# Bayesian Optimisation for Bayesian Evidence (BOBE)

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#### Broad Strokes: What is the problem?



Parameter	Plik best fit	Plik[1]	CamSpec [2]	$([2] - [1])/\sigma_1$	Combined
Ω <sub>b</sub> h <sup>2</sup>	0.022383	0.02237 ± 0.00015	$0.02229 \pm 0.00015$	-0.5	0.02233 ± 0.00015
$\Omega_{c}h^{2}$	0.12011	$0.1200 \pm 0.0012$	$0.1197 \pm 0.0012$	-0.3	$0.1198 \pm 0.0012$
1000mc	1.040909	$1.04092 \pm 0.00031$	$1.04087 \pm 0.00031$	-0.2	$1.04089 \pm 0.00031$
τ	0.0543	$0.0544 \pm 0.0073$	0.0536+0.0069	-0.1	$0.0540 \pm 0.0074$
ln(10 <sup>10</sup> A <sub>2</sub> )	3.0448	$3.044 \pm 0.014$	$3.041 \pm 0.015$	-0.3	$3.043 \pm 0.014$
n <sub>s</sub>	0.96605	$0.9649 \pm 0.0042$	$0.9656 \pm 0.0042$	+0.2	$0.9652 \pm 0.0042$
Ω <i>h</i> <sup>2</sup>	0.14314	$0.1430 \pm 0.0011$	$0.1426 \pm 0.0011$	-0.3	$0.1428 \pm 0.0011$
$H_0$ [ km s <sup>-1</sup> Mpc <sup>-1</sup> ]	67.32	67.36 ± 0.54	$67.39 \pm 0.54$	+0.1	$67.37 \pm 0.54$
Ω <sub>m</sub>	0.3158	$0.3153 \pm 0.0073$	$0.3142 \pm 0.0074$	-0.2	$0.3147 \pm 0.0074$
Age [Gvr]	13,7971	13.797 ± 0.023	$13.805 \pm 0.023$	+0.4	$13.801 \pm 0.024$
σ <sub>8</sub>	0.8120	$0.8111 \pm 0.0060$	$0.8091 \pm 0.0060$	-0.3	$0.8101 \pm 0.0061$
$S_8 \equiv \sigma_8 (\Omega_m / 0.3)^{0.5}$	0.8331	$0.832 \pm 0.013$	$0.828 \pm 0.013$	-0.3	$0.830 \pm 0.013$
Z	7.68	$7.67 \pm 0.73$	$7.61 \pm 0.75$	-0.1	$7.64 \pm 0.74$
1009.	1.041085	$1.04110 \pm 0.00031$	$1.04106 \pm 0.00031$	-0.1	$1.04108 \pm 0.00031$
r <sub>drag</sub> [Mpc]	147.049	$147.09 \pm 0.26$	$147.26 \pm 0.28$	+0.6	$147.18\pm0.29$

[NASA / WMAP Science Team, Planck2018 results: VI. Cosmological parameters]

#### How to Infer a Parameter

- $\bullet\,$  Given a Cosmological Model  ${\cal M}$  with parameters  $\theta$ 
  - Standard ACDM ( $n_s$ ,  $t_0$ ,  $\tau$ ,  $H_0$ ,  $\Omega_b h^2$ ,  $\Omega_c h^2$ ,  $\Omega_m h^2$ )
  - ACDM with primordial features (ACDM + Amplitude, Frequency, Phase)
- Use Boltzmann Code (such as CLASS or CAMB)
- Get Theoretical Prediction



[ESA and the Planck Collaboration]

• Given some observational data  $\mathcal D$  we can calculate a likelihood  $\mathcal L(\mathcal D|\theta,\mathcal M)$ 

• This is the probability of the data given our model  ${\mathcal M}$  and specific values of parameters  $\theta$ 

• Bayes Theorem:

$$P(B|A) = rac{P(A|B)P(B)}{P(A)}$$
 $Posterior = rac{Prior imes Likelihood}{Evidence}$ 

With our Likelihood *L* and a prior π(θ|*M*) we can calculate a posterior probability *P*(θ|*D*, *M*)

 This allows us to update our likelihood and posterior probability based on new data

#### The circle of parameter inference



#### What do we get out of this?

• We will eventually reach a 'termination criterion'

 This should indicate we are confident in our predictions (to some level of accuracy)



#### The Good the Bad and the Ugly of MCMC

Good:

- Easy to implement
- Easy to parallelise
- Very litte extra computation required
- Scales mildly with number of dimensions
- Works great for most of cosmology (near-Gaussian posteriors)

Bad:

- Not very good at finding the maximum
- Requires a lot of function evaluations  $(\mathcal{O}(10^4)$  for  $N=\mathcal{O}(10))$
- Ignores most of the information collected

Ugly:

• Struggles with not-nice posteriors (multi-modal, non-Gaussian, etc..)

• If the likelihood is not easy to evaluate or not a "nice" shape MCMC doesn't work as well

 In these cases, MCMC's random sampling methodology can be problematic

• Can we do better by not ignoring the information from previous samples?

• Our goal is to design a more efficient method that learns the shape of the posterior

• We want to take advantage of what we already know by deterministically selecting our next "sample"

• Bayesian Optimisation may present a possible solution

Fundamentally consists of two steps:

Regression - Guess the shape of the function based on the data

Next Step Selection - Decide at which point to evaluate the next function value

#### 1) Gaussian Process Regression (GPR)

• A Gaussian Process is specified by a Kernel:

 $\mathcal{N}(\mu(x), \mathcal{K}(x, x'))$ 

Where K(x, x') is the covariance function and  $\mu(x)$  is the mean of the  $\mathcal{N}$ .

• A simple example covariance function:

$$K(x, x') = Ae^{\left[\frac{-(x-x')^2}{L}\right]}$$

- Hyperparameters (h):
  - Covariance of the data: K(x, x') (relationship between parameters)
  - Prior Width: A (Certainty in the prediction)
  - Correlation Width: L (How much structure we expect in a given distance)

### 1) GPR Hyperparameters



- Hyperparameters:
  - Prior Width: A (Certainty in the prediction)
  - Correlation Width: L (How much structure we expect in a given distance)

#### 1) GPR Linear Algebra Machine



- $E(h, \underline{x}|y)$  is the probability of the model given the data
- Maximising E as a function of hyperparameters allows us to let the data decide the most appropriate GP!



• We are replacing an expensive/complex posterior with a cheap and easy GPR interpolation

### Regression - Guess the shape of the function based on the data

Next Step Selection - Decide at which point to evaluate the next function value

- GPR gives us a guess based on data
  - This helps quantify uncertainty, but we want to calculate the next sample in a smart way not randomly
- We want to recreate how an informed human agent would fit a function
- Define some function dependant on GPR mean and uncertainty that optimises a metric
  - i.e Expected Improvement (Largest Value)

#### What is Bayesian Optimisation?

- Bayesian Optimisation is a decision making framework
- The range of possible acquisition functions makes Bayesian Optimisation very adaptable



Advantages:

- Efficiency
- Good at finding global maxima
- Good at determing shape of posterior
- Works for not-nice functions
- Does not require fine tuning of settings by user

Disadvantages:

- Falls into the trap of dimensionality
- Significant extra computation

• This computational overhead means there is a threshold in terms of complexity and cost of posterior evaluations

• Above this threshold, Bayesian Optimisation will peform better than MCMC (Limit:  $\cos t \to \infty$ )

- To do parameter inference, we must choose a model
- How do we choose this model?
- How can we compare different models?
- Bayesian Optimisation may also provide a solution to this!

#### Model Selection with Bayesian Evidence

• Bayes Theorem:

$$\mathcal{P}(\mathcal{M}|\mathcal{D}) = rac{\mathcal{L}(\mathcal{D}|\mathcal{M})\pi( heta|\mathcal{M})}{P(\mathcal{D})}$$

• Probability of  $\mathcal{M}$  given  $\mathcal{D}$ :

$$P(\mathcal{M}|\mathcal{D}) = rac{P(\mathcal{D}|\mathcal{M})P(\mathcal{M})}{P(\mathcal{D})}$$

• Bayesian Evidence (Bayes Factor):

$$B = \mathcal{P}(\mathcal{D}|\mathcal{M}) = \int d heta \mathcal{L}(\mathcal{D}| heta,\mathcal{M}) \pi( heta|\mathcal{M})$$

• To compare two different models  $\mathcal{M}_1$  and  $\mathcal{M}_2$ :

$$B_{12} = \frac{P(\mathcal{D}|\mathcal{M}_1)}{P(\mathcal{D}|\mathcal{M}_2)}$$

•  $\mathcal{M}_1$  is  $\mathcal{B}_{12}$  times more probable than  $\mathcal{M}_2$ 

#### Model Selection with Bayesian Evidence

• Bayesian Evidence:

$$P(\mathcal{D}|\mathcal{M}) = \int d heta \mathcal{L}(\mathcal{D}| heta,\mathcal{M}) \pi( heta|\mathcal{M})$$

• This is the integral over our posterior prediction  $\mathcal{P}(\mathcal{M}|\mathcal{D})$ 

 Rewards models with accurate 'risky' predictions over generic ones - Occam's Razor

- The aim of this optimisation is to find the Bayesian Evidence to some level of precision.
- The evidence is a relative measure of how well a model fits the data.
- That is to say, we can use the evidence to compare how well two models fit the data.

#### Bayesian Optimisation for Bayesian Evidence

• Integration over mutli-dimensional posteriors is still hard

- Typical methods require  $O(10^5 10^6)$  evaluations for  $\Lambda CDM$  with primordial features
  - Nested Sampling [Skilling 2004, Feroz et al. 2013, Handley et al. 2015]

• We can use the same methodology as parameter inference - replace the actual posterior with the GPR interpolation

#### Optimising for the Evidence

- We want our acquisition function 'metric' to be reducing the uncertainty in the integral over parameter space
- We use the Weighted Negative Integrated Posterior Variance

$$WNIPV( heta) = \int d heta^{'} \sigma \, \widehat{GP( heta)}( heta^{'})$$

- $\widehat{GP(\theta)}$  posterior if we pretend to take a sample at a new point  $\theta$
- $\sigma$  GP Uncertainty

#### Bayesian Optimisation (iteration 1)



#### Bayesian Optimisation (iteration 2)



#### Bayesian Optimisation (iteration 3)



#### Bayesian Optimisation (iteration 4)



#### Bayesian Optimisation (iteration 5)



#### Bayesian Optimisation (iteration 6)



- Through each iteration the overal uncertainty of the GP goes down
- The maximum value of the acquisition function also reduces
- This allows us to define a threshold or termination criteria on the precision of the evidence

- We use some smart priors to help deal with higher dimensional parameter spaces (Sparse Axis-Aligned Subspaces) [Eriksson & Jankowiak (2021)]
  - All dimensions are innocent until proven guilty
- We use Nested Sampling to get both a direct numerical estimate of the uncertainty on the evidence as well as uncorrelated samples of the posterior
- We sample from hyperparameter space PDF with NUTS (modified HMC) instead of optimising over it

#### Preliminary Results - Annular Function

- Efficient ML algorithm for model selection and parameter inference
  - Preliminary improvement of 100x fewer samples

• Best for difficult to obtain, expensive to calculate and complicated likelihoods

• Benchmark is to take fewer samples than other methods with the same precision on evidence

• Paper and code coming soon!

#### Extra! Extra! Read all about it!

#### Metropolis-Hastings MCMC



[Metropolis et al. (1953)]

#### **BOBE** Algorithm



## Example Application: ACDM with modulated Primordial Fluctuations [Jan Hamann & Julius Wons, 2021]



- Using Nested Sampling :  $\mathcal{O}(10^5)$  samples
- Our results with BO: 2 orders of magnitude improvement in # evaluations

## Example Application: ACDM with modulated Primordial Fluctuations [Jan Hamann & Julius Wons, 2021]



- Also learns the global shape of the function!
- 1700 samples, 8 frequency bins