

Spontaneous Symmetry Breaking in the Bulk as a Localization Mechanism on the Brane

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Orfeu Bertolami and Carla Carvalho, arXiv:0705.1923 [hep-th]

Motivation (... Why not?)

Motivated by the dimensional asymmetry characteristic of braneworlds:

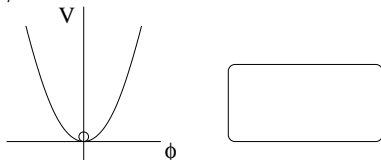
- Populate the bulk with matter fields and couple them non-minimally to gravity.
- Break Lorentz symmetry in the bulk and study consequences in the brane:
 - Reverberate [[G.D. Starkman, D. Stojkovic, M. Trodden, Phys.Rev.Lett.87, 231303 \(2001\)](#)]: Must have a bulk with the same symmetries as those of the brane, since perturbations in the bulk propagate into perturbations in the brane.
 - Phenomenology still to be worked out...
- Use spontaneous symmetry breaking to relate the mass on the brane with a bulk mechanism to look for signatures of extra dimensions. A localization mechanism on the brane?

Spontaneous symmetry breaking (or a clever, little man's view)

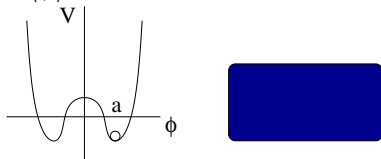
Symmetry of the system (Lagrangian) is not shared by the ground state (vacuum) solution: degenerate vacua.

For a scalar field ϕ , let $V(\phi) = \mu^2(\phi^2/2) + \lambda(\phi^4/4)$.

If $\mu^2 > 0$, then $\langle \phi \rangle = 0$.

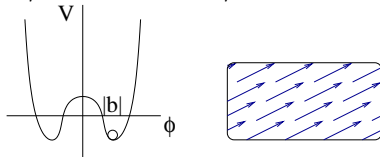


If $\mu^2 < 0$, then $V(\phi) = (\lambda/4)(\phi^2 - a^2)^2 - (\lambda/4)a^4$, with $a^2 = -(\mu^2/\lambda)$ and $\langle \phi \rangle = \pm a$.



Lorentz symmetry breaking

For a vector field B_μ , let $V(\mathbf{B}) = B_\mu B^\mu - b^2$. Then $\langle B_\mu \rangle = \pm b_\mu$.



The vacuum contains an intrinsic direction defined by $\langle B_\mu \rangle$, thus violating rotation invariance and consequently violating Lorentz invariance.

[O. Bertolami, C.C., Phys.Rev.D74, 084020 (2006)]

[O. Bertolami, C.C., gr-qc/0612129]

Bulk scalar field with non-minimal coupling

Consider a bulk scalar field Φ non-minimally coupled to gravity

$$\mathcal{L} = \frac{R}{\kappa_{(5)}^2} - 2\Lambda_{(5)} + \xi\Phi^2 R - \frac{g^{\mu\nu}}{2}(\nabla_\mu\Phi)(\nabla_\nu\Phi) - V(\phi^2),$$

where $\kappa_{(5)}^2 = 8\pi G_N = 1/M_P^3$ is the 5-dimensional gravitational coupling constant, $\Lambda = \Lambda_{(5)} + \Lambda_{(4)}$ is the 5-dimensional cosmological constant with $\Lambda_{(4)} = \sigma\delta(N)$.

Why the non-minimal coupling: the simplest interaction with a canonical kinetical term, reduces to Brans-Dicke up to a field transformation for a vanishing vev.

Bulk equations of motion

In the gravitational sector,

$$\left(\frac{1}{\kappa_{(5)}^2} + \xi \Phi^2 \right) G_{\mu\nu} + \Lambda g_{\mu\nu} - \xi \Sigma_{\mu\nu}^{(\Phi)} = \frac{1}{2} T_{\mu\nu}^{(\Phi)},$$

with

$$\Sigma_{\mu\nu}^{(\Phi)} = \nabla_{\mu} \nabla_{\nu} \Phi^2 - g_{\mu\nu} g^{\alpha\beta} \nabla_{\alpha} \nabla_{\beta} \Phi^2,$$

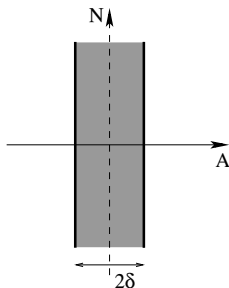
$$T_{\mu\nu}^{(\Phi)} = (\nabla_{\mu} \Phi)(\nabla_{\nu} \Phi) + g_{\mu\nu} \left[-\frac{1}{2} g^{\alpha\beta} (\nabla_{\alpha} \Phi)(\nabla_{\beta} \Phi) - V(\Phi^2) \right].$$

In the scalar field sector,

$$g^{\mu\nu} \nabla_{\mu} \nabla_{\nu} \Phi - \frac{\partial V}{\partial \Phi} + 2\xi \Phi R = 0.$$

Boundary conditions

We regard the brane as a Z_2 -symmetric shell of thickness 2δ in the limit $\delta \rightarrow 0$.



- The Z_2 -symmetry establishes the continuity conditions of the fields across the brane,
- Derivatives of quantities discontinuous across the brane generate singular distributions on the brane. Integration of these terms in N relates the induced geometry with the localization of the induced stress-energy.

Junction conditions

The junction condition from the ϕ equation:

$$\int dN(\Phi_{\text{eqn}}) \Rightarrow \nabla_N \Phi - 4\xi K \Phi = 0.$$

The junction condition from the (AB) Einstein equation:

$$\left(\frac{1}{\kappa_{(5)}^2} + \xi \phi^2 \right) (-K_{AB} + g_{AB} K) = -g_{AB} (\sigma + \xi \nabla_N \phi^2).$$

Two separable conditions on the boundary ($N = 0$):

$$\nabla_N \Phi = -4\xi \Phi \sigma \frac{d/(d-1)}{\lambda/\kappa_{(5)}^2 + \xi \Phi^2 (1 + 8\xi/(d-1))},$$

$$K_{AB} = -g_{AB} \sigma \frac{1/(d-1)}{\lambda/\kappa_{(5)}^2 + \xi \Phi^2 (1 + 8\xi/(d-1))}.$$

Induced equations on the brane (α')

We obtain for the propagation of ϕ on the brane that

$$g^{AB}\nabla_A\nabla_B\phi - \frac{\partial V}{\partial\phi} + 2\xi\phi \left[R^{(ind)} - K_{AB}K_{BA} + (1 + 8\xi) K^2 \right] = 0.$$

From the (NN) Einstein equation we extract the induced intrinsic curvature $R^{(ind)}$

$$\begin{aligned} & \left(\frac{1}{\kappa_{(5)}^2} + \xi\phi^2(1 + 4\xi) \right) R^{(ind)} = \\ & = \left(\frac{1}{4} + 2\xi \right) (\nabla_C\phi)^2 + \frac{1}{2}V + 2\xi\phi \frac{\partial V}{\partial\phi} + \Lambda_{(5)} \\ & - K_{CD}K_{CD} \left(\frac{1}{\kappa_{(5)}^2} + \xi\phi^2(1 - 4\xi) \right) + K^2 \left(\frac{1}{\kappa_{(5)}^2} + \xi\phi^2(1 - 32\xi^2) \right), \end{aligned}$$

Induced equations on the brane (β')

from the (AN) Einstein equation we find the conservation of the induced stress-energy tensor \mathcal{T}_{AB}

$$\begin{aligned} G_{AN} &= K_{AB;B} - K_{;A} \\ &= -\nabla_B \left(\int_{-\delta}^{+\delta} dN G_{AB} \right) = -\kappa_{(5)}^2 \nabla_B \mathcal{T}_{AB} = 0, \end{aligned}$$

and from the (AB) Einstein equation

$$\begin{aligned} G_{AB}^{(ind)} &= \left(\frac{1}{\kappa_{(5)}^2} + \xi \phi^2 \right)^{-1} \left[\left(\frac{1}{2} + 2\xi \right) (\nabla_A \phi)(\nabla_B \phi) + 2\xi \phi \nabla_A \nabla_B \phi \right] \\ &- g_{AB} \left[R^{(ind)} + K^2 \frac{-d^2 + d + 4}{2d^2} \right]. \end{aligned}$$

Effective cosmological constant Λ_{eff}

Read off of the induced Einstein equation the Λ_{eff} , comprising all the terms proportional to g_{AB} when all the contributions from the matter fields are evaluated at the vev's.

- For $\langle \phi \rangle = 0$,

$$\Lambda_{\text{eff}} \frac{1}{\kappa_{(5)}^2} = \Lambda_{(5)} + K^2 \frac{1}{\kappa_{(5)}^2} \frac{d^2 - d + 4}{2d^2};$$

- When $\langle \phi \rangle \neq 0$, with $\nabla_A \langle \phi \rangle = 0$,

$$\begin{aligned} \Lambda_{\text{eff}} \left(\frac{1}{\kappa_{(5)}^2} + \xi \langle \phi \rangle^2 (1 + 4\xi) \right) &= \Lambda_{(5)} + K^2 \frac{1}{\kappa_{(5)}^2} \frac{d^2 - d + 4}{2d^2} \\ + \frac{1}{2} V(\langle \phi \rangle^2) + 2\xi \langle \phi \rangle \left. \frac{\partial V}{\partial \phi} \right|_{\langle \phi \rangle} \\ + \xi \langle \phi \rangle^2 K^2 \left[\frac{d^2 - d + 4}{2d^2} + 4\xi \left(\frac{-d^2 + 3d + 4}{2d^2} - 8\xi \right) \right]. \end{aligned}$$

Effective potential

The evolution of ϕ on the brane

$$g^{AB}\nabla_A\nabla_B\phi - \frac{\partial V_{\text{eff}}}{\partial\phi} + \left(\frac{1}{4} + 2\xi\right) (\nabla_C\phi)^2 \frac{2\xi\phi}{1/\kappa_{(5)}^2 + \xi\phi^2(1 + 4\xi)} = 0$$

subject to the effective potential induced on the brane

$$\begin{aligned} -\frac{\partial V_{\text{eff}}}{\partial\phi} &= \frac{1}{1/\kappa_{(5)}^2 + \xi\phi^2(1 + 4\xi)} \left\{ -\frac{\partial V}{\partial\phi} \left(\frac{1}{\kappa_{(5)}^2} + \xi\phi^2 \right) \right. \\ &+ 2\xi\phi \left[\frac{1}{2} V + \Lambda_{(5)} \right] \\ &\left. + 2\xi\phi K^2 \left[\frac{1}{\kappa_{(5)}^2} \left(-\frac{2}{d} + 2 + 8\xi \right) + \xi\phi^2 \left(-\frac{2}{d} + 2 + 12\xi \right) \right] \right\} \end{aligned}$$

where $V(\phi^2)$ is the bulk potential, assumed to have a Higgs type form $V(\phi) = \mu_{(5)}^2(\phi^2/2) + \lambda_{(5)}(\phi^4/4)$ with $\lambda_{(5)} > 0$.

Conditions of the parameters

Integrate to find V_{eff} . Expect that there exists a hierarchy of scales depending on whether $\langle\phi\rangle$ is related to the Standard Model or the grand unified theory scale: $|\xi\phi^2| \ll 1/\kappa_{(5)}^2$. Then

$$V_{\text{eff}} = \mu_{\text{eff}}^2(\phi^2/2) + \lambda_{\text{eff}}(\phi^4/4) + O[\phi^6\sigma^2\kappa_{(5)}^2(\xi\kappa_{(5)}^2)^3],$$

up to sub-dominant logarithmic terms in ξ , where

$$\begin{aligned}\mu_{\text{eff}}^2 &\sim \mu_{(5)}^2 + \lambda_{(5)} \frac{1}{\xi\kappa_{(5)}^2} - \sigma^2\xi\kappa_{(5)}^4, \\ \lambda_{\text{eff}} &\sim \lambda_{(5)} + \sigma^2\xi^2\kappa_{(5)}^6.\end{aligned}$$

The parameters of the potential will influence the magnitude of its minimum and thus the mass of ϕ measured on the brane.

$$\begin{aligned}\lambda_{\text{eff}} > 0 &\Rightarrow \lambda_{(5)} > -\sigma^2\xi^2\kappa_{(5)}^6 \\ \mu_{\text{eff}}^2 < 0 &\Rightarrow \mu_{(5)}^2 < -2\sigma^2\xi\kappa_{(5)}^4.\end{aligned}$$

A note on the dimensionality of ϕ

The bulk scalar field ϕ , being a five-dimensional field, has dimensions $[\phi] = M^{3/2}$. Accordingly, $[\mu_{(5)}] = M$ and $[\lambda_{(5)}] = M^{-1}$. In order to recover characteristically four-dimensional quantities, we define the four-dimensional scalar field Φ as the rescaling of ϕ by an appropriate mass scale M_ϕ .

[An observation: In the mode expansion of a bulk field, this mass can be identified with the mode function dependent on the direction N evaluated at the position of the brane in the bulk.]

For $\phi = M_\phi^{1/2} \Phi$, the induced equation of motion for Φ on the brane becomes

$$\square \Phi - \frac{1}{M_\phi} \frac{\partial V_{\text{eff}}}{\partial \Phi} + \left(\frac{1}{4} + 2\xi \right) (\nabla_C \Phi)^2 \frac{2\xi M_\phi \Phi}{1/\kappa_{(5)}^2 + \xi M_\phi \Phi^2 (1 + 4\xi)} = 0.$$

The effective scalar field Φ

The parameters of the effective potential will scale as

$$\frac{1}{M_\phi} V_{\text{eff}}(\Phi^2) = \mu_{\text{eff}}^2 \Phi^2 + \lambda_{\text{eff}} M_\phi \Phi^4 + M_\phi^2 \mathcal{O}[\Phi^6]$$

with

$$\begin{aligned}\mu_{\text{eff}}^2 &\sim \mu^2 - 2\sigma^2 \xi \kappa_{(5)}^4, \\ M_\phi \lambda_{\text{eff}} &\sim \lambda + M_\phi \xi^2 \sigma^2 \kappa_{(5)}^6,\end{aligned}$$

where $\mu = \mu_{(5)}$ and $\lambda = M_\phi \lambda_{(5)}$. For $\xi > 0$ we have two possible mechanisms for the generation of a non-vanishing vev:

- the canonical way, via the potential associated with the scalar field,
- the braneworld way, via the interaction of the scalar field with the brane tension.

Constraints from the standard model: back to reality

- For the latter to be viable in the context of the SM, then $\langle \Phi \rangle = 246$ GeV, as constrained by the interaction with the boson vectors, and

$$\left| \frac{\mu_{\text{eff}}^2}{M_\phi \lambda_{\text{eff}}} \right| \sim \frac{1}{\xi M_\phi} \frac{1}{\kappa_{(5)}^2}$$

must be of order TeV^2 .

- In order to recover the four-dimensional gravitational coupling constant in the induced Einstein equation, we find from the ϕ contribution that $M_{Pl(4)}^{-2} = \kappa_{(5)}^2 M_\phi$. It follows that

$$M_{Pl}^3 \equiv \frac{1}{\kappa_{(5)}^2} \sim M_{Pl(4)}^2 M_\phi.$$

Then,

$$\left| \frac{\mu_{\text{eff}}^2}{M_\phi \lambda_{\text{eff}}} \right| \sim \frac{1}{\xi} M_{Pl(4)}^2 \gg \text{TeV}^2.$$

Since $\langle \Phi^2 \rangle \sim \left| \frac{\mu_{\text{eff}}^2}{M_\phi \lambda_{\text{eff}}} \right|$, then $\langle \Phi \rangle \sim M_{Pl(4)}$. The brane mediated mechanism of SSB is rendered unviable by the phenomenological hierarchy between the SM typical energy scale of TeV and the Planck scale of the induced dynamics of Φ on the brane. Moreover, for $\Lambda_{\text{eff}} = 0$, then

$$\sigma^2 \sim M_{Pl(4)}^2 M_\phi \left[-\Lambda_{(5)} - M_\phi V(\langle \Phi \rangle^2) \right].$$

- $|\mu_{\text{eff}}^2|$ dominates: the scalar field becomes a short range field about the brane $|\mu_{\text{eff}}| \sim M_{Pl(4)}$ and therefore strongly localized therein $\delta \sim 1/|\mu_{\text{eff}}|$; then $M_\phi \lambda_{\text{eff}} \sim \xi$ and for
 - $M_\phi \sim M_{Pl(4)}$, then $\sigma \sim (10^{16})^4 TeV^4$ and $\Lambda_5 \sim (10^{16})^5 TeV^5$;
 - $M_\phi \sim TeV$, then $\sigma \sim (10^{16})^3 TeV^4$ and $\Lambda_5 \sim (10^{16})^4 TeV^5$.
- $1/M_\phi \lambda_{\text{eff}}$ dominates: for $|\mu_{\text{eff}}| \sim TeV$ and
 - $M_\phi \sim TeV$, then $\sigma \sim (10^{16})^2 TeV^4$ and $\Lambda_5 \sim (10^{16})^2 TeV^5$;
 - $M_\phi \sim 10^{-16} TeV$, then $\sigma \sim 10^{16} TeV^4$ and $\Lambda_5 \sim 10^{16} TeV^5$;

Conclusions

- Populated the bulk a real scalar field, non-minimally coupled to gravity.
- Upon spontaneous symmetry breaking, found the induced mass of the scalar field to be of order $M_{Pl(4)}$, thus strongly localized about the locus of the brane.
- Localization is achieved but at a scale way off the standard model scale, hence no sign of a spontaneous symmetry breaking mechanism in the bulk would be manifest in the induced masses on the brane.
- For $\xi = 0$, $V_{eff} = V + 3\sigma^2\kappa_{(5)}^2 d^2 / [(d + 1)(d - 1)]$:
 σ contributes to V_{eff} but no longer to the mechanism of spontaneous symmetry breaking or to the mixing of the discontinuity in \mathbf{K} with that in the normal derivative of the bulk field.
- Matter localized on the brane seems to interact with bulk matter fields through gravity only when a non-minimal coupling exists.

Bulk potential point of view:

- use finiteness of integration of the action along the normal direction:
[B. Bajc, G. Gabadadze, Phys.Lett.B474 (2000)]
[I. Oda, Phys.Lett.B571 (2000)];
- study the dynamics along the direction normal to the surface of the brane:
[K. Farakos, P. Pasipoularides, Phys.Lett.B621 (2005)];

Effective potential point of view:

- analyse the constraints from the effective mass:
[A. Kehagias, K. Tamvakis, Phys.Lett.B628 (2005)];
- consider the inverse of the mass inside the brane as a measure of the confinement to the brane:
[M. Laine, H.B. Meyer, K. Rummukainen, M. Shaposhnikov, JHEP 0404 (2004)] ;