
UV sensitivity of dark matter production

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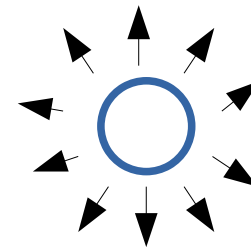
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- *non-thermal dark matter*
- *Planck-suppressed operators*
- *particle production after inflation*
- *relic abundance*

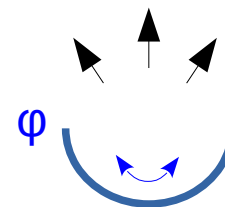
Non-thermal DM has memory !

Production mechanisms (all add up):

- during inflation



- via inflaton oscillations



- inflaton decay



- thermal emission (freeze-in)

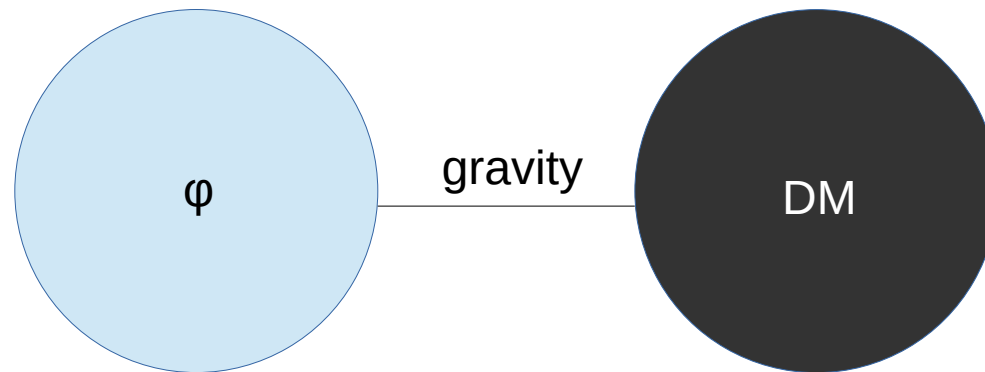


Focus:

inflaton oscillation epoch (preheating)

Assume:

no renormalizable inflaton-dark matter coupling

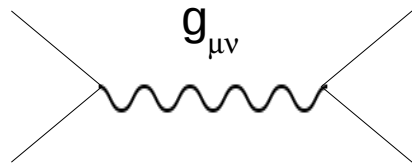


Dim-6 gravity-induced couplings:

$$\Delta\mathcal{L}_6 = \frac{C_1}{M_{\text{Pl}}^2} (\partial_\mu\phi)^2 s^2 + \frac{C_2}{M_{\text{Pl}}^2} (\phi\partial_\mu\phi)(s\partial^\mu s) + \frac{C_3}{M_{\text{Pl}}^2} (\partial_\mu s)^2 \phi^2 - \frac{C_4}{M_{\text{Pl}}^2} \phi^4 s^2 - \frac{C_5}{M_{\text{Pl}}^2} \phi^2 s^4$$

Simplest examples:

graviton exchange



frame change

Jordan frame \rightarrow Einstein frame

$$\xi s^2 R \quad \rightarrow \quad \frac{1}{M_{\text{Pl}}^2} s^2 V(\phi)$$

Many more sources: *non-perturbative, string states, ...*



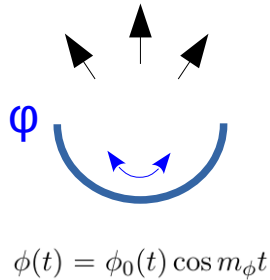
**unknown
coefficients**

On shell, 2 independent dim-6 operators contribute to DM pair-production:

$$\mathcal{O}_3 = \frac{1}{M_{\text{Pl}}^2} (\partial_\mu s)^2 \phi^2 \quad , \quad \mathcal{O}_4 = \frac{1}{M_{\text{Pl}}^2} \phi^4 s^2$$

(supplemented with dim-4 $\mathcal{O}_{\text{renorm}} = \frac{m_\phi^2}{M_{\text{Pl}}^2} \phi^2 s^2$ and 4-DM op $\frac{C_5}{M_{\text{Pl}}^2} \phi^2 s^4$)

Perturbative DM production rates:



$$\mathcal{O}_3 / \mathcal{O}_4 \sim \text{energy}^2 / \phi^2 \sim m_\phi^2 / \phi^2 \ll 1$$

$$\frac{\Gamma[\mathcal{O}_3]}{\Gamma[\mathcal{O}_4]} \sim \frac{C_3^2}{C_4^2} \frac{\omega^4}{\phi_0^4} \sim 10^{-20}$$



“graviton exchange” operator much **suppressed**

E.g. $\frac{\tilde{C}_6}{M_{\text{Pl}}^4} \phi^6 s^2 + \frac{\tilde{C}_8}{M_{\text{Pl}}^6} \phi^8 s^2$ more important

Resonant DM production:

Semiclassical analysis:

$$S = \int d^4x \sqrt{|g|} \left(\frac{1}{2} \mathcal{K}(\phi) g^{\mu\nu} \partial_\mu s \partial_\nu s - \mathcal{V} \right)$$

$\phi(t) = \phi_0(t) \cos m_\phi t$

$$\ddot{s} - \frac{1}{a^2} \partial_i \partial_i s + \left(\frac{\dot{\mathcal{K}}}{\mathcal{K}} + 3H \right) \dot{s} + \frac{\mathcal{V}'_s}{\mathcal{K}} = 0$$

Change of variables
(3-momentum k-mode):

$$s_k = a^{-3/2} \mathcal{K}^{-1/2} X_k$$

$(\partial_\mu s)^2 \phi^2$

$$\ddot{X}_k + \left[\frac{2C_3 m_\phi^2 \phi_0^2}{M_{\text{Pl}}^2} \cos 2m_\phi t + \frac{k^2}{a^2} \right] X_k = 0$$

no resonant production ($q \sim 1$)

$\phi^4 s^2$

$$\ddot{X}_k + \left[\frac{2C_4 \phi_0^4}{M_{\text{Pl}}^2} \cos^4 m_\phi t + \frac{k^2}{a^2} \right] X_k = 0$$

resonant production ($q \gg 1$)

$$\underline{(\partial_\mu s)^2 \phi^2} \ll \phi^4 s^2$$

On effect of $\frac{C_5}{M_{\text{Pl}}^2} \phi^2 s^4$:

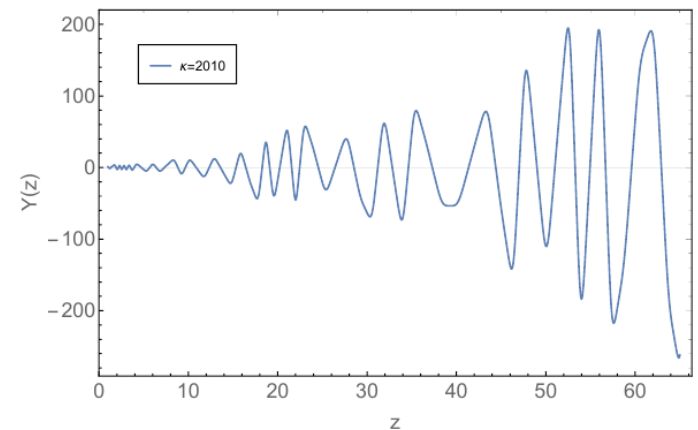
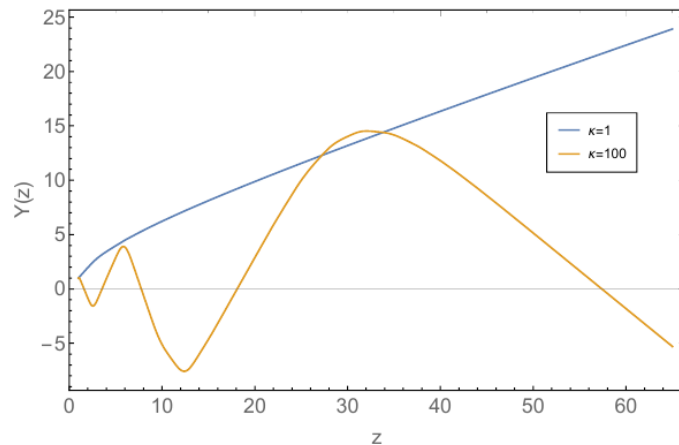
$$\ddot{s} - \frac{1}{a^2} \partial_i \partial_i s + 3H \dot{s} + \frac{4C_5 \phi^2(t)}{M_{\text{Pl}}^2} s^3 = 0$$

Does not belong to the **Hill type** equations!

Zero mode:

$$Y'' + \kappa \frac{\cos^2 z}{z^4} Y^3 = 0$$

$$z = m_\phi t, \quad \kappa \simeq \frac{14C_5 \varphi_0^2}{M_{\text{Pl}}^2 a_0^3} \lesssim \mathcal{O}(1)$$



no resonant production for $\kappa \sim 1$

DM abundance via Planck-suppressed operators

Focus on the most important op $\frac{C_4}{M_{\text{Pl}}^2} \phi^4 s^2$

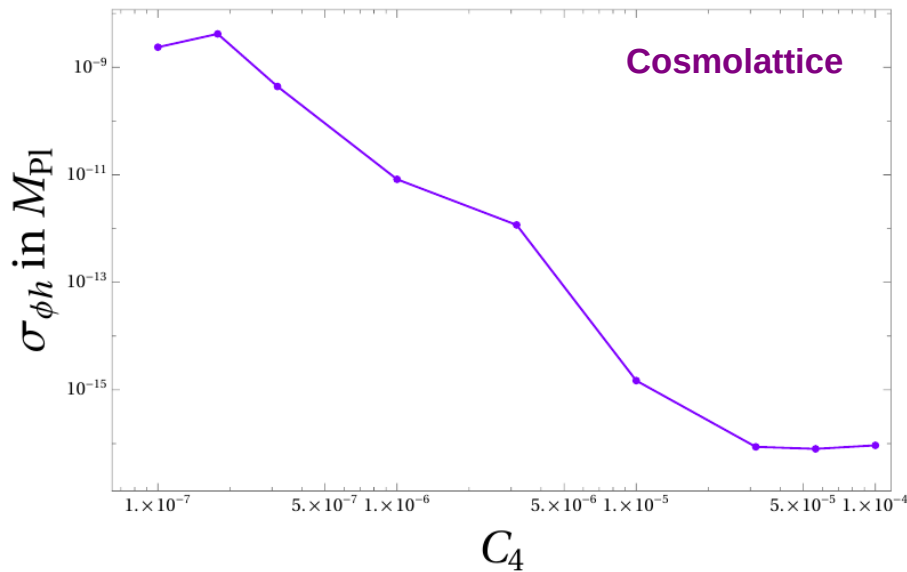
Reheating:

$$V_{\phi h} = \sigma_{\phi h} \phi H^\dagger H \quad \rightarrow \quad \phi \rightarrow hh$$

Correct DM abundance requires:

$$\sigma_{\phi h} \simeq 5 \times 10^{-9} \frac{m_\phi^{3/2}}{M_{\text{Pl}}^{1/2}} \left(\frac{n_e(\phi)}{n_e(s)} \right) \left(\frac{\text{GeV}}{m_s} \right)$$

simulation output



$$m_\phi = 10^{13} \text{ GeV}$$

$$m_s = 1 \text{ GeV},$$

$$\varphi_0 \simeq M_{\text{Pl}}$$



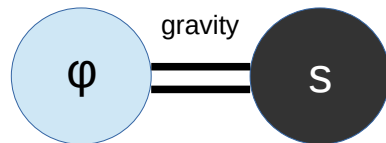
$$C_4 \ll 1 !$$

Main observation :

*Planck—suppressed (“gravity--induced”) operators
with small Wilson coefficients
can account for all of the dark matter !*

$$\frac{C_4}{M_{\text{Pl}}^2} \phi^4 s^2, \text{ etc. (dim-8, ...)}$$

Non-thermal DM model building is highly **UV sensitive** :



- abundance is additive (“memory”)
- need to control quantum gravity
- predictivity ?

CONCLUSION

- *non-thermal DM production at preheating*
- *(small) Planck—suppressed operators produce enough DM*
- *strong UV sensitivity of non-thermal DM (quantum gravity)*