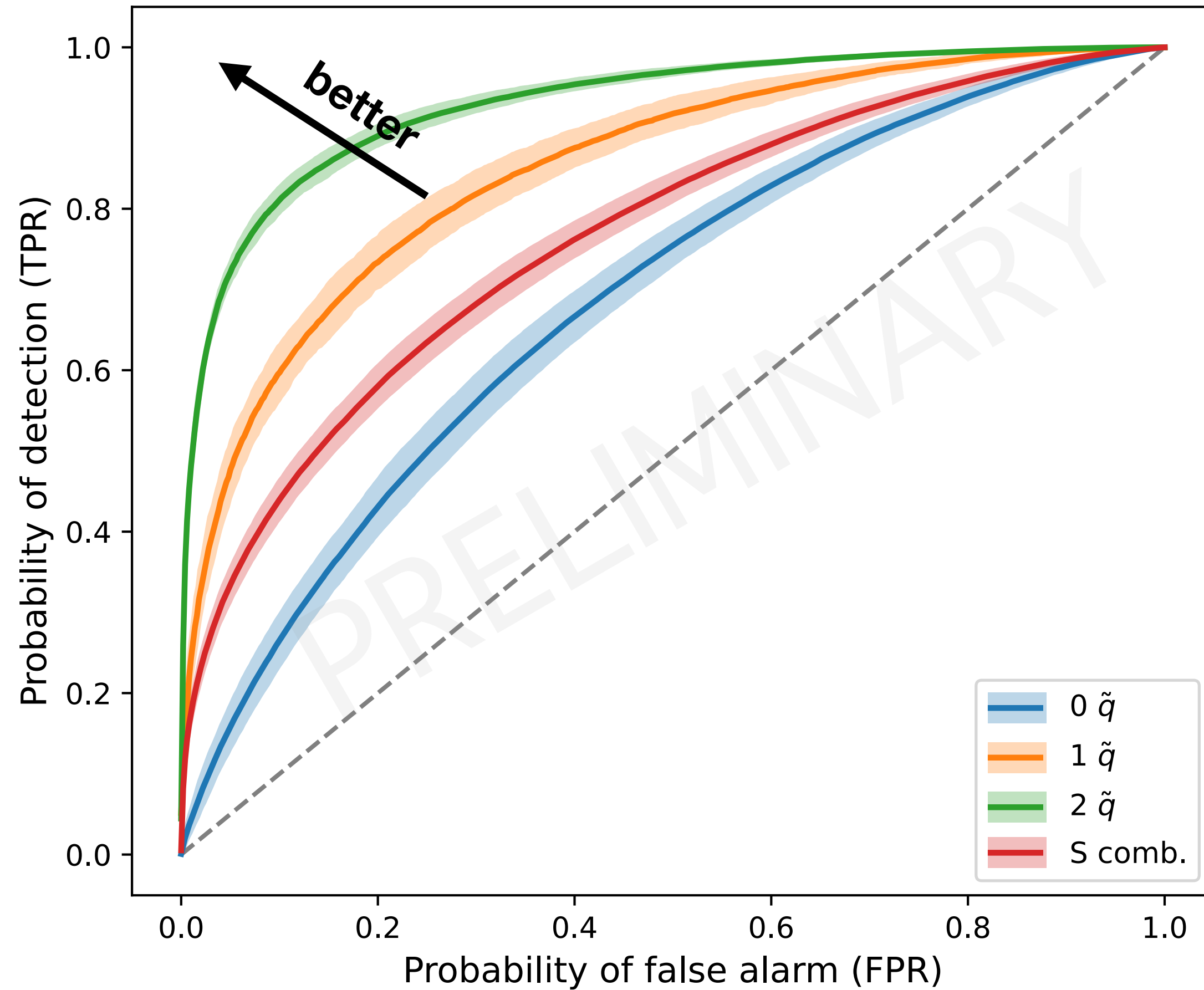


# Evaluation

**winos**

$m_{\tilde{\chi}} = 300 \text{ GeV}, m_{\tilde{q}} = 2.2 \text{ TeV}$

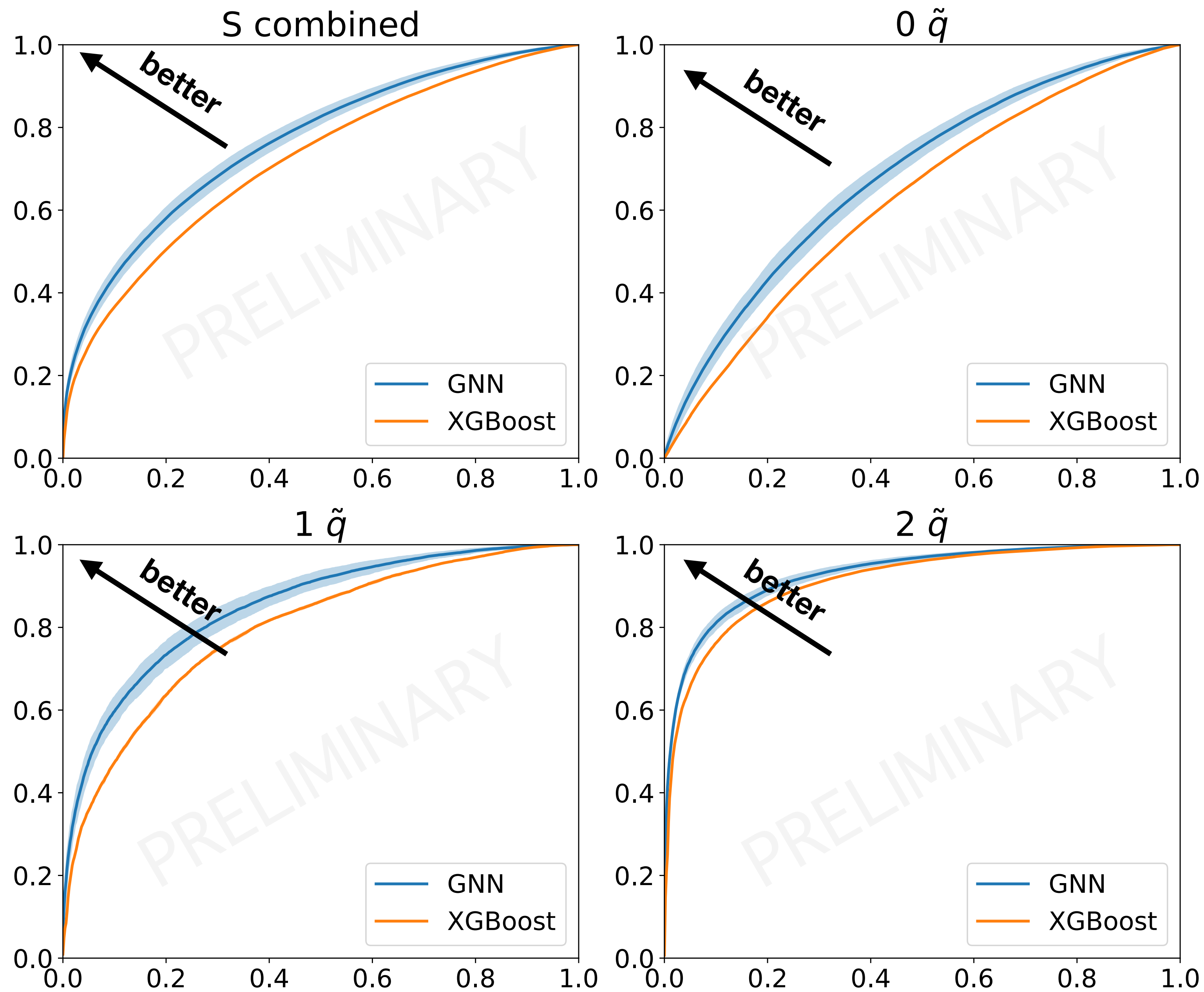
**higgsinos**



Is it better than BDTs?

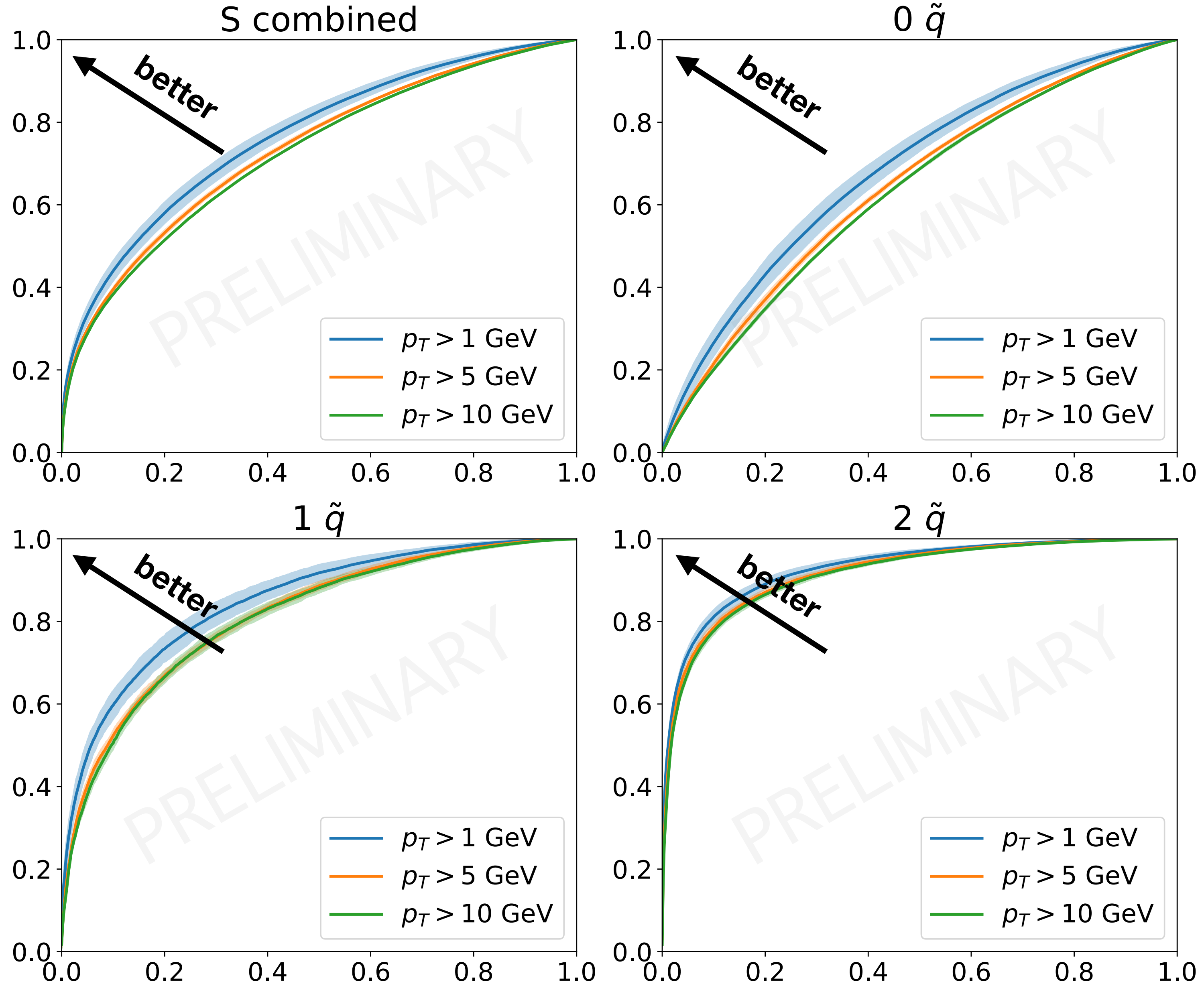


# GNN vs. BDT



What about the  $p_T$  cut?

# particles' $p_T$ cut

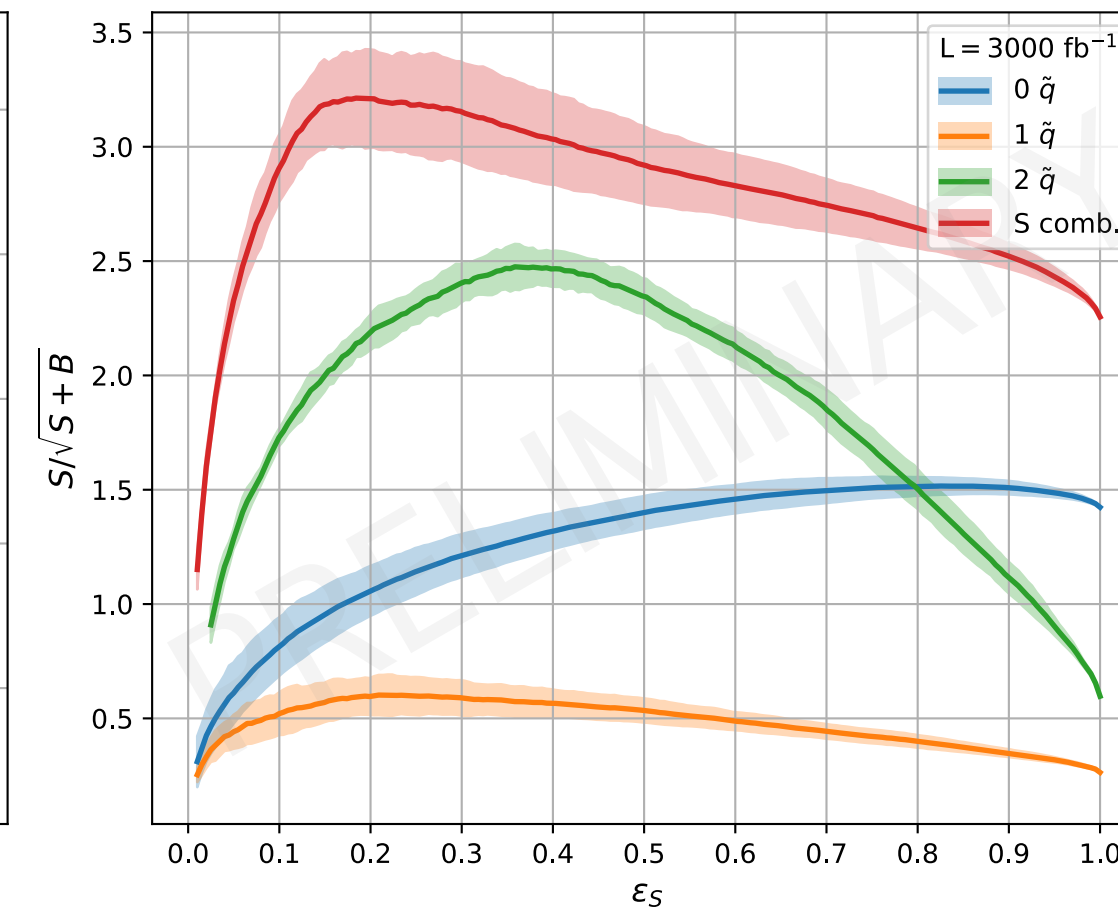
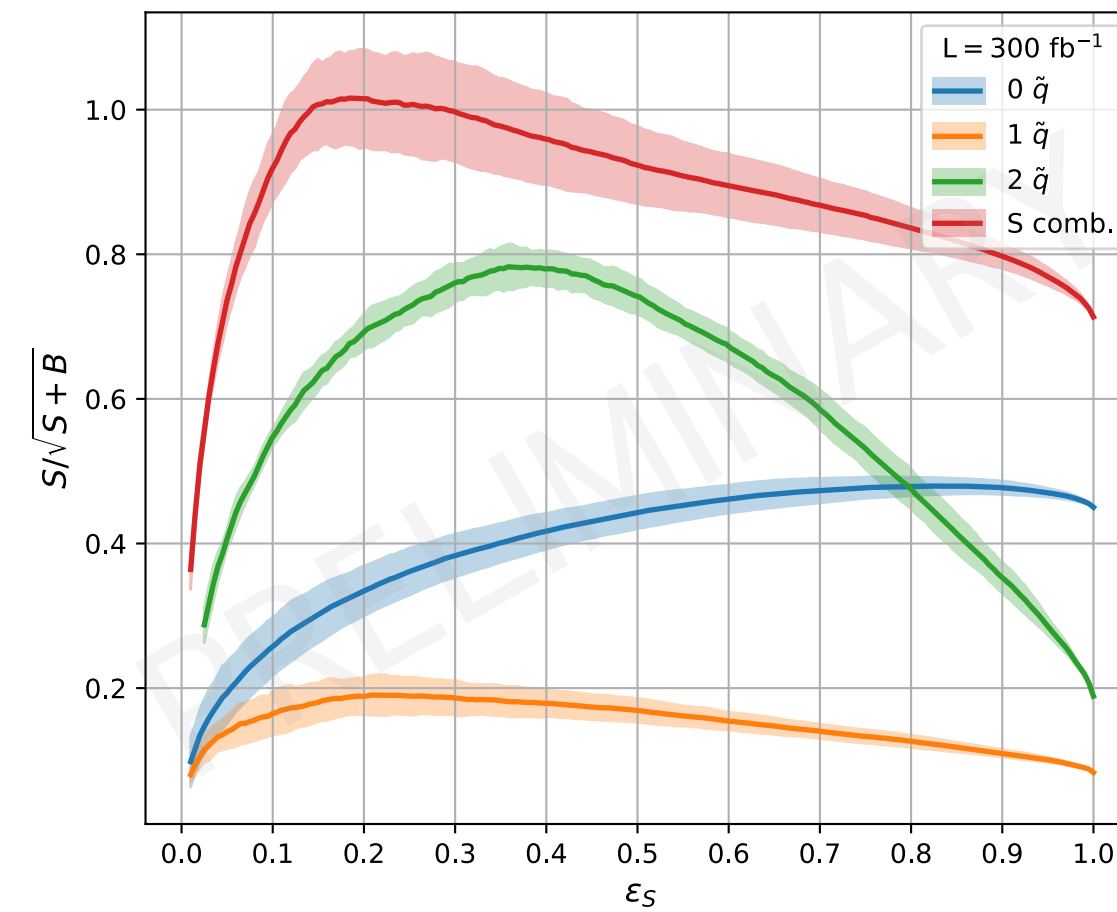


# Signal efficiency vs. naive significance

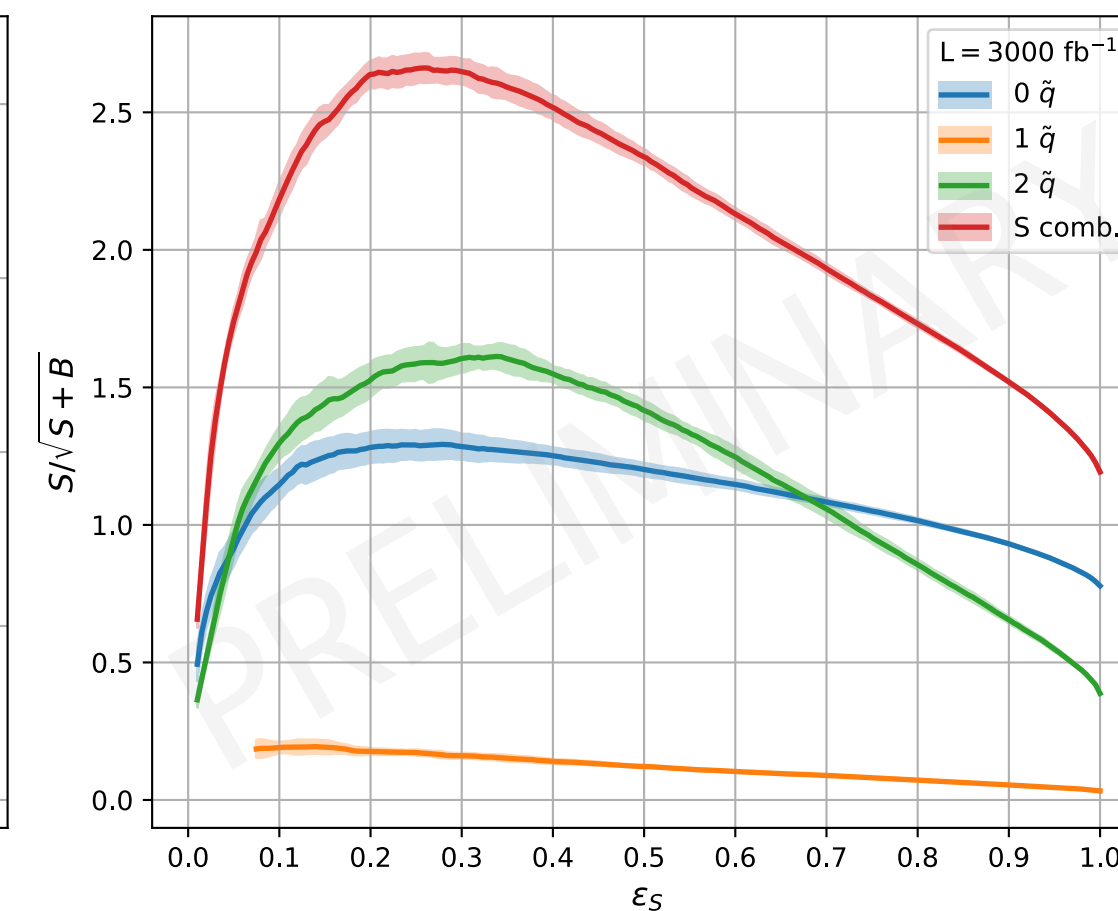
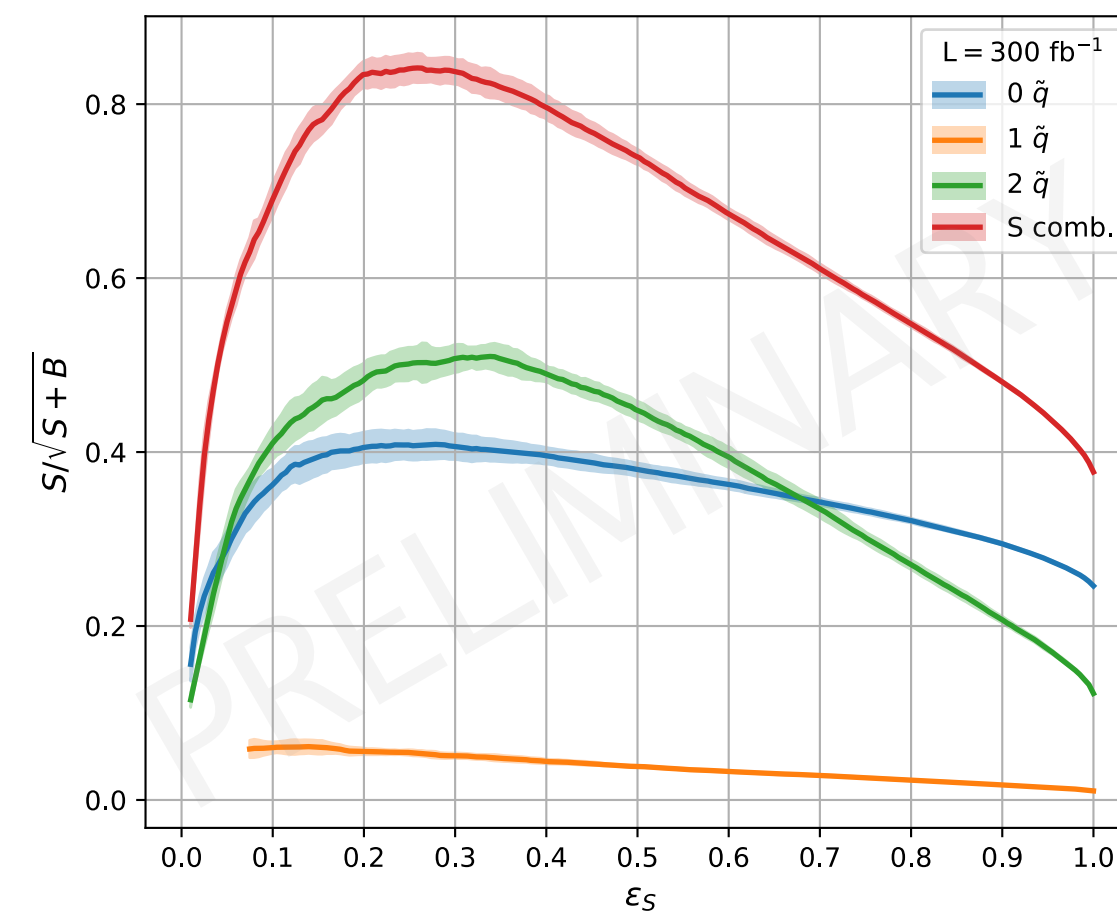
Run-3

HL-LHC

winos



higgsinos

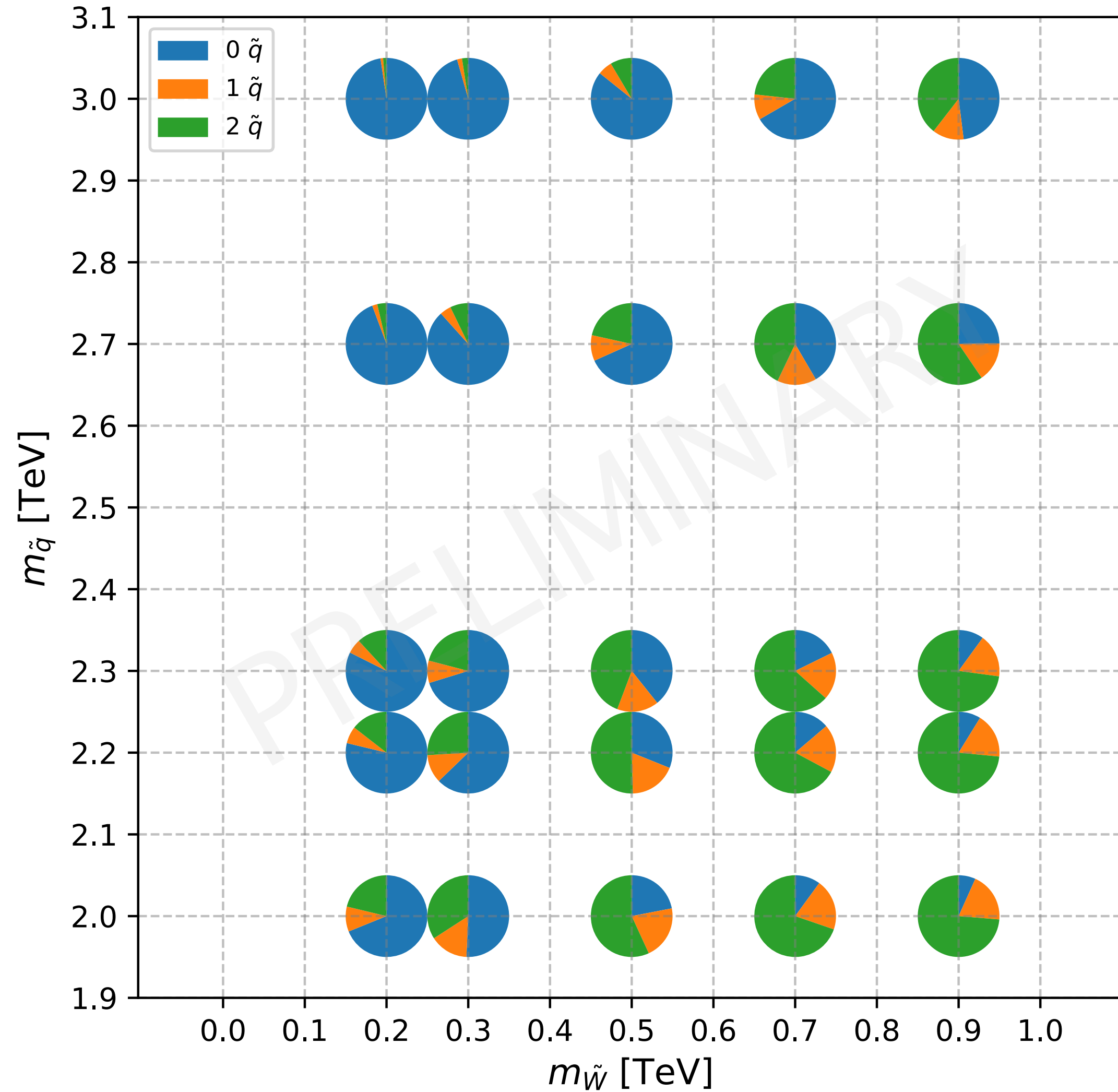


$$Z = S/\sqrt{S+B}$$

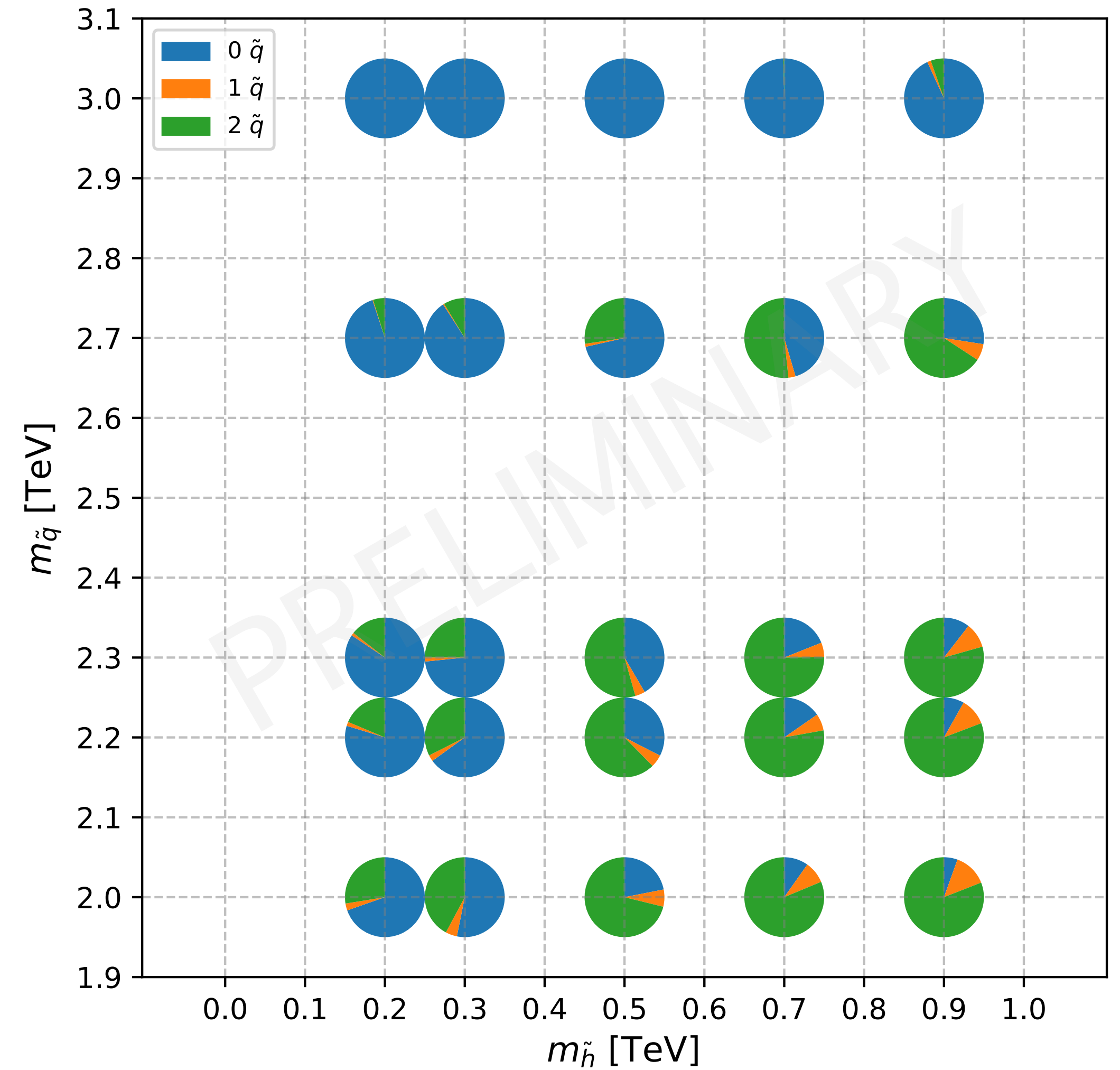
$m_{\tilde{\chi}} = 300 \text{ GeV}, m_{\tilde{q}} = 2.2 \text{ TeV}$

# Sample composition

## wino

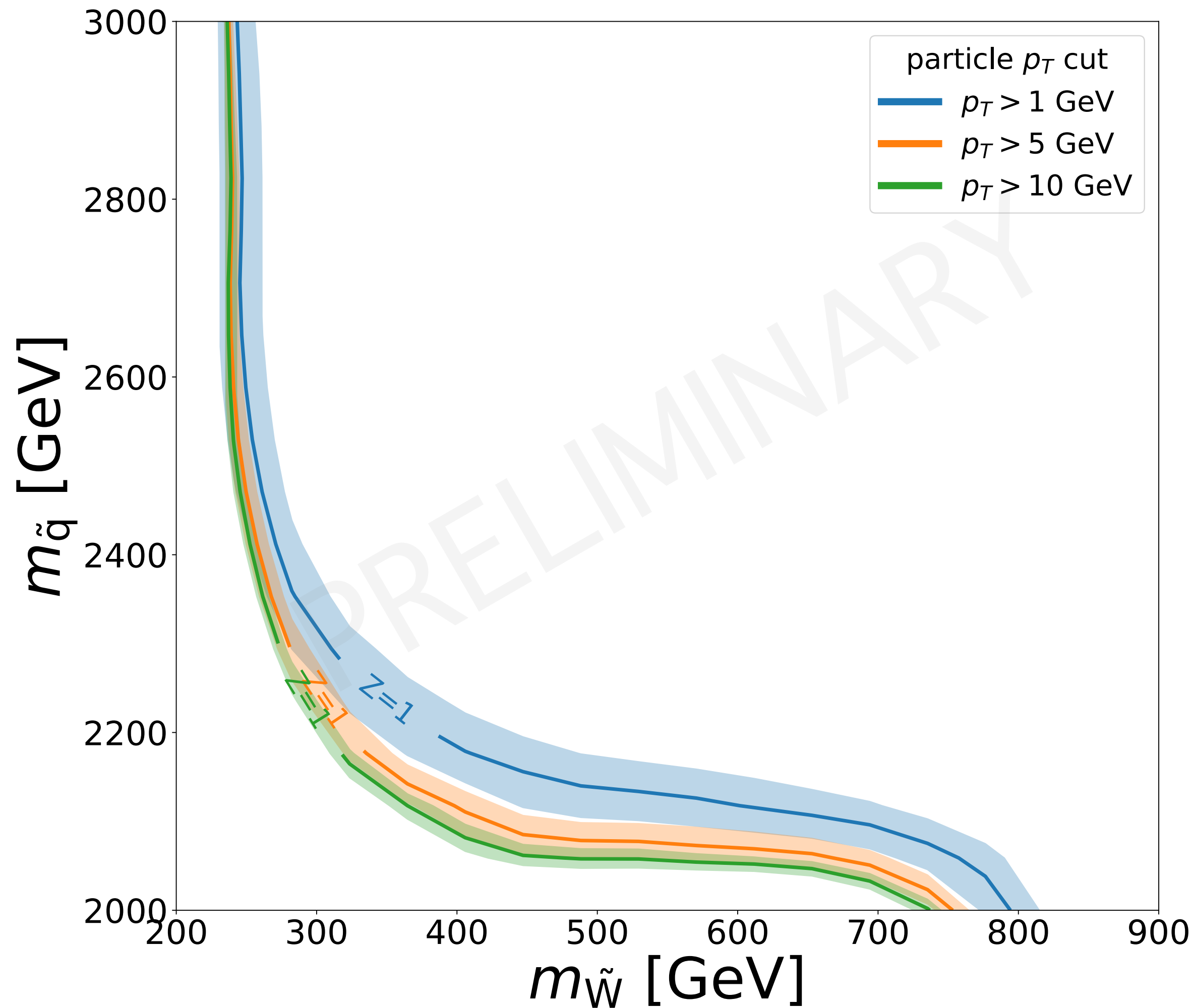


## higgsino

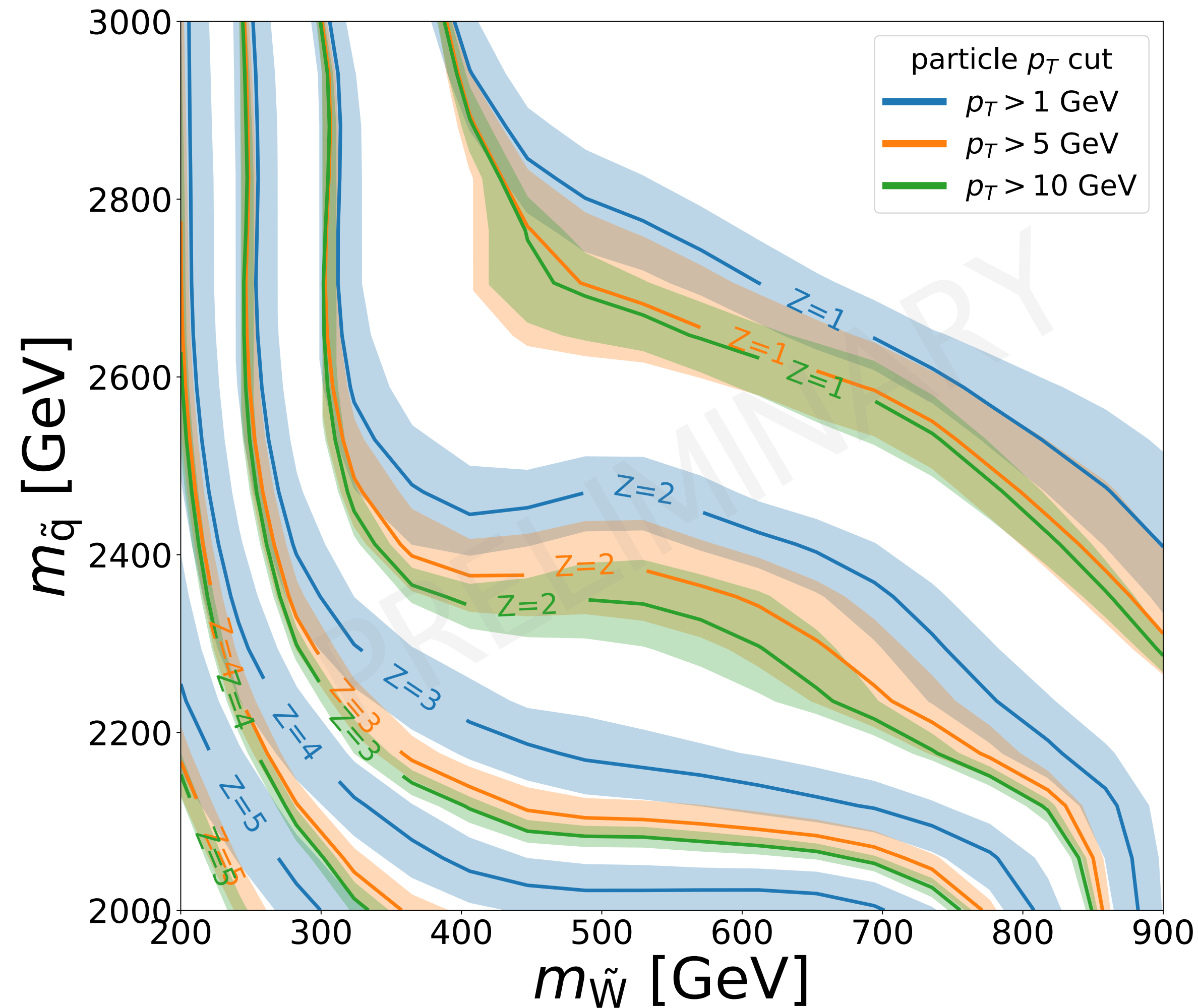


# Naive significance for winos

Run-3

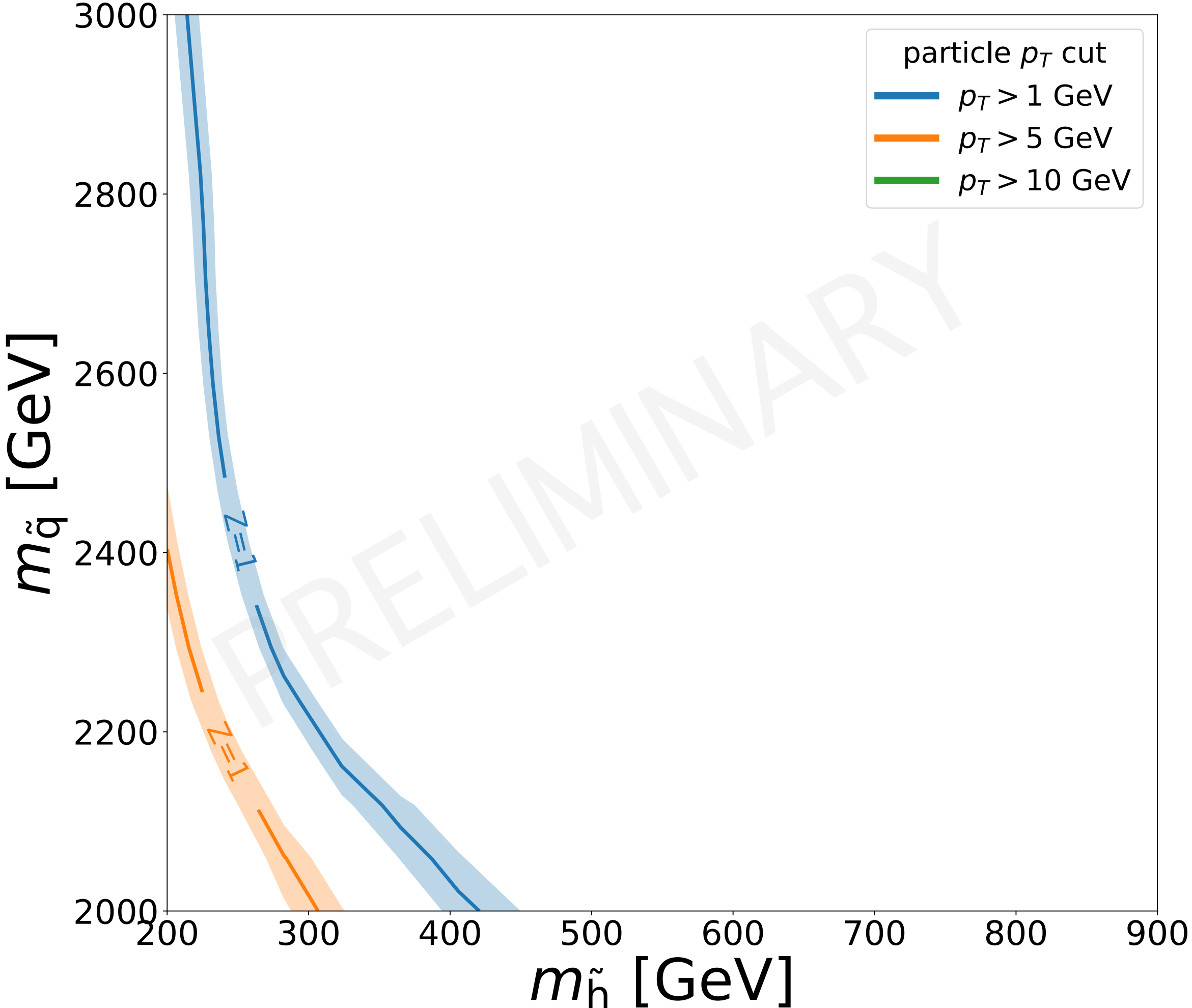


HL-LHC

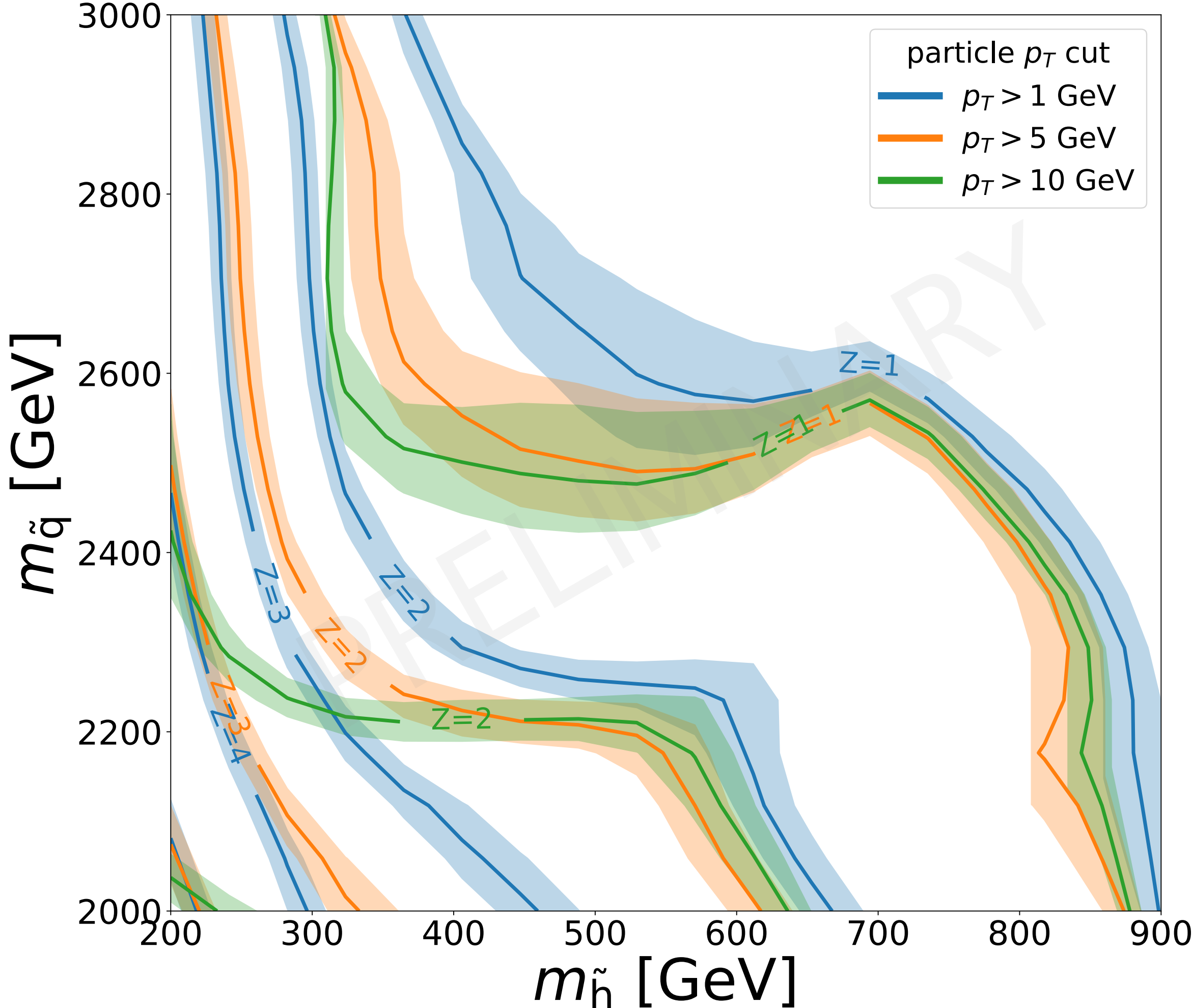


# Naive significance for higgsinos

Run-3

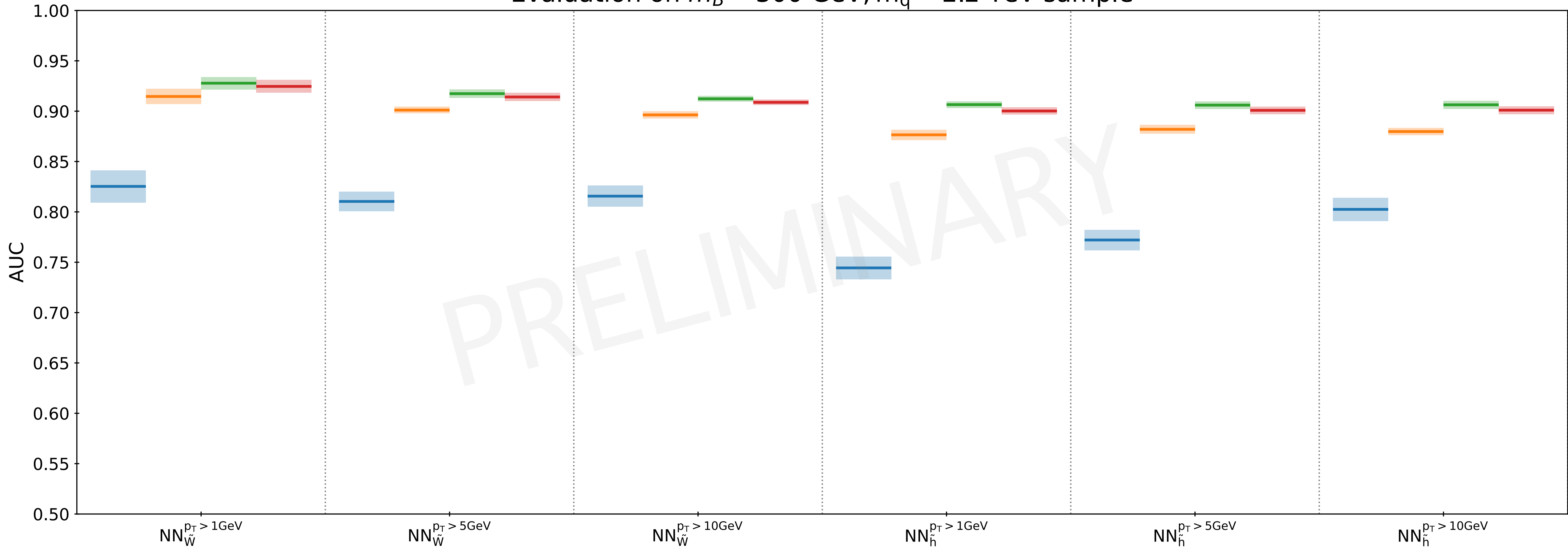


HL-LHC



# cross-evaluation

Evaluation on  $m_{\tilde{B}} = 300$  GeV,  $m_{\tilde{q}} = 2.2$  TeV sample





# Upcoming paper

- new data set
- new architecture
- evaluation
  - ROC curves
  - AUC
  - comparison with BDT
  - impact of cut on particles' pT
  - cross-evaluation
- limits
  - wino
  - higgsino
  - bino
- interpretation
  - input and output correlations
  - feature importance
  - jet characteristics



# Summary

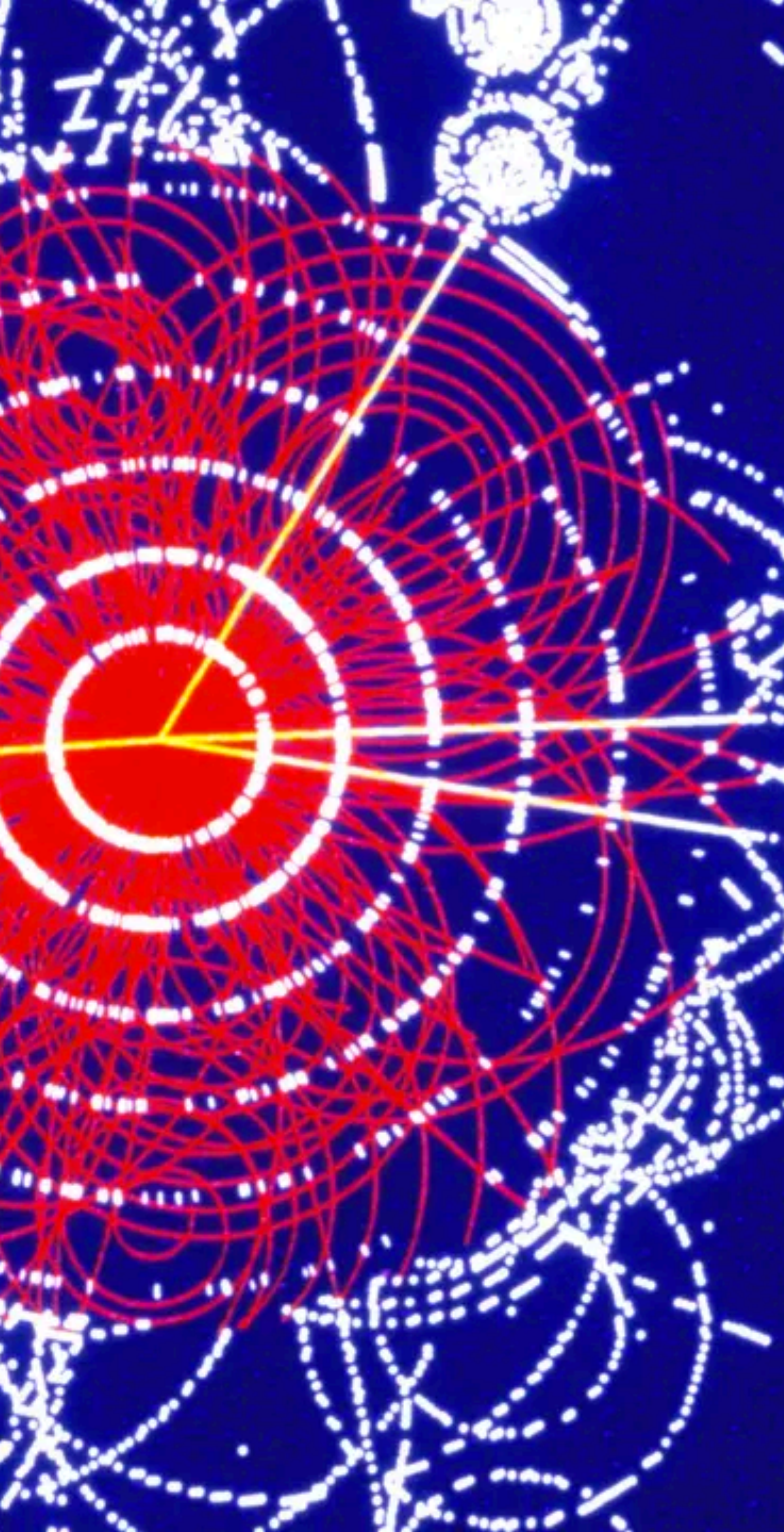
- ⊗ Dark Matter can be searched @ LHC
- ⊗ We introduce new analysis based on GNN
- ⊗ SUSY as benchmark model:
  - ⊗ EWKino pair production
  - ⊗ EWKino-squark associated production
  - ⊗ squark-pair production
- ⊗ We evaluate our model:
  - ⊗ sample composition is crucial: the more squarks the easier classification
  - ⊗ high robustness against a change of the EWKino type
  - ⊗ high robustness against change of the masses of sparticles
- ⊗ We derive the limits:
  - ⊗ For Run-3 LHC, there is always  $Z < 2$
  - ⊗ For HL-LHC,  $Z=5$  for light sparticles, and  $Z=2$  even for  $m_{\tilde{\chi}} = 900$  GeV
  - ⊗ Limits on higgsinos are a little weaker
- ⊗ Paper should be on arXiv soon.





Thank you for attention!

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# Backup slides

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DSU2024 12-09-2024

# AUC values

Table 4: Area Under Curve for ensemble of NN models trained and evaluated on  $m_{\tilde{W}} = 300$  GeV,  $m_{\tilde{q}} = 2.2$  TeV, and  $m_{\tilde{g}} = 10$  TeV.

Signal class	mean AUC	standard deviation
0 $\tilde{q}$	0.6844	0.0215
1 $\tilde{q}$	0.8494	0.0173
2 $\tilde{q}$	0.9299	0.0058
combined	0.7646	0.0156

Table 5: Area Under Curve for ensemble of NN models trained and evaluated on  $m_{\tilde{h}} = 300$  GeV and  $m_{\tilde{q}} = 2.2$  TeV.

Signal class	mean AUC	standard deviation
0 $\tilde{q}$	0.8018	0.0038
1 $\tilde{q}$	0.9146	0.0057
2 $\tilde{q}$	0.9157	0.0028
combined	0.8410	0.0019