



UNIVERSITÀ DEGLI STUDI DI NAPOLI
FEDERICO II

INTERPLAY BETWEEN SPACETIME CURVATURE, SPEED OF LIGHT AND QUANTUM DEFORMATIONS OF RELATIVISTIC SYMMETRIES

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Università di Napoli "Federico II"

Noncommutative and generalized geometry in string theory, gauge theory and related physical models

Corfu, September 2022

References

- ◆ Ballesteros, GG, Mercati, *Interplay between Spacetime Curvature, Speed of Light and Quantum Deformations of Relativistic Symmetries*, Symmetry 2021
- ◆ Ballesteros, GG, Gutierrez-Sagredo, Herranz, *The κ -Newtonian and κ -Carrollian algebras and their noncommutative spacetimes*, PLB 2020
- ◆ Ballesteros, GG, Herranz, *Lorentzian Snyder spacetimes and their Galilei and Carroll limits from projective geometry*, CQG 2020
- ◆ Ballesteros, GG, Gutierrez-Sagredo, Herranz, *Curved momentum spaces from quantum (anti-)de Sitter groups in (3+1) dimensions*, PRD 2018
- ◆ Ballesteros, GG, Gutierrez-Sagredo, Herranz, *Curved momentum spaces from quantum groups with cosmological constant*, PLB 2017

The Poincaré algebra of special relativity

- ◆ Special relativistic symmetries are described by the Poincaré Lie algebra $\mathfrak{p}(3+1)$ generated by time translations P_0 , spatial translation P_a , boosts K_a , and rotations J_a

$$\begin{array}{lll} [J_a, J_b] = \epsilon_{abc} J_c & [K_a, P_0] = P_a & [P_0, P_a] = 0 \\ [J_a, P_b] = \epsilon_{abc} P_c & [K_a, P_b] = \delta_{ab} P_0 & [P_a, P_b] = 0 \\ [J_a, K_b] = \epsilon_{abc} K_c & [K_a, K_b] = -\epsilon_{abc} J_c & [P_0, J_a] = 0 \end{array}$$

Casimir:

$$\mathcal{C} = P_0^2 - \mathbf{P}^2$$

- ◆ Minkowski spacetime \mathbf{M}^{3+1} is obtained as the homogeneous space of the Poincaré Lie group with isotropy subgroup given by Lorentz

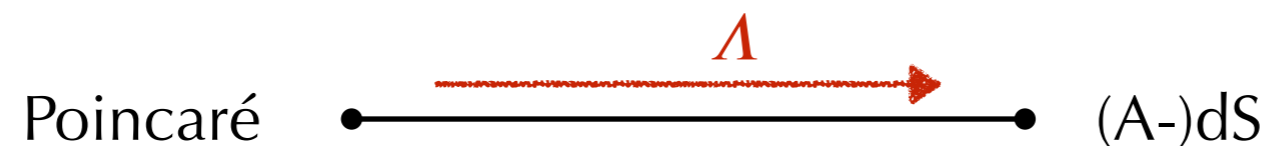
$$\mathbf{M}^{3+1} \equiv \text{ISO}(3,1)/\text{SO}(3,1)$$

Classical deformation of the Poincaré algebra: (A)dS symmetries

◆ **Classical** deformations of the Poincaré algebra

Governed by the curvature parameter Λ : (Anti-)de Sitter symmetries

Squared-length scale: the induced modifications are relevant at length scales comparable to $\sqrt{\Lambda}$



Classical deformation of the Poincaré algebra: (A)dS symmetries

♦ Classical deformations of the Poincaré algebra

Linearly deformed commutation rules:

$$\begin{array}{lll}
 [J_a, J_b] = \epsilon_{abc} J_c & [K_a, P_0] = P_a & [P_0, P_a] = 0 \\
 [J_a, P_b] = \epsilon_{abc} P_c & [K_a, P_b] = \delta_{ab} P_0 & [P_a, P_b] = 0 \\
 [J_a, K_b] = \epsilon_{abc} K_c & [K_a, K_b] = -\epsilon_{abc} J_c & [P_0, J_a] = 0
 \end{array}$$

$$\mathcal{C} = P_0^2 - \mathbf{P}^2$$



$$\begin{array}{lll}
 [J_a, J_b] = \epsilon_{abc} J_c & [K_a, P_0] = P_a & [P_0, P_a] = -\Lambda K_a \\
 [J_a, P_b] = \epsilon_{abc} P_c & [K_a, P_b] = \delta_{ab} P_0 & [P_a, P_b] = \Lambda \epsilon_{abc} J_c \\
 [J_a, K_b] = \epsilon_{abc} K_c & [K_a, K_b] = -\epsilon_{abc} J_c & [P_0, J_a] = 0
 \end{array}$$

$$\mathcal{C} = P_0^2 - \mathbf{P}^2 - \Lambda (\mathbf{J}^2 - \mathbf{K}^2)$$

Classical deformation of the Poincaré algebra: (A)dS symmetries

- ◆ (Anti-)de Sitter algebra in the kinematical basis

$$\begin{array}{lll} [J_a, J_b] = \epsilon_{abc} J_c & [K_a, P_0] = P_a & [P_0, P_a] = -\Lambda K_a \\ [J_a, P_b] = \epsilon_{abc} P_c & [K_a, P_b] = \delta_{ab} P_0 & [P_a, P_b] = \Lambda \epsilon_{abc} J_c \\ [J_a, K_b] = \epsilon_{abc} K_c & [K_a, K_b] = -\epsilon_{abc} J_c & [P_0, J_a] = 0 \end{array}$$

Casimir:

$$\mathcal{C} = P_0^2 - \mathbf{P}^2 - \Lambda (\mathbf{J}^2 - \mathbf{K}^2)$$

- ◆ (anti-)de Sitter spacetime $(\mathbf{A})\mathbf{dS}^{3+1}$ is obtained as a homogeneous space:

$$\Lambda < 0: \mathbf{AdS}^{3+1} \equiv SO(3,2)/SO(3,1)$$

$$\Lambda > 0: \mathbf{dS}^{3+1} \equiv SO(4,1)/SO(3,1)$$

$$\Lambda = 0: \mathbf{M}^{3+1} \equiv ISO(3,1)/SO(3,1)$$

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de Sitter spacetime, phase space and particle kinematics

$$[H = \sqrt{\Lambda}]$$

- ♦ line element in comoving coordinates (1+1 dimensions)

$$ds^2 = (dx^0)^2 - e^{2Hx^0} (dx^1)^2$$

- ♦ algebra of symmetries in the comoving basis (Poisson brackets!)

$$\begin{aligned} \{\mathcal{P}_0, \mathcal{P}_1\} &= H \mathcal{P}_1 \\ \{\mathcal{P}_0, \mathcal{N}\} &= \mathcal{P}_1 - H \mathcal{N} \\ \{\mathcal{P}_1, \mathcal{N}\} &= \mathcal{P}_0 \end{aligned}$$

- ♦ mass Casimir $\mathcal{C}_{dS} = \mathcal{P}_0^2 - \mathcal{P}_1^2 + 2H\mathcal{N}\mathcal{P}_1$

- ♦ representation of symmetry generators:

$$\begin{aligned} \{x^\mu, x^\nu\} &= 0, \\ \{x^\mu, p_\nu\} &= -\delta_\nu^\mu, \\ \{p_\mu, p_\nu\} &= 0. \end{aligned} \quad \longrightarrow \quad \begin{aligned} \mathcal{P}_0 &= p_0 - Hx^1 p_1 \\ \mathcal{P}_1 &= p_1 \\ \mathcal{N} &= x^1 p_0 + p_1 \left(\frac{1 - e^{-2Hx^0}}{2H} - \frac{H}{2} (x^1)^2 \right) \end{aligned}$$

- ♦ the massless condition $\mathcal{C}_{dS} = 0$ relates energy and spatial momentum, encoding energy redshift

$$p_0 = |p_1| e^{-Hx^0}$$

- ♦ particles worldline (using Hamiltonian formalism)

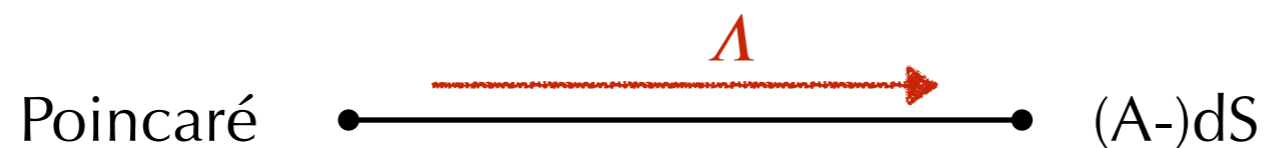
$$\begin{aligned} \dot{x}^1 &= \{\mathcal{C}_{dS}, x^1\} = -2e^{-2Hx^0} p_1 \\ \dot{x}^0 &= \{\mathcal{C}_{dS}, x^0\} = 2p_0 \end{aligned} \quad \longrightarrow \quad x^1(x^0) - \bar{x}^1 \equiv \int_{\bar{x}^0}^{x^0} \frac{\dot{x}^1}{\dot{x}^0} dx^0 = \left(\frac{e^{-H\bar{x}^0} - e^{-Hx^0}}{H} \right)$$

Deforming the Poincaré algebra of special relativity

◆ **Classical** deformations of the Poincaré algebra

Governed by the curvature parameter Λ : (Anti-)de Sitter symmetries

Squared-length scale: the induced modifications are relevant at length scales comparable to $\sqrt{\Lambda}$



Deforming the Poincaré algebra of special relativity

◆ **Classical** deformations of the Poincaré algebra

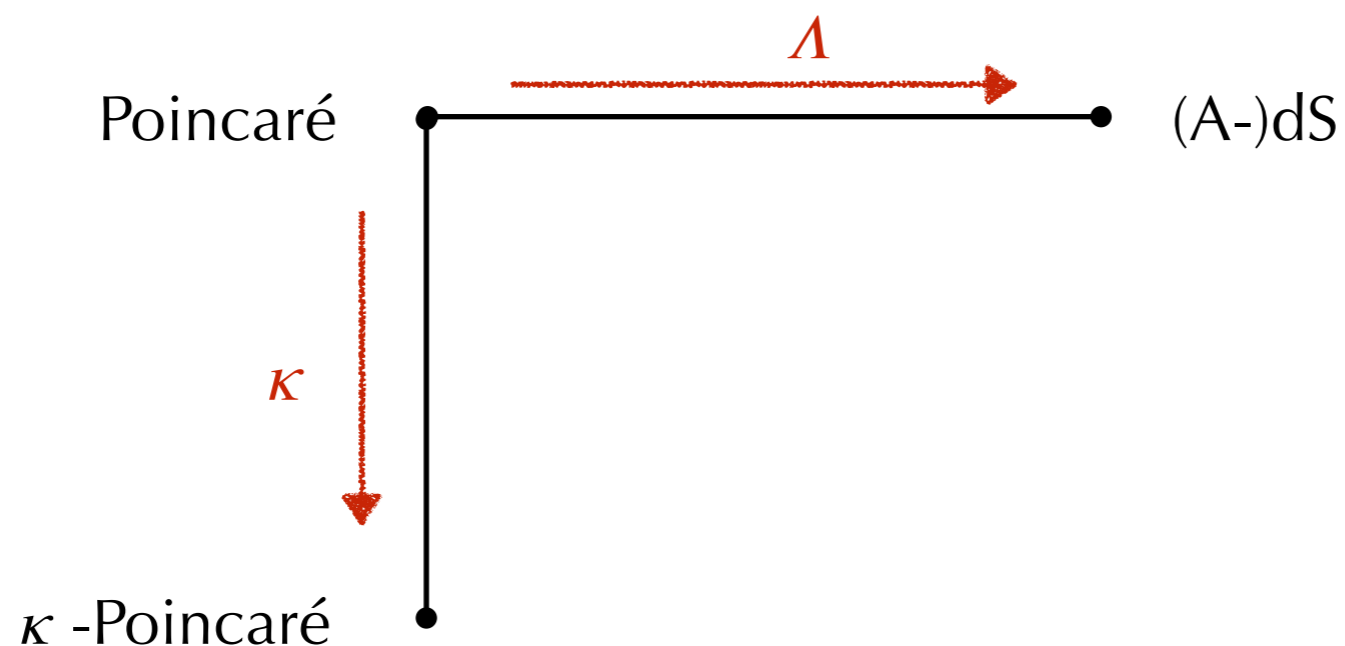
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◆ **Quantum** deformations of the Poincaré algebra

Governed by the quantum deformation parameter κ

Energy scale: the induced modifications are relevant at energies comparable to κ



Quantum deformation of the Poincaré algebra: κ -Poincaré

- ♦ **Quantum** deformations of the Poincaré algebra \Rightarrow Hopf algebra
(we show the κ -Poincaré algebra in the bicrossproduct basis)

Nonlinearly deformed commutation rules:

$$\begin{array}{lll}
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 \end{array}$$

$$\mathcal{C} = P_0^2 - \mathbf{P}^2$$



$$[K_a, P_b] = \delta_{ab} \left(\frac{\kappa}{2} (1 - e^{-2P_0/\kappa}) + \frac{1}{2\kappa} \mathbf{P}^2 \right) - \frac{1}{\kappa} P_a P_b$$

Casimir: $\mathcal{C}_\kappa = 4\kappa^2 \sinh^2(P_0/2\kappa) - e^{P_0/\kappa} \mathbf{P}^2$

- Lukierski, Nowicki, Ruegg, PLB 1992
- Lukierski, Ruegg, Nowicki, Tolstoi, PLB 1991
- Majid, Ruegg, PLB 1994

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Quantum deformation of the Poincaré algebra: κ -Poincaré

- ♦ **Quantum** deformations of the Poincaré algebra \Rightarrow Hopf algebra
(we show the κ -Poincaré algebra in the bicrossproduct basis)

Deformed action on multi-particle states (coproducts)
(and conservation rules of energy-momentum in interactions)

$$\Delta(X) = X \otimes 1 + 1 \otimes X$$



$$\Delta(P_0) = P_0 \otimes 1 + 1 \otimes P_0,$$

$$\Delta(J_a) = J_a \otimes 1 + 1 \otimes J_a,$$

$$\Delta(P_a) = P_a \otimes 1 + e^{-P_0/\kappa} \otimes P_a,$$

$$\Delta(K_a) = K_a \otimes 1 + e^{-P_0/\kappa} \otimes K_a + \frac{1}{\kappa} \epsilon_{abc} P_b \otimes J_c.$$

- Lukierski, Nowicki, Ruegg, *PLB* 1992
- Lukierski, Ruegg, Nowicki, Tolstoi, *PLB* 1991
- Majid, Ruegg, *PLB* 1994

Quantum deformation of the Poincaré algebra: κ -Poincaré

◆ Quantum deformations of the Poincaré algebra

The κ -Poincaré algebra (as a Lie bialgebra) is generated by the r-matrix

$$r_{\Lambda} = \frac{1}{\kappa}(K_1 \wedge P_1 + K_2 \wedge P_2 + K_3 \wedge P_3)$$

Which produces the cocommutators via $\delta(X) = [X \otimes 1 + 1 \otimes X, r_{\Lambda}]$, $X \in \{P_0, P_a, J_a, K_a\}$

$$\delta(P_0) = \delta(J_a) = 0$$

$$\delta(P_a) = \frac{1}{\kappa}P_a \wedge P_0$$

$$\delta(K_1) = \frac{1}{\kappa}(K_1 \wedge P_0 + P_2 \wedge J_3 - P_3 \wedge J_2),$$

$$\delta(K_2) = \frac{1}{\kappa}(K_2 \wedge P_0 - P_1 \wedge J_3 + P_3 \wedge J_1),$$

$$\delta(K_3) = \frac{1}{\kappa}(K_3 \wedge P_0 + P_1 \wedge J_2 - P_2 \wedge J_1).$$

κ -Poincaré phase space: free particle worldlines

$$[\ell \equiv 1/\kappa]$$

- ♦ representation of the algebra of symmetries on phase space (Poisson brackets)

$$\{p_1, p_0\} = 0$$

$$\{x^1, x^0\} = 0$$

$$\{x^\mu, p_\nu\} = \delta_\nu^\mu$$



$$\mathcal{P}_0 = p_0$$

$$\mathcal{P}_1 = p_1$$

$$\mathcal{N} = p_1 x^0 + \left(\frac{1 - e^{-2\ell p_0}}{2\ell} - \frac{\ell}{2} p_1^2 \right) x^1$$

- ♦ evolution of phase space coordinates is given by the Hamilton equations with the κ -Poincaré Casimir as Hamiltonian

$$\dot{x}^0 = \{\mathcal{C}_\ell, x^0\} = \frac{1}{\ell} (e^{\ell p_0} - e^{-\ell p_0}) - \ell p_1^2 e^{\ell p_0}$$

$$\dot{x}^1 = \{\mathcal{C}_\ell, x^1\} = 2 p_1 e^{\ell p_0}.$$

- ♦ Then the free particle worldline reads $x^1(x^0) = x^1(0) + v(p)x^0$



$$v(p) = \frac{e^{p_0/\kappa} \sqrt{e^{2p_0/\kappa} + 1 - 2 e^{p_0/\kappa} \cosh(m/\kappa)}}{1 - e^{p_0/\kappa} \cosh(m/\kappa)} \rightarrow e^{p_0/\kappa} \quad \text{for } m=0$$

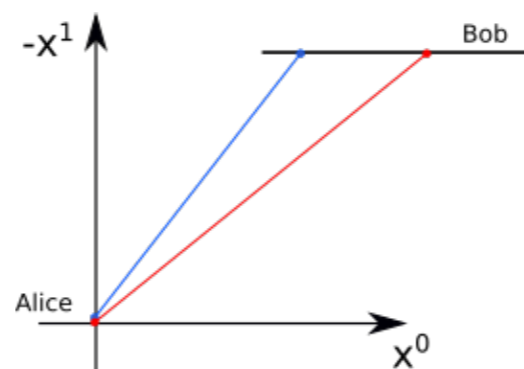
Particle worldlines and relativity of locality

$$[\ell \equiv 1/\kappa]$$

worldlines of two massless particles emitted simultaneously with different energies

- ♦ using coordinates dual to momenta:
(energy dependent worldlines)

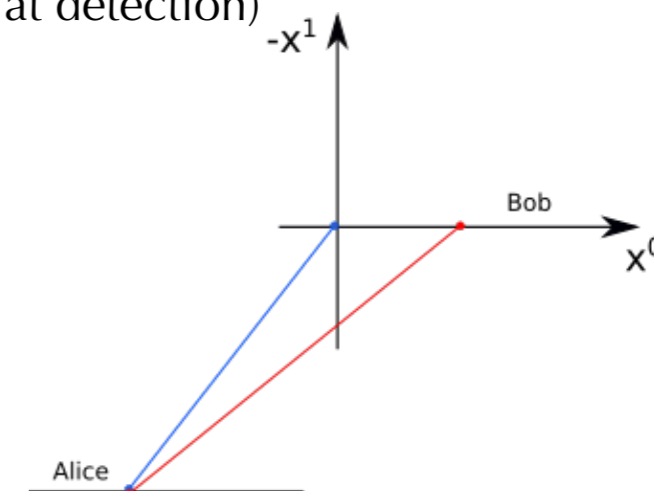
worldlines seen by Alice
(local at emission)



$$x_A^1 = -e^{\ell p_0} x_A^0$$

$$x_A^1 = -e^{\ell \tilde{p}_0} x_A^0$$

worldlines seen by Bob
(local at detection)



$$x_B^0 = \mathcal{T}_a \triangleright x_A^0 = x_A^0 - a^0$$

$$x_B^1 = \mathcal{T}_a \triangleright x_A^1 = x_A^1 - a^1$$

Particle worldlines and relativity of locality

$$[\ell \equiv 1/\kappa]$$

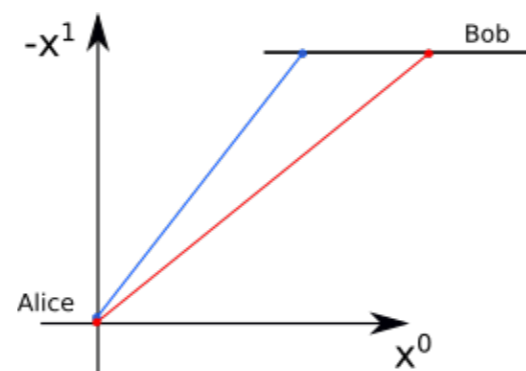
worldlines of two massless particles emitted simultaneously with different energies

♦ using coordinates dual to momenta:
(energy dependent worldlines)

♦ using 'κ-Minkowski coordinates':
(energy independent worldlines
but deformed action of translations)

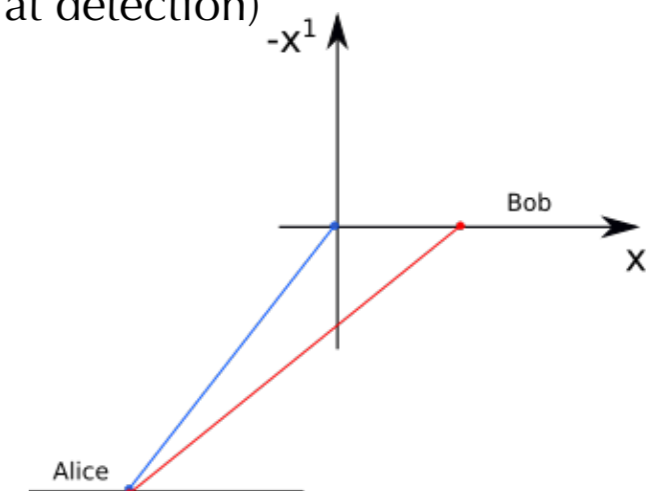
$$\begin{aligned} \chi^1 &= x^1 & \{\chi^0, \chi^1\} &= \ell \chi^1 \\ \chi^0 &= x^0 - \ell x^1 p_1 & \{\chi^0, p_1\} &= -\ell p_1 \\ & & \{\chi^0, p_0\} &= 0 \end{aligned}$$

worldlines seen by Alice
(local at emission)

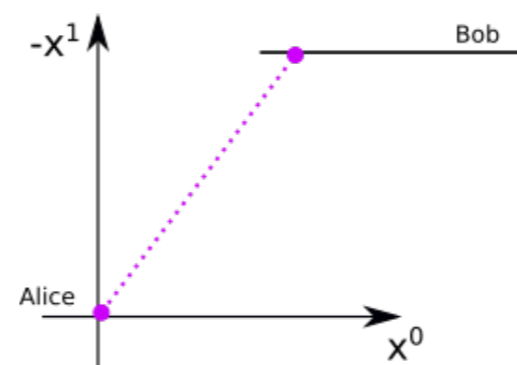


$$\begin{aligned} x_A^1 &= -e^{\ell p_0} x_A^0 \\ x_A^1 &= -e^{\ell \tilde{p}_0} x_A^0 \end{aligned}$$

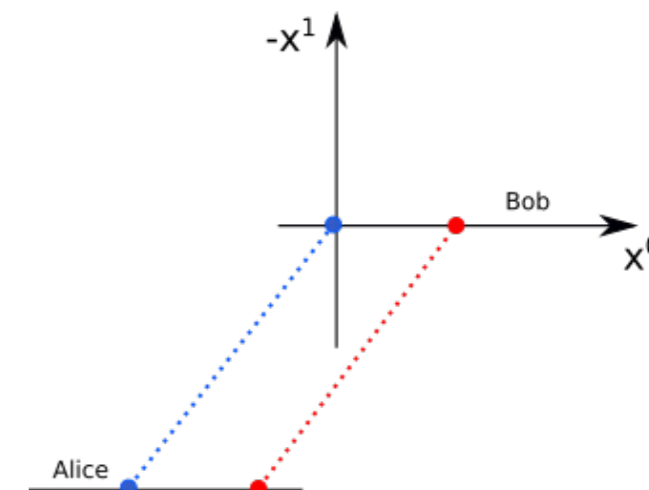
worldlines seen by Bob
(local at detection)



$$\begin{aligned} x_B^0 &= \mathcal{T}_a \triangleright x_A^0 = x_A^0 - a^0 \\ x_B^1 &= \mathcal{T}_a \triangleright x_A^1 = x_A^1 - a^1 \end{aligned}$$



$$\chi^1 - \bar{\chi}^1 = \chi^0 - \bar{\chi}^0$$



$$\begin{aligned} \chi_B^0 &= \mathcal{T}_a \triangleright \chi_A^0 = \chi_A^0 - a^0 + a^1 \ell p_1 \\ \chi_B^1 &= \mathcal{T}_a \triangleright \chi_A^1 = \chi_A^1 - a^1 \end{aligned}$$

Particle worldlines and relativity of locality

$$[\ell \equiv 1/\kappa]$$

worldlines of two massless particles emitted simultaneously with different energies

♦ using coordinates dual to momenta:
(energy dependent worldlines)

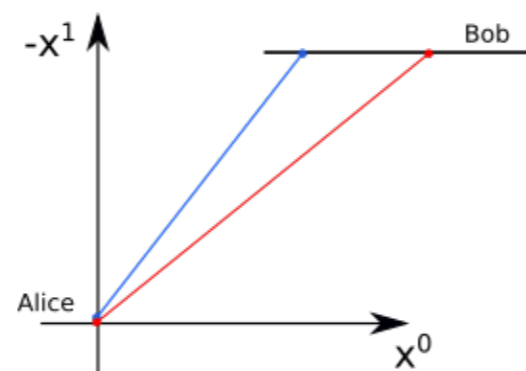
♦ using 'κ-Minkowski coordinates':
(energy independent worldlines
but deformed action of translations)

$$\begin{aligned} \chi^1 &= x^1 & \{\chi^0, \chi^1\} &= \ell \chi^1 \\ \chi^0 &= x^0 - \ell x^1 p_1 & \{\chi^0, p_1\} &= -\ell p_1 \\ & & \{\chi^0, p_0\} &= 0 \end{aligned}$$

♦ measurements done locally (i.e. at spatial origin of each observer) do not depend on choice of coordinates: Alice emits the particles at the same time and Bob detects them with a time delay

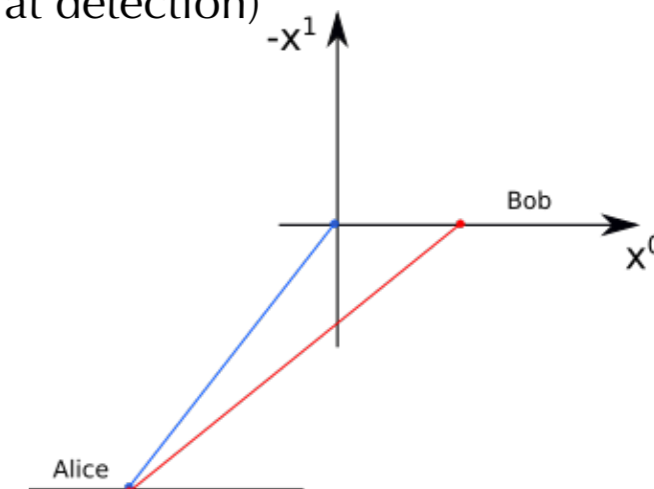
$$\Delta x^0 = a^0 (e^{-\ell \Delta p_0} - 1)$$

worldlines seen by Alice
(local at emission)

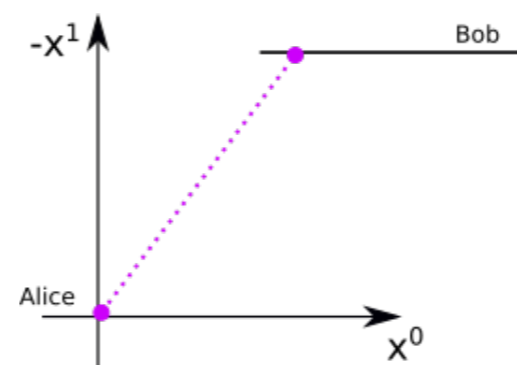


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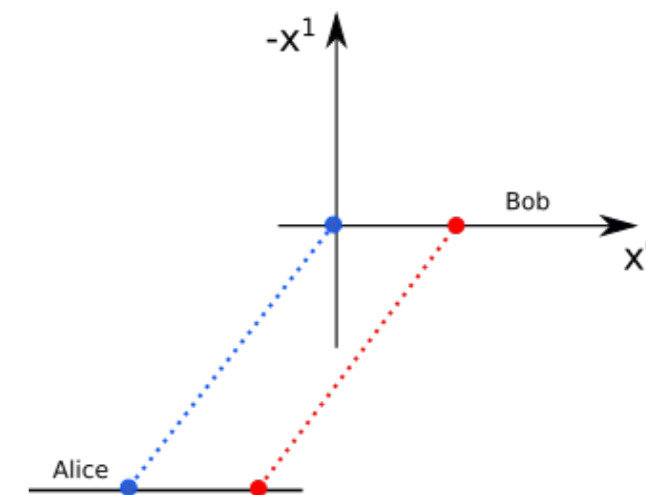
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$$\chi^1 - \bar{\chi}^1 = \chi^0 - \bar{\chi}^0$$



$$\begin{aligned} \chi_B^0 &= \mathcal{T}_a \triangleright \chi_A^0 = \chi_A^0 - a^0 + a^1 \ell p_1 \\ \chi_B^1 &= \mathcal{T}_a \triangleright \chi_A^1 = \chi_A^1 - a^1 \end{aligned}$$

Duality between classical and quantum deformation

$$H \equiv \sqrt{\Lambda}$$

de Sitter phase space

$$\begin{aligned} \{x^\mu, p_\nu\} &= -\delta_\nu^\mu \\ \{x^1, x^0\} &= 0 \\ \{\mathcal{N}, x^0\} &= x^1 \\ \{\mathcal{N}, x^1\} &= \frac{1-e^{-2Hx^0}}{2H} - \frac{H}{2}(x^1)^2 \\ \{p_0, p_1\} &= 0 \\ \{p_0, \mathcal{N}\} &= p_1 e^{-2Hx^0} \\ \{p_1, \mathcal{N}\} &= p_0 - Hp_1 x^1 \\ \{\mathcal{P}_0, \mathcal{P}_1\} &= H\mathcal{P}_1 \\ \{\mathcal{P}_0, \mathcal{N}\} &= \mathcal{P}_1 - H\mathcal{N} \\ \{\mathcal{P}_1, \mathcal{N}\} &= \mathcal{P}_0 \end{aligned}$$

the Casimir of the algebra $\{\mathcal{N}, \mathcal{P}_0, \mathcal{P}_1\}$
determines the dispersion relation

the Casimir of the algebra $\{\mathcal{N}, x^0, x^1\}$
determines the worldline
(and is the Newton-Wigner operator)

κ -Poincaré phase space

$$\begin{aligned} \{p_\nu, x^\mu\} &= -\delta_\nu^\mu & \ell \equiv 1/\kappa \\ \{\mathcal{P}_1, \mathcal{P}_0\} &= 0 \\ \{\mathcal{N}, \mathcal{P}_0\} &= \mathcal{P}_1 \\ \{\mathcal{N}, \mathcal{P}_1\} &= \frac{1-e^{-2\ell\mathcal{P}_0}}{2\ell} - \frac{\ell}{2}\mathcal{P}_1^2 \\ \{x^0, x^1\} &= 0 \\ \{x^0, \mathcal{N}\} &= x^1 e^{-2\ell p_0} \\ \{x^1, \mathcal{N}\} &= x^0 - \ell x^1 p_1 \\ \{\chi^0, \chi^1\} &= \ell\chi^1 \\ \{\chi^0, \mathcal{N}\} &= \chi^1 - \ell\mathcal{N} \\ \{\chi^1, \mathcal{N}\} &= \chi^0 \end{aligned}$$

the Casimir of the algebra $\{\mathcal{N}, \chi^0, \chi^1\}$
determines the worldline
(and is the Newton-Wigner operator)

the Casimir of the algebra $\{\mathcal{N}, p_0, p_1\}$
determines the dispersion relation

Duality between classical and quantum deformation

$$H \equiv \sqrt{\Lambda}$$

de Sitter spacetime

spacetime metric

$$ds^2 = (dx^0)^2 - e^{2Hx^0} (dx^1)^2$$

worldline

$$x^1 = \frac{1 - e^{-Hx^0}}{H}$$

dispersion relation

$$p_1 = -e^{Hx^0} p_0$$

generators of translations

$$\begin{aligned} \mathcal{P}_1 &= p_1 \\ \mathcal{P}_0 &= p_0 - Hx^1 p_1 \end{aligned}$$

de Sitter momentum space

$$\ell \equiv 1/\kappa$$

momentum space metric

$$ds_p^2 = dp_0^2 - e^{2\ell p_0} dp_1^2$$

dispersion relation

$$p_1 = \frac{1 - e^{-\ell p_0}}{\ell}$$

worldline

$$x^1 = -e^{\ell p_0} x^0$$

'k-Minkowski coordinates'

$$\begin{aligned} \chi^1 &= x^1 \\ \chi^0 &= x^0 - \ell x^1 p_1 \end{aligned}$$

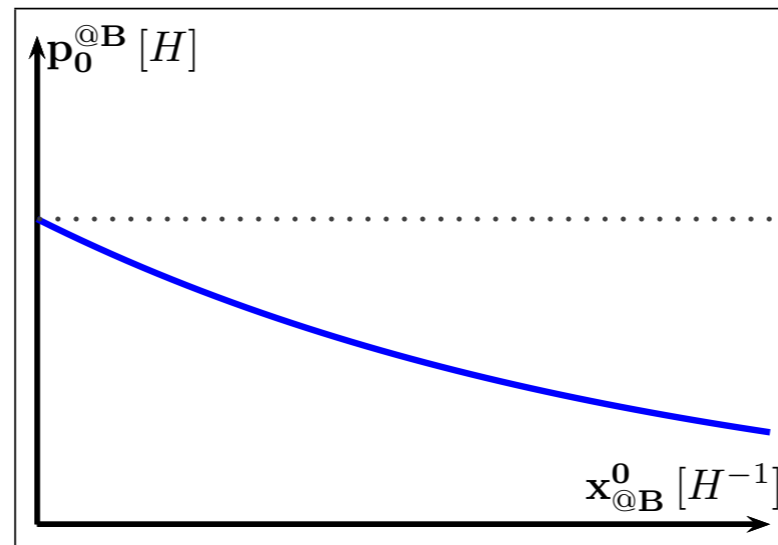
•Amelino-Camelia, Barcaroli, Gubitosi, Loreti, CQG 2013

♦ related to the fact that in Hopf algebras noncommutativity induces curvature in the dual space, and viceversa

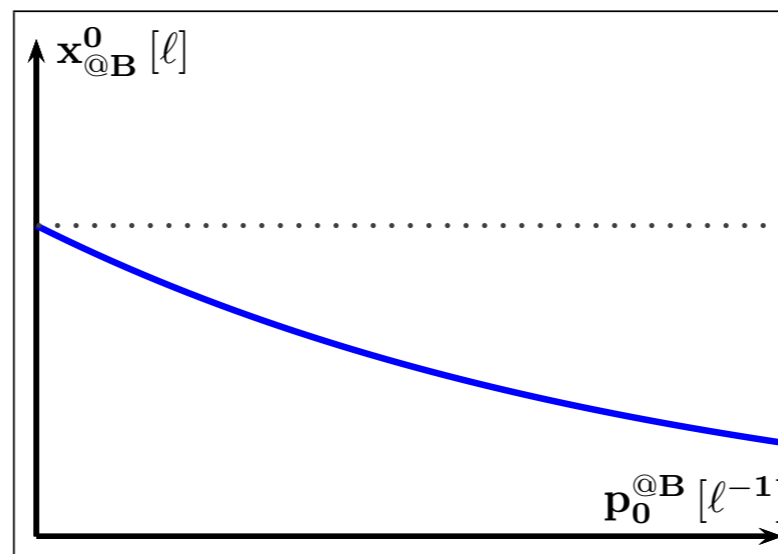
•Majid ArXiv: hep-th/0604130

Duality between de Sitter spacetime and de Sitter momentum space

- ♦ correlation between time of detection and energy, for fixed energy of emission (de Sitter spacetime) - *redshift*



- ♦ correlation between time of detection and energy, for fixed time of emission (de Sitter momentum space) - *time shift*



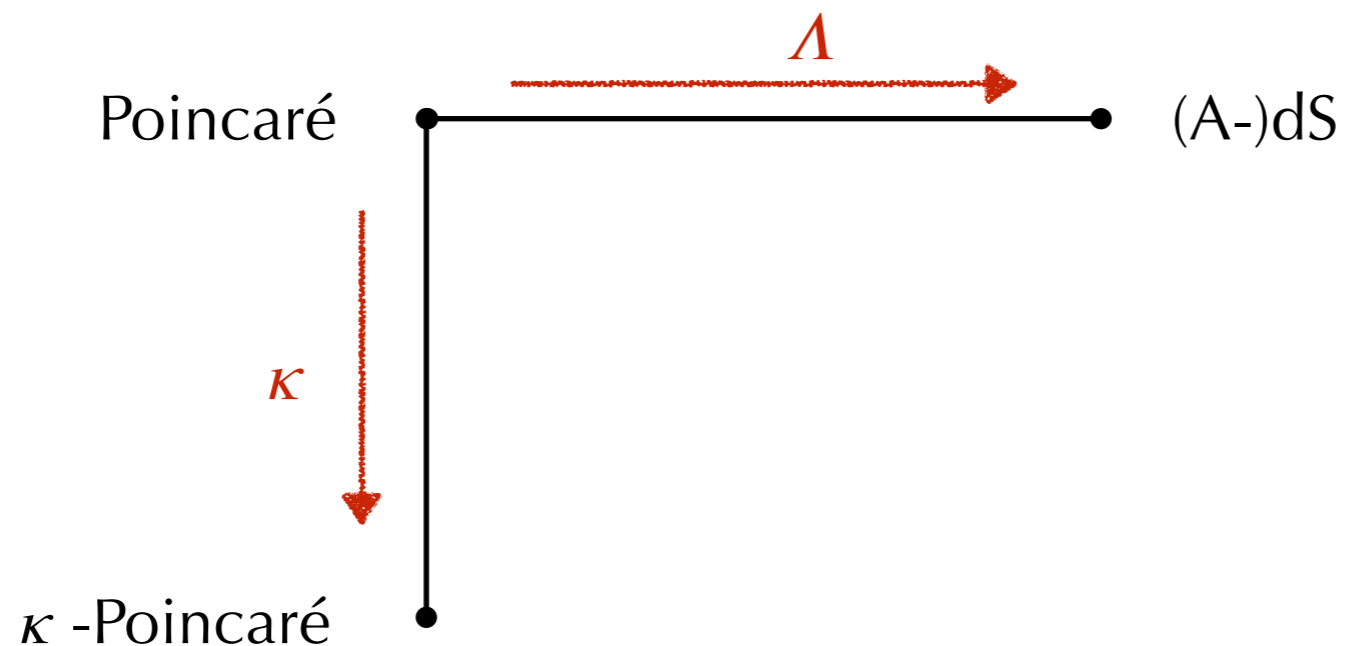
Deforming the Poincaré algebra of special relativity

- ◆ **Classical** deformations of the Poincaré algebra

Governed by the curvature parameter Λ : (Anti-)de Sitter symmetries

- ◆ **Quantum** deformations of the Poincaré algebra

Governed by the quantum deformation parameter κ



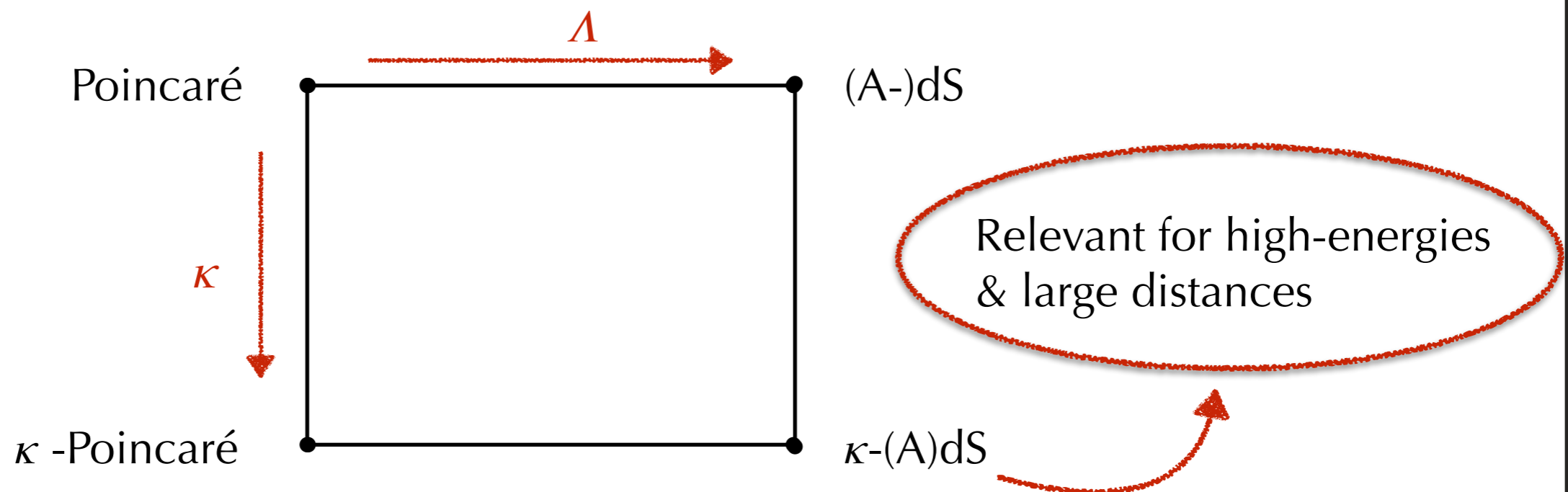
Deforming the Poincaré algebra of special relativity

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Classical+Quantum deformation of the Poincaré algebra: κ -(A)dS

♦ **Classical+Quantum** deformations of the Poincaré algebra

$$[\eta \equiv -\Lambda^2]$$

The 3+1 κ -(A)dS (Poisson) algebra is generated by the r-matrix

$$r_\Lambda = \frac{1}{\kappa}(K_1 \wedge P_1 + K_2 \wedge P_2 + K_3 \wedge P_3 + \eta J_1 \wedge J_2)$$

Which produces the cocommutators via $\delta(X) = [X \otimes 1 + 1 \otimes X, r_\Lambda]$, $X \in (A)dS$

$$\delta(P_0) = \delta(J_3) = 0, \quad \delta(J_1) = \frac{\eta}{\kappa} J_1 \wedge J_3, \quad \delta(J_2) = \frac{\eta}{\kappa} J_2 \wedge J_3,$$

$$\delta(P_1) = \frac{1}{\kappa}(P_1 \wedge P_0 - \eta P_3 \wedge J_1 - \eta^2 K_2 \wedge J_3 + \eta^2 K_3 \wedge J_2),$$

$$\delta(P_2) = \frac{1}{\kappa}(P_2 \wedge P_0 - \eta P_3 \wedge J_2 + \eta^2 K_1 \wedge J_3 - \eta^2 K_3 \wedge J_1),$$

$$\delta(P_3) = \frac{1}{\kappa}(P_3 \wedge P_0 + \eta P_1 \wedge J_1 + \eta P_2 \wedge J_2 - \eta^2 K_1 \wedge J_2 + \eta^2 K_2 \wedge J_1),$$

$$\delta(K_1) = \frac{1}{\kappa}(K_1 \wedge P_0 + P_2 \wedge J_3 - P_3 \wedge J_2 - \eta K_3 \wedge J_1),$$

$$\delta(K_2) = \frac{1}{\kappa}(K_2 \wedge P_0 - P_1 \wedge J_3 + P_3 \wedge J_1 - \eta K_3 \wedge J_2),$$

$$\delta(K_3) = \frac{1}{\kappa}(K_3 \wedge P_0 + P_1 \wedge J_2 - P_2 \wedge J_1 + \eta K_1 \wedge J_1 + \eta K_2 \wedge J_2).$$

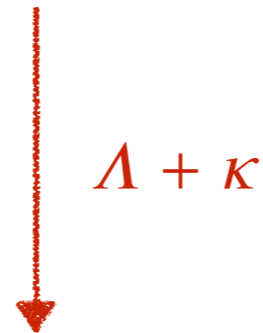
The r-matrix is unique if one asks that P_0 has trivial coproduct

Classical+Quantum deformation of the Poincaré algebra: κ -(A)dS

◆ **Classical+Quantum** deformations of the Poincaré algebra

$$\begin{array}{lll}
 [J_a, J_b] = \epsilon_{abc} J_c & [K_a, P_0] = P_a & [P_0, P_a] = 0 \\
 [J_a, P_b] = \epsilon_{abc} P_c & [K_a, P_b] = \delta_{ab} P_0 & [P_a, P_b] = 0 \\
 [J_a, K_b] = \epsilon_{abc} K_c & [K_a, K_b] = -\epsilon_{abc} J_c & [P_0, J_a] = 0
 \end{array}$$

$$\mathcal{C} = P_0^2 - \mathbf{P}^2$$



Translations sector:

noncommutativity is induced by curvature, nonlinearity by quantum deformation

$$\{P_1, P_2\} = -\eta^2 \frac{\sinh(2\frac{\eta}{\kappa} J_3)}{2\eta/\kappa} - \frac{\eta}{2\kappa} (2P_3^2 + \eta^2(J_1^2 + J_2^2)) - \frac{\eta^5}{4\kappa^3} e^{-2\frac{\eta}{\kappa} J_3} (J_1^2 + J_2^2)^2$$

$$\{P_1, P_3\} = \frac{1}{2} \eta^2 J_2 \left(1 + e^{-2\frac{\eta}{\kappa} J_3} \left[1 + \frac{\eta^2}{\kappa^2} (J_1^2 + J_2^2) \right] \right) + \frac{\eta}{\kappa} P_2 P_3$$

$$\{P_2, P_3\} = -\frac{1}{2} \eta^2 J_1 \left(1 + e^{-2\frac{\eta}{\kappa} J_3} \left[1 + \frac{\eta^2}{\kappa^2} (J_1^2 + J_2^2) \right] \right) - \frac{\eta}{\kappa} P_1 P_3$$

$$[\eta \equiv -\Lambda^2]$$

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Rotations sector:

$$[\eta \equiv -\Lambda^2]$$

$$\{J_1, J_2\} = \frac{e^{2\frac{\eta}{\kappa}J_3} - 1}{2\eta/\kappa} - \frac{\eta}{2\kappa} (J_1^2 + J_2^2), \quad \{J_1, J_3\} = -J_2, \quad \{J_2, J_3\} = J_1$$

$$\Delta(J_3) = J_3 \otimes 1 + 1 \otimes J_3 \quad \Delta(J_1) = J_1 \otimes e^{\frac{\eta}{\kappa}J_3} + 1 \otimes J_1 \quad \Delta(J_2) = J_2 \otimes e^{\frac{\eta}{\kappa}J_3} + 1 \otimes J_2$$

Deformed rotation sector! This is a pure “ $\Lambda + \kappa$ ” effect and only present in 3+1D

Classical+Quantum effects on particle propagation

- ♦ The interplay of curvature and quantum deformation can be seen at the level of redshift and worldlines in the 1+1 case (3+1 case is more complicated because of the nontrivial rotation sector)

the dispersion relation and the worldlines show explicitly the interplay, already at first order in the deformation parameters:

$$p_0 = -p_1 \left(1 - Hx^0 - \ell p_1 \left(\frac{1}{2} - Hx^0 \right) \right)$$

$$\begin{aligned} [H = \sqrt{\Lambda}] \\ [\ell \equiv 1/\kappa] \end{aligned}$$

$$x^1 - \bar{x}^1 = (x^0 - \bar{x}^0)(1 - \ell p_1) - \frac{1}{2}H ((x^0)^2 - (\bar{x}^0)^2) (1 - 2\ell p_1)$$

these features are inherited by observable properties, such as the time delay in the travel time of photons with different energies and the energy redshift of a photon traveling between far away observers:

$$\Delta x^0 = \ell a^0 \Delta p_0 (1 + H a^0)$$

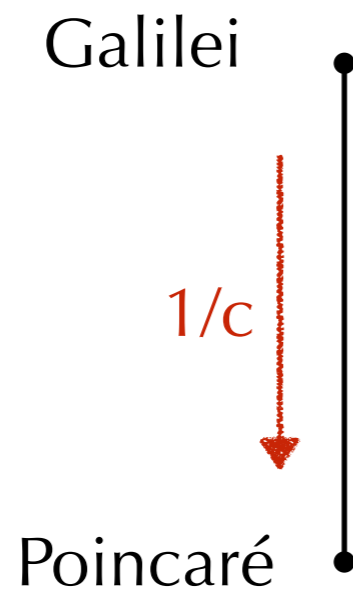
$$\Delta p_0 = -H p_0 x^0 \left(1 + \frac{\ell}{2} p_0 \right)$$

Speed-of-light contractions of the Poincaré algebra

- ◆ **Galilean** contraction (speed/space contraction) $1/c \rightarrow 0$

Small-velocities/large time intervals

'Absolute time': the light cone opens along $t=0$, spatial leaves at constant time

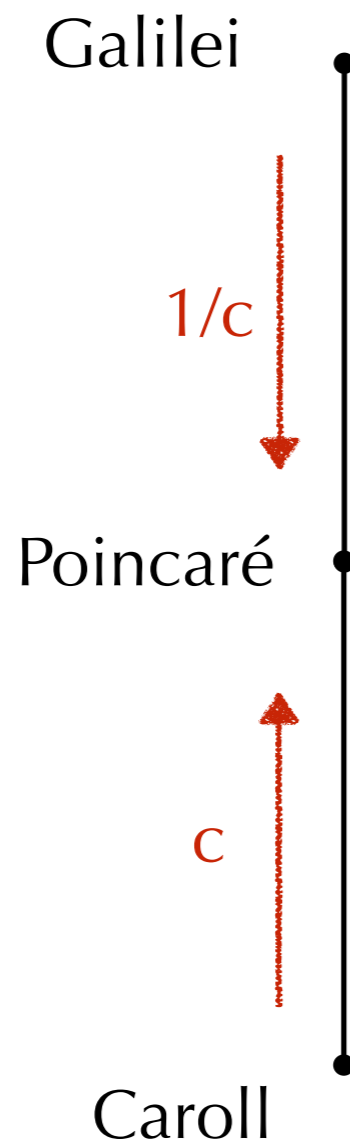


Speed-of-light contractions of the Poincaré algebra

♦ **Carroll** contraction (speed/time contraction) $c \rightarrow 0$

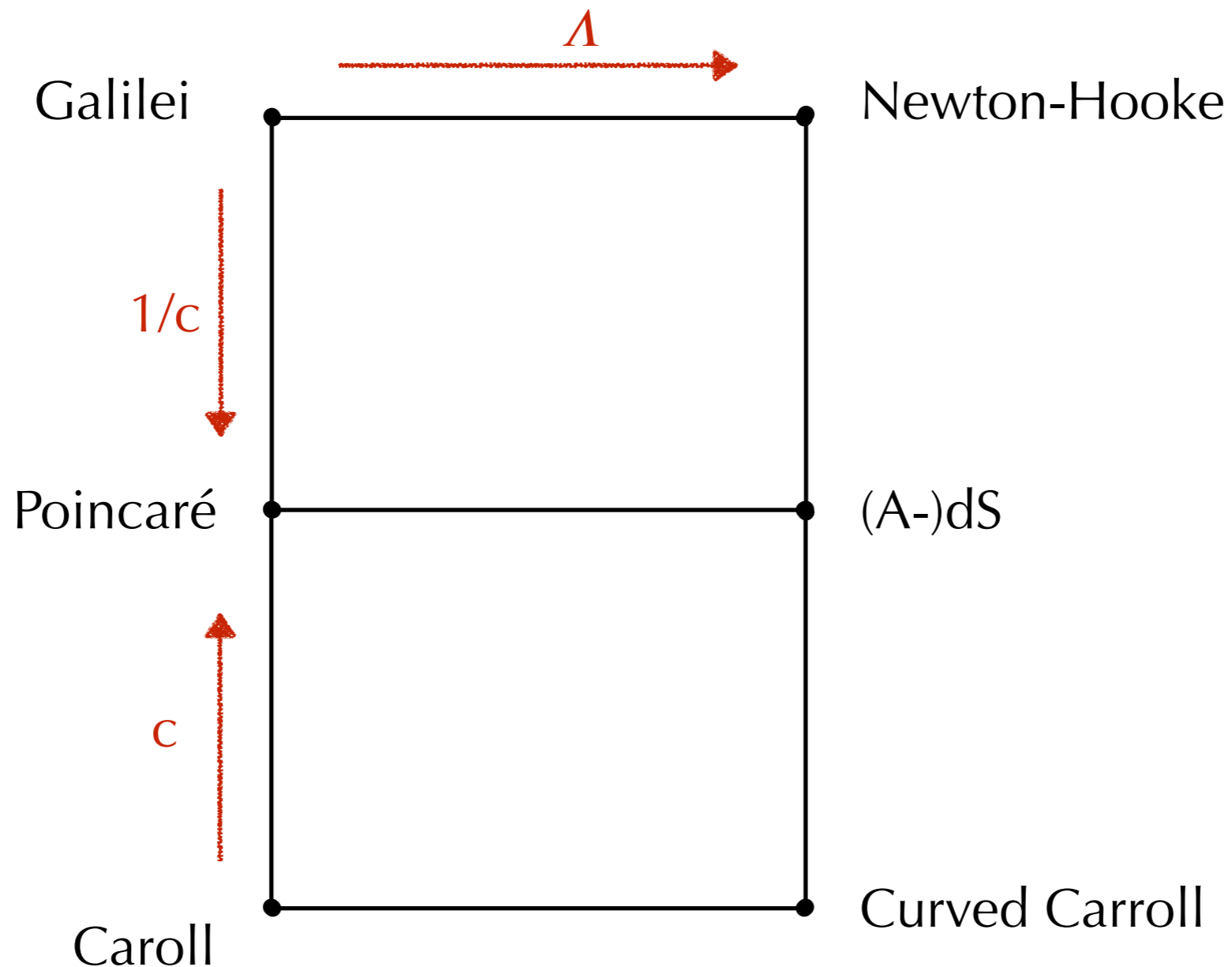
Large space intervals

'Absolute space': the light cone closes along t , time leaves at constant space



Speed-of-light contractions of the (A)dS algebra

- ◆ **Galilei and Carroll** contractions can be performed also in the presence of curvature



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◆ Galilei: Inönü–Wigner contraction, induced by the algebra automorphism $\mathcal{P}(P_0, P_a, K_a, J_a) = (P_0, -P_a, -K_a, J_a)$ (speed-space contraction)

This allows us to perform the rescaling $P_a = \frac{\bar{P}_a}{c}, K_a = \frac{\bar{K}_a}{c}$

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$$[K_a, K_b] = -\epsilon_{abc} \frac{1}{c^2} J_c$$

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$1/c \rightarrow 0$



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$$\mathcal{C} = \mathbf{P}^2 - \Lambda \mathbf{K}^2$$

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This allows us to perform the rescaling $P_0 = c\bar{P}_0, K_a = c\bar{K}_a$

$$[K_a, K_b] = -\epsilon_{abc}c^2J_c$$

$$[K_a, P_0] = c^2P_a$$

$$[P_a, P_b] = \Lambda\epsilon_{abc}J_c$$

$c \rightarrow 0$



$$[K_a, K_b] = 0$$

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$$\mathcal{C} = \frac{1}{c^2}P_0^2 - \mathbf{P}^2 - \Lambda \left(\mathbf{J}^2 - \frac{1}{c^2}\mathbf{K}^2 \right)$$

$$\mathcal{C} = P_0^2 + \Lambda \mathbf{K}^2$$

◆ Carrollian spacetimes \mathbf{C}_Λ^{3+1} are obtained as the homogeneous space

$$\mathbf{C}_\Lambda^{3+1} = \mathbf{C}_\Lambda / \mathbf{H}, \quad \mathbf{H} = \text{ISO}(3) = \{\mathbf{K}, \mathbf{J}\}$$

Speed-of-light contractions of the (A)dS algebra

◆ **Galilei and Carroll** contractions can be performed also in the presence of curvature

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Speed-of-light contractions of the (A)dS algebra

	Galilean limit	(A)dS	Carrollian limit
$[J_a, J_b]$		$\epsilon_{abc} J_c$	
$[J_a, P_b]$		$\epsilon_{abc} P_c$	
$[J_a, K_b]$		$\epsilon_{abc} K_c$	
$[J_a, P_0]$		0	
$[K_a, K_b]$	0	$-\epsilon_{abc} J_c$	0
$[K_a, P_b]$	0	$\delta_{ab} P_0$	$\delta_{ab} P_0$
$[K_a, P_0]$	P_a	P_a	0
$[P_a, P_b]$	0	$\Lambda \epsilon_{abc} J_c$	$\Lambda \epsilon_{abc} J_c$
$[P_a, P_0]$	ΛK_a	ΛK_a	ΛK_a

Speed-of-light contractions of the (A)dS algebra

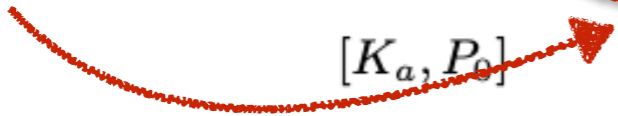
	Galilean limit	(A)dS	Carrollian limit
$[J_a, J_b]$		$\epsilon_{abc} J_c$	
$[J_a, P_b]$		$\epsilon_{abc} P_c$	
$[J_a, K_b]$		$\epsilon_{abc} K_c$	
$[J_a, P_0]$		0	
$[K_a, K_b]$	0	$-\epsilon_{abc} J_c$	0
$[K_a, P_b]$	0	$\delta_{ab} P_0$	$\delta_{ab} P_0$
$[K_a, P_0]$	P_a	P_a	0
$[P_a, P_b]$	0	$\Lambda \epsilon_{abc} J_c$	$\Lambda \epsilon_{abc} J_c$
$[P_a, P_0]$	ΛK_a	ΛK_a	ΛK_a

Nonrelativistic boost

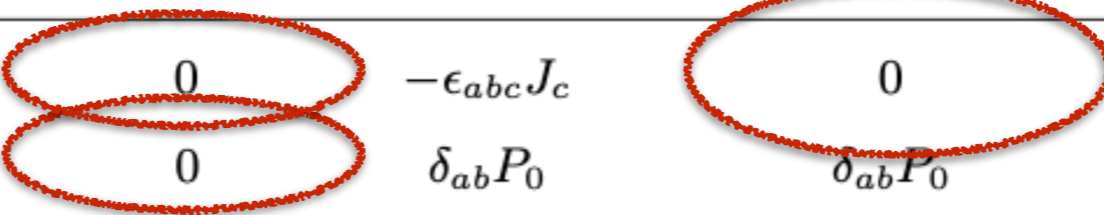
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$[J_a, P_0]$		0	
$[K_a, K_b]$	0	$-\epsilon_{abc} J_c$	0
$[K_a, P_b]$	0	$\delta_{ab} P_0$	$\delta_{ab} P_0$
$[K_a, P_0]$	P_a	P_a	0
$[P_a, P_b]$	0	$\Lambda \epsilon_{abc} J_c$	$\Lambda \epsilon_{abc} J_c$
$[P_a, P_0]$	ΛK_a	ΛK_a	ΛK_a

Absolute time



Nonrelativistic boost



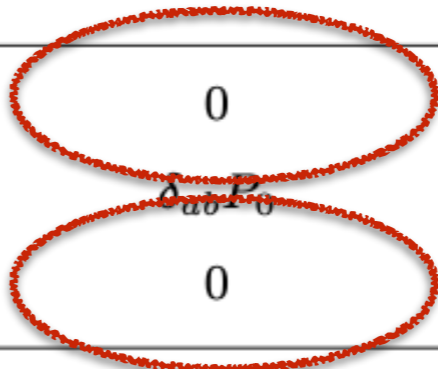
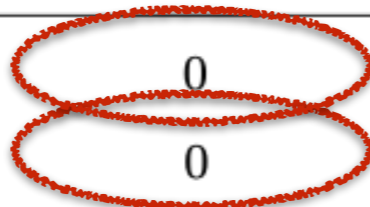
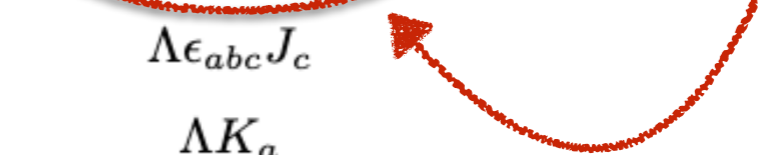
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Absolute time

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Absolute space

Flat spatial slices even when curvature is nonzero

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$[K_a, P_b]$	0	$\delta_{ab} P_0$	$\delta_{ab} P_0$
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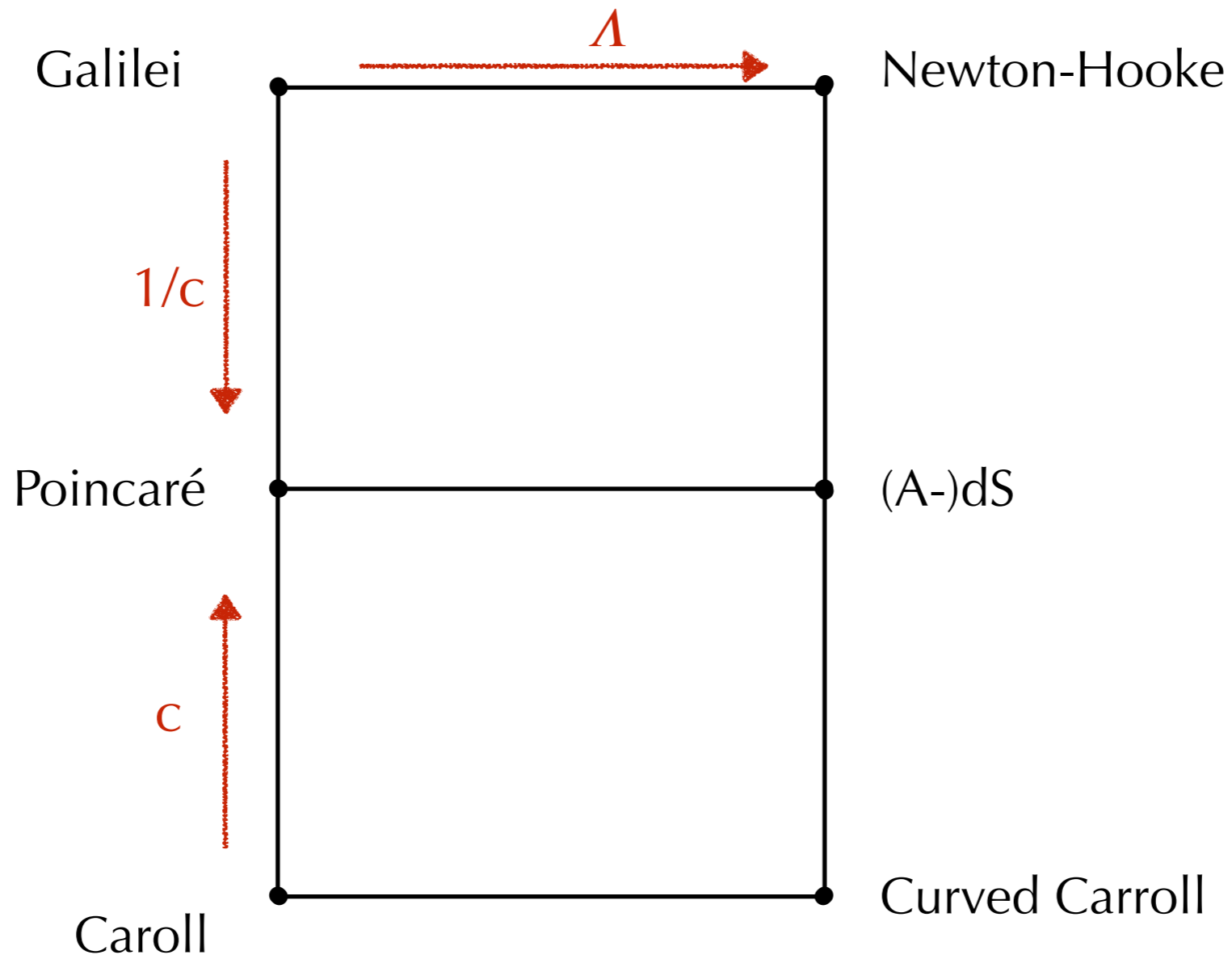
Absolute space

Flat spatial slices even when curvature is nonzero

- ◆ Can we import this classical intuition into the realm of quantum symmetries?

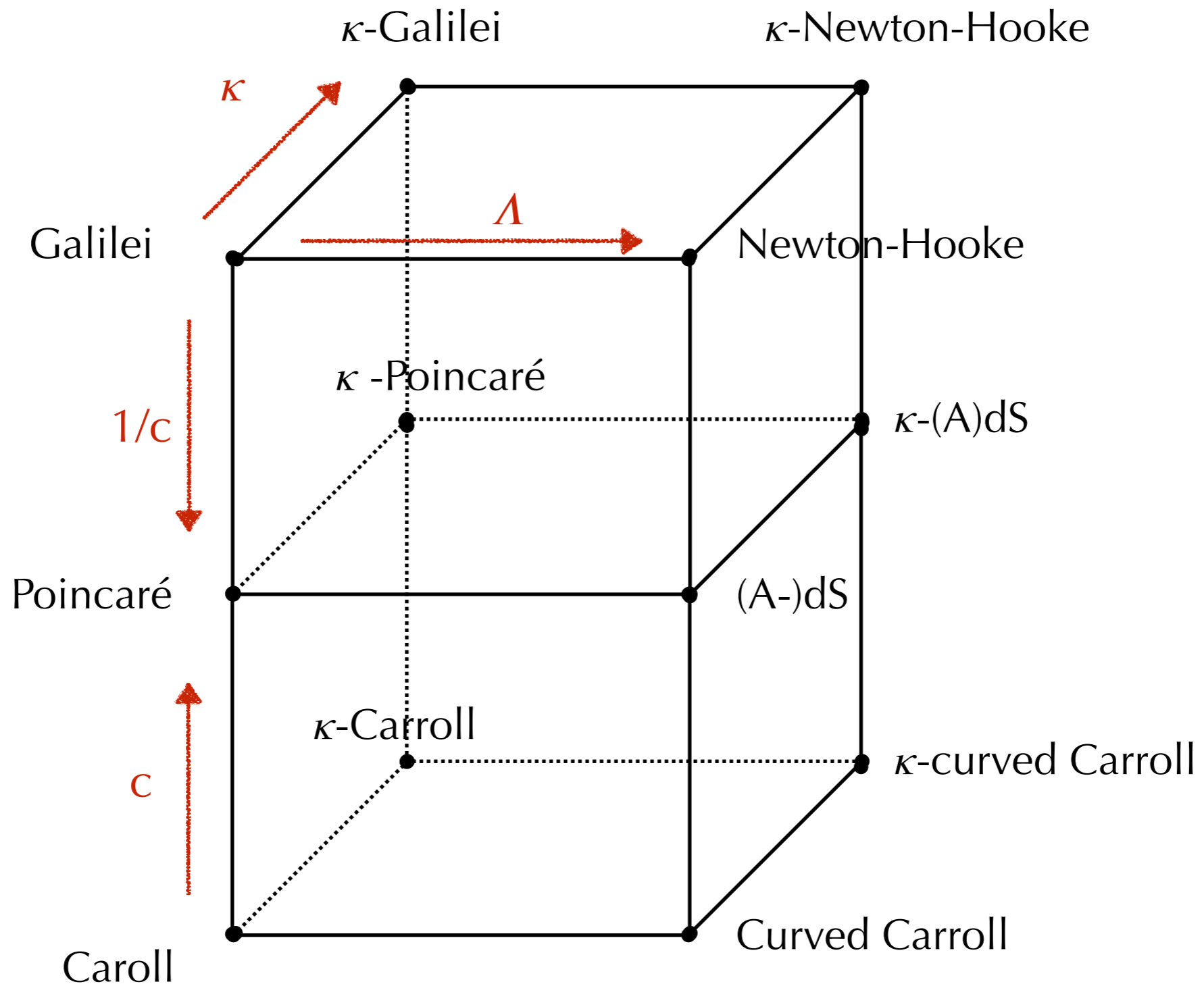
Possible relativistic models

- ◆ One can consider quantum deformations of each of these classical relativistic models



Possible relativistic models

- ◆ One can consider quantum deformations of each of these classical relativistic models



Speed-of-light contraction of quantum-deformed algebras

- ◆ The easiest way to work out the Galilean and Carrollian quantum algebras is via contraction of the κ -(A)dS algebra
- ◆ This contraction might require a rescaling of the quantum deformation parameter, along with the generators, in order to obtain meaningful structures, either at the level of the r-matrix or at the level of the cocommutators

Galilei contraction (flat case):

Contraction at the level of the r-matrix $r = \frac{c^2}{\kappa} (K_1 \wedge P_1 + K_2 \wedge P_2 + K_3 \wedge P_3)$ requires a rescaling of κ , however this produces trivial cocommutators

Contraction can be performed directly at the level of the cocommutators and does not require to rescale κ

$$\begin{aligned}
 [K_a, P_b] &= \frac{\delta_{ab}}{c^2} \left[\frac{\kappa}{2} (1 - e^{-2P_0/\kappa}) + c^2 \frac{\mathbf{P}^2}{2\kappa} \right] - \frac{P_a P_b}{\kappa} \\
 [K_a, K_b] &= -\frac{\epsilon_{abc}}{c^2} J_c
 \end{aligned}
 \xrightarrow{1/c \rightarrow 0}
 \begin{aligned}
 [K_a, P_b] &= \delta_{ab} \frac{\mathbf{P}^2}{2\kappa} - \frac{P_a P_b}{\kappa} \\
 [K_a, K_b] &= 0 \\
 \mathcal{C}_\kappa &= e^{P_0/\kappa} \mathbf{P}^2
 \end{aligned}$$

The coproducts are the same as in κ -Poincaré

- Ballesteros, GG, Gutierrez-Sagredo, Herranz, PLB 2020
- Giller, Kosinski, Majewski, Maslanka, Kunz, PLB 1992

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$$[K_a, P_0] = P_a c^2$$

$$[K_a, P_b] = c\delta_{ab} \left[\frac{\kappa}{2c} (1 - e^{-2P_0/\kappa}) + \frac{\mathbf{P}^2 c}{2\kappa} \right] - c \frac{P_a P_b}{\kappa}$$

$$[K_a, K_b] = -c^2 \epsilon_{abc} J_c$$

$c \rightarrow 0$ 

$$[K_a, P_0] = 0$$

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$$[K_a, K_b] = 0$$

$$\mathcal{C}_\kappa = 2\kappa^2 (\cosh(P_0/\kappa) - 1)$$

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The coproducts are the same as in κ -Poincaré

Speed-of-light contraction of quantum-deformed algebras

$$[\eta \equiv -\Lambda^2]$$

◆ Galilei contraction of the κ -(A)dS algebra

$$\begin{aligned} \{J_1, J_2\} &= \frac{e^{2\eta J_3/\kappa} - 1}{2\eta/\kappa} - \frac{\eta}{2\kappa} (J_1^2 + J_2^2), & \{J_1, J_3\} &= -J_2, & \{J_2, J_3\} &= J_1, \\ \{J_1, P_1\} &= \frac{\eta}{\kappa} J_1 P_2, & \{J_1, P_2\} &= P_3 - \frac{\eta}{\kappa} J_1 P_1, & \{J_1, P_3\} &= -P_2, \\ \{J_2, P_1\} &= -P_3 + \frac{\eta}{\kappa} J_2 P_2, & \{J_2, P_2\} &= -\frac{\eta}{\kappa} J_2 P_1, & \{J_2, P_3\} &= P_1, \\ \{J_3, P_1\} &= P_2, & \{J_3, P_2\} &= -P_1, & \{J_3, P_3\} &= 0, \\ \{J_1, K_1\} &= \frac{\eta}{\kappa} J_1 K_2, & \{J_1, K_2\} &= K_3 - \frac{\eta}{\kappa} J_1 K_1, & \{J_1, K_3\} &= -K_2, \\ \{J_2, K_1\} &= -K_3 + \frac{\eta}{\kappa} J_2 K_2, & \{J_2, K_2\} &= -\frac{\eta}{\kappa} J_2 K_1, & \{J_2, K_3\} &= K_1, \\ \{J_3, K_1\} &= K_2, & \{J_3, K_2\} &= -K_1, & \{J_3, K_3\} &= 0, \\ \{K_a, P_0\} &= P_a, & \{P_0, P_a\} &= \eta^2 K_a, & \{P_0, J_a\} &= 0, \end{aligned}$$

$$\begin{aligned} \{K_1, P_1\} &= \frac{1}{2\kappa} (P_2^2 + P_3^2 - P_1^2) + \frac{\eta^2}{2\kappa} (K_2^2 + K_3^2 - K_1^2), & \{P_1, K_3\} &= \frac{1}{\kappa} (P_1 P_3 + \eta^2 K_1 K_3 + \eta K_2 P_3), \\ \{K_2, P_2\} &= \frac{1}{2\kappa} (P_1^2 + P_3^2 - P_2^2) + \frac{\eta^2}{2\kappa} (K_1^2 + K_3^2 - K_2^2), & \{P_3, K_1\} &= \frac{1}{\kappa} (P_1 P_3 + \eta^2 K_1 K_3 - \eta P_2 K_3), \\ \{K_3, P_3\} &= \frac{1}{2\kappa} [(P_1 + \eta K_2)^2 + (P_2 - \eta K_1)^2 - P_3^2 - \eta^2 K_3^2], & \{P_2, K_3\} &= \frac{1}{\kappa} (P_2 P_3 + \eta^2 K_2 K_3 - \eta K_1 P_3), \\ \{P_1, K_2\} &= \frac{1}{\kappa} (P_1 P_2 + \eta^2 K_1 K_2 - \eta P_3 K_3), & \{P_3, K_2\} &= \frac{1}{\kappa} (P_2 P_3 + \eta^2 K_2 K_3 + \eta P_1 K_3), \\ \{P_2, K_1\} &= \frac{1}{\kappa} (P_1 P_2 + \eta^2 K_1 K_2 + \eta P_3 K_3), & \{K_a, K_b\} &= -\frac{\eta}{\kappa} \epsilon_{abc} K_c K_3, & \{P_a, P_b\} &= -\frac{\eta}{\kappa} \epsilon_{abc} P_c P_3. \end{aligned}$$

$$\begin{aligned} \text{Casimir: } \mathcal{C}_\kappa &= e^{P_0/\kappa} (\mathbf{P}^2 + \eta^2 \mathbf{K}^2) \left[\cosh(\eta J_3/\kappa) + \frac{\eta^2}{2\kappa^2} (J_1^2 + J_2^2) e^{-\eta J_3/\kappa} \right] \\ &\quad - 2\eta^2 e^{P_0/\kappa} \left[\frac{\sinh(\eta J_3/\kappa)}{\eta} R_3 + \frac{1}{\kappa} \left(J_1 R_1 + J_2 R_2 + \frac{\eta}{2\kappa} (J_1^2 + J_2^2) R_3 \right) e^{-\eta J_3/\kappa} \right] \end{aligned}$$

The coproducts are the same as in κ -(A)dS

Speed-of-light contraction of quantum-deformed algebras

$$[\eta \equiv -\Lambda^2]$$

◆ Galilei contraction of the κ -(A)dS algebra

$$\{J_1, J_2\} = \frac{e^{2\eta J_3/\kappa} - 1}{2\eta/\kappa} - \frac{\eta}{2\kappa} (J_1^2 + J_2^2), \quad \{J_1, J_3\} = -J_2, \quad \{J_2, J_3\} = J_1,$$

$$\{J_1, P_1\} = \frac{\eta}{\kappa} J_1 P_2, \quad \{J_1, P_2\} = P_3 - \frac{\eta}{\kappa} J_1 P_1, \quad \{J_1, P_3\} = -P_2,$$

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$$\{J_3, P_1\} = P_2, \quad \{J_3, P_2\} = -P_1, \quad \{J_3, P_3\} = 0,$$

$$\{J_1, K_1\} = \frac{\eta}{\kappa} J_1 K_2, \quad \{J_1, K_2\} = K_3 - \frac{\eta}{\kappa} J_1 K_1, \quad \{J_1, K_3\} = -K_2,$$

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$$\{J_3, K_1\} = K_2, \quad \{J_3, K_2\} = -K_1, \quad \{J_3, K_3\} = 0,$$

$$\{K_a, P_0\} = P_a, \quad \{P_0, P_a\} = \eta^2 K_a, \quad \{P_0, J_a\} = 0,$$

$$\{K_1, P_1\} = \frac{1}{2\kappa} (P_2^2 + P_3^2 - P_1^2) + \frac{\eta^2}{2\kappa} (K_2^2 + K_3^2 - K_1^2),$$

$$\{P_1, K_3\} = \frac{1}{\kappa} (P_1 P_3 + \eta^2 K_1 K_3 + \eta K_2 P_3),$$

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$$\{P_3, K_1\} = \frac{1}{\kappa} (P_1 P_3 + \eta^2 K_1 K_3 - \eta P_2 K_3),$$

$$\{K_3, P_3\} = \frac{1}{2\kappa} [(P_1 + \eta K_2)^2 + (P_2 - \eta K_1)^2 - P_3^2 - \eta^2 K_3^2],$$

$$\{P_2, K_3\} = \frac{1}{\kappa} (P_2 P_3 + \eta^2 K_2 K_3 - \eta K_1 P_3),$$

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$$\{P_3, K_2\} = \frac{1}{\kappa} (P_2 P_3 + \eta^2 K_2 K_3 + \eta P_1 K_3),$$

$$\{P_2, K_1\} = \frac{1}{\kappa} (P_1 P_2 + \eta^2 K_1 K_2 + \eta P_3 K_3),$$

$$\{K_a, K_b\} = -\frac{\eta}{\kappa} \epsilon_{abc} K_c K_3, \quad \{P_a, P_b\} = -\frac{\eta}{\kappa} \epsilon_{abc} P_c P_3.$$

$$\text{Casimir: } \mathcal{C}_\kappa = e^{P_0/\kappa} (\mathbf{P}^2 + \eta^2 \mathbf{K}^2) \left[\cosh(\eta J_3/\kappa) + \frac{\eta^2}{2\kappa^2} (J_1^2 + J_2^2) e^{-\eta J_3/\kappa} \right] \\ - 2\eta^2 e^{P_0/\kappa} \left[\frac{\sinh(\eta J_3/\kappa)}{\eta} R_3 + \frac{1}{\kappa} \left(J_1 R_1 + J_2 R_2 + \frac{\eta}{2\kappa} (J_1^2 + J_2^2) R_3 \right) e^{-\eta J_3/\kappa} \right]$$

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$$\begin{aligned} \{K_1, P_1\} &= \frac{1}{2\kappa} (P_2^2 + P_3^2 - P_1^2) + \frac{\eta^2}{2\kappa} (K_2^2 + K_3^2 - K_1^2), & \{P_1, K_3\} &= \frac{1}{\kappa} (P_1 P_3 + \eta^2 K_1 K_3 + \eta K_2 P_3), \\ \{K_2, P_2\} &= \frac{1}{2\kappa} (P_1^2 + P_3^2 - P_2^2) + \frac{\eta^2}{2\kappa} (K_1^2 + K_3^2 - K_2^2), & \{P_3, K_1\} &= \frac{1}{\kappa} (P_1 P_3 + \eta^2 K_1 K_3 - \eta P_2 K_3), \\ \{K_3, P_3\} &= \frac{1}{2\kappa} [(P_1 + \eta K_2)^2 + (P_2 - \eta K_1)^2 - P_3^2 - \eta^2 K_3^2], & \{P_2, K_3\} &= \frac{1}{\kappa} (P_2 P_3 + \eta^2 K_2 K_3 - \eta K_1 P_3), \\ \{P_1, K_2\} &= \frac{1}{\kappa} (P_1 P_2 + \eta^2 K_1 K_2 - \eta P_3 K_3), & \{P_3, K_2\} &= \frac{1}{\kappa} (P_2 P_3 + \eta^2 K_2 K_3 + \eta P_1 K_3), \\ \{P_2, K_1\} &= \frac{1}{\kappa} (P_1 P_2 + \eta^2 K_1 K_2 + \eta P_3 K_3), & \{K_a, K_b\} &= -\frac{\eta}{\kappa} \epsilon_{abc} K_c K_3, & \{P_a, P_b\} &= -\frac{\eta}{\kappa} \epsilon_{abc} P_c P_3. \end{aligned}$$

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$$\begin{aligned} \{J_1, J_2\} &= \frac{e^{2\eta J_3/\kappa} - 1}{2\eta/\kappa} - \frac{\eta}{2\kappa} (J_1^2 + J_2^2), & \{J_1, J_3\} &= -J_2, & \{J_2, J_3\} &= J_1, \\ \{J_1, P_1\} &= \frac{\eta}{\kappa} J_1 P_2, & \{J_1, P_2\} &= P_3 - \frac{\eta}{\kappa} J_1 P_1, & \{J_1, P_3\} &= -P_2, \\ \{J_2, P_1\} &= -P_3 + \frac{\eta}{\kappa} J_2 P_2, & \{J_2, P_2\} &= -\frac{\eta}{\kappa} J_2 P_1, & \{J_2, P_3\} &= P_1, \\ \{J_3, P_1\} &= P_2, & \{J_3, P_2\} &= -P_1, & \{J_3, P_3\} &= 0, \\ \{J_1, K_1\} &= \frac{\eta}{\kappa} J_1 K_2, & \{J_1, K_2\} &= K_3 - \frac{\eta}{\kappa} J_1 K_1, & \{J_1, K_3\} &= -K_2, \\ \{J_2, K_1\} &= -K_3 + \frac{\eta}{\kappa} J_2 K_2, & \{J_2, K_2\} &= -\frac{\eta}{\kappa} J_2 K_1, & \{J_2, K_3\} &= K_1, \\ \{J_3, K_1\} &= K_2, & \{J_3, K_2\} &= -K_1, & \{J_3, K_3\} &= 0, \\ \{K_a, P_0\} &= P_a, & \{P_0, P_a\} &= \eta^2 K_a, & \{P_0, J_a\} &= 0, \end{aligned}$$

$$\begin{aligned} \{K_1, P_1\} &= \frac{1}{2\kappa} (P_2^2 + P_3^2 - P_1^2) + \frac{\eta^2}{2\kappa} (K_2^2 + K_3^2 - K_1^2), & \{P_1, K_3\} &= \frac{1}{\kappa} (P_1 P_3 + \eta^2 K_1 K_3 + \eta K_2 P_3), \\ \{K_2, P_2\} &= \frac{1}{2\kappa} (P_1^2 + P_3^2 - P_2^2) + \frac{\eta^2}{2\kappa} (K_1^2 + K_3^2 - K_2^2), & \{P_3, K_1\} &= \frac{1}{\kappa} (P_1 P_3 + \eta^2 K_1 K_3 - \eta P_2 K_3), \\ \{K_3, P_3\} &= \frac{1}{2\kappa} [(P_1 + \eta K_2)^2 + (P_2 - \eta K_1)^2 - P_3^2 - \eta^2 K_3^2], & \{P_2, K_3\} &= \frac{1}{\kappa} (P_2 P_3 + \eta^2 K_2 K_3 - \eta K_1 P_3), \\ \{P_1, K_2\} &= \frac{1}{\kappa} (P_1 P_2 + \eta^2 K_1 K_2 - \eta P_3 K_3), & \{P_3, K_2\} &= \frac{1}{\kappa} (P_2 P_3 + \eta^2 K_2 K_3 + \eta P_1 K_3), \\ \{P_2, K_1\} &= \frac{1}{\kappa} (P_1 P_2 + \eta^2 K_1 K_2 + \eta P_3 K_3), & \{K_a, K_b\} &= -\frac{\eta}{\kappa} \epsilon_{abc} K_c K_3, & \{P_a, P_b\} &= -\frac{\eta}{\kappa} \epsilon_{abc} P_c P_3. \end{aligned}$$

$$\begin{aligned} \text{Casimir: } \mathcal{C}_\kappa &= e^{P_0/\kappa} (\mathbf{P}^2 + \eta^2 \mathbf{K}^2) \left[\cosh(\eta J_3/\kappa) + \frac{\eta^2}{2\kappa^2} (J_1^2 + J_2^2) e^{-\eta J_3/\kappa} \right] \\ &\quad - 2\eta^2 e^{P_0/\kappa} \left[\frac{\sinh(\eta J_3/\kappa)}{\eta} R_3 + \frac{1}{\kappa} \left(J_1 R_1 + J_2 R_2 + \frac{\eta}{2\kappa} (J_1^2 + J_2^2) R_3 \right) e^{-\eta J_3/\kappa} \right] \end{aligned}$$

The coproducts are the same as in κ -(A)dS

Speed-of-light contraction of quantum-deformed algebras

◆ Carroll contraction of the κ -(A)dS algebra

$$[\eta \equiv -\Lambda^2]$$

$$\begin{aligned} \{J_a, J_b\} &= \epsilon_{abc} J_c, & \{J_a, P_b\} &= \epsilon_{abc} P_c, & \{J_a, K_b\} &= \epsilon_{abc} K_c, \\ \{K_a, P_0\} &= 0, & \{K_a, K_b\} &= 0, & \{P_0, J_a\} &= 0, \\ \{P_0, P_a\} &= \eta^2 K_a, & \{P_a, P_b\} &= -\eta^2 \epsilon_{abc} J_c, \\ \{K_a, P_b\} &= \delta_{ab} \left(\frac{1 - e^{-2P_0/\kappa}}{2/\kappa} + \frac{\eta^2}{2\kappa} \mathbf{K}^2 \right) - \frac{\eta^2}{\kappa} K_a K_b. \end{aligned}$$

Casimir:
$$\mathcal{C}_\kappa = 2\kappa^2 (\cosh(P_0/\kappa) - 1) - \eta^2 e^{P_0/\kappa} \mathbf{K}^2.$$

Coproducts:
$$\begin{aligned} \Delta(P_0) &= P_0 \otimes 1 + 1 \otimes P_0, & \Delta(J_a) &= J_a \otimes 1 + 1 \otimes J_a, \\ \Delta(P_a) &= P_a \otimes 1 + e^{-P_0/\kappa} \otimes P_a - \frac{\eta^2}{\kappa} \epsilon_{abc} K_b \otimes J_c, \\ \Delta(K_a) &= K_a \otimes 1 + e^{-P_0/\kappa} \otimes K_a, \end{aligned}$$

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Possible relativistic models

	(curved) κ -Galilei	κ -(A)dS	(curved) κ -Carroll
$[J_a, J_b]$			
$[J_a, P_b]$	anisotropy $\sim \frac{\Lambda}{\kappa}$	anisotropy $\sim \frac{\Lambda}{\kappa}$	isotropy
$[J_a, K_b]$			
$[J_a, P_0]$	0		
$[K_a, K_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa})$	$-\epsilon_{abc}J_c + O(\frac{\sqrt{\Lambda}}{\kappa})$	0
$[K_a, P_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa}, \frac{1}{\kappa})$	$\delta_{ab}P_0 + O(\frac{\Lambda}{\kappa}, \frac{1}{\kappa})$	$\delta_{ab}P_0 + O(\frac{\Lambda}{\kappa}, \frac{1}{\kappa})$
$[K_a, P_0]$	P_a	P_a	0
$[P_a, P_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa})$	$\Lambda\epsilon_{abc}J_c + O(\frac{\sqrt{\Lambda}}{\kappa})$	$\Lambda\epsilon_{abc}J_c$
$[P_a, P_0]$	ΛK_a	ΛK_a	ΛK_a

The “ $\Lambda + \kappa$ ” effects are milder in the Carroll models

Possible relativistic models

	(curved) κ -Galilei	κ -(Λ)dS	(curved) κ -Carroll
$[J_a, J_b]$			
$[J_a, P_b]$	anisotropy $\sim \frac{\Lambda}{\kappa}$	anisotropy $\sim \frac{\Lambda}{\kappa}$	isotropy
$[J_a, K_b]$			
$[J_a, P_0]$	0		
$[K_a, K_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa})$	$-\epsilon_{abc}J_c + O(\frac{\sqrt{\Lambda}}{\kappa})$	0
$[K_a, P_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa}, \frac{1}{\kappa})$	$\delta_{ab}P_0 + O(\frac{\Lambda}{\kappa}, \frac{1}{\kappa})$	$\delta_{ab}P_0 + O(\frac{\Lambda}{\kappa}, \frac{1}{\kappa})$
$[K_a, P_0]$	P_a	P_a	0
$[P_a, P_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa})$	$\Lambda\epsilon_{abc}J_c + O(\frac{\sqrt{\Lambda}}{\kappa})$	$\Lambda\epsilon_{abc}J_c$
$[P_a, P_0]$	ΛK_a	ΛK_a	ΛK_a

Carroll contraction restores isotropy

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Possible relativistic models

	(curved) κ -Galilei	κ -(A)dS	(curved) κ -Carroll
$[J_a, J_b]$			
$[J_a, P_b]$	anisotropy $\sim \frac{\Lambda}{\kappa}$	anisotropy $\sim \frac{\Lambda}{\kappa}$	isotropy
$[J_a, K_b]$			
$[J_a, P_0]$		0	
$[K_a, K_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa})$	$-\epsilon_{abc}J_c + O(\frac{\sqrt{\Lambda}}{\kappa})$	0
$[K_a, P_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa}, \frac{1}{\kappa})$	$\delta_{ab}P_0 + O(\frac{\Lambda}{\kappa}, \frac{1}{\kappa})$	$\delta_{ab}P_0 + O(\frac{\Lambda}{\kappa}, \frac{1}{\kappa})$
$[K_a, P_0]$	P_a	P_a	0
$[P_a, P_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa})$	$\Lambda\epsilon_{abc}J_c + O(\frac{\sqrt{\Lambda}}{\kappa})$	$\Lambda\epsilon_{abc}J_c$
$[P_a, P_0]$	ΛK_a	ΛK_a	ΛK_a

Carroll contraction restores isotropy

Non-commuting boosts

The " $\Lambda + \kappa$ " effects are milder in the Carroll models

Possible relativistic models

	(curved) κ -Galilei	κ -(Λ)dS	(curved) κ -Carroll
$[J_a, J_b]$			
$[J_a, P_b]$	anisotropy $\sim \frac{\Lambda}{\kappa}$	anisotropy $\sim \frac{\Lambda}{\kappa}$	isotropy
$[J_a, K_b]$			
$[J_a, P_0]$		0	
$[K_a, K_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa})$	$-\epsilon_{abc}J_c + O(\frac{\sqrt{\Lambda}}{\kappa})$	0
$[K_a, P_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa}, \frac{1}{\kappa})$	$\delta_{ab}P_0 + O(\frac{\Lambda}{\kappa}, \frac{1}{\kappa})$	$\delta_{ab}P_0 + O(\frac{\Lambda}{\kappa}, \frac{1}{\kappa})$
$[K_a, P_0]$	P_a	P_a	0
$[P_a, P_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa})$	$\Lambda\epsilon_{abc}J_c + O(\frac{\sqrt{\Lambda}}{\kappa})$	$\Lambda\epsilon_{abc}J_c$
$[P_a, P_0]$	ΛK_a	ΛK_a	ΛK_a

Carroll contraction restores isotropy

Non-commuting boosts

No absolute time

The " $\Lambda + \kappa$ " effects are milder in the Carroll models

Possible relativistic models

	(curved) κ -Galilei	κ -(Λ)dS	(curved) κ -Carroll	
$[J_a, J_b]$				Carroll contraction restores isotropy
$[J_a, P_b]$	anisotropy $\sim \frac{\Lambda}{\kappa}$	anisotropy $\sim \frac{\Lambda}{\kappa}$	isotropy	
$[J_a, K_b]$				
Non-commuting boosts $[J_a, P_0]$		0		
$[K_a, K_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa})$	$-\epsilon_{abc}J_c + O(\frac{\sqrt{\Lambda}}{\kappa})$	0	
No absolute time $[K_a, P_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa}, \frac{1}{\kappa})$	$\delta_{ab}P_0 + O(\frac{\Lambda}{\kappa}, \frac{1}{\kappa})$	$\delta_{ab}P_0 + O(\frac{\Lambda}{\kappa}, \frac{1}{\kappa})$	Absolute space
$[K_a, P_0]$	P_a	P_a	0	
$[P_a, P_b]$	$O(\frac{\sqrt{\Lambda}}{\kappa})$	$\Lambda\epsilon_{abc}J_c + O(\frac{\sqrt{\Lambda}}{\kappa})$	$\Lambda\epsilon_{abc}J_c$	
$[P_a, P_0]$	ΛK_a	ΛK_a	ΛK_a	

The " $\Lambda + \kappa$ " effects are milder in the Carroll models

Thanks!