

Noncommutativity and curved momentum space

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1 Momentum space and vacuum energies





Outline

1 Momentum space and vacuum energies

Ø Momentum space and fermions

Conclusions

Noncommutativity

$$[\hat{x}_a, \hat{x}_b] = 2 i \Theta_{ab} \sim \ell_P^2$$

The idea was conceived in the '40s as a minimum distance could lead to the cure of ultraviolet divergencies (Snyder, 1946).



LIV vs. DSR

Lorentz invariance violation

- There is a loss of the relativity principle.
- There is a privileged observer \rightarrow physical laws depending on the observer.

Doubly Special Relativity (DSR)

- There is a relativity principle.
- Two fundamental constants: speed of light c and Planck's length ℓ_P .

Snyder

anti-Snyder phase-space:

$$\begin{split} & [\hat{J}_{ij}, \hat{J}_{kl}] = i \left(\delta_{ik} \hat{J}_{jl} - \delta_{il} \hat{J}_{jk} - \delta_{jk} \hat{J}_{il} + \delta_{lj} \hat{J}_{ik} \right) \\ & [\hat{J}_{ij}, \hat{\rho}_k] = i \left(\delta_{ik} \hat{\rho}_j - \delta_{jk} \hat{\rho}_i \right), \\ & [\hat{J}_{ij}, \hat{x}_k] = i \left(\delta_{ik} \hat{x}_j - \delta_{jk} \hat{x}_i \right), \\ & [\hat{x}_i, \hat{\rho}_j] = i \left(\delta_{ij} - \beta^2 \hat{\rho}_i \hat{\rho}_j \right), \\ & [\hat{x}_i, \hat{x}_j] = -i \beta^2 \hat{J}_{ij}, \qquad [\hat{\rho}_i, \hat{\rho}_j] = 0, \end{split}$$

,

Lorentz generators $(\hat{J}_{ij}=\hat{x}_i\hat{p}_j-\hat{x}_j\hat{p}_i).$

Snyder

anti-Snyder phase-space:

$$\begin{split} & [\hat{J}_{ij}, \hat{J}_{kl}] = i \left(\delta_{ik} \hat{J}_{jl} - \delta_{il} \hat{J}_{jk} - \delta_{jk} \hat{J}_{il} + \delta_{lj} \hat{J}_{ik} \right) \\ & [\hat{J}_{ij}, \hat{p}_k] = i \left(\delta_{ik} \hat{p}_j - \delta_{jk} \hat{p}_l \right), \\ & [\hat{J}_{ij}, \hat{x}_k] = i \left(\delta_{ik} \hat{x}_j - \delta_{jk} \hat{x}_l \right), \\ & [\hat{x}_i, \hat{p}_j] = i \left(\delta_{ij} - \beta^2 \hat{p}_i \hat{p}_j \right), \\ & [\hat{x}_i, \hat{x}_j] = -i \beta^2 \hat{J}_{ij}, \qquad [\hat{p}_i, \hat{p}_j] = 0, \end{split}$$

Lorentz generators $(\hat{J}_{ij} = \hat{x}_i \hat{p}_j - \hat{x}_j \hat{p}_i).$



Experiments



meaning but the difference between these sums in the two situations, $\frac{1}{2}(\Sigma h \omega)_{ll} - \frac{1}{2}(\Sigma h \omega)_{ll}$, will be shown to have a well defined value and this value will be interpreted as the interaction between the plate and the xy face.

The possible vibrations of a cavity defined by

 $0 \leqslant x \leqslant L$, $0 \leqslant y \leqslant L$, $0 \leqslant z \leqslant a$

have wave numbers

$$k_x = \frac{\pi}{L} n_x$$
, $k_y = \frac{\pi}{L} n_y$, $k_z = \frac{\pi}{a} n_z$,

$$\frac{F}{A} = \frac{1}{2} \sum_{n} \omega_n = \frac{\pi^2}{240} \frac{\hbar c}{a^4}$$
$$\sim \frac{0.016}{a^4} \frac{\mathrm{dyn}\,\mu\mathrm{m}^4}{\mathrm{cm}^2}$$



Experiments



Capasso et al (2007)



$$H_V = \hat{p}^2 + VH(\hat{x}_{\perp} - L) + VH(-\hat{x}_{\perp} - L).$$

For example, choose a realization of the algebra,

$$\hat{p}_i = rac{p_i}{\sqrt{1+eta^2p^2}}, \qquad \hat{x}_i = i\sqrt{1+eta^2p^2}rac{\partial}{\partial p_i}.$$

$$\mathcal{F} = \frac{\Omega_{D-2}}{2(2\pi)^{D-1}} \int_0^{1/\beta} \frac{d\rho \, p^{D-2}}{(1-\beta^2 \rho^2)^{D/2-1/2}} \left(\sum_{n=1}^\infty f_n(\rho) - \int_0^\infty dn \, f_n(\rho) \right).$$

$$f_n(p) = \frac{k_n}{\beta L} \frac{\tanh(\beta k_n)}{\cosh^2(\beta k_n) \sqrt{p^2 + \beta^{-2} \tanh^2(\beta k_n) (1 - \beta^2 p^2) + m^2}}, \quad k_n := \frac{n\pi}{2L}.$$
(1)



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 \Downarrow scalar field, time \times anti-Snyder_{D-1}, sum over frequencies

$$\mathcal{F} = \frac{\Omega_{D-2}}{2(2\pi)^{D-1}} \int_0^{1/\beta} \frac{dp \, p^{D-2}}{(1-\beta^2 p^2)^{D/2-1/2}} \left(\sum_{n=1}^\infty f_n(p) - \int_0^\infty dn \, f_n(p) \right),$$



Momentum space

$$f_n(p) = \frac{k_n}{\beta L} \frac{\tanh(\beta k_n)}{\cosh^2(\beta k_n)\sqrt{p^2 + \beta^{-2} \tanh^2(\beta k_n)(1 - \beta^2 p^2) + m^2}}, \quad k_n := \frac{n\pi}{2L}.$$
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- Momentum space
- Divergences

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- Momentum space
- Divergences
- Invariant under realizations

$$f_n(p) = \frac{k_n}{\beta L} \frac{\tanh(\beta k_n)}{\cosh^2(\beta k_n)\sqrt{p^2 + \beta^{-2} \tanh^2(\beta k_n)(1 - \beta^2 p^2) + m^2}}, \quad k_n := \frac{n\pi}{2L}.$$
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- Momentum space
- Divergences
- Invariant under realizations
- Spectral geometry

$$f_n(p) = \frac{k_n}{\beta L} \frac{\tanh(\beta k_n)}{\cosh^2(\beta k_n)\sqrt{p^2 + \beta^{-2} \tanh^2(\beta k_n)(1 - \beta^2 p^2) + m^2}}, \quad k_n := \frac{n\pi}{2L}.$$
(1)

For example, the vacuum energy can be expanded as

$$\mathcal{E} = \sqrt{1 + \beta^2 m^2} \left[\frac{2L}{(2\pi)^D \beta} \operatorname{Vol}(\mathbb{H}_D) + \frac{1}{4(2\pi)^{D-1}\beta} \operatorname{Vol}(\mathbb{H}_{D,w=0}) + \cdots \right].$$
(2)

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(2)

Can one hear the shape of momentum space?

Experiments

- Noncommutativity, see previous slide, $\mathcal{F} = -\frac{\pi^2}{7680} \frac{1}{L^4} \frac{\pi^4}{48384} \frac{\beta^2}{L^6} + \mathcal{O}(\beta^4).$
- First graviton corrections (Donoghue et al), $V(r) = -\frac{Gm_1m_2}{r} \left(1 \frac{G(m_1+m_2)}{rc^2} \frac{127}{30\pi^2} \frac{G\hbar}{c^3r^2}\right)$
- Dark scalar ϕ (Brax et al., 2021), $\mathcal{O}_1 = \Lambda^{-3} \bar{N} N \phi^2$, $\mathcal{O}_2 = \Lambda^{-2} \bar{N} \gamma^{\mu} N \phi^* i \partial_{\mu} \phi$, ...): UV $r^{-\alpha}$, IR $e^{-2mr} r^{-\beta/2}$
- Axion a (Klimchitskaya et al., 2021) , $-ig_a\bar{\psi}\gamma^5\psi a(x)$, $V(r) = -\frac{g_{an}^4}{32\pi^3m^2}\frac{m_a}{r^2}K_1(2m_ar)$



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Conclusions

Kinematics in DSR

• Free particles described by a (modified) dispersion relation

$$C(k) = k_0^2 - \mathbf{k}^2 + \frac{k_0^3}{\Lambda} + \dots = m^2.$$

• A (modified) conservation of total momentum holds when interaction takes place

Total momentum =
$$(p \oplus q)_{\mu} = p_{\mu} + q_{\mu} + \frac{p_{\mu}q_0}{\Lambda} + \cdots$$

Dispersion relation and conservation law compatible with relativity principle → deformed Lorentz transformations.

Kinematics in DSR

Most known, κ-Poincaré, customarily studied in the context of Hopf algebras [Majid1994]. (κ-Minkowski, [x₀, x_i] = ix_i)

Product $\nabla = [\cdot, \cdot]$, unit η , coproduct Δ , counit ε , antipode *S*.



• Particular example: symmetric basis [Lukierski:1992dt]

Deformed dispersion relation

$$C_{A}^{(5)}(P) = \left(2\Lambda \sinh\left(\frac{P_{0}}{2\Lambda}\right)\right)^{2} - P_{i}P_{i}.$$

Deformed conservation law

$$(P \oplus Q)_0 = P_0 + Q_0, \qquad (P \oplus Q)_i = P_i e^{Q_0/2\Lambda} + Q_i e^{-P_0/2\Lambda}.$$

For the usual Poincaré symmetry, one introduces fermions by

$$J_{AB} = M_{AB} \otimes I + I \otimes \Sigma_{AB}. \tag{3}$$

(-)

In the *k*-Poincaré setup, one can take into account the nontrivial coproduct (symmetric basis)

$$\Delta L_{i} = L_{i} \otimes e^{\frac{P_{0}}{2\kappa}} + e^{-\frac{P_{0}}{2\kappa}} \otimes L_{i} + \frac{1}{2\kappa} \epsilon_{ijk} \left(P_{j} \otimes M_{k} e^{\frac{P_{0}}{2\kappa}} + e^{-\frac{P_{0}}{2\kappa}} M_{k} \otimes P_{j} \right),$$
(4)

so that replacing the generators on the right side of the tensor products by a finite dimensional representation one gets [Nowicki:1992if]

$$\mathcal{L}_{i} = L_{i} + e^{\frac{-P_{0}}{2\kappa}} l_{i} - \frac{1}{2\kappa} \epsilon_{ijk} m_{j} P_{k}.$$
(5)

The Dirac operator is

$$\mathcal{D}_{\mathsf{Hopf}} := \gamma^0 \left(\Lambda \sinh\left(rac{P_0}{\Lambda}
ight) - rac{\mathbf{P}^2}{2\Lambda}
ight) + e^{-P_0/2\Lambda} P_i \gamma^i.$$

Our perspective [AmelinoCamelia:2011bm, Carmona:2019fwf], see also [Kowalski-Glikman:2002oyi], [Majid:1999td]:

- Dispersion relation \rightarrow Squared distance from the origin to k.
- To have (modified) Poincaré invariance: translations ⊕ (or T_μ), Lorentz generators J_{μν}, all acting on momentum space → 10 isometries of the metric

 \rightarrow maximally symmetric momentum space (Minkowski, dS, AdS).

• Connection determined by the composition law

$$\partial_{p_{\mu}}\partial_{q_{\nu}}(p\oplus q)_{\sigma}\Big|_{q=p=0} = -\Gamma_{\sigma}^{\mu\nu}(0).$$
(6)

Geometry in momentum space

• Consider a momentum metric

$$g_{00}(p) = 1$$
, $g_{0i}(p) = g_{i0}(p) = rac{p_i}{2\Lambda}$, $g_{ij}(p) = -\delta_{ij}e^{-
ho_0/\Lambda} + rac{p_ip_j}{4\Lambda^2}$

• Compute the composition law, defined from the translation symmetry

$$g_{\mu
u}\left(p \oplus q
ight) \, = \, rac{\partial \left(p \oplus q
ight)_{\mu}}{\partial q_{
ho}} g_{
ho \sigma}(q) rac{\partial \left(p \oplus q
ight)_{
u}}{\partial q_{\sigma}},$$

$$(p \oplus q)_0 = p_0 + q_0, \qquad (p \oplus q)_i = p_i e^{q_0/2\Lambda} + q_i e^{-p_0/2\Lambda}.$$
 (7)

• This metric is related to the kinematics of κ -Poincaré in the symmetric basis.

Deformed relativistic wave equations

• Klein-Gordon equation derived from the Casimir (squared distance) (SF and J. Relancio, 2022).

$$(C_{
m D}-m^2)\phi = \left[\Lambda^2 \operatorname{arccosh}^2\left(\cosh\left(rac{p_0}{\Lambda}
ight)-rac{{f p}^2}{2\Lambda^2}
ight)-m^2
ight]\phi({m p})\,=\,0$$

• Klein–Gordon equation derived from the Casimir (Hopf algebra) [Lukierski:1992dt]

$$(C_{\rm A} - m^2)\phi = \left[\left(2\Lambda\sinh\left(\frac{p_0}{2\Lambda}\right)\right)^2 - \mathbf{p}^2 - m^2\right]\phi(\mathbf{p}) = 0$$

• $C_{\rm D} = f(C_{\rm A})$

Remarks

 Different kinematics can be obtained from the same metric: same dispersion relation but different composition laws Example:

$$g_{\mu
u}(p) = \eta_{\mu
u} + p_{\mu}p_{
u}/\Lambda^2.$$

Snyder kinematics [Battisti:2010sr]

$$(p\oplus q)^{
m Snyder}_\mu \,=\, p_\mu \left(\sqrt{1+rac{q^2}{\Lambda^2}}+rac{p_\mu\eta^{\mu
u}q_
u}{\Lambda^2\left(1+\sqrt{1+p^2/\Lambda^2}
ight)}
ight)+q_\mu.$$

κ-Poincaré in classical basis [Borowiec2010]

$$(p\oplus q)^{\kappa-{ ext{Poincare}}}_{\mu}= p_{\mu}\left(\sqrt{1+rac{q^2}{\Lambda^2}}+rac{q_0}{\Lambda}
ight)+q_{\mu}+n_{\mu}\left[rac{\sqrt{1+p^2/\Lambda^2}-p_0/\Lambda}{1-p^2/\Lambda^2}\left(q_0+rac{q_lpha\eta^{lphaeta}p_eta}{\Lambda}
ight)-q_0
ight],$$

where $n_{\mu} := (1, 0, 0, 0)$.

• Action in momentum space

$$S_{
m KG} \, := \, \int {
m d}^4 p \, \sqrt{-g} \, \phi^*(p) \left({\cal C}_{
m D}(p) - m^2
ight) \phi(p).$$

- Covariance.
- Invariance under deformed Lorentz transformations of the metric assuming the field transforms as a scalar

$$\phi'(p') = \phi(p),$$

since

$$C_{\rm D}(p) = C_{\rm D}(p').$$

• We want to find the equivalent of the Dirac equation in this framework:

"
$$(\gamma^{\mu} \mathcal{P}_{\mu} - m) \psi(p) = 0."$$

• We want to find the equivalent of the Dirac equation in this framework:

"
$$(\gamma^{\mu} \mathcal{P}_{\mu} - m) \psi(p) = 0.$$
"

• What is \mathcal{P}_{μ} ?

$$\mathcal{P}_{\mu}=f_{\mu}(p):=g_{\mu\nu}(p)f^{\nu}(p)=rac{1}{2}g_{\mu\nu}(p)rac{\partial \mathcal{C}_{\mathrm{D}}(p)}{\partial p_{\nu}}.$$

• We want to find the equivalent of the Dirac equation in this framework:

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$$(\gamma^{\mu} \mathcal{P}_{\mu} - m) \psi(p) = 0.$$
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• What is \mathcal{P}_{μ} ?

$$\mathcal{P}_{\mu}=f_{\mu}(p):=g_{\mu
u}(p)f^{
u}(p)=rac{1}{2}g_{\mu
u}(p)rac{\partial\mathcal{C}_{\mathrm{D}}(p)}{\partial p_{
u}}.$$

• Our guess is "local" Lorentz representation

$$\underline{\gamma}^{\mu} := \gamma^{a} e^{\mu}{}_{a}(p).$$

The gamma matrices satisfy

$$\{\underline{\gamma}^{\mu},\underline{\gamma}^{\nu}\} = 2g^{\mu\nu}(p)\mathbb{1}.$$

 There exist different tetrads that lead to the same metric; which one should we use? The composition law identifies one and only one tetrad:

$$g_{\mu
u}\left(p\oplus q
ight) = \, rac{\partial\left(p\oplus q
ight) _{\mu}}{\partial q_{
ho}}g_{
ho\sigma}\left(q
ight) rac{\partial\left(p\oplus q
ight) _{
u}}{\partial q_{\sigma}},$$

so for
$$q \rightarrow 0$$

$$egin{aligned} \mathbf{g}_{\mu
u}\left(\mathbf{p}
ight) = \left. rac{\partial \left(\mathbf{p} \oplus \mathbf{q}
ight)_{\mu}}{\partial \mathbf{q}_{
ho}}
ight|_{q o 0} \eta_{
ho \sigma} \left. rac{\partial \left(\mathbf{p} \oplus \mathbf{q}
ight)_{
u}}{\partial \mathbf{q}_{\sigma}}
ight|_{q o 0}, \end{aligned}$$

i.e.

$$e_{\mu}{}^{s}(p) := \left. \delta_{\nu}^{s} \left. rac{\partial \left(p \oplus q
ight)_{\mu}}{\partial q_{
u}} \right|_{q o 0}
ight.$$

• For the symmetric basis we find

$$\mathcal{D}_{\mathrm{D}}^{(S)} := rac{\sqrt{rac{C_{\mathrm{D}}^{(S)}(p)}{\Lambda^2}}}{2\Lambda\sinh\left(\sqrt{rac{C_{\mathrm{D}}^{(S)}(p)}{\Lambda^2}}
ight)} \left[2\Lambda e^{-rac{p_0}{2\Lambda}}\gamma^i p_i + \gamma^0\left(2\Lambda^2\sinh\left(rac{p_0}{\Lambda}
ight) - \mathbf{p}^2
ight)
ight].$$

• If we use instead $C_A^{(S)}(p)$

$$\mathcal{D}_{\mathsf{A}}^{(\mathsf{S})} := \gamma^{0} \left(\Lambda \sinh \left(\frac{p_{0}}{\Lambda} \right) - \frac{\mathbf{p}^{2}}{2\Lambda} \right) + e^{-p_{0}/2\Lambda} p_{i} \gamma^{i},$$

which is the same result as that obtained within the Hopf-algebraic approach [Nowicki:1992if].

Our construction leads to

$$\left(\mathcal{D}_{\mathrm{D}}^{(\mathcal{S})}\right)^2 = C_{\mathrm{D}}^{(\mathcal{S})}.$$



• Klein–Gordon equation is obtained straightforwardly

$$\left(\underline{\gamma}^{\nu}f_{\nu}(p)-m\right)\left(\underline{\gamma}^{\nu}f_{\nu}(p)+m\right) = C_{\mathrm{D}}(p)-m^{2}.$$

• The Dirac equation can be obtained from the action

$$\mathcal{S}_{ ext{Dirac}} \, := \, \int \mathrm{d}^4 p \, \sqrt{-g} ar{\psi}(-p) \left(\underline{\gamma}^\mu f_\mu(p) - m
ight) \psi(p).$$

• Spinors in Snyder

Remarks

- Invariant under deformed Lorentz transformations.
- Covariant.
- Introduce discrete symmetries

$$\begin{split} \mathcal{P}_0 &:= \mathrm{i}\gamma^0 \,, & \tilde{\psi}_{\mathcal{P}} := \mathrm{i}\gamma^0 \tilde{\psi}(\boldsymbol{p}_0, -\mathbf{p}) \,, \\ \mathcal{T}_0 &:= \mathrm{i}\gamma^1 \gamma^3 \mathcal{K} \,, & \tilde{\psi}_{\mathcal{T}} := \mathrm{i}\gamma^1 \gamma^3 \tilde{\psi}^*(\boldsymbol{p}_0, -\mathbf{p}) \,, \\ \mathcal{C}_0 &:= \mathrm{i}\gamma^2 \mathcal{K} \,, & \tilde{\psi}_{\mathcal{C}} := \mathrm{i}\gamma^2 \tilde{\psi}^*(-\boldsymbol{p}) \,. \end{split}$$

• Invariant under \mathcal{P} and \mathcal{T} ,

$$e^{\mu}{}_{a}(p_{0},-\mathbf{p})f_{\mu}(p_{0},-\mathbf{p}) = -e^{\mu}{}_{a}(p)f_{\mu}(p), \quad a = 1,2,3.$$
 (8)

• Invariant under C when $\Lambda \to -\Lambda$,

$$e^{\mu}{}_{a}(-p)f_{\mu}(-p)\,=\,-e^{\mu}{}_{a}(p)f_{\mu}(p)\,,\quad a\,=\,0,1,2,3\,.$$

(9)

Outline

Momentum space and vacuum energies

Ø Momentum space and fermions

3 Conclusions

• DSR as curved momentum space.

• Vacuum energies in Snyder spaces, spectral properties.

• DSR Fermions as local realizations of the Lorentz group in momentum space.

• . . .

Snyder representation [Snyder:1946qz, Lu:2011fh], is defined by the operators

$$\hat{p}_i = p_i, \quad \hat{x}_i = i \left(\delta_{ij} - \beta^2 p_i p_j \right) \frac{\partial}{\partial p_j}.$$
(10)

acting on a Hilbert space of functions $\psi(p)$ with measure $d\mu = \frac{d^{D}p}{(1-\beta^{2}p^{2})^{(D+1)/2}}$ [Lu:2011fh]. A different realization was introduced in [Mignemi:2011gr]:

$$\hat{p}_i = \frac{p_i}{\sqrt{1+\beta^2 p^2}}, \qquad \hat{x}_i = i\sqrt{1+\beta^2 p^2} \frac{\partial}{\partial p_i}.$$
(11)

The measure on the Hilbert space is in this case $d\mu = rac{d^D p}{\sqrt{1+\beta^2 p^2}}$. transformation of variables $p = rac{q}{\sqrt{1+\beta^2 q^2}}$

$$\mathcal{E} = \frac{\Omega_{D-2}}{2\beta} \sum_{n=1}^{\infty} \int_{0}^{\infty} \frac{dq}{(2\pi)^{D-1}} q^{D-2} \sqrt{1 + \beta^{2} m^{2} - \frac{1}{(1 + \beta^{2} q^{2}) \cosh^{2}(\beta k_{n})}}$$

$$= \frac{\Omega_{D-2}}{2\beta} \sum_{n=1}^{\infty} \int_{0}^{\infty} \frac{dq}{(2\pi)^{D-1}} q^{D-2} \sqrt{1 + \beta^{2} m^{2}} \left[1 - \frac{1}{2u} - \frac{1}{8u^{2}} + \cdots\right],$$
(12)

where $u = (1 + \beta^2 m^2)(1 + \beta^2 q^2) \cosh^2(\beta k_n)$. After the expansion, there exists only a finite number of divergent terms in expression (18) for a fixed dimension D. Using an adequate regularization one can make use of Abel-Plana formula to approximate the series with an integral plus a constant contribution, that in conjunction with the change of variables (??) gives

$$\frac{\Omega_{D-2}}{2\beta} \sum_{n=1}^{\infty} \int_{0}^{\infty} \frac{dq}{(2\pi)^{D-1}} q^{D-2} = \frac{2L}{(2\pi)^{D\beta}} \text{Vol}(\mathbb{H}_{D}) + \frac{1}{4(2\pi)^{D-1}\beta} \text{Vol}(\mathbb{H}_{D,w=0}).$$
(13)

$$\mathcal{F} = \frac{\Omega_{D-2}}{2(2\pi)^{D-1}} \int_0^{1/\beta} \frac{dp \, p^{D-2}}{(1-\beta^2 p^2)^{D/2-1/2}} \left(\sum_{n=1}^\infty f_n(p) - \int_0^\infty dn \, f_n(p) \right), \tag{14}$$

where we have introduced the functions

$$f_n(p) = \frac{k_n}{\beta L} \frac{\tanh(\beta k_n)}{\cosh^2(\beta k_n) \sqrt{p^2 + \beta^{-2} \tanh^2(\beta k_n) (1 - \beta^2 p^2) + m^2}}.$$
 (15)

Snyder spaces

Noncommutative coordinates \bar{x}_{μ} and p_{μ} satisfy the following commutation relations¹

$$\begin{split} &[\bar{\mathbf{x}}_{\mu}, \bar{\mathbf{x}}_{\nu}] = i\beta M_{\mu\nu} \psi(\beta \boldsymbol{p}^{2}), \\ &[\boldsymbol{p}_{\mu}, \boldsymbol{p}_{\nu}] = 0, \\ &[\boldsymbol{p}_{\mu}, \bar{\mathbf{x}}_{\nu}] = -i\phi_{\mu\nu}(\beta \boldsymbol{p}^{2}), \end{split}$$
(16)

while the Lorentz generators $M_{\mu\nu}$ have the same commutation relations as in the usual case, i.e.

$$\begin{split} [M_{\mu\nu}, M_{\rho\sigma}] &= i(\eta_{\mu\rho}M_{\nu\sigma} - \eta_{\mu\sigma}M_{\nu\rho} + \eta_{\nu\rho}M_{\mu\sigma} - \eta_{\nu\sigma}M_{\mu\rho}), \\ [M_{\mu\nu}, \mathbf{p}_{\lambda}] &= i(\eta_{\mu\lambda}\mathbf{p}_{\nu} - \eta_{\lambda\nu}\mathbf{p}_{\mu}), \\ [M_{\mu\nu}, \bar{\mathbf{x}}_{\lambda}] &= i(\eta_{\mu\lambda}\bar{\mathbf{x}}_{\nu} - \eta_{\lambda\nu}\bar{\mathbf{x}}_{\mu}). \end{split}$$

$$\end{split}$$

¹Mignemi et al.: 2017

Expansion of Casimir:

$$\mathcal{E} = \frac{\Omega_{D-2}}{2\beta} \sum_{n=1}^{\infty} \int_{0}^{\infty} \frac{dq}{(2\pi)^{D-1}} q^{D-2} \sqrt{1+\beta^2 m^2} \left[1 - \frac{1}{2u} - \frac{1}{8u^2} + \cdots \right],$$
(18)

 $u = (1 + \beta^2 m^2)(1 + \beta^2 q^2) \cosh^2(\beta k_n)$

 $T_
ho:=T'_
ho+c_
ho^{\mu
u}J_{\mu
u}\Longrightarrow$

$$[T^{0}, T^{i}] = \frac{c_{1}}{\Lambda} T^{i} + \frac{c_{2}}{\Lambda^{2}} J^{0i}, \qquad [T^{i}, T^{j}] = \frac{c_{2}}{\Lambda^{2}} J^{ij}.$$
(19)

$$(\phi,\psi)_{\mathrm{KG}} = \int \frac{\mathrm{d}^{3}\boldsymbol{p}}{\omega_{\boldsymbol{p}}} \left[\phi_{+}^{*}(\boldsymbol{p})\psi_{+}(\boldsymbol{p}) - \phi_{-}^{*}(\boldsymbol{p})\psi_{-}(\boldsymbol{p}) \right], \qquad (20)$$

$$(\phi,\psi)_{C_{\mathrm{D}}} := 2 \int \mathrm{d}^4 p \, \sqrt{-g} \, \delta\left(C_{\mathrm{D}} - m^2\right) \, \Theta(f^{\nu} t_{\nu}) \, \phi^*(p) \psi(p) \,, \tag{21}$$

$$\psi'(p') = S(\Lambda(p))\psi(p).$$
⁽²²⁾

$$\left(\mathcal{S}_{\underline{\gamma}}^{\mu}\mathcal{S}^{-1}f_{\mu}(p)-m\right)\mathcal{S}\psi(p) = \left(\mathcal{S}\gamma^{a}\mathcal{S}^{-1}e^{\prime\rho}{}_{b}(p^{\prime})f_{\rho}^{\prime}(p^{\prime})-m\right)\psi^{\prime}(p^{\prime}) = 0, \qquad (23)$$

where we have used appropriate transformations for the vielbein and the functions f^{μ} , cf. the definition (??) (all the quantities in the new system of coordinates are denoted with a prime).

Then, since the proposed diffeomorphism is an isometry of the metric, the vielbein satisfies

$$e^{\rho_{a}}(p') = e^{\rho_{b}}(p') \Lambda^{b_{a}}(p'), \qquad (24)$$

where $\Lambda^{b}{}_{a}(p')$ is a (local) Lorentz transformation that may depend on the point p'. Note that this transformation is not the deformed Lorentz transformation obtained from the isometries of the metric. Considering (24) and the symmetry of f_{μ} (cf. the definition (??)), we obtain

$$\left(S\gamma^{a}S^{-1}\Lambda^{b}{}_{a}(p')e^{\rho}{}_{b}(p')f_{\rho}(p')-m\right)\psi'(p')=0.$$
(25)

Of course, local Lorentz transformations near the identity can be expanded in terms of antisymmetric parameters $\epsilon_{ab} := \eta_{ac} \epsilon^c{}_b$ as customarily,

$$\Lambda^{b}{}_{a}(p') = \delta^{b}_{a} + \epsilon^{b}{}_{a}(p') + \cdots .$$
⁽²⁶⁾

Thus, we can also expand the matrix S in Eq. (23) for transformations around the identity; this allows us to determine the infinitesimal form of S as a function of the Lorentz coefficients $\epsilon_{ab}(p')$:

$$S = \mathbb{I} - \frac{\mathrm{i}}{4} \sigma^{ab} \epsilon_{ab}(p') + \cdots, \sigma^{ab} := \frac{\mathrm{i}}{2} \left[\gamma^{a}, \gamma^{b} \right] .$$
⁽²⁷⁾

More generally for an arbitrary metric in momentum space, the expression (??) is covariant once we postulate

NC QFT

Consider the simpler situation (Moyal plane)

$$[\hat{x}_a, \hat{x}_b] = 2i\Theta_{ab}$$

Functions are promoted to operators by means of Weyl's transform:

$$\hat{f} := \Omega(f)(\hat{x}) := \int d\tilde{p} \,\tilde{f}(p) \,e^{ip\hat{x}}.$$
(31)

The product of these operators

$$egin{aligned} \hat{f}\hat{g} &= \int d ilde{p}d ilde{q}\, ilde{f}(p)\,e^{ip\hat{\lambda}} ilde{g}(q)\,e^{iq\hat{\lambda}} \ &= \int d ilde{p}\left(\int d ilde{q}\, ilde{f}(p-q) ilde{g}(q)\,e^{ip\Theta q}
ight)\,e^{ip\hat{\lambda}}, \end{aligned}$$

is such that its Weyl's antitransformation is nothing but the Moyal product:

$$\Omega^{-1}(\hat{f}\hat{g}) = \int d\tilde{p}d\tilde{q}\,\tilde{f}(p-q)\tilde{g}(q)\,e^{ip\Theta q}\,e^{ipx} =:(f\star g)(x).$$

NC QFT

An approach to NC QFT is to introduce the \star noncommutative product (Groenewold-Moyal) to deform the algebra of functions (coordinates are not operators anymore):

$$\begin{split} f(\hat{x})g(\hat{x}) &= (f \star g)(x) = e^{i\Theta_{ab}\partial_a^f\partial_b^g} f(x) g(x) \\ &= f(x_a + i\Theta_{ab}\partial_b^g) g(x) \\ &= f(x) g(x_a - i\Theta_{ab}\partial_b^f). \end{split}$$

The \star product is NC and associative; of course

$$[x_a, x_b]_{\star} = 2i\Theta_{ab}.$$

Snyder spaces

One can find a *-product; for two plane waves (Mignemi et al.: 2017):

$$e^{ik \cdot x} \star e^{iq \cdot x} = e^{iD(k,q) \cdot x + iG(k,q)}, \tag{32}$$

where

$$D^{\mu}(k,q) = k^{\mu} + q^{\mu} + \beta \left[k^{\mu} \left(s_{1}q^{2} + \left(s_{1} + \frac{s_{2}}{2} \right) k \cdot q \right) + q^{\mu} s_{2} \left(k \cdot q + \frac{k^{2}}{2} \right) \right] + \mathcal{O}(\beta^{2}),$$

$$G(k,q) = -i\beta \left(s_{1} + \frac{D+1}{2} s_{2} \right) k \cdot q + \mathcal{O}(\beta^{2}).$$
(33)

One can show that

$$\int f(x) \star g(x) = \int f(x)g(x). \tag{34}$$

The *-product is nonassociative (but the Jacobi identities are still satisfied).

Snyder $\lambda \phi_{\star}^4$

Consider the action (Meljanac, Mingemi et al., 2017; SAF and Mignemi: 2018)

$$S[\varphi] = \int \frac{1}{2} \partial_{\mu} \varphi \star \partial^{\mu} \varphi + \frac{m^2}{2} \varphi \star \varphi - \frac{\lambda}{4!} \varphi \star (\varphi \star (\varphi \star \varphi)).$$
(35)

so that

$$\Gamma_{1-loop}[\phi] = S[\phi] + \frac{\mu^{-\epsilon}}{2} \operatorname{Tr} \log A, \quad Af(x) = \int dy \frac{\delta^2 S}{\delta \varphi(x) \delta \varphi(y)}[\phi] f(y).$$
(36)

The interaction term gives a contribution

$$\frac{\delta^{2} S_{l}}{\delta \varphi_{x} \delta \varphi_{y}} \approx -\frac{\lambda}{4!} \frac{1}{2} \int \left(\prod_{m=1}^{3} \frac{dq_{m}}{(2\pi)^{D}} \right) \tilde{\phi}_{1} \tilde{\phi}_{2} e^{-iq_{3}y+i(q_{1}+q_{2}+q_{3})x} \left[4! + 4\beta(s_{1}+s_{2}) \right] \times \left(-2i \sum_{k} q_{k}^{2} q_{k} \cdot x + (2+D) \left(-\sum_{i=1}^{3} q_{i}^{2} + q_{4}^{2} \right) \right) \right]_{q_{4}=-(q_{1}+q_{1}+q_{3})}$$
(37)

Snyder-de Sitter

The commutation relations are (SAF and Mignemi: 2019, 2021)

$$\begin{aligned} & [\hat{x}_i, \hat{x}_j] = i\beta^2 J_{ij}, \qquad [\hat{p}_i, \hat{p}_j] = i\alpha^2 J_{ij}, \\ & [\hat{x}_i, \hat{p}_j] = i[\delta_{ij} + \alpha^2 \hat{x}_i \hat{x}_j + \beta^2 \hat{p}_j \hat{p}_i + \alpha\beta(\hat{x}_j \hat{p}_i + \hat{p}_i \hat{x}_j)]. \end{aligned}$$

with the "usual" Lorentz generators

$$J_{ij} = \frac{1}{2} (\hat{x}_i \hat{p}_j - \hat{x}_j \hat{p}_i + \hat{p}_j \hat{x}_i - \hat{p}_i \hat{x}_j).$$
(39)

An interacting theory with action

$$S_{5d5} = \frac{1}{2} \int d^D x \ \varphi \left(\hat{p}^2 + m^2 \right) \varphi + \frac{\lambda}{4!} \int d^D x \ \varphi(\hat{x}) \Big[\varphi(\hat{x}) \Big(\varphi^2(\hat{x}) \Big) \Big], \tag{40}$$

so that the relevant operator at the one-loop level is

$$\delta^2 S_W \approx p^2 + \omega^2 x^2 + \alpha_{\text{eff}}^2 (x_i x_j p_j p_i)_S + m^2 + V_W, \quad \omega^2 = \frac{\alpha^2}{\beta^2}.$$
(41)

Snyder-de Sitter

The terms to be renormalized in $D=4-\epsilon$ dimensions are

$$\Gamma_{\rm div}^{(2)} = -\frac{\lambda}{192\pi^{2}\varepsilon} \int d^{D}x \,\phi \left[-26\alpha_{\rm eff}^{2} + 6m^{2} + 12m^{4}\beta^{2} + \frac{6\alpha_{\rm eff}^{2}m^{4}}{\omega^{2}} + 48\beta^{2}\omega^{2} \right.$$

$$\left. + x^{2}(15\alpha_{\rm eff}^{2}m^{2} + 6\omega^{2} + 36m^{2}\beta^{2}\omega^{2} - 4\beta^{2}\omega^{2}\partial^{2}) - 8\beta^{2}\omega^{2}x_{\mu}x_{\nu}\partial^{\mu}\partial^{\nu} \right.$$

$$\left. + x^{4}(9\alpha_{\rm eff}^{2}\omega^{2} + 24\beta^{2}\omega^{4}) \right] \phi,$$

$$\Gamma_{\rm div}^{(4)} = \frac{1}{4!} \frac{3\lambda^{2}}{16\pi^{2}\varepsilon} \int d^{D}x \,\phi^{2} \left[\frac{\alpha_{\rm eff}^{2}\partial^{2}}{2\omega^{2}} - \frac{2m^{2}\left(\alpha_{\rm eff}t^{2} + 2\beta^{2}\omega^{2}\right) + \omega^{2}}{\omega^{2}} - \frac{x^{2}\left(5\alpha_{\rm eff}^{2} + 16\beta^{2}\omega^{2}\right)}{2} \right] \phi^{2}$$

$$\left. -\beta^{2}\phi_{\star,(1)}, \right\}$$

$$\left. \left. + \frac{1}{2} \frac{3\lambda^{2}}{2\omega^{2}} + \frac{1}{2} \frac{3\lambda^{2}}{2$$

Recall that in de Sitter

$$\sqrt{\det g} = \frac{1}{(1 + \alpha^2 x^2)^{(D+1)/2}}.$$
(44)

 SdS



Numerical solutions for the running of λ , β^2 and α^2 .

What is WLF? To do QFT (or to compute HK) using QM in first quantization.

Consider a real scalar field φ with action S in a d-dimensional Euclidean space. The effective action is

$$\Gamma = S + \frac{\hbar}{2} \log \det \delta^2 S + \dots \tag{45}$$

For

$$S=\int dx \; rac{1}{2} (\partial_x arphi)^2 + rac{m^2}{2} arphi^2 + U(arphi)$$

the quantum corrections can be written by using Schwinger proper time in the following form

$$\Gamma_{1-\text{loop}} = -\frac{1}{2} \int_{0}^{\infty} \frac{d\beta}{\beta} e^{-\beta m^{2}} \underbrace{\text{Tr}\left(e^{-\beta\left(-\partial^{2} + U''(\phi)\right)}\right)}_{\text{Heat-Kernel trace} = \int dx \ K(x,x,\beta)}$$
(46)