

The Weak Gravity Conjecture with Scalar Fields in AdS

Marco Michel

work in collaboration with
Stefano Andriolo and Eran Palti

Ben-Gurion University of the Negev

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אוניברסיטת בן-גוריון בנגב
Ben-Gurion University of the Negev



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Introduction: The Positive Binding Conjecture

The Weak Gravity Conjecture

- The Swampland program is motivated by the identification and verification of universal features of quantum gravity.
- A prominent example is the *Weak Gravity Conjecture* [Akani-Hamed, Motl, Nicolis, Vafa '06]

$$\sqrt{2}gqM_P \geq m$$

- An equivalent formulation is the *Repulsive-Force Conjecture* [Palti '17; Heidenreich, Reece, Rudelius '19], which requires a consistent gravitational theory, with a $U(1)$ gauge symmetry, must have a self-repulsive charged particle.
- The presence of an additional scalar field leads to a modification in the presence of additional massless scalar fields:

$$2g^2q^2M_p^2 \geq m^2 + \mu^2M_p^2.$$

Positive Binding Conjecture

- A more appropriate formulation in AdS is the

Positive Binding Conjecture [Aharony, Palti '21]

For a (weakly coupled) gravitational theory with a $U(1)$ gauge field, there should exist at least one charged particle in the theory, with charge of order one, to have a non-negative self-binding energy.

- Maps to certain convexity properties of charged operators in the holographically dual CFTs
- Demanding positive self-binding energy places constraints on the effective theory

Goal

We determine what those constraints are by calculating the self-binding energy for a charged particle in terms of its couplings.

Setup

- As a prototype model we consider following theory consisting of a charged scalar ϕ , a set of photons A_i , a graviton $h_{\mu\nu}$:

$$S[\phi, A^i, h] = \int d^5x \sqrt{-g} \left[\frac{1}{\kappa^2} \left(\frac{R}{2} + \frac{6}{L^2} \right) - \sum_{i=1}^N \frac{F_i^2}{4} - |D\phi|^2 - m^2|\phi|^2 - V(\phi) \right], \quad \text{with} \quad V = a|\phi|^4 + b|\phi|^2|D\phi|^2.$$

- We use global AdS coordinates to describe our background.

$$ds^2 = \frac{L^2}{y^2} \left(-dt^2 + \frac{dy^2}{1-y^2} + (1-y^2)d\Omega_3^2 \right).$$

- Under the assumption that

$$\frac{g_i}{L^{1/2}} \ll 1, \quad \frac{\kappa}{L^{3/2}} \ll 1, \quad \frac{a}{L} \ll 1, \quad \frac{b}{L^3} \ll 1,$$

we will find that $\mathcal{O}(\kappa^2), \mathcal{O}(g_i^2), \mathcal{O}(a), \mathcal{O}(b)$ are the leading contributions to the binding energy. In the context of AdS supergravity:

$g_i \propto \frac{\kappa}{L}, a \propto \frac{\kappa^2}{L}, b \propto \kappa^2$, and thus this is ensured by $\frac{\kappa}{L^{3/2}} \ll 1$.

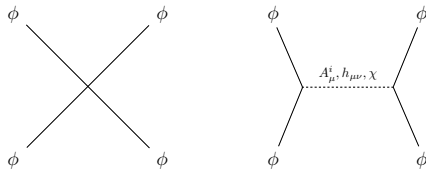
Binding energy from effective potentials

Binding energy

- The binding energy is defined as the difference between the energy of the bound state and twice the energy of the single (free) state $|\phi\rangle$:

$$\gamma \equiv E_{\phi\phi} - 2E_{\phi} = \langle \phi\phi | H | \phi\phi \rangle - 2 \langle \phi | H | \phi \rangle .$$

- States can be either $\gamma > 0$ self-repulsive or $\gamma < 0$ self-attractive. A state is BPS if in an interacting theory $\gamma = 0$.
- We are interested in computing the binding energy at tree level, in which case γ is merely determined by quartic interactions ϕ^4 .
- In order to avoid going to second order perturbation theory we first integrating out classically all additional fields [Fitzpatrick, Shih '11].



Basic scalar field in AdS

- Field decomposition in creation and annihilation operators:

$$\phi = \sum_{n,l,J} (a_{n,l,J} \psi_{n,l,J}^*(x) + b_{n,l,J}^\dagger \psi_{n,l,J}(x)),$$

- Free Hamiltonian is then given by

$$H_{free} = \sum_{n,l,J} E_{n,l} (a_{n,l,J}^\dagger a_{n,l,J} + b_{n,l,J}^\dagger b_{n,l,J}),$$

$$E_{n,l} = \Delta + 2n + l, \quad n, l = 0, 1, 2, \dots \quad \Delta = 2 \pm \sqrt{4 + m^2 L^2}$$

- We are interested in the state with lowest energy $n = l = J = 0$,

$$|\phi\rangle = b_0^\dagger |0\rangle, \quad |\phi\phi\rangle = \frac{1}{\sqrt{2}} b_0^\dagger b_0^\dagger |0\rangle, \quad b_0 \equiv b_{0,0,0},$$

corresponding to the eigenfunction

$$\psi_0(x) \equiv \psi_{0,0,0}(x) = L^{-3/2} N_\Delta e^{it\Delta} y^\Delta, \quad N_\Delta = \sqrt{\frac{\Delta - 1}{2\pi^2}}.$$

Integrating out additional fields

- Expand

$$A_i = A_i^{(0)} + g_i A_i^{(1)} \quad g_{\mu\nu} = g_{\mu\nu}^{(0)} + \kappa h_{\mu\nu}$$

- Integrating out additional field (photon and graviton) classically by using their equations of motion:

$$\frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} F_i^{\mu\nu}) = g_i q_i J^\nu, \quad \Delta_{\mu\nu}^{\rho\sigma} h_{\rho\sigma} = \frac{\kappa}{2} T_{\mu\nu},$$

with sources

$$J_\mu[\phi, \phi^\dagger] = i\phi^\dagger \partial_\mu \phi + h.c.,$$

$$T_{\mu\nu}[\phi, \phi^\dagger] = g_{\mu\nu} (-|\partial\phi|^2) - m^2|\phi|^2 + (\partial_\mu \phi^\dagger \partial_\nu \phi + h.c.).$$

- Resulting in an non-local effective potential that only depends on ϕ :

$$S_{\text{eff}} = \int d^5x \sqrt{-g} [-|\partial\phi|^2 - m^2\phi^2 - V_{\text{eff}}[\phi, \phi^\dagger]],$$

$$V_{\text{eff}}[\phi, \phi^\dagger] = V[\phi, \phi^\dagger] + \frac{1}{2} \sum_i A_{i\mu}^{(1)}[\phi, \phi^\dagger] J_i^\mu[\phi, \phi^\dagger] - \frac{\kappa}{4} h_{\mu\nu}[\phi, \phi^\dagger] T^{\mu\nu}[\phi, \phi^\dagger]$$

Contributions to binding energy

- With the Hamiltonian

$$H = H_{\text{free}} + \delta H_{\text{eff}}, \quad \delta H_{\text{eff}} = \int d^4x \sqrt{-g} V_{\text{eff}}.$$

- Now we can use first order perturbation theory:

$$\begin{aligned} E_{\phi\phi} &\equiv \langle \phi\phi | H | \phi\phi \rangle = E_{\phi\phi}^{\text{free}} + \langle \phi\phi | \delta H_{\text{eff}} | \phi\phi \rangle \\ &= 2(E_{\phi}^{\text{free}} + \langle \phi | \delta H_{\text{eff}}^{\text{quadratic}} | \phi \rangle) + \langle \phi\phi | \delta H_{\text{eff}}^{\text{quartic}} | \phi\phi \rangle, \end{aligned}$$

- The binding energy is therefore

$$\gamma = \int d^4x \sqrt{-g} \langle \phi\phi | V_{\text{eff}} | \phi\phi \rangle = 2 \int d^4x \sqrt{-g} V_{\text{eff}}[\phi(x)^{(\dagger)} = \psi_0(x)^{(\dagger)}],$$

- Contribution to binding energy [Fitzpatrick, Shih '11]:

$$\gamma^{(a,b)} = \frac{\pi^2 N_{\Delta}^4 (aL^2 - b(\Delta - 2)\Delta)}{(\Delta - 1)(2\Delta - 1)L^3},$$

$$\gamma_i^{(A_i)} = \frac{\pi^2 g_i^2 q_i^2 N_{\Delta}^4}{L(2\Delta - 1)}, \quad \gamma^{(h)} = -\frac{2\pi^2 (\Delta - 2)\Delta^2 \kappa^2 N_{\Delta}^4}{3(\Delta - 1)(2\Delta - 1)L^3}.$$

Contribution for scalar fields

Additional neutral scalar field

- Next we add an additional neutral scalar field χ .

$$S[\phi, A^i, h, \chi] = S[\phi, A, h] - \int d^5x \sqrt{-g} \left[\frac{1}{2}(\partial\chi)^2 + \frac{M^2}{2}\chi^2 + Y\chi|\phi|^2 \right],$$

- The equation of motion for χ is

$$\square\chi - M^2\chi = Y|\phi|^2 \quad \text{with} \quad YL^{1/2} \ll 1$$

- Expand

$$\chi = \chi_0 + Y\chi_1$$

- χ contributes to the effective potential with

$$V^{(\chi)}[\phi, \phi^\dagger] = \frac{Y^2}{2}\chi_1[\phi, \phi^\dagger]|\phi|^2,$$

- Note, requiring boundary terms to vanish leads to restrictions of the solution. Alternatively, start with modified action [Klebanov, Witten '99].
- For ϕ in the ground state :

$$N_\Delta^2 y^{2\Delta} + M^2\chi_1 + y((y^2 + 3)\chi_1' + y(y^2 - 1)\chi_1'') = 0,$$

Massless case

- For $M = 0$, we can solve the eom analytically:

$$\chi_1 = \frac{N_\Delta^2}{4(\Delta - 2)(\Delta - 1)} \left[\frac{y^2}{y^2 - 1} - \log(1 - y^2) + f_\Delta(y) \right],$$

where the integration constants are fixed by requiring that the solution is smooth at $y = 1$ and vanishes at $y = 0$.

$$f_\Delta(y) = \frac{\Delta - 1}{\Delta} y^{2\Delta} + B_{y^2}(\Delta + 1, -1) + (\Delta - 1)B_{y^2}(\Delta + 1, 0),$$

$$B_x(a, b) \equiv \int_0^x dt t^{a-1} (1-t)^{b-1},$$

where $B_x(a, b)$ are incomplete Beta functions.

- Contributing to the binding energy with

$$\begin{aligned} \gamma^{(x)} = & \frac{Y^2 \pi^2 N_\Delta^4}{8(\Delta - 2)^2 (\Delta - 1)^2} \left[1 - \Delta + \frac{1}{\Delta} + \frac{4}{\Delta - 1} \right. \\ & \left. + \frac{2}{2\Delta - 1} + 4H_{\Delta-2} - 2H_{2\Delta} \right], \end{aligned}$$

where H_n are the harmonic numbers.

Breitenlohner-Freedman bound saturation - $M^2 = -4$

- For $M = -4$, we can solve the eom analytically:

$$\chi_1 = \frac{N_\Delta^2}{4L(\Delta - 1)^2} \left(\frac{-y^{2\Delta} + y^2}{y^2 - 1} \right).$$

- Note that here, in contrast to the $M = 0$ case, the second integration constant is fixed by requiring a sufficiently fast fall off towards the boundary.
- Computing the binding energy one finds

$$\gamma^{(x)} = -\frac{Y^2 \pi^2 N_\Delta^4}{8(\Delta - 1)^3}.$$

Numerical result - general case

- In case for general mass M and scaling dimension Δ an analytic solution is not available and one has to rely on numerical methods.

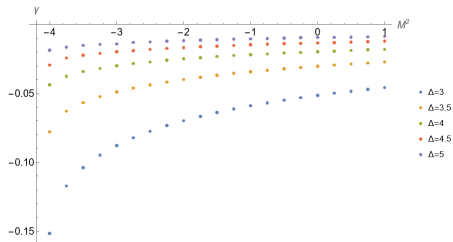


Figure: Binding energy for different scaling dimensions Δ and mass M^2 .

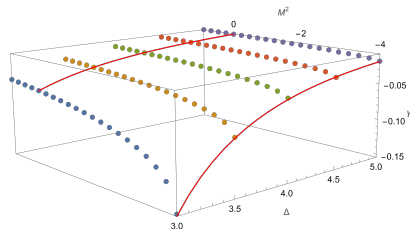


Figure: Binding energy for different scaling dimensions Δ and mass M^2 .

Tests in special cases

Flat Space Limit and BPS State

- In the case of massless χ we can compare to the flat space case.
- Taking $L \rightarrow \infty, \Delta \rightarrow \infty$ while keeping $m = \frac{\sqrt{\Delta(\Delta-4)}}{L}$ fixed, reproduces the WGC bound:

$$\sum_i q_i^2 g_i^2 - \frac{2}{3} m^2 \kappa^2 - \frac{Y^2}{4m^2} + \frac{1}{L} \left(\frac{a}{m} - bm \right) \geq 0.$$

¹Analysed in detail in [Ceresole, Dall'Agata, Kallosh, Van Proeyen '01]

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- We consider a $\mathcal{N} = 2$ 5d gauged supergravity¹ with a gravity multiplet, an hypermultiplet and a vector multiplet.
- The moduli space is given by $\frac{SU(2,1)}{SU(2) \times U(1)} \times O(1,1)$
- Using relations between the couplings:

$$\gamma^{(A_0)} + \gamma^{(A_1)} + \gamma^{(h)} + \gamma^{(a,b)} = -\gamma^{(\chi)}$$

and thus the total self-binding energy $\gamma_{z_1 z_1} = 0$ as expected for a BPS state.

¹Analysed in detail in [Ceresole, Dall'Agata, Kallosh, Van Proeyen '01]

Summary

Summary

- Positive binding energy conjecture as a natural formulation of the WGC applicable to AdS.
- Requirement of the existence of at least one particle with non-negative binding energy

$$\gamma \geq 0$$

leads to constraints on couplings for effective theories.

- Our results contribute towards sharpening quantum gravity constraints on effective theories in Anti-de Sitter spacetime.
- Outlook: Extending analysis to AdS_4 and AdS_3 , fermions, ...

Thank you!