Starobinsky Inflation and the Swampland

Joaquin Masias

Workshop on Quantum Gravity, Strings and the Swampland, Corfu, September 4th, 2024

Based on [2312.13210, D. Lüst, M. Scalisi, B. Muntz, **JM**]

MAX-PLANCK-INS

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- The Swampland Program aims to determine general features of theories consistent with quantum gravity.

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 $p\rightarrow\infty$ String excitations $p\simeq\mathcal{O}(1)\;$ KK modes $\Lambda_{\rm sp} \simeq M_{\rm Pl, d+p}$ $\Lambda_{\rm sp} \simeq M_s$ d+p dim Planck mass string scale

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• Hard to motivate from string theory, even harder to construct

Cosmological Argument Starobinsky Inflation and the Swampland

• The transformation between frames acts as

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• This acts as a correction to the momentum dependence of the tree level graviton propagator.

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Perturbative Argument Starobinsky Inflation and the Swampland

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\gamma = \frac{\Lambda_{\rm s}^{\prime}}{\Lambda_{\rm s}} \qquad \qquad \frac{1}{\sqrt{6}} \leq \Big|\frac{\Lambda_{\rm s}^{\prime}}{\Lambda_{\rm s}}\Big| \leq \frac{1}{\sqrt{2}}
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S_{\text{het}} \sim \frac{M_s^8}{g_s^2} \mathcal{V}_s \ell_s^6 \int d^4x \sqrt{-g_4^{(E)}} \left[R_4 + \alpha' R_4^2 + \cdots \right] .
$$

$$
S_{\text{IIB}} \sim M_{P,4}^2 \int d^4x \sqrt{-g_4^{(E)}} \left[R_4 + g_s^{-2} \mathcal{V}_s^{1/3} \frac{R_4^2}{M_{P,4}^2} + \cdots \right]
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\nType IIB: $M \simeq g_s \frac{M_P}{\mathcal{V}_s^{1/6}}$ KK tower: $\begin{cases} \Lambda_s \simeq M_{P,10} \simeq \frac{M_P}{\sqrt{\tau_2}} \\ N \simeq \tau_2 \simeq \mathcal{V}^{\frac{1}{3}} \end{cases}$

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• Not protected from higher curvature corrections!

- The Starobinsky model of inflation can be interpreted as a QG correction to EH gravity.
- In particular, it can be generated by the renormalization effects of a tower of light species.
- Top-down arguments can be used to identify $M \simeq \Lambda_s$.
- Identifying $M \simeq \Lambda_s$, leads to $\Lambda_s \simeq 10^{14}$ GeV, $N_{sp} \simeq 10^{10}$.
- Starobinsky Inflation is spoiled by an exponential scaling $|\gamma| \gtrsim \mathcal{O}(10^{-3})$.