

Holography in anomaly flow in orbifold gauge theory

Yutaka Hosotani, Osaka Univ

Universality in anomaly flow

by Y. Hosotani

PTEP 2022, 073B01 (2205.00154)

Anomaly flow by an Aharonov-Bohm phase

by Funatsu, Hatanaka, Hosotani, Orikasa, Yamatsu

PTEP 2022, 043B04 (2202.01393)

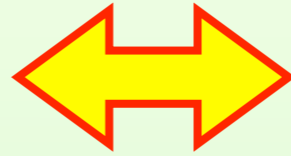
Corfu Summer Institute 2022

Workshop on Standard Model and Beyond

Corfu, Greece, 7 September 2022

Introduction

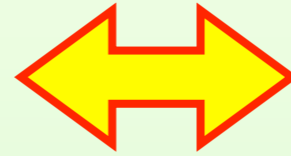
Chiral fermion



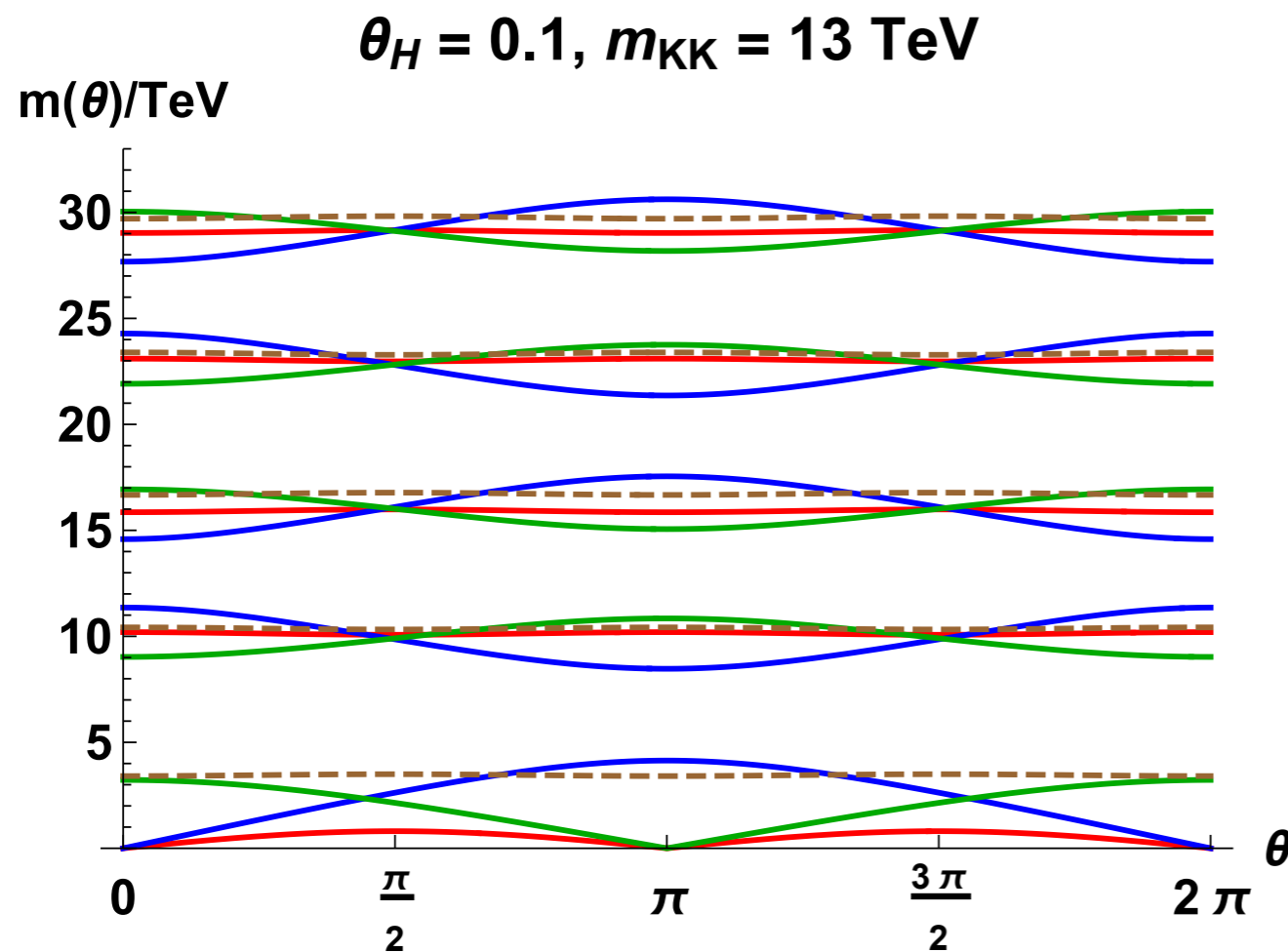
Chiral anomaly

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Chiral fermion



Chiral anomaly



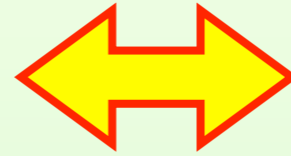
GUT inspired
 $SU(3) \times SO(5) \times U(1)$
GHU in RS

— W
— top
— darkF
- - - darkV

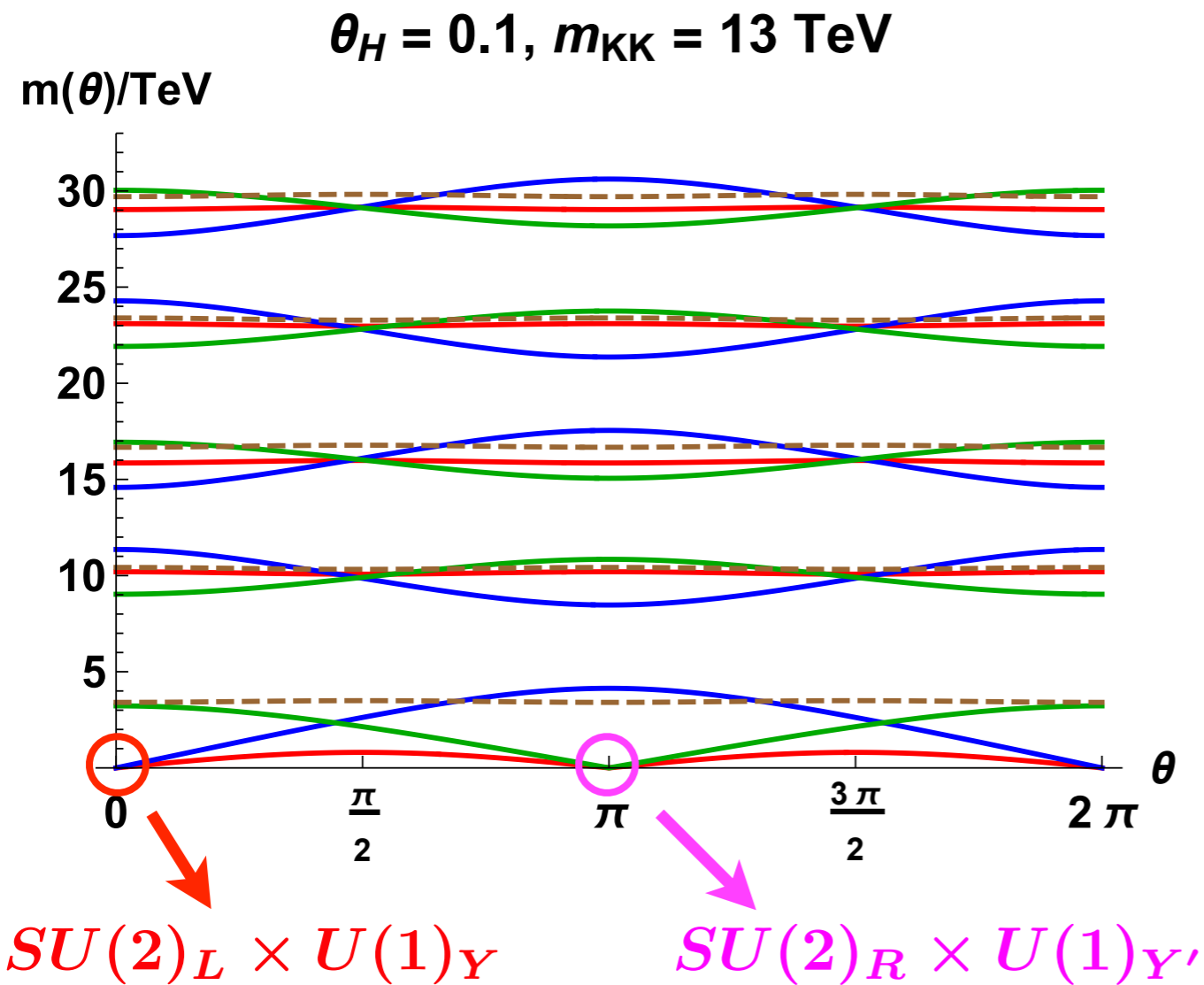
PRD.104.115018 (2021)

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Chiral fermion



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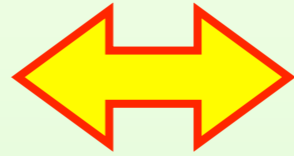
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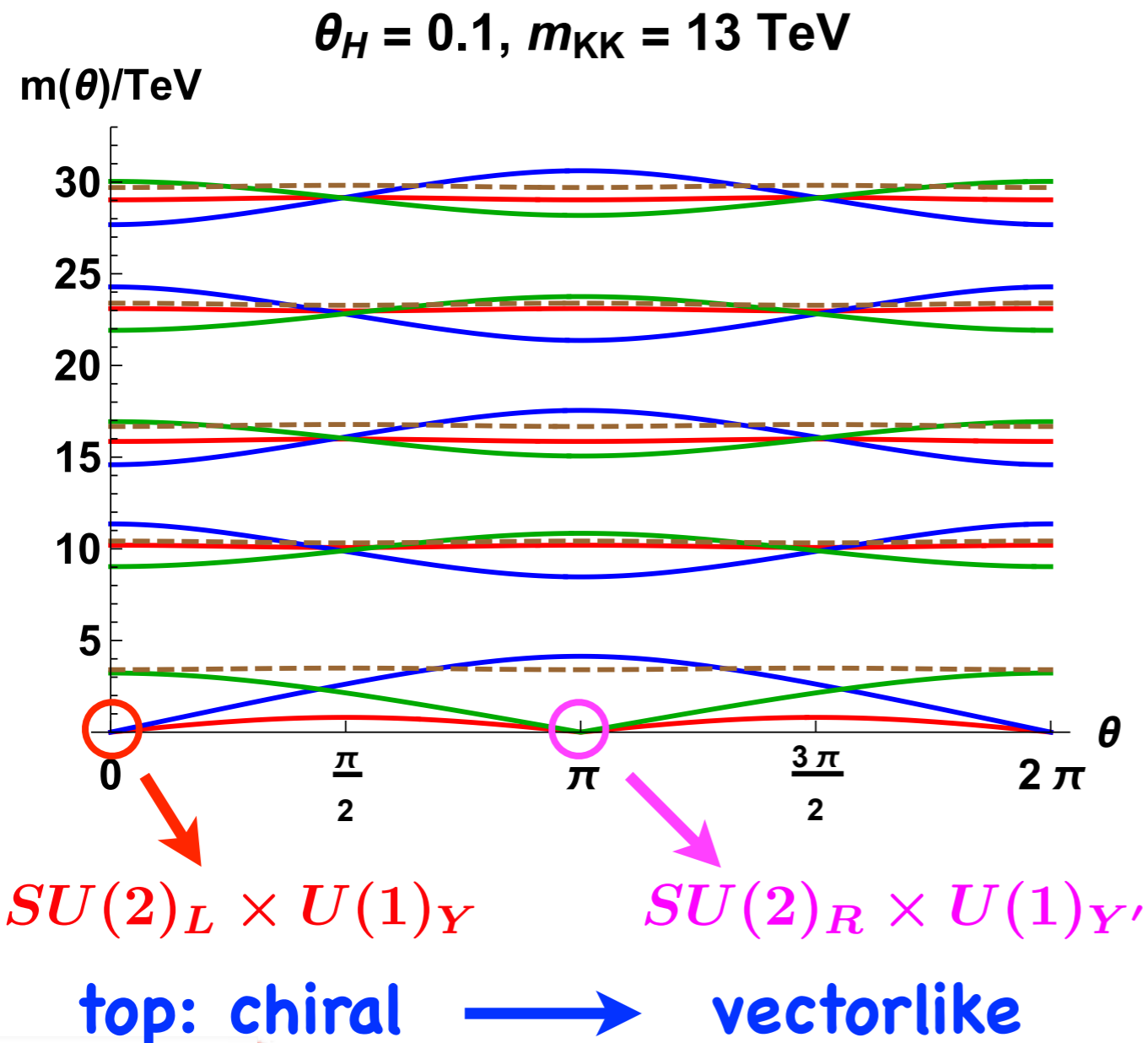
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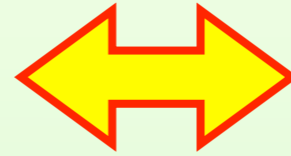


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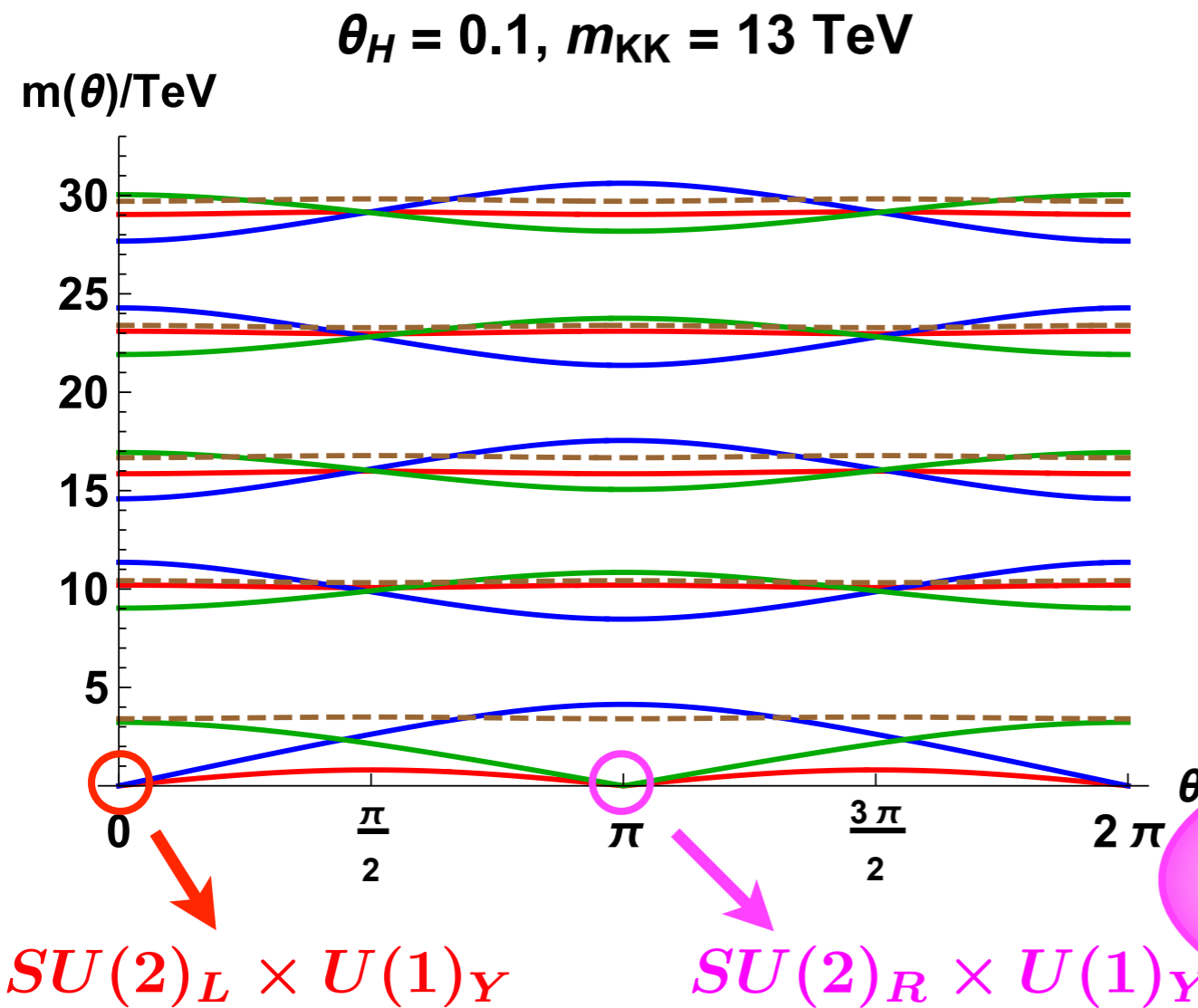
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Chiral anomaly



GUT inspired
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PRD.104.115018 (2021)

- W
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What is the fate of chiral anomaly?

top: chiral → vectorlike

$SU(2)$ GHU in $M^4 \times (S^1/Z_2)$

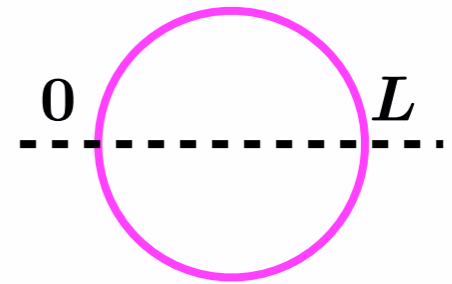
$$\begin{pmatrix} A_\mu \\ A_y \end{pmatrix} (x, y_j - y) = P_j \begin{pmatrix} A_\mu \\ -A_y \end{pmatrix} (x, y_j + y) P_j^{-1}$$

$$\Psi(x, y_j - y) = P_j \gamma^5 \Psi(x, y_j + y)$$

$$P_0 = P_1 = \tau^3$$

$$SU(2) \Rightarrow U(1)$$

$$(y_0, y_1) = (0, L)$$



$SU(2)$ GHU in $M^4 \times (S^1/Z_2)$

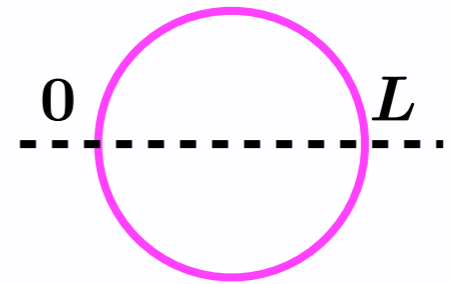
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AB phase θ_H

$$P e^{ig_A \int_0^{2L} dy \langle A_y \rangle} = e^{i\theta_H \tau^2}$$

$SU(2)$ GHU in $M^4 \times (S^1/Z_2)$

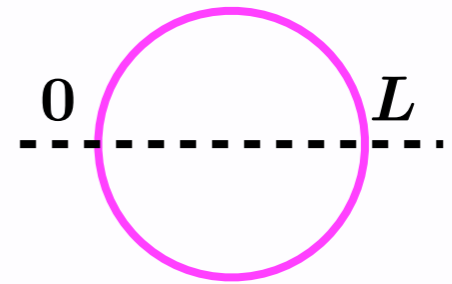
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$$\begin{pmatrix} A_\mu^1 \\ A_\mu^3 \end{pmatrix} = \sum_{n=-\infty}^{\infty} B_\mu^{(n)}(x) \frac{1}{\sqrt{\pi R}} \begin{pmatrix} \sin \frac{ny}{R} \\ \cos \frac{ny}{R} \end{pmatrix}$$

$$m_{B^{(n)}} = \frac{1}{R} \left| n + \frac{\theta_H}{\pi} \right|$$

Massless mode

$\theta_H = 0$	$B_\mu^{(0)}$
π	$B_\mu^{(-1)}$
2π	$B_\mu^{(-2)}$

Spectrum

$$\begin{pmatrix} u_R \\ d_R \end{pmatrix} = \sum_n \psi_R^{(n)}(x) \frac{1}{\sqrt{\pi R}} \begin{pmatrix} \cos \frac{ny}{R} \\ \sin \frac{ny}{R} \end{pmatrix}$$

$$\begin{pmatrix} u_L \\ d_L \end{pmatrix} = \sum_n \psi_L^{(n)}(x) \frac{1}{\sqrt{\pi R}} \begin{pmatrix} \sin \frac{ny}{R} \\ -\cos \frac{ny}{R} \end{pmatrix}$$

$$m_{\psi^{(n)}} = \frac{1}{R} \left| n + \frac{\theta_H}{2\pi} \right|$$

Massless mode

$\theta_H = 0$: u_R , d_L
chiral

Spectrum

$$\begin{pmatrix} u_R \\ d_R \end{pmatrix} = \sum_n \psi_R^{(n)}(x) \frac{1}{\sqrt{\pi R}} \begin{pmatrix} \cos \frac{ny}{R} \\ \sin \frac{ny}{R} \end{pmatrix}$$

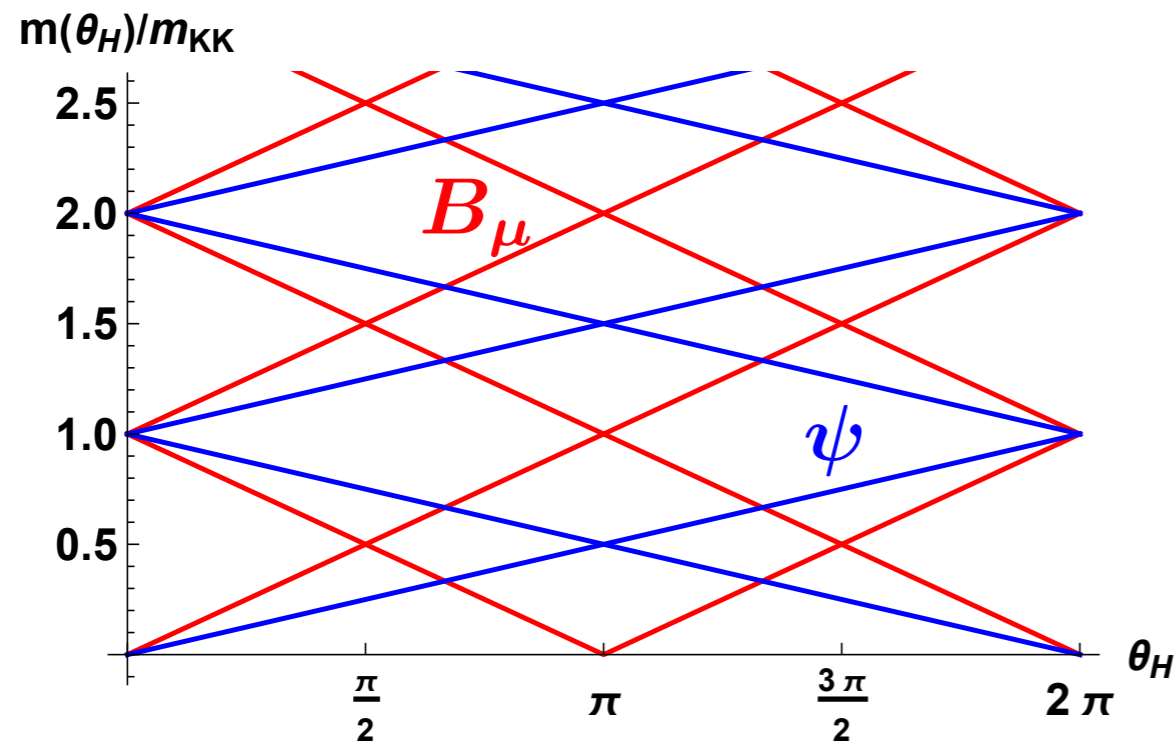
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Massless mode

$\theta_H = 0$: u_R , d_L
chiral

Spectrum in flat space



Spectrum

$$\begin{pmatrix} u_R \\ d_R \end{pmatrix} = \sum_n \psi_R^{(n)}(x) \frac{1}{\sqrt{\pi R}} \begin{pmatrix} \cos \frac{ny}{R} \\ \sin \frac{ny}{R} \end{pmatrix}$$

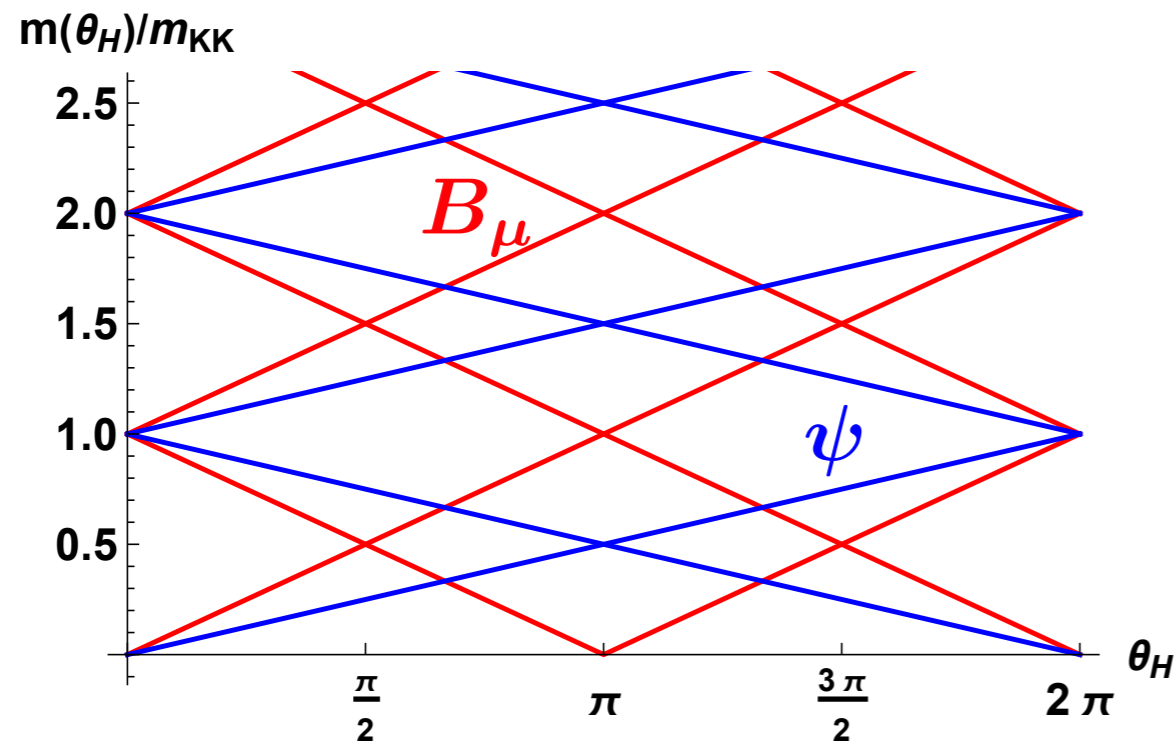
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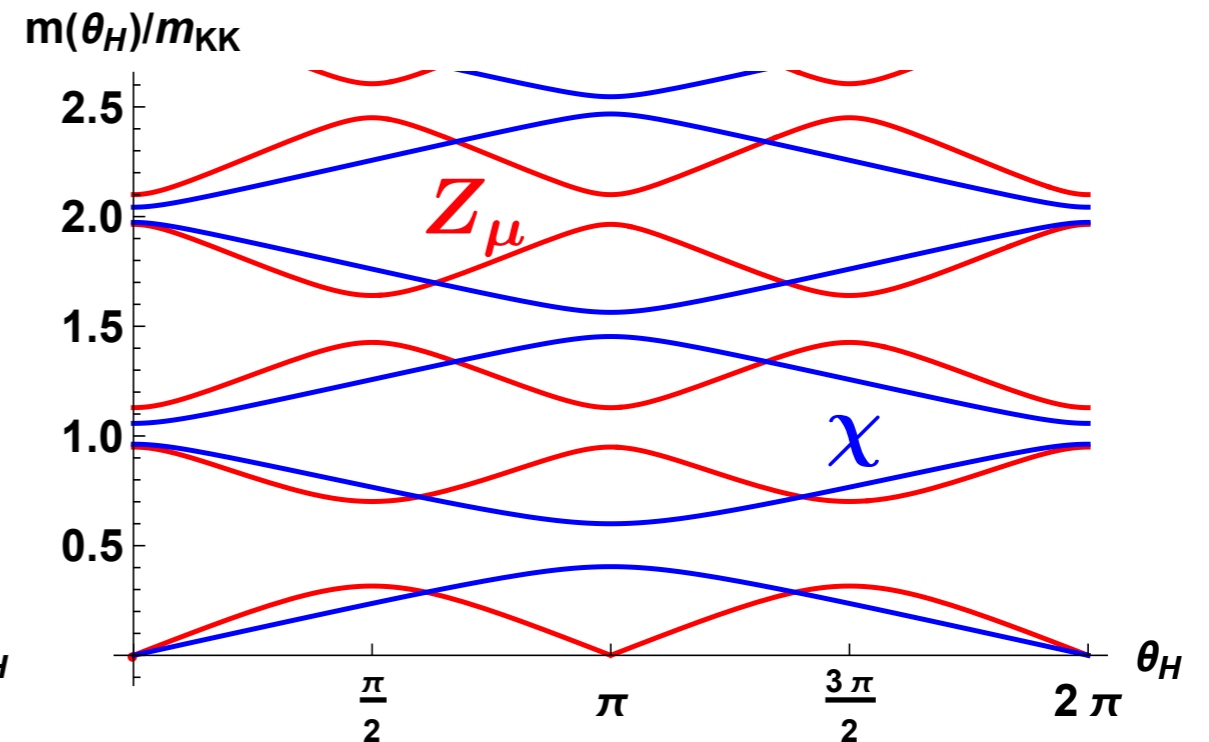
Massless mode

$\theta_H = 0 : u_R, d_L$
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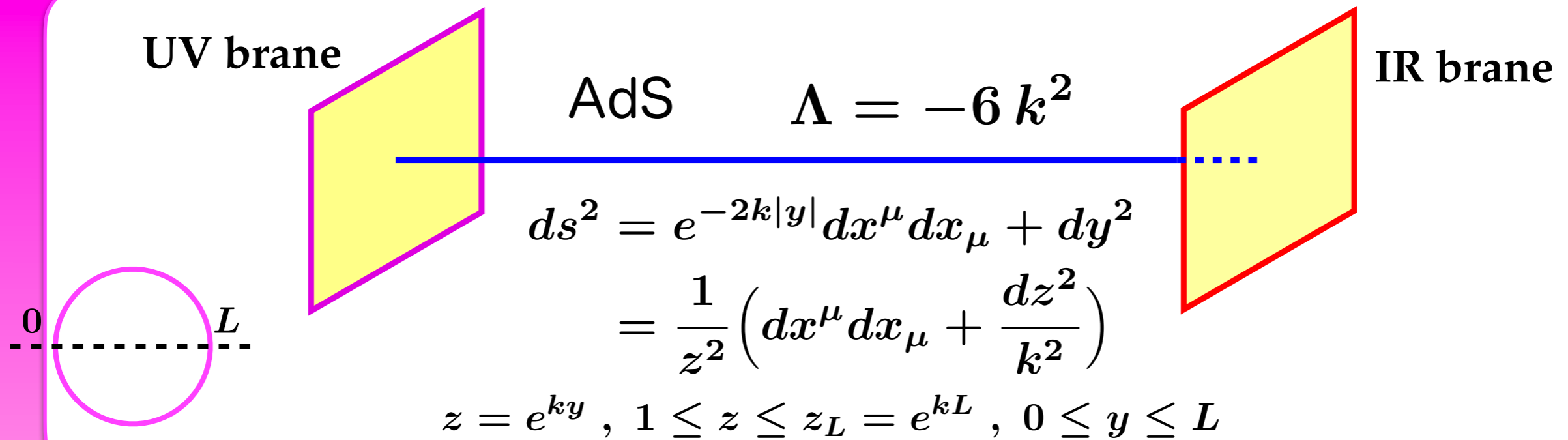
Spectrum in flat space



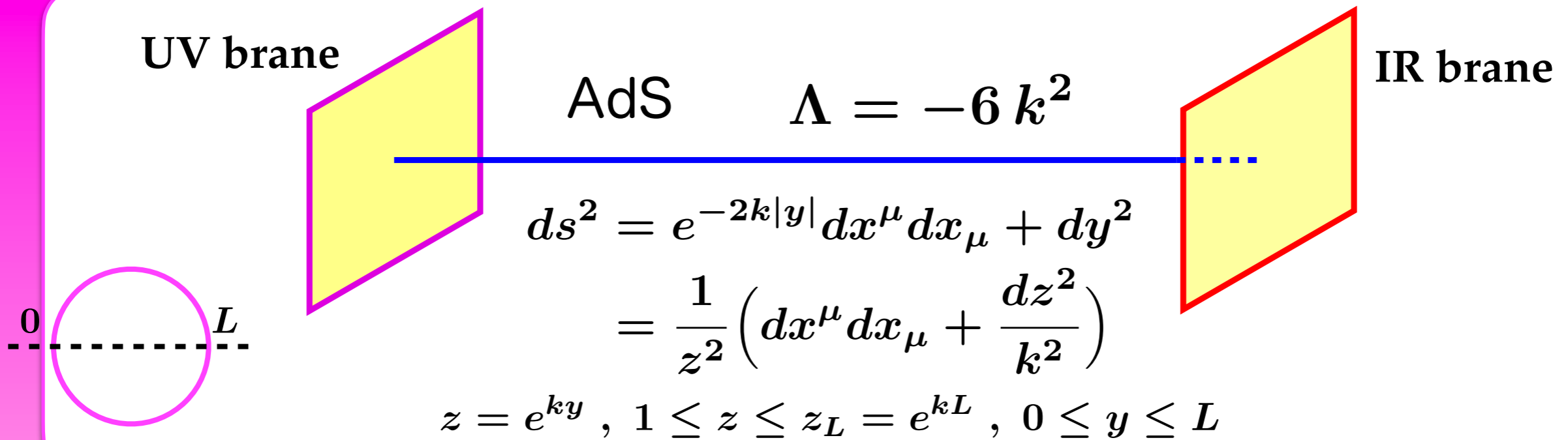
Spectrum in RS, $z_L=10$



$SU(2)$ GHU in Randall-Sundrum



$SU(2)$ GHU in Randall-Sundrum



$$\begin{pmatrix} A_\mu^1 \\ A_\mu^3 \end{pmatrix} = \frac{1}{\sqrt{L}} \sum_{n=0}^{\infty} Z_\mu^{(n)}(x) \begin{pmatrix} h_n(y) \\ k_n(y) \end{pmatrix}$$

$$\tilde{\Psi} = \frac{1}{z^4} \Psi \quad \begin{pmatrix} \check{u}_R \\ \check{d}_R \end{pmatrix} = \sqrt{k} \sum_{n=0}^{\infty} \chi_R^{(n)}(x) \begin{pmatrix} f_{Rn}(y) \\ g_{Rn}(y) \end{pmatrix}$$

$$\begin{pmatrix} \check{u}_L \\ \check{d}_L \end{pmatrix} = \sqrt{k} \sum_{n=0}^{\infty} \chi_L^{(n)}(x) \begin{pmatrix} f_{Ln}(y) \\ g_{Ln}(y) \end{pmatrix}$$

Spectrum

$$Z_{\mu}^{(n)} : SC'(1; \lambda_n) + \lambda_n \sin^2 \theta_H = 0$$

$$C(z; \lambda) = \frac{\pi}{2} \lambda z z_L F_{1,0}(\lambda z, \lambda z_L)$$

$$S(z; \lambda) = -\frac{\pi}{2} \lambda z F_{1,1}(\lambda z, \lambda z_L)$$

$$F_{\alpha,\beta}(u, v) \equiv J_{\alpha}(u)Y_{\beta}(v) - Y_{\alpha}(u)J_{\beta}(v)$$

Spectrum

$$Z_{\mu}^{(n)} : SC'(1; \lambda_n) + \lambda_n \sin^2 \theta_H = 0$$

$$\chi^{(n)} : S_L S_R(1; \lambda_n, c) + \sin^2 \frac{\theta_H}{2} = 0$$

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$$F_{\alpha,\beta}(u, v) \equiv J_{\alpha}(u) Y_{\beta}(v) - Y_{\alpha}(u) J_{\beta}(v)$$

$$\begin{pmatrix} C_L \\ S_L \end{pmatrix} (z; \lambda, c) = \pm \frac{\pi}{2} \lambda \sqrt{z z_L} F_{c+\frac{1}{2}, c \mp \frac{1}{2}}(\lambda z, \lambda z_L)$$

Spectrum

$$Z_{\mu}^{(n)} : SC'(1; \lambda_n) + \lambda_n \sin^2 \theta_H = 0$$

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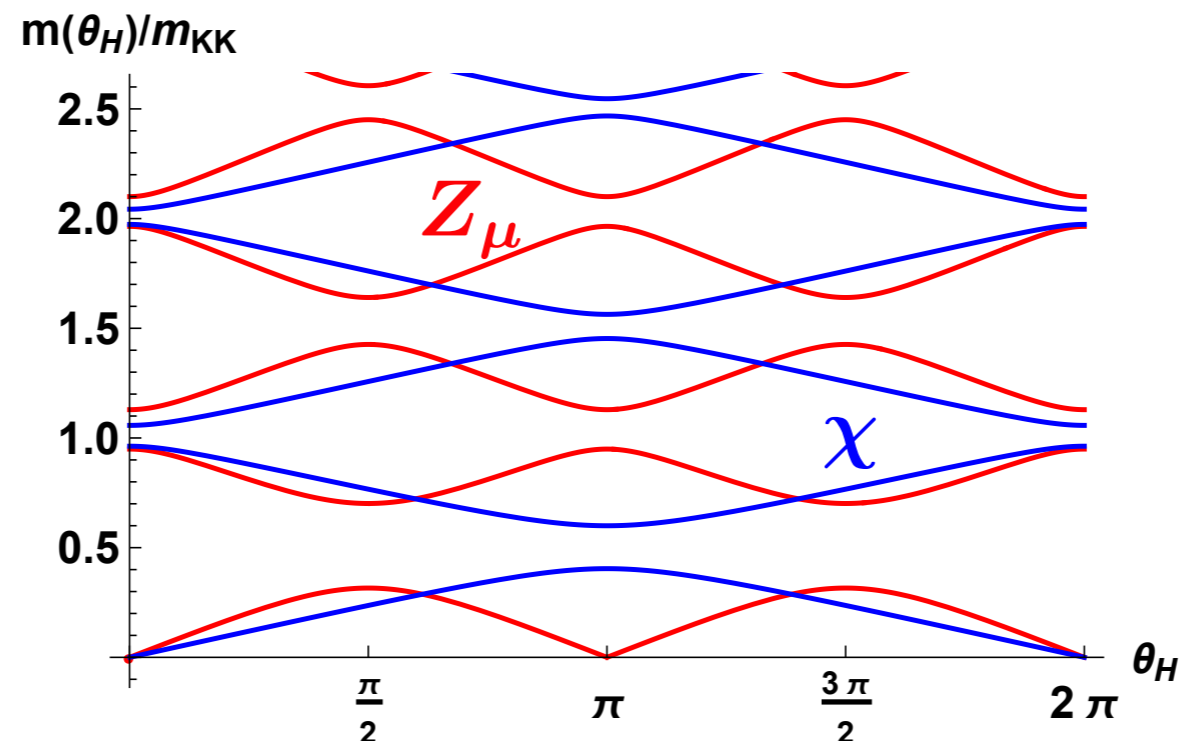
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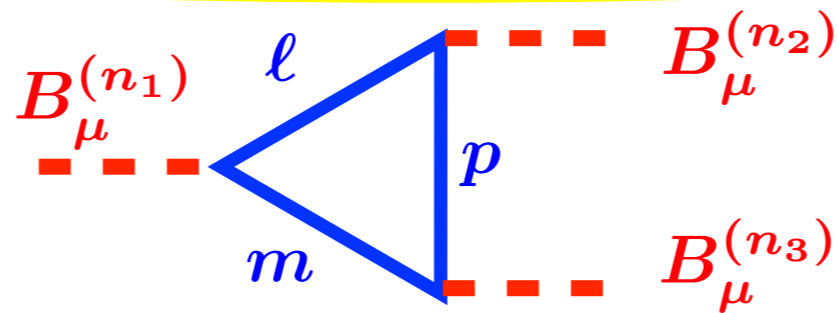
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Spectrum in RS, $z_L=10$

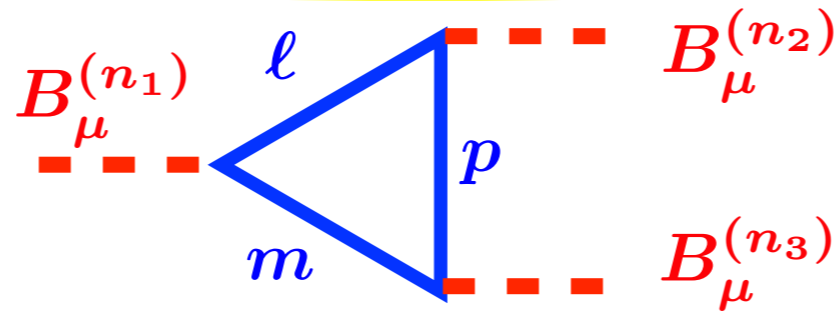


Gauge couplings and anomaly in flat space



$$\frac{g_4}{2} \sum_{n,m,\ell=-\infty}^{\infty} B_\mu^{(n)} \left\{ \underset{\delta_{n,m+\ell}}{s_{nml}^R} \psi_R^{(m)\dagger} \bar{\sigma}^\mu \psi_R^{(\ell)} + \underset{\delta_{n,m+\ell}}{s_{nml}^L} \psi_L^{(m)\dagger} \sigma^\mu \psi_L^{(\ell)} \right\}$$

Gauge couplings and anomaly in flat space



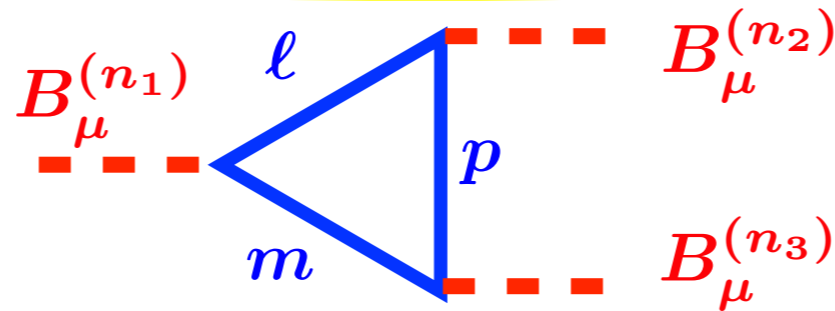
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Anomaly coefficient

$$\partial_\mu j_{(n)}^\mu \Rightarrow - \left(\frac{g_4}{2} \right)^3 \sum_\ell \sum_m \frac{b_{n\ell m}}{16\pi^2} B_{\mu\nu}^{(\ell)} \tilde{B}^{(m)\mu\nu}$$

$$b_{n_1 n_2 n_3} = \sum_{m,\ell,p} \left\{ s_{n_1 m \ell}^R s_{n_2 \ell p}^R s_{n_3 p m}^R + s_{n_1 m \ell}^L s_{n_2 \ell p}^L s_{n_3 p m}^L \right\}$$

Gauge couplings and anomaly in flat space



$$\frac{g_4}{2} \sum_{n,m,\ell=-\infty}^{\infty} B_\mu^{(n)} \left\{ s_{nml}^R \psi_R^{(m)\dagger} \bar{\sigma}^\mu \psi_R^{(\ell)} + s_{nml}^L \psi_L^{(m)\dagger} \sigma^\mu \psi_L^{(\ell)} \right\}$$

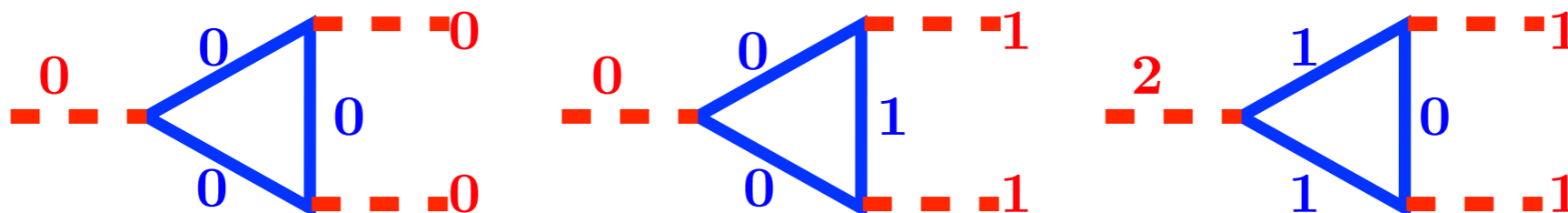
$\delta_{n,m+\ell}$ $\delta_{n,m+\ell}$

Anomaly coefficient

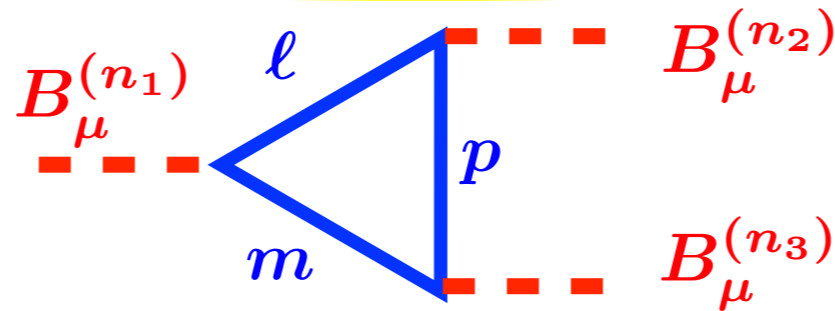
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$$b_{n_1 n_2 n_3} = \sum_{m,\ell,p} \left\{ s_{n_1 m \ell}^R s_{n_2 \ell p}^R s_{n_3 p m}^R + s_{n_1 m \ell}^L s_{n_2 \ell p}^L s_{n_3 p m}^L \right\}$$

$$= \begin{cases} 2 & \text{for } n_1 + n_2 + n_3 = \text{even} \\ 0 & \text{for } n_1 + n_2 + n_3 = \text{odd} \end{cases}$$



Gauge couplings and anomaly in flat space



$$\frac{g_4}{2} \sum_{n,m,\ell=-\infty}^{\infty} B_\mu^{(n)} \left\{ s_{nml}^R \psi_R^{(m)\dagger} \bar{\sigma}^\mu \psi_R^{(\ell)} + s_{nml}^L \psi_L^{(m)\dagger} \sigma^\mu \psi_L^{(\ell)} \right\}$$

$\delta_{n,m+\ell}$ $\delta_{n,m+\ell}$

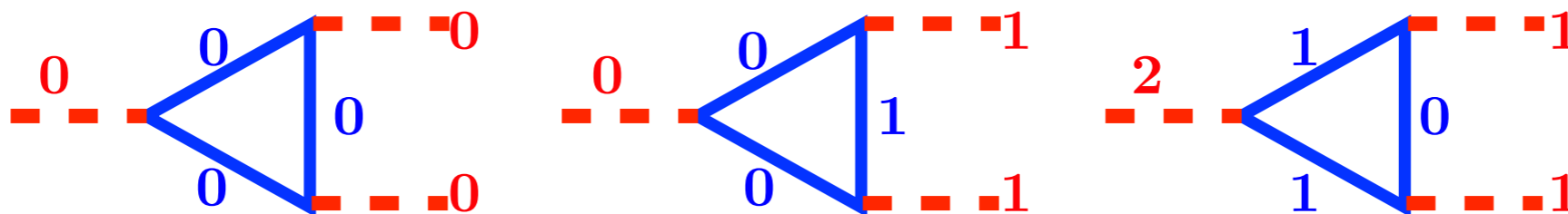
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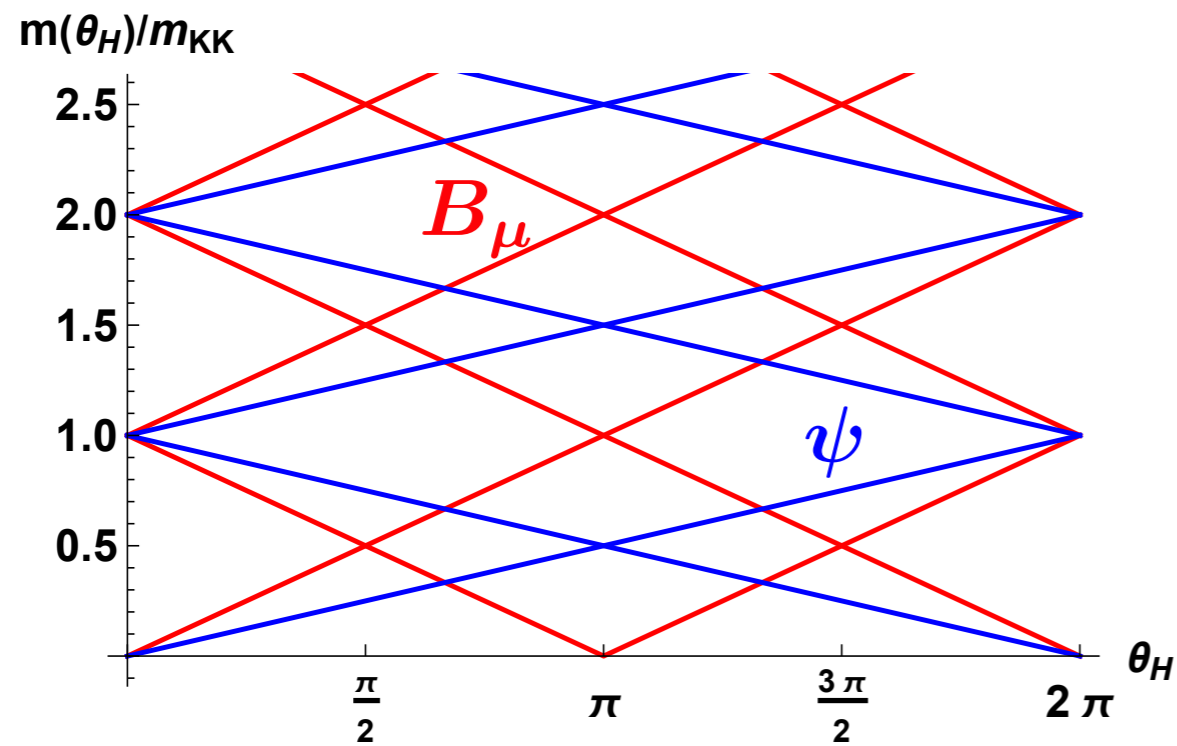
$$= \begin{cases} 2 \\ 0 \end{cases} \text{ for } n_1 + n_2 + n_3 = \begin{cases} \text{even} \\ \text{odd} \end{cases}$$

θ_H -independent?

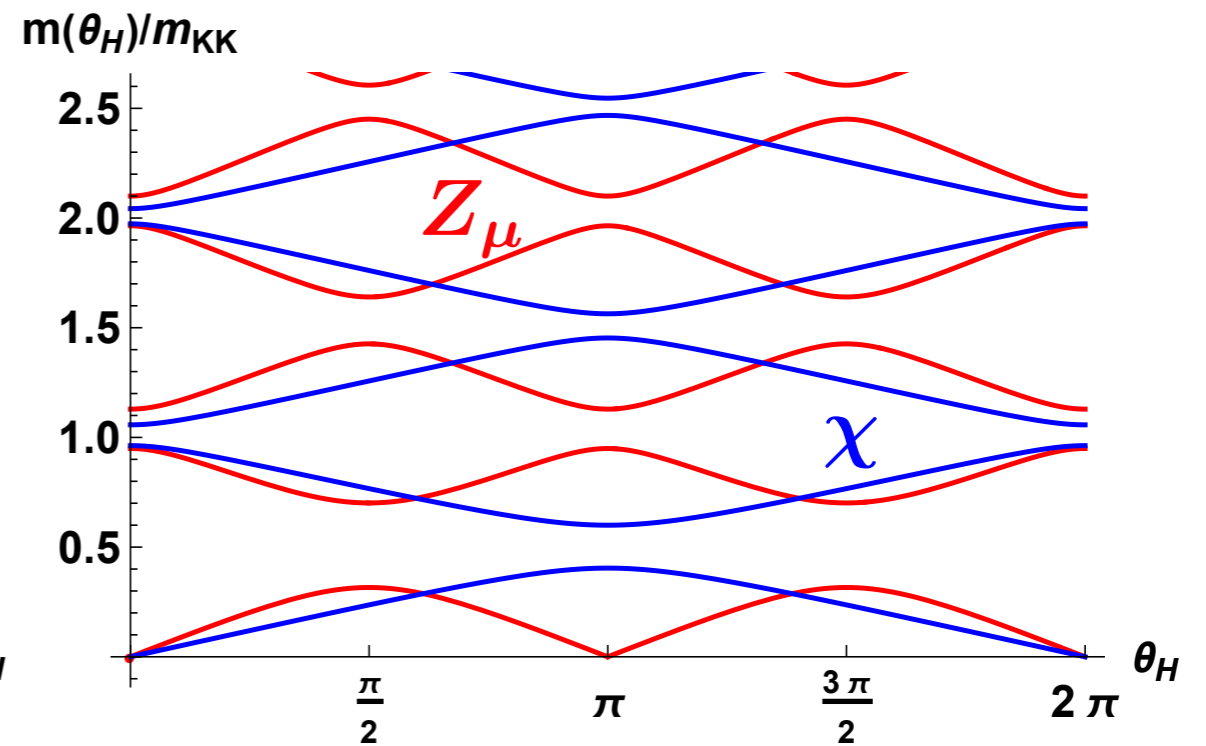


Anomaly flow

Spectrum in flat space

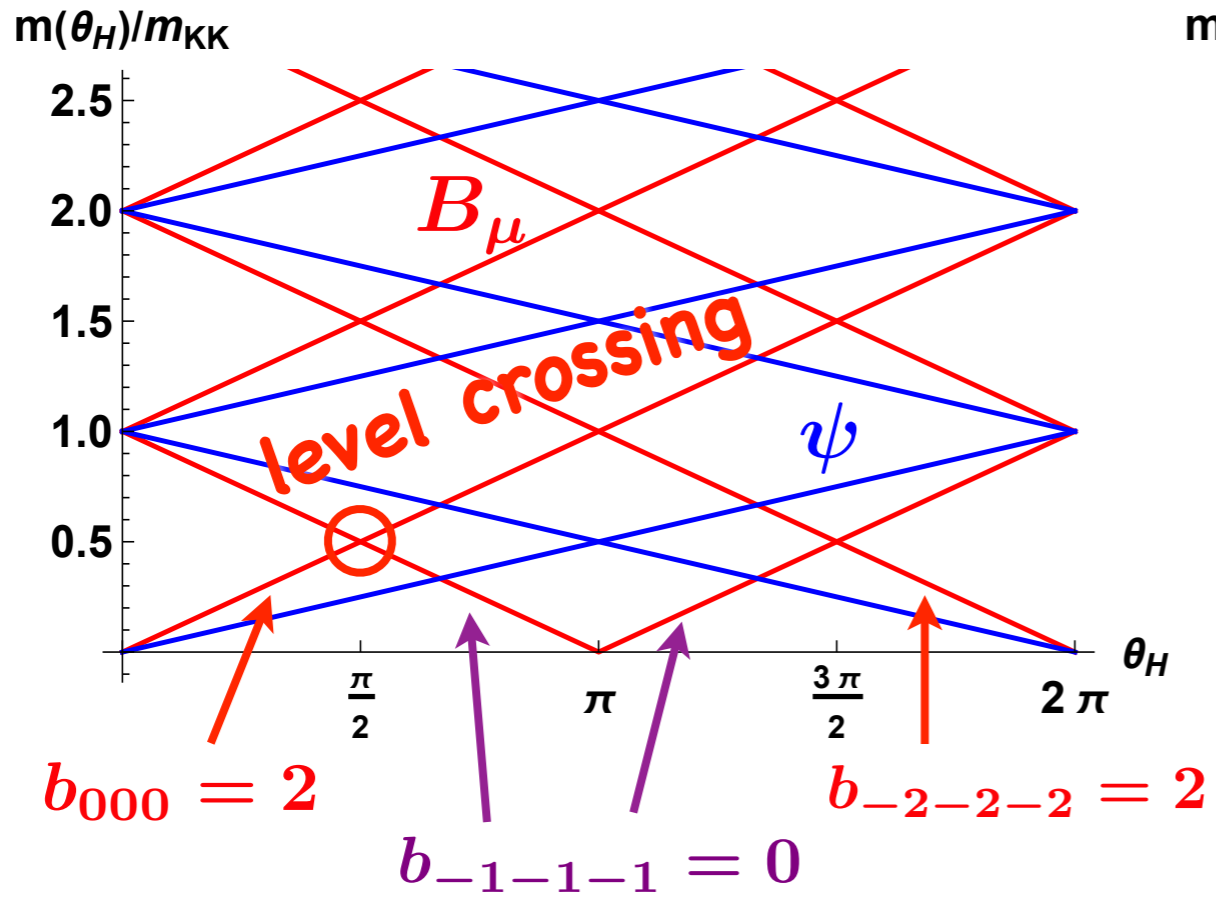


Spectrum in RS, $z_L=10$

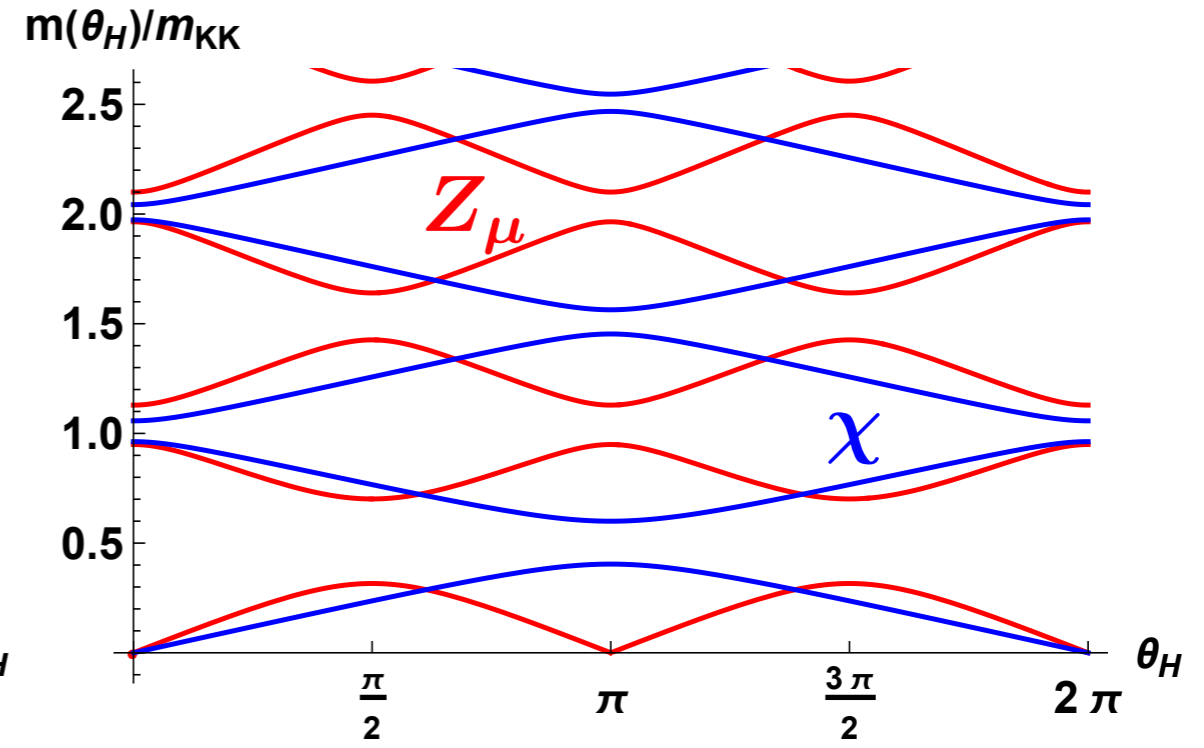


Anomaly flow

Spectrum in flat space

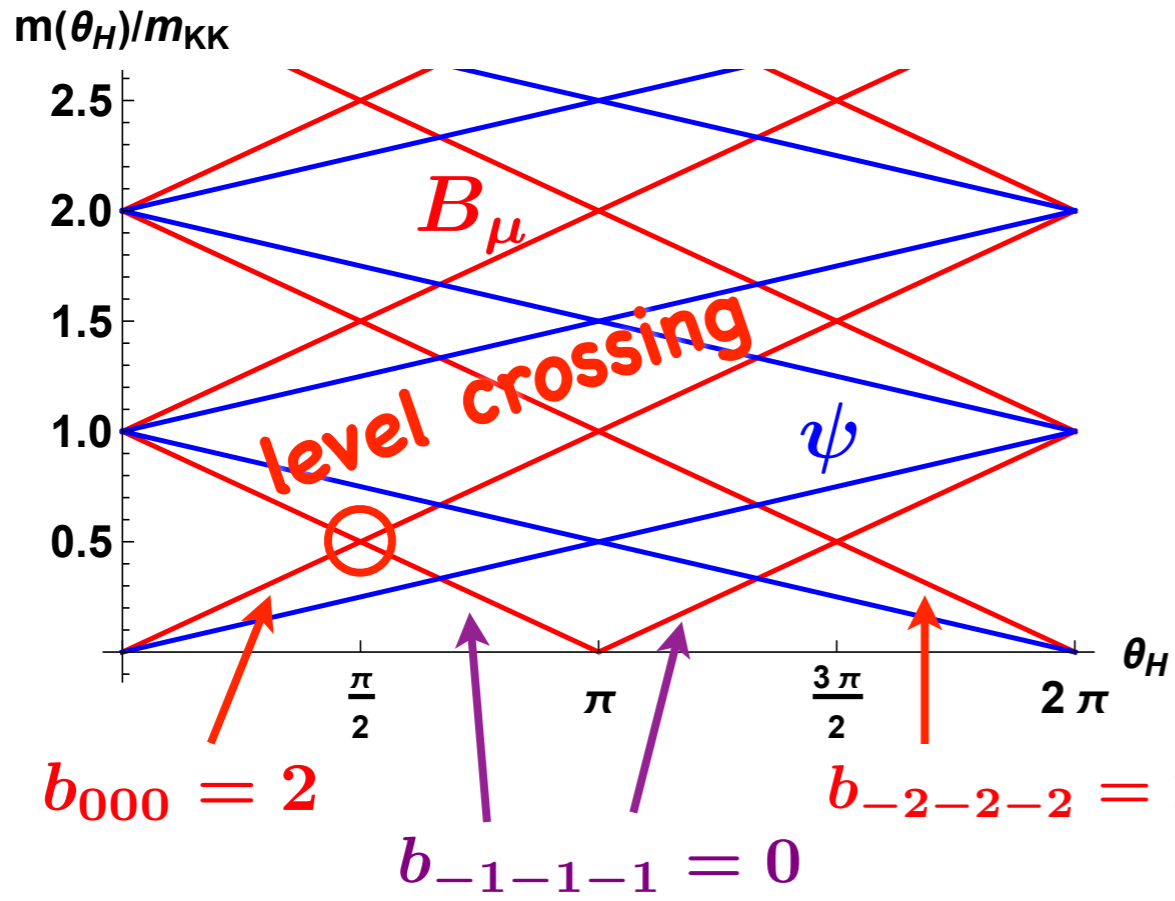


Spectrum in RS, $z_L=10$

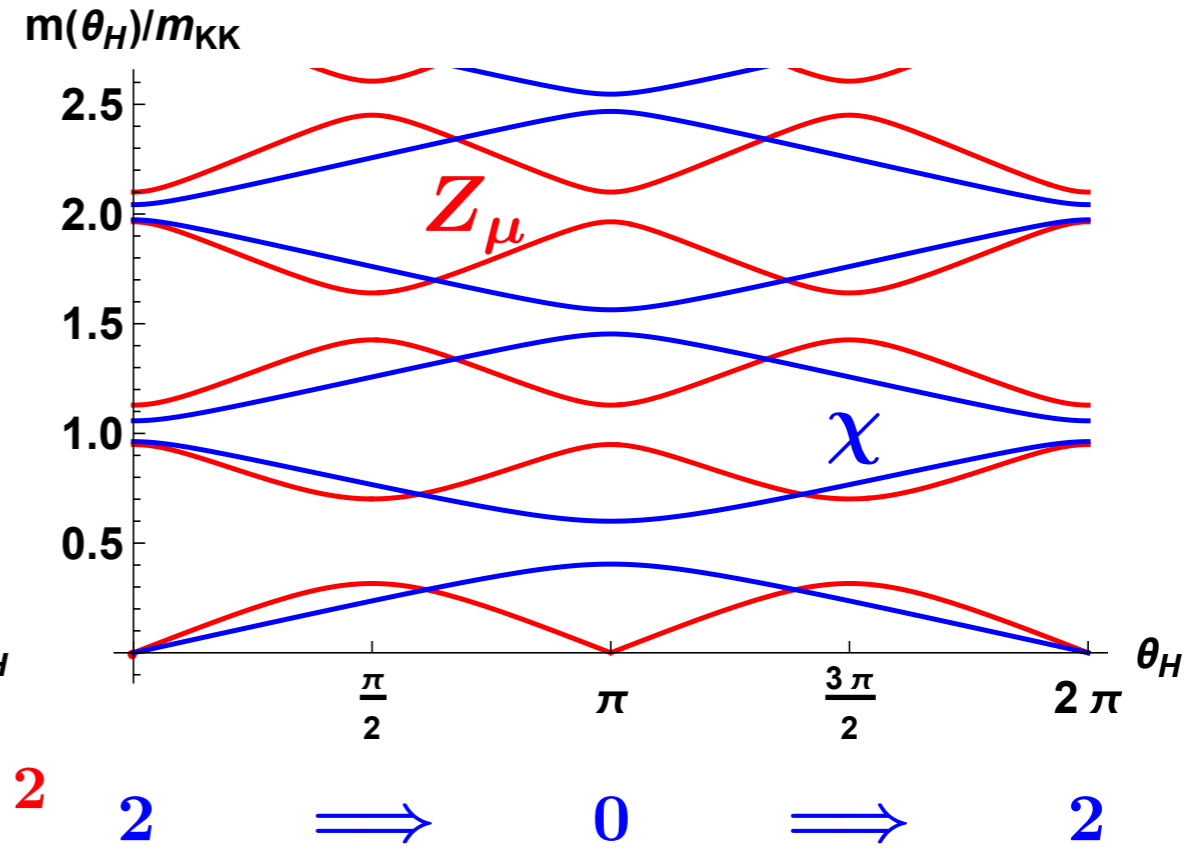


Anomaly flow

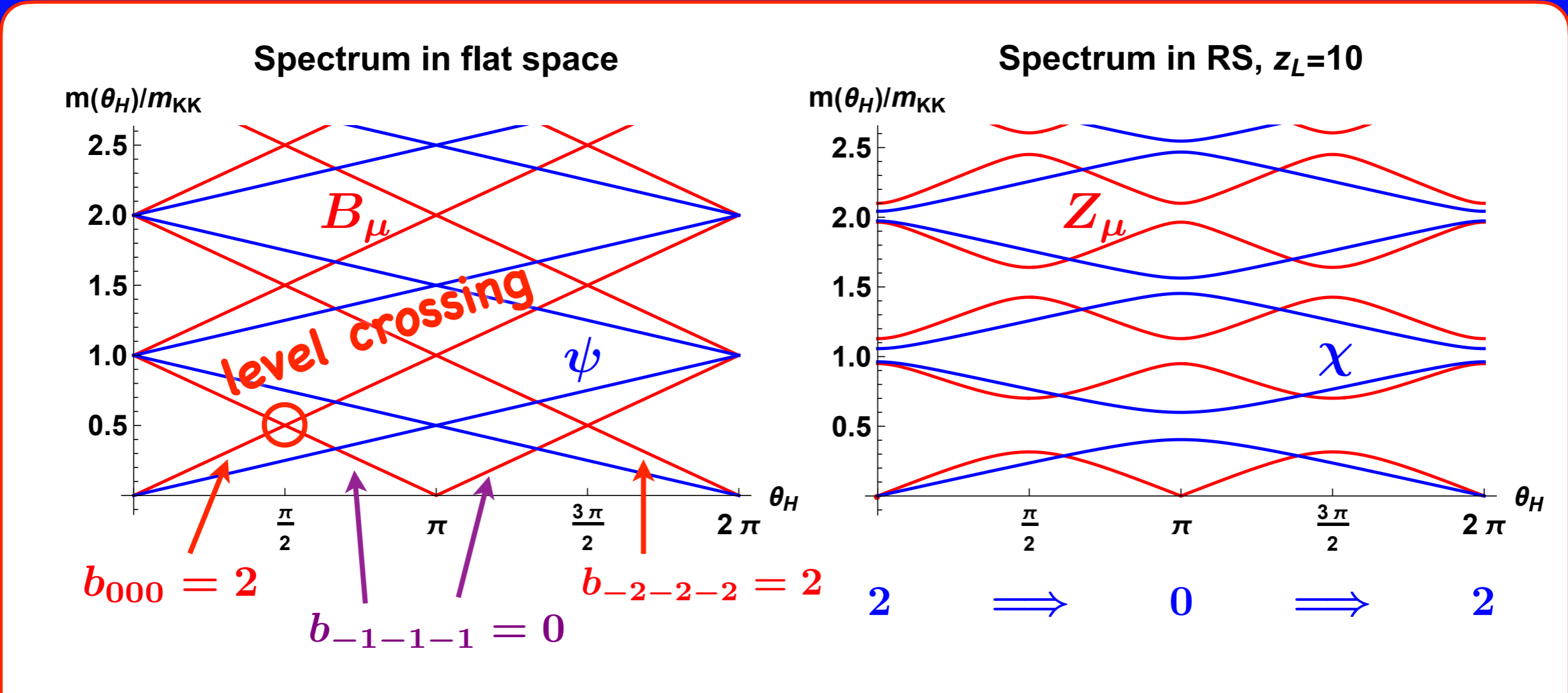
Spectrum in flat space



Spectrum in RS, $z_L=10$



Anomaly flow



Anomaly must flow with θ_H !

Couplings in RS

$$\frac{g_4}{2} \sum_{n,\ell,m=0}^{\infty} Z_{\mu}^{(n)} \left\{ t_{n\ell m}^R \chi_R^{(\ell)\dagger} \bar{\sigma}^{\mu} \chi_R^{(m)} + t_{n\ell m}^L \chi_L^{(\ell)\dagger} \sigma^{\mu} \chi_L^{(m)} \right\}$$

$$t_{n\ell m}^R = \frac{k}{2} \int_{-L}^L dy e^{k|y|} \left\{ h_n (f_{R\ell}^* g_{Rm} + g_{R\ell}^* f_{Rm}) + k_n (f_{R\ell}^* f_{Rm} - g_{R\ell}^* g_{Rm}) \right\}$$

$$t_{n\ell m}^L = -\frac{k}{2} \int_{-L}^L dy e^{k|y|} \left\{ f_{Rm} \rightarrow f_{Lm}, g_{Rm} \rightarrow g_{Lm} \right\}$$

Couplings in RS

$$\frac{g_4}{2} \sum_{n,\ell,m=0}^{\infty} Z_{\mu}^{(n)} \left\{ t_{n\ell m}^R \chi_R^{(\ell)\dagger} \bar{\sigma}^{\mu} \chi_R^{(m)} + t_{n\ell m}^L \chi_L^{(\ell)\dagger} \sigma^{\mu} \chi_L^{(m)} \right\}$$

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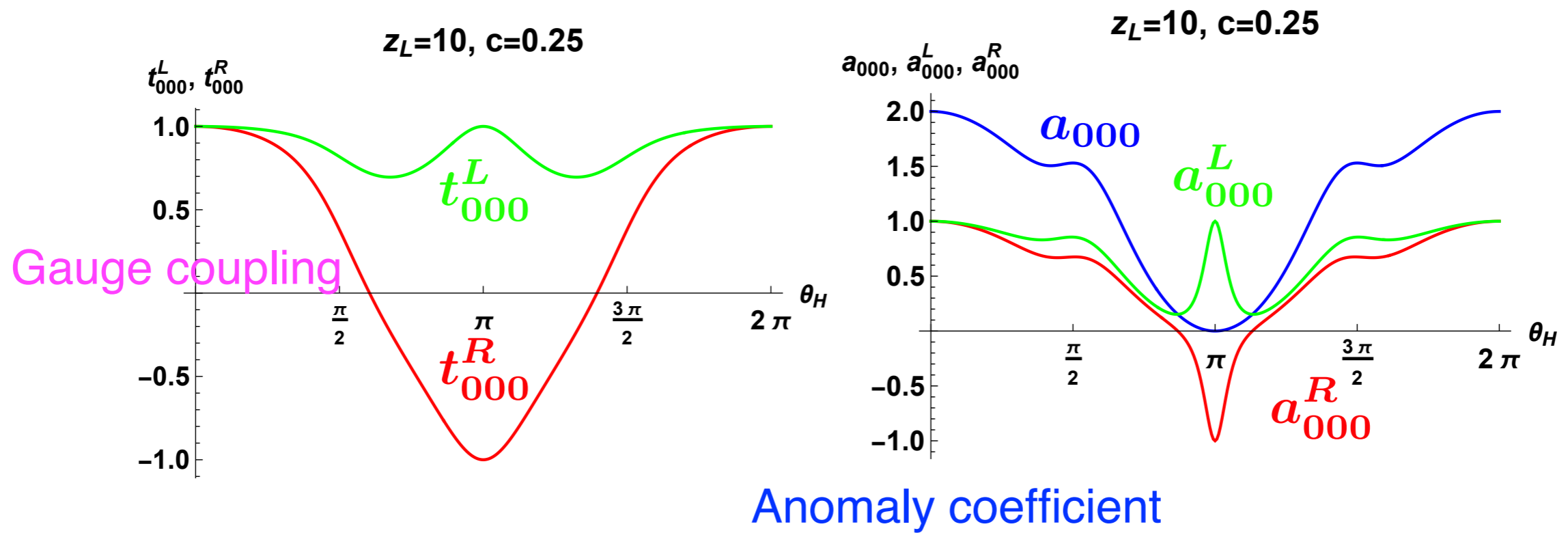
$$t_{n\ell m}^L = -\frac{k}{2} \int_{-L}^L dy e^{k|y|} \left\{ f_{Rm} \rightarrow f_{Lm}, g_{Rm} \rightarrow g_{Lm} \right\}$$

Anomaly coefficient

$$a_{n_1 n_2 n_3} = \sum_{m,\ell,p} \left\{ t_{n_1 m \ell}^R t_{n_2 \ell p}^R t_{n_3 p m}^R + t_{n_1 m \ell}^L t_{n_2 \ell p}^L t_{n_3 p m}^L \right\}$$

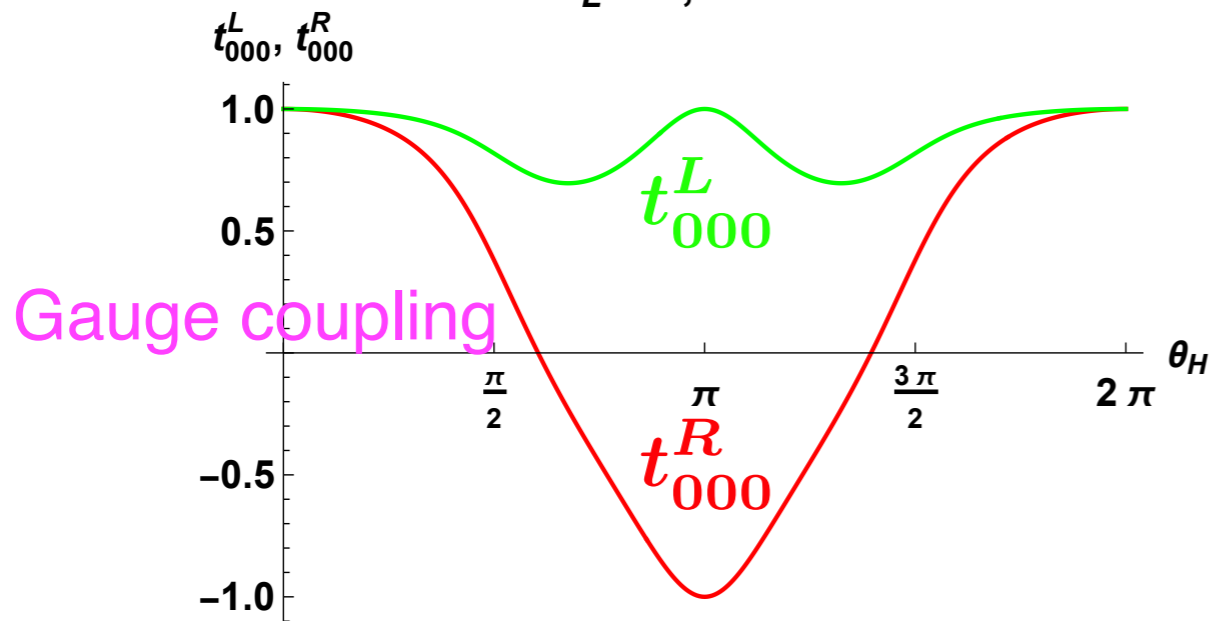
Note: $t_{n\ell m}^{R/L} = t_{n\ell m}^{R/L}(\theta_H, z_L, c)$

Anomaly flows !

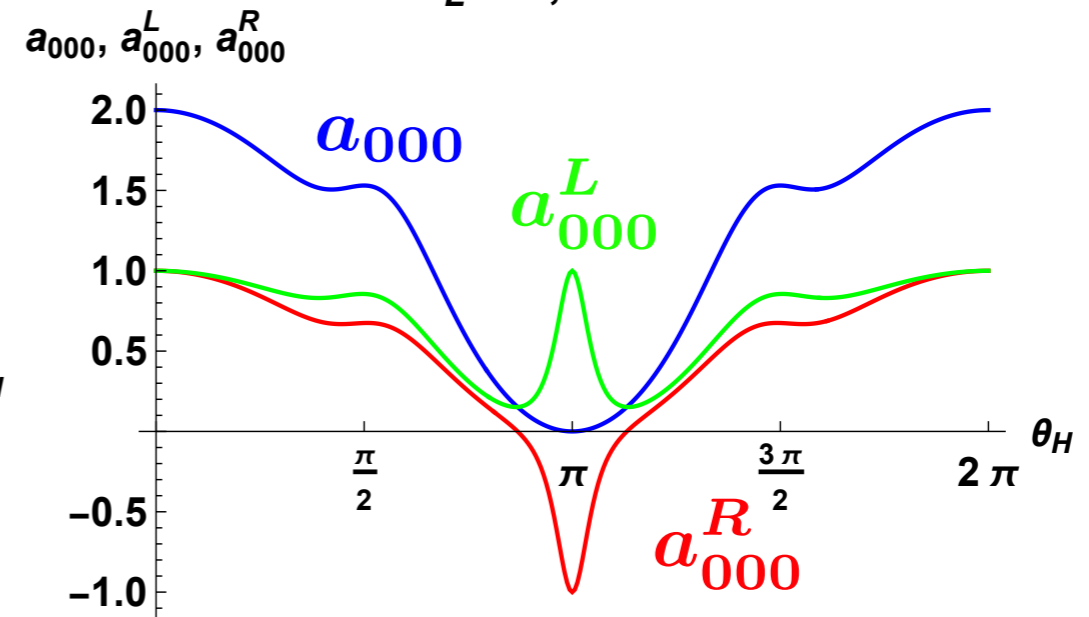


Anomaly flows !

$z_L=10, c=0.25$

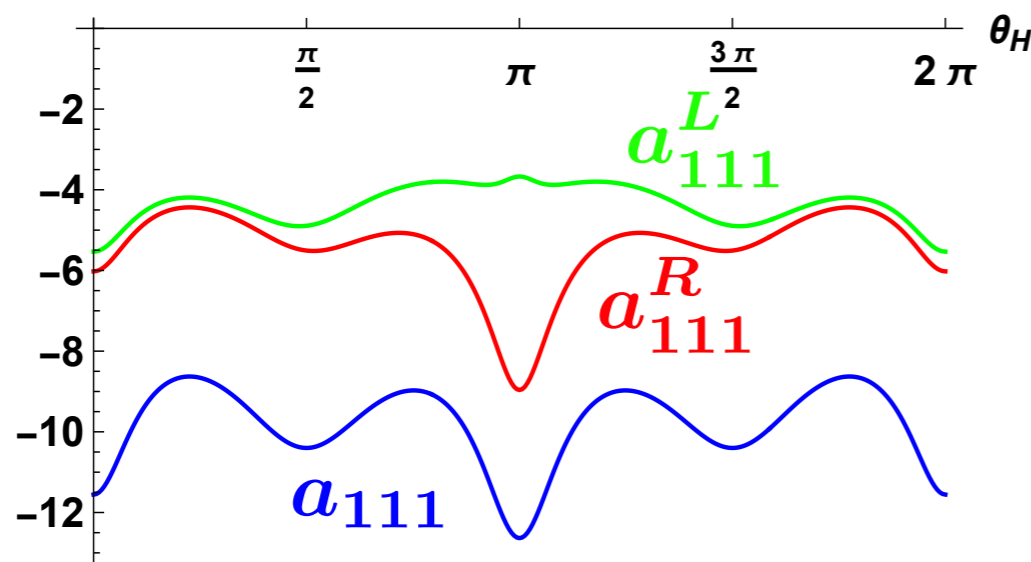


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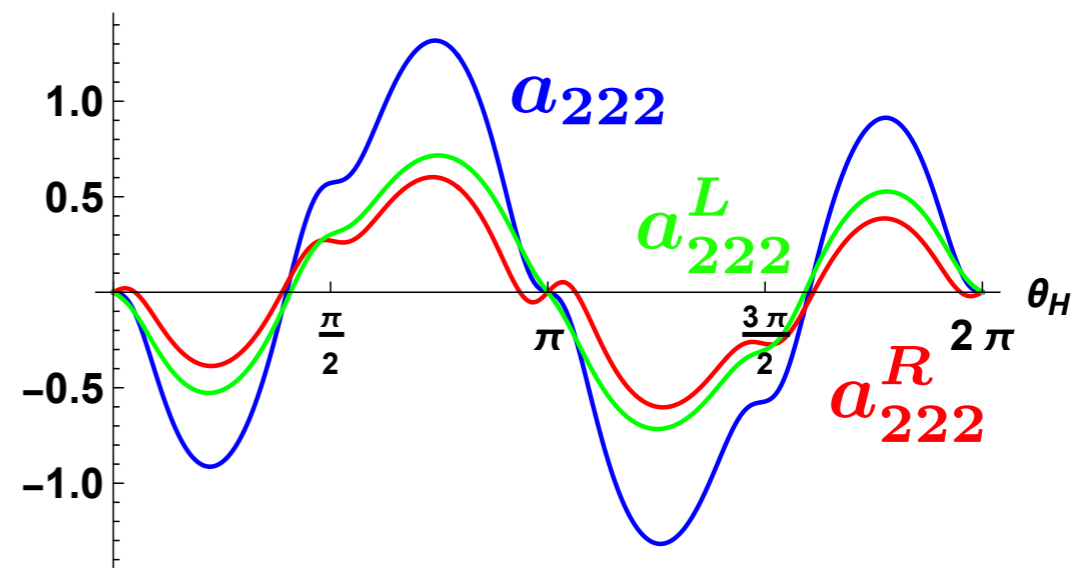


Anomaly coefficient

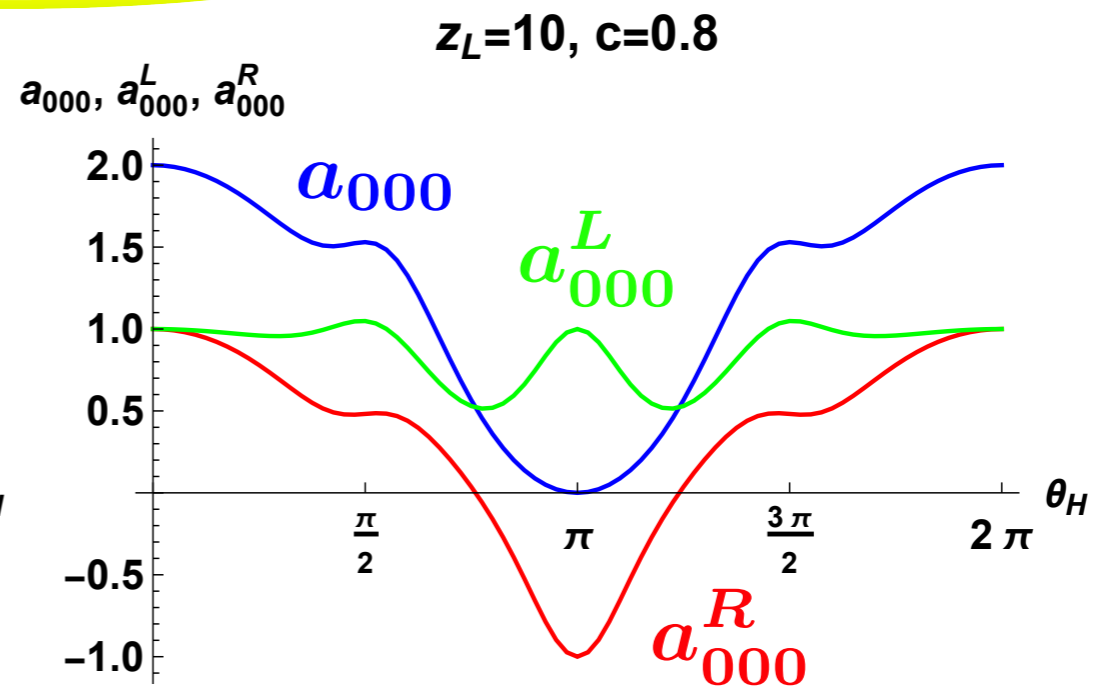
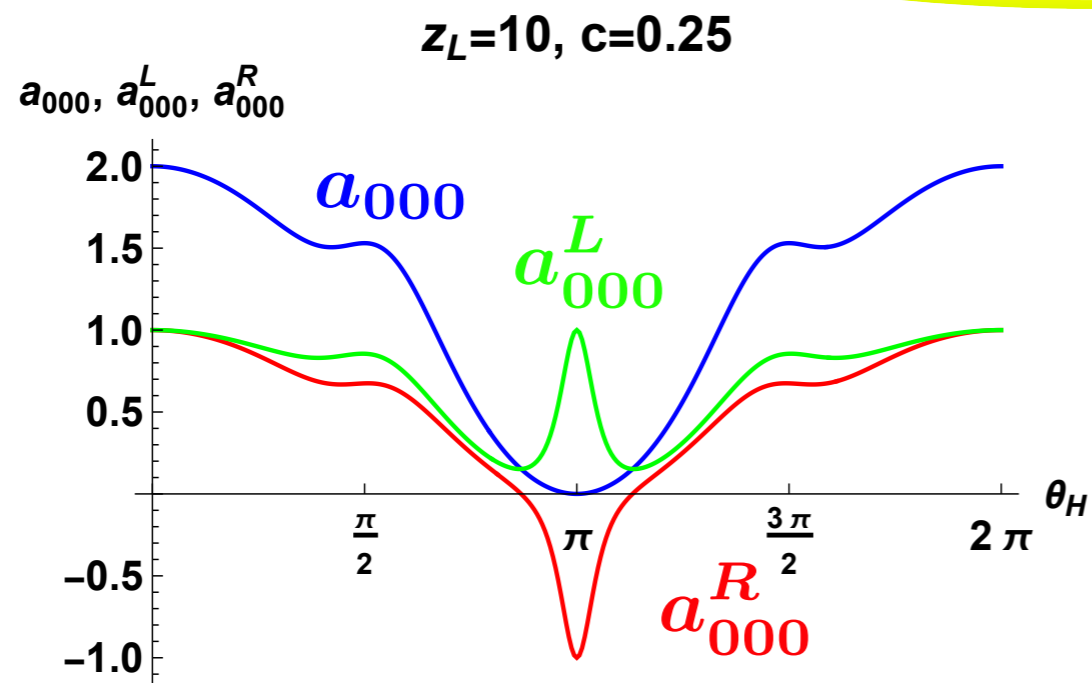
$a_{111}, a_{111}^L, a_{111}^R$



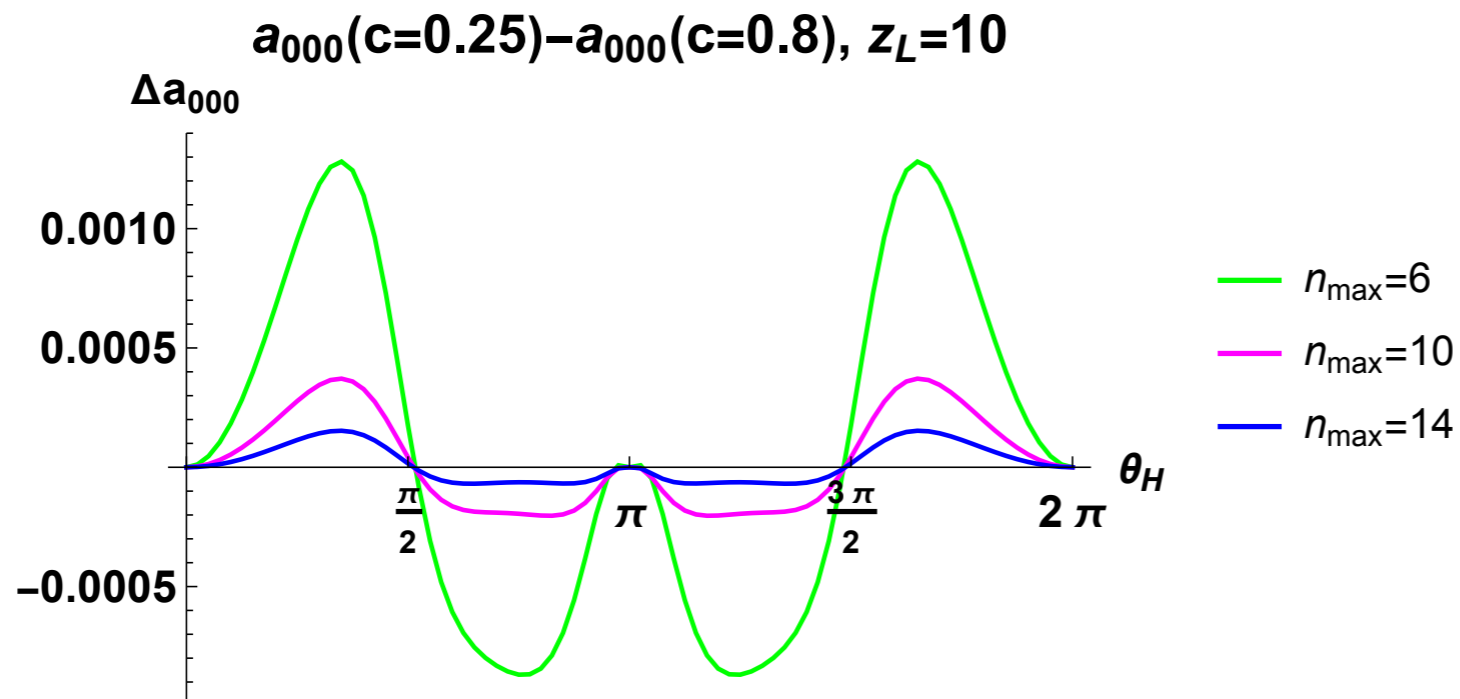
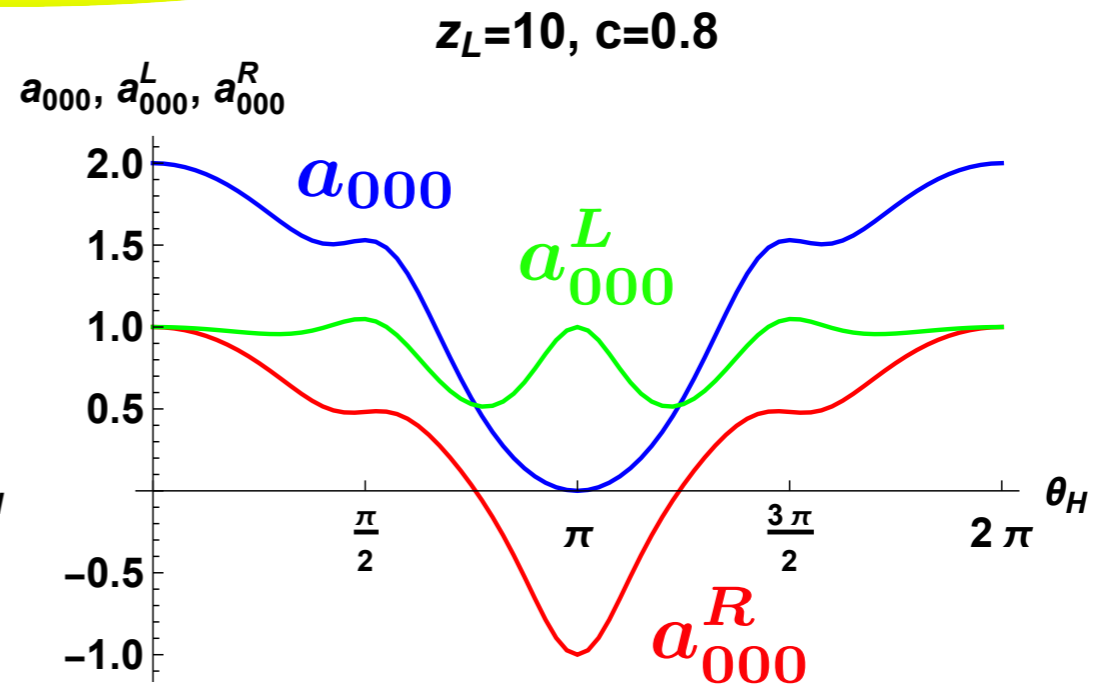
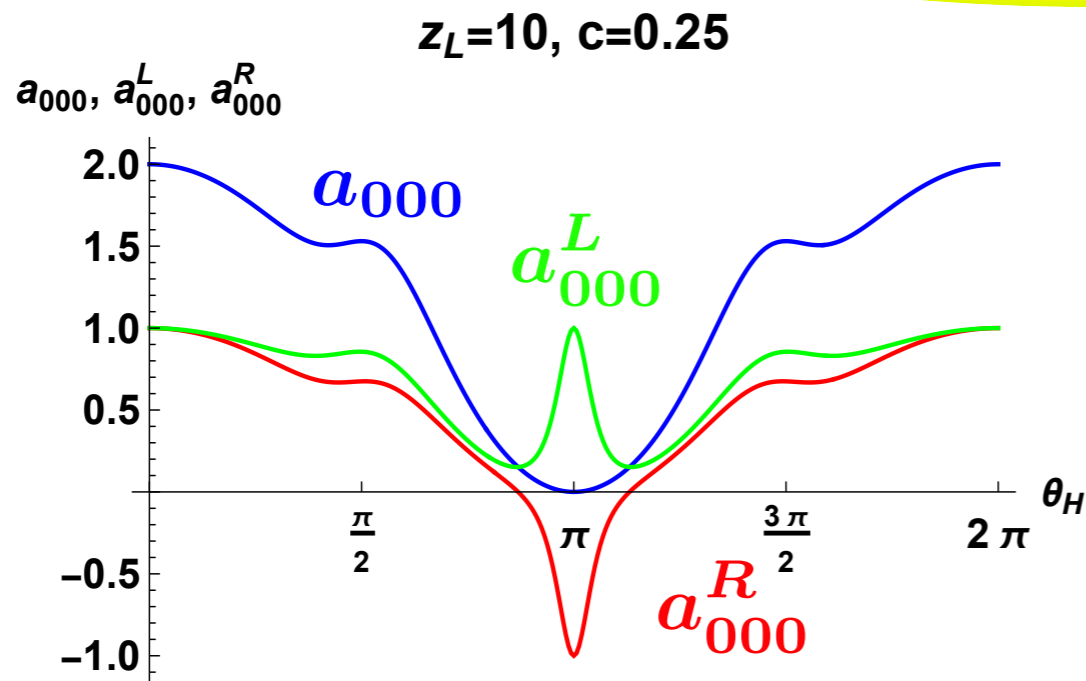
$a_{222}, a_{222}^L, a_{222}^R$



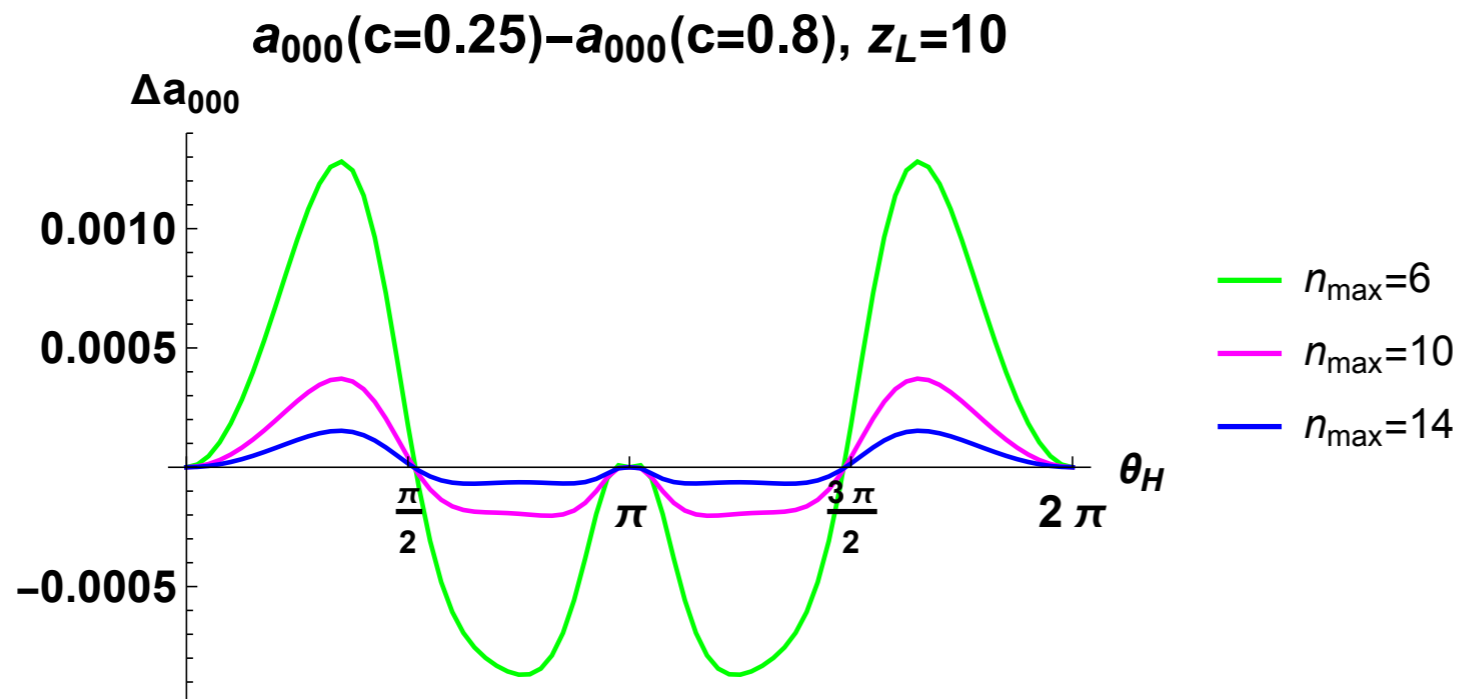
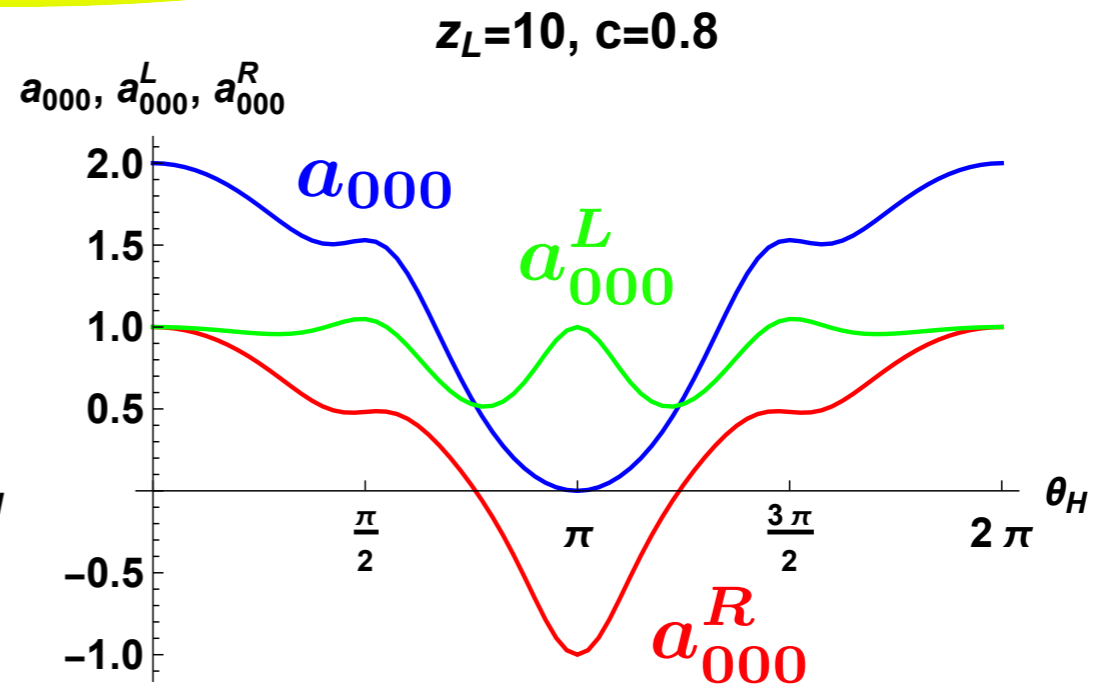
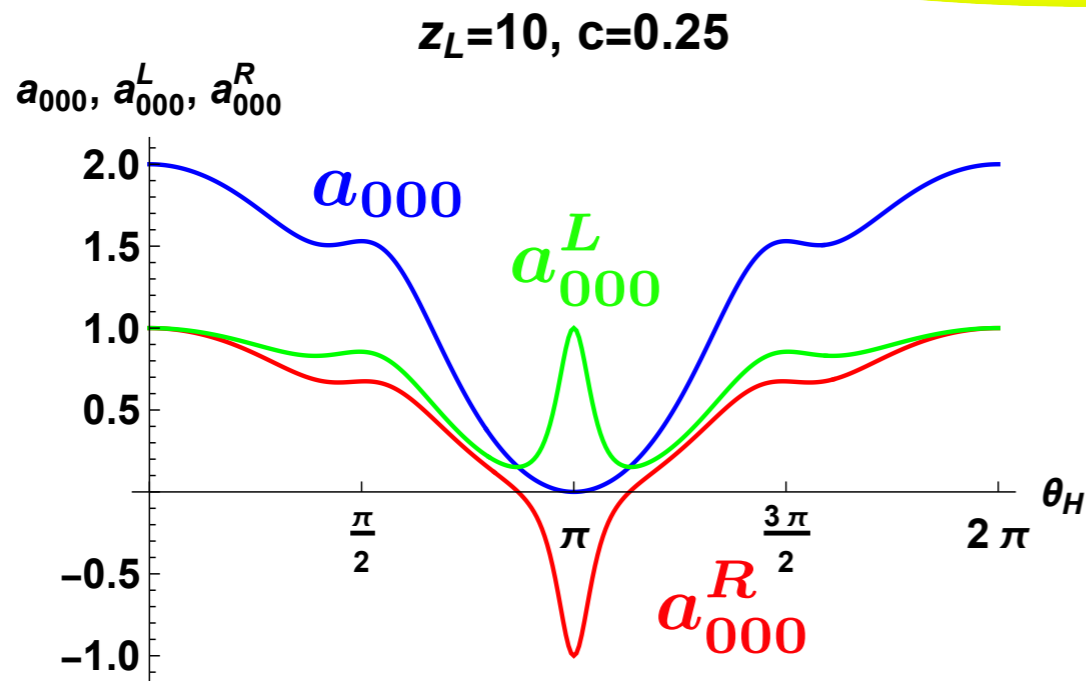
c-independence



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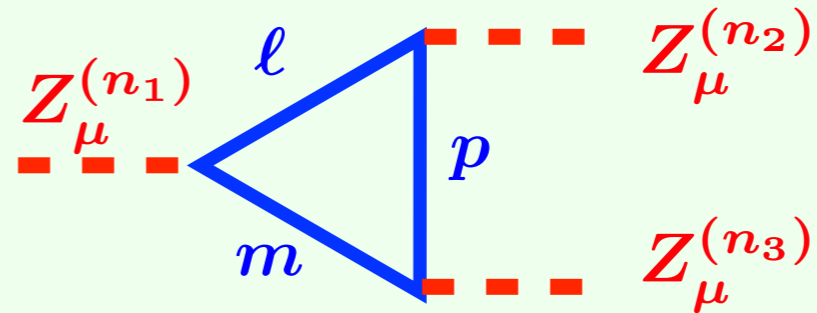
c-independence



Why?

Holography in anomaly flow

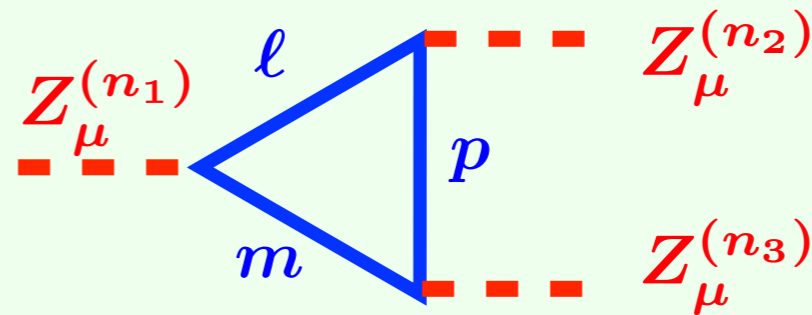
Anomaly coefficient



$$a_{n_1 n_2 n_3} = \sum_{m, l, p} \left\{ t_{n_1 m l}^R t_{n_2 l p}^R t_{n_3 p m}^R + t_{n_1 m l}^L t_{n_2 l p}^L t_{n_3 p m}^L \right\}$$

Holography in anomaly flow

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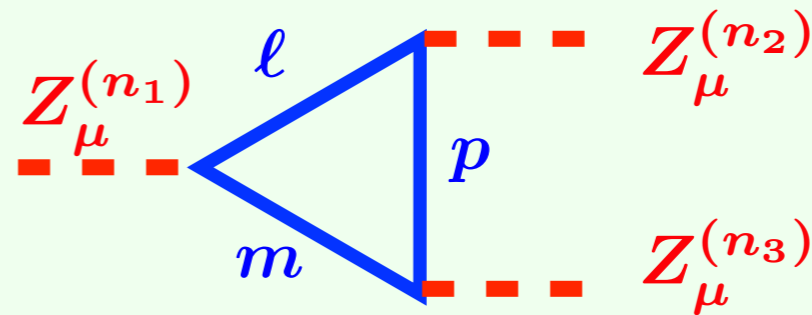


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Method 1 (i) Evaluate $t_{nml}^{R/L}$. (ii) Do $\sum_{m, l, p}$.

Method 2 (i) Do $\sum_{m, l, p}$. (ii) Do $\int dy_1 dy_2 dy_3$.

$$\begin{aligned}
a_{n_1 n_2 n_3} &= \left(\frac{k}{2}\right)^3 \int \int \int_{-a}^{2L-a} dy_1 dy_2 dy_3 e^{\sigma(y_1) + \sigma(y_2) + \sigma(y_3)} \\
&\times \left[k_1 k_2 k_3 \{ A_R(1, 2) A_R(2, 3) A_R(3, 1) - B_R(1, 2) B_R(2, 3) B_R(3, 1) \right. \\
&\quad \left. + B_L(1, 2) B_L(2, 3) B_L(3, 1) - A_L(1, 2) A_L(2, 3) A_L(3, 1) \right\} \\
&\quad \left. + k_1 h_2 h_3 \{ \dots \} + h_1 k_2 h_3 \{ \dots \} + h_1 h_2 k_3 \{ \dots \} \right]
\end{aligned}$$

$$k_j = k_{n_j}(y_j), \quad h_j = h_{n_j}(y_j)$$


$$\begin{pmatrix} A_R(j, k) \\ B_R(j, k) \end{pmatrix} = \sum_{n=0}^{\infty} \begin{pmatrix} f_{Rn}(y_j) f_{Rn}^*(y_k) \\ g_{Rn}(y_j) g_{Rn}^*(y_k) \end{pmatrix}$$

Anomaly formula

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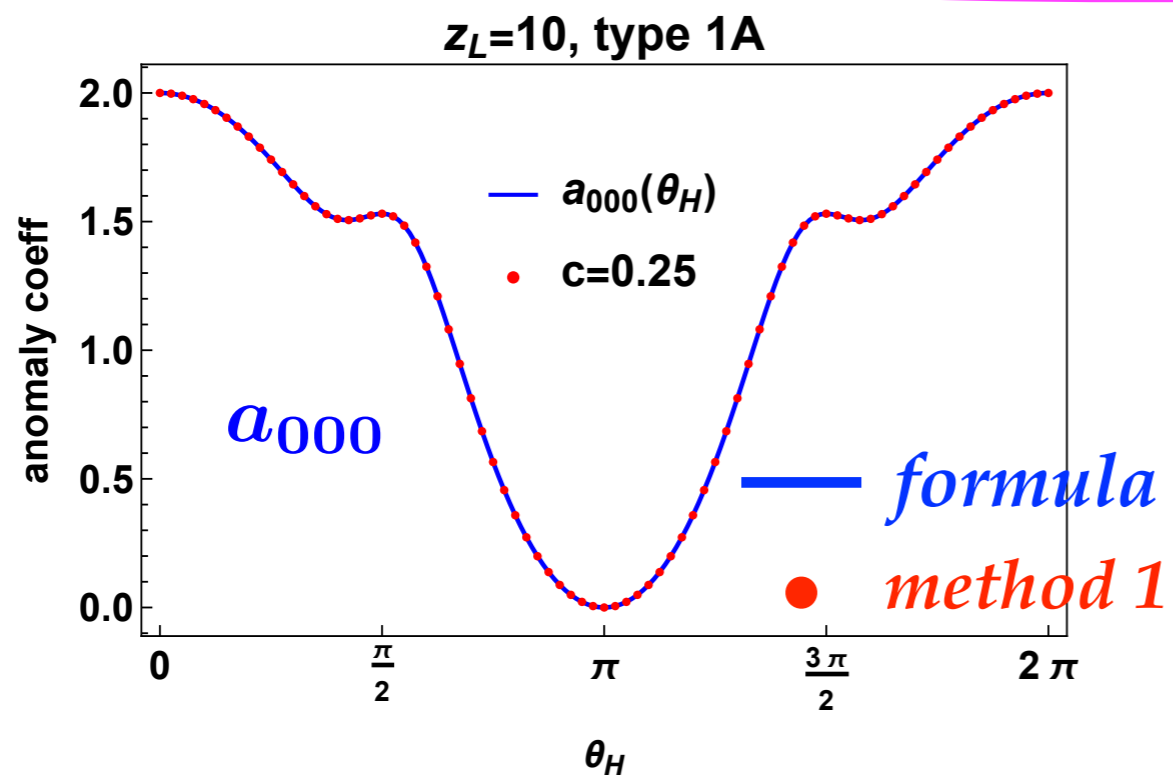
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Holography

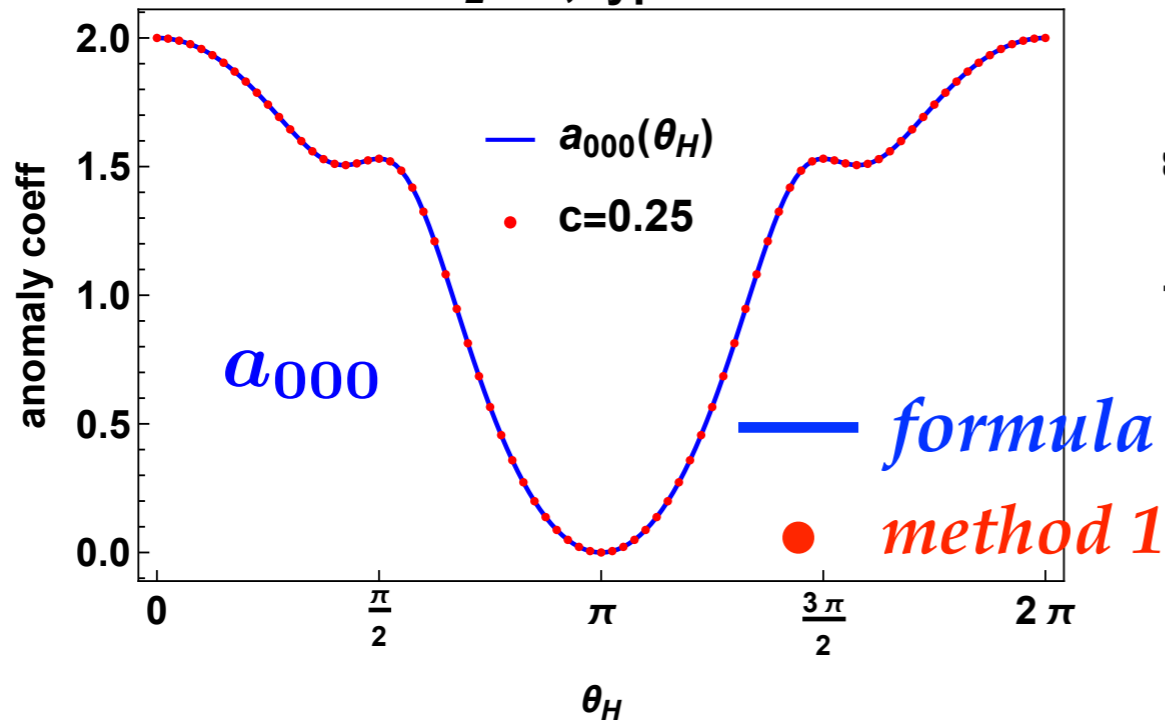
Anomaly flow and holography

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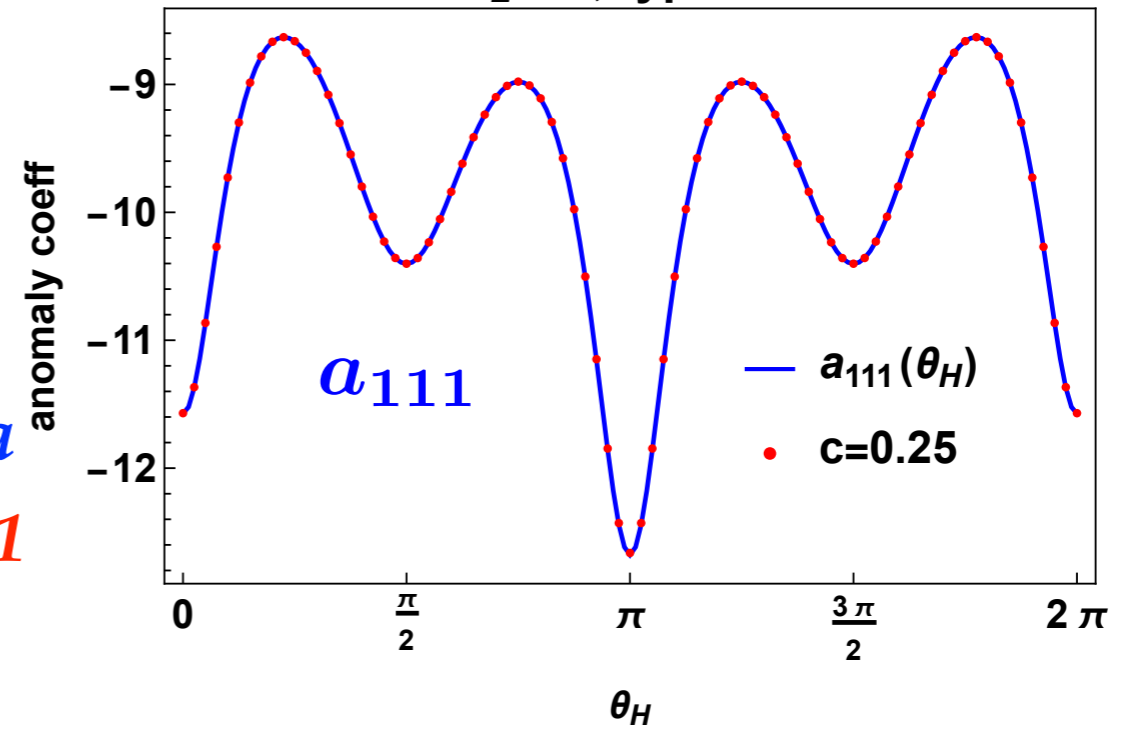


Anomaly flow and holography

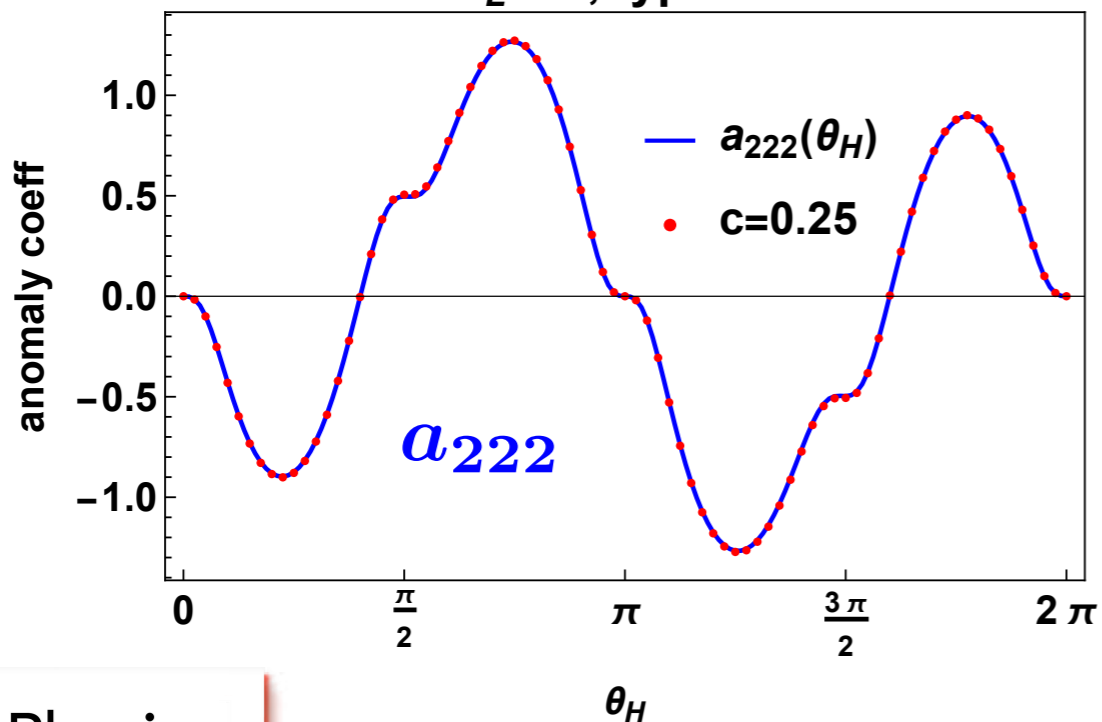
$z_L=10$, type 1A



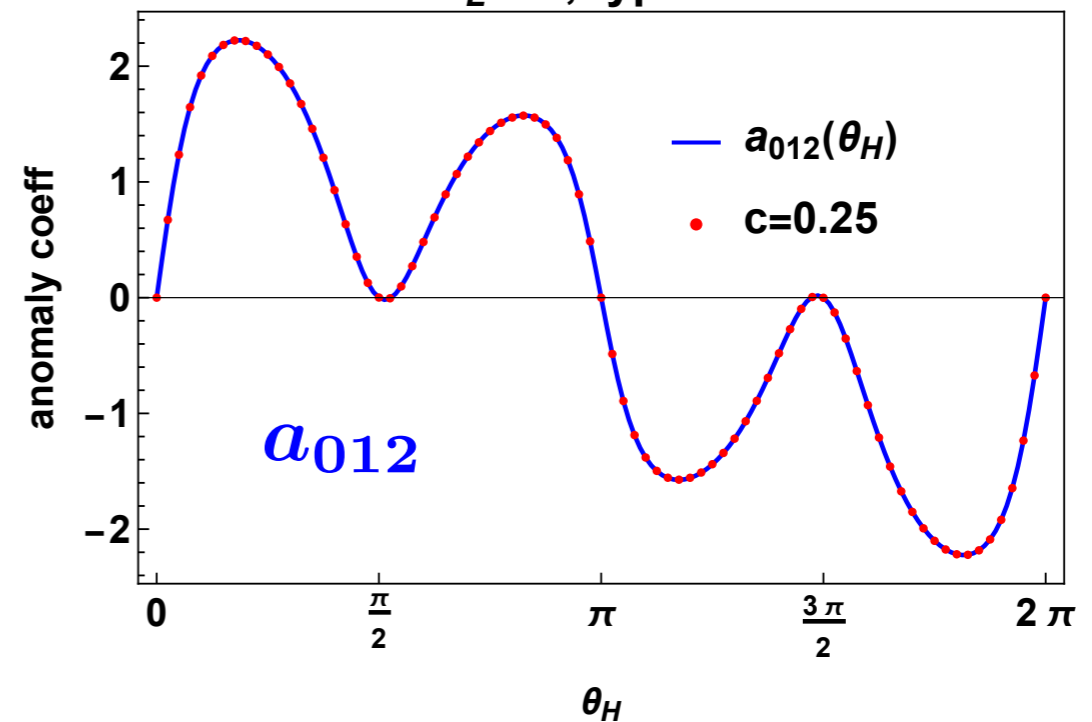
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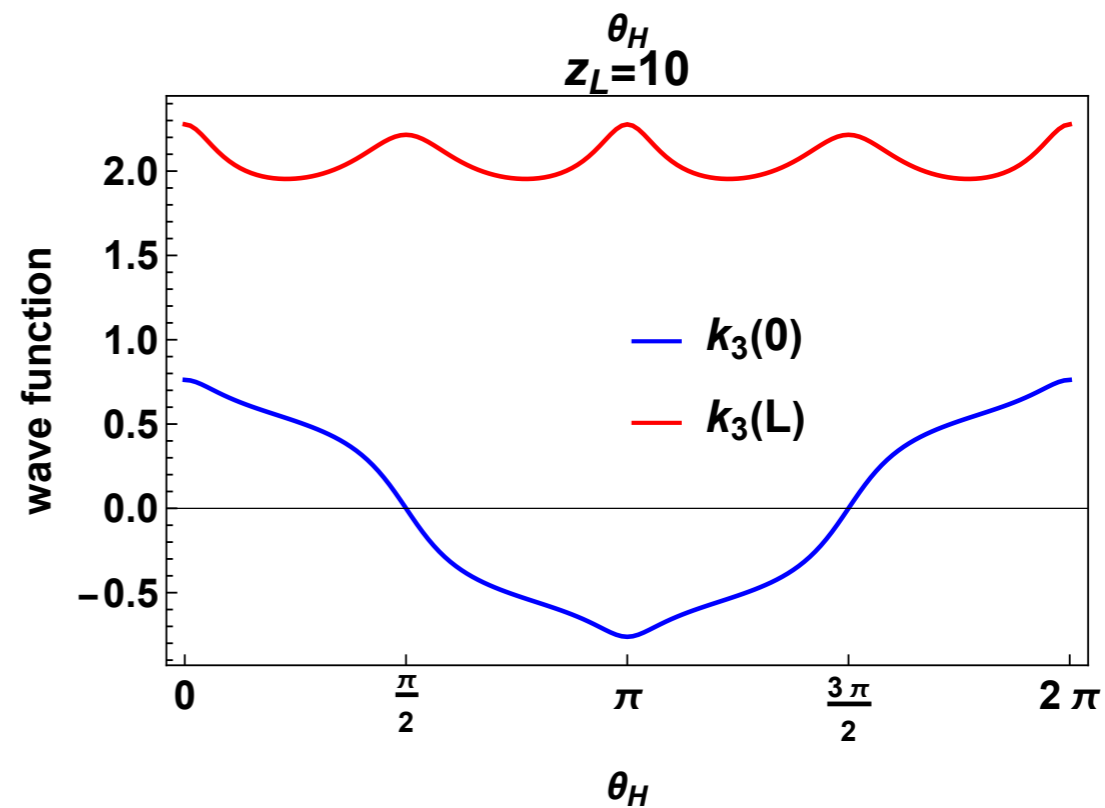
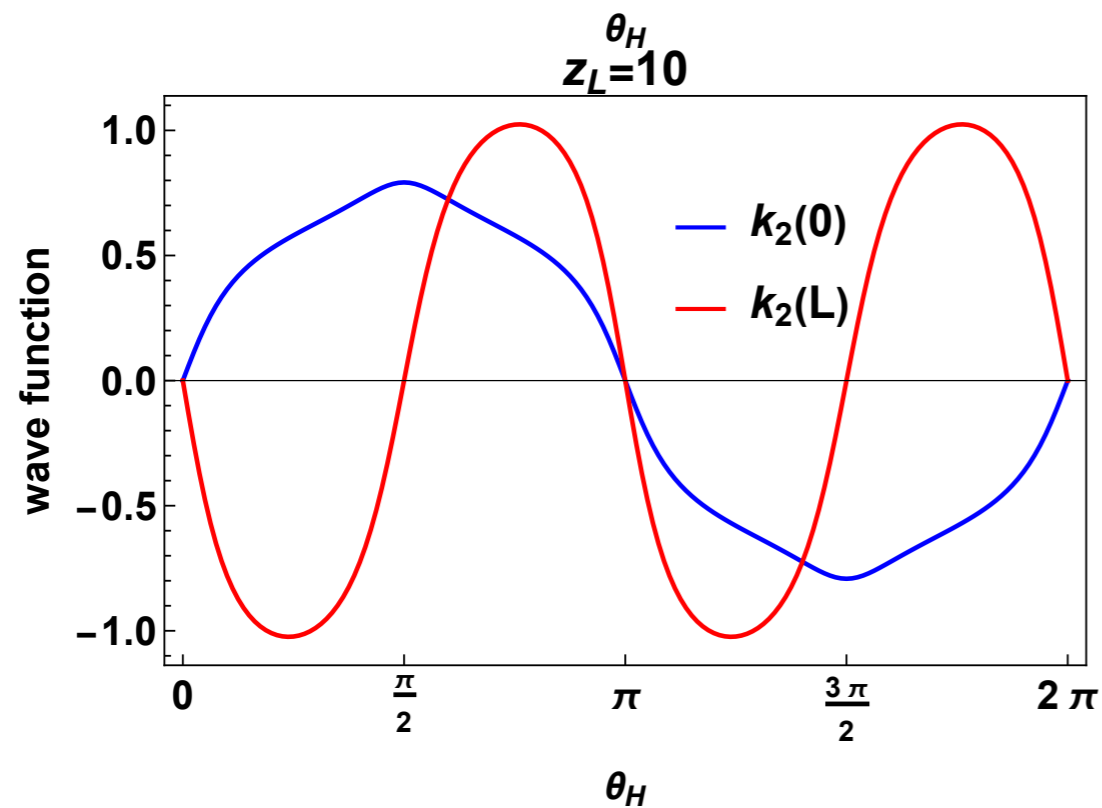
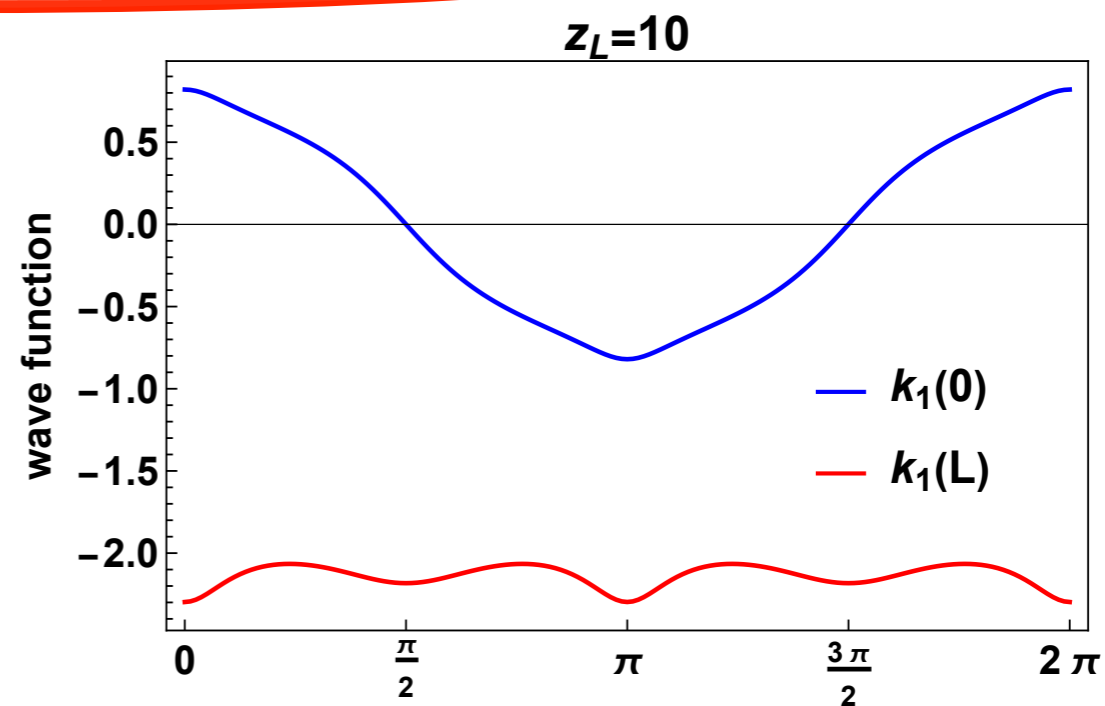
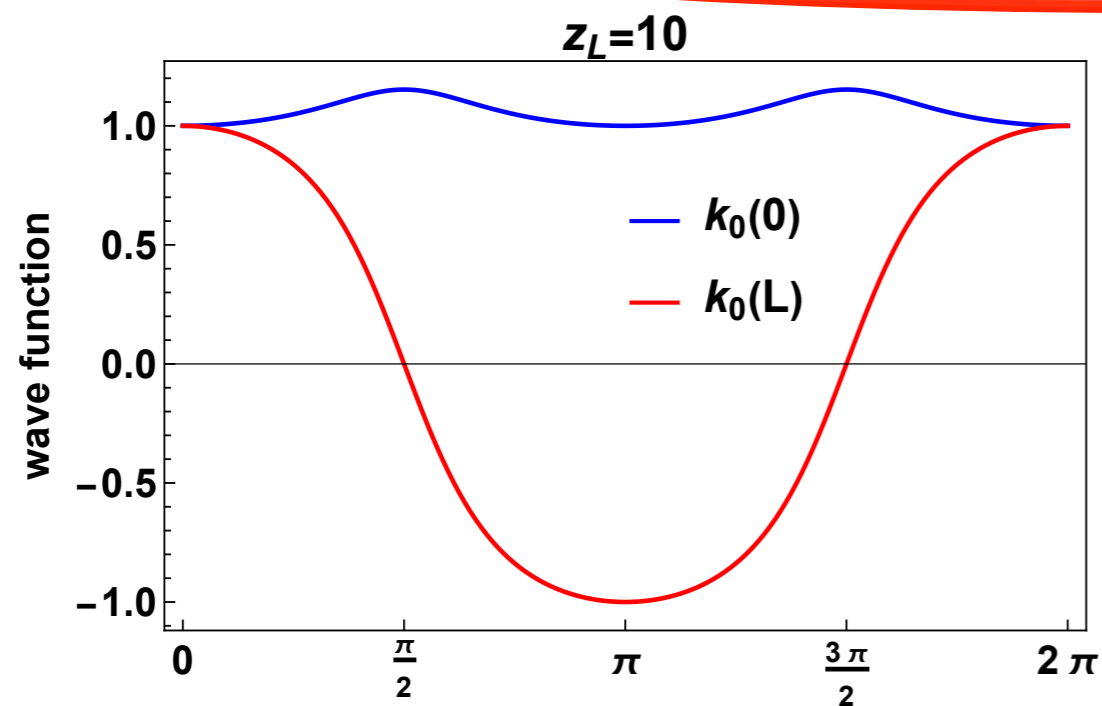
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$z_L=10$, type 1A



Gauge wave functions $k_n(0), k_n(L)$



Anomaly cancellation

Anomaly cancellation

$$\partial_\mu j_{(n)}^\mu \Rightarrow -\left(\frac{g_4}{2}\right)^3 \sum_\ell \sum_m \frac{a_{n\ell m}}{16\pi^2} Z_{\mu\nu}^{(\ell)} \tilde{Z}^{(m)\mu\nu}$$

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More generally if

$$(n_{1A} - n_{1B}) + (n_{2A} - n_{2B}) + (\hat{n}_R - \hat{n}_L) = 0,$$

$$(n_{1A} - n_{1B}) - (n_{2A} - n_{2B}) = 0$$



brane fermions at $y = 0$

Summary

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Chiral anomaly flows by θ_H

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In RS, no level crossing,

smooth variation of $a_{nlm}(\theta_H)$

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Anomaly cancellation in GHU

Flat space limit

$$\text{flat } M^4 \times (S^1/Z_2) \quad 0 \leq y \leq L = \pi R$$

$$\text{RS: } ds^2 = \frac{1}{z^2} \left(\eta_{\mu\nu} dx^\mu dx^\nu + \frac{dz^2}{k^2} \right) \quad (1 \leq z \leq z_L = e^{kL})$$

$$z = e^{ky}, \quad \text{AdS curvature} = -6k^2 \quad m_{\text{KK}} = \frac{\pi k}{z_L - 1} \implies \frac{1}{R}$$

$$b_{n\ell m} = Q_0 k_n^{\text{flat}}(0) k_\ell^{\text{flat}}(0) k_m^{\text{flat}}(0) + Q_1 k_n^{\text{flat}}(L) k_\ell^{\text{flat}}(L) k_m^{\text{flat}}(L)$$

$$k_n^{\text{flat}}(y) = \cos \frac{n\pi y}{L} \implies b_{n\ell m} = Q_0 + (-1)^{n+\ell+m} Q_1$$

