NON-ABELIAN INFRARED DIVERGENCES ON THE CELESTIAL SPHERE

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Caveats

- A (somewhat) different perspective
- Strictly perturbative (but all-orders)
- Leading-power only
- No Mellin (but see González-Rojas)

Outline

- Infrared factorisation of scattering amplitudes
- Infrared factorisation on the celestial sphere
- A Lie-algebra conformal field theory
- Many open questions

INFRARED VISIONS



A. Sen, A.H. Mueller, J. Collins, G. Sterman, J. Botts, LM, S. Catani, L. Dixon, E. Gardi, T. Becher, M. Neubert, I. Feige, M. Schwartz, O. Erdogan, Y. Ma, ...

The factorised amplitude

Infrared divergences in fixed-angle multi-particle scattering amplitudes factorise

$$\mathcal{A}_n\left(\frac{p_i}{\mu},\alpha_s(\mu^2),\epsilon\right) = \mathcal{Z}_n\left(\frac{p_i}{\mu},\alpha_s(\mu^2),\epsilon\right)\mathcal{F}_n\left(\frac{p_i}{\mu},\alpha_s(\mu^2),\epsilon\right),$$

The infrared factor is a colour operator determined by a finite anomalous dimension matrix

$$\mathcal{Z}_n\left(\frac{p_i}{\mu},\alpha_s(\mu^2),\epsilon\right) = \mathcal{P}\exp\left[\frac{1}{2}\int_0^{\mu^2}\frac{d\lambda^2}{\lambda^2}\,\Gamma_n\left(\frac{p_i}{\lambda},\alpha_s(\lambda^2,\epsilon)\right)\right]\,,$$

All infrared poles arise from the scale integration, through the d-dimensional running coupling

$$\lambda \frac{\partial \alpha_s}{\partial \lambda} \equiv \beta(\alpha_s, \epsilon) = -2\epsilon \alpha_s - \frac{\alpha_s^2}{2\pi} \sum_{k=0}^{\infty} \left(\frac{\alpha_s}{\pi}\right)^k b_k \,.$$

For massless theories, the all-order structure of the anomalous dimension in known, up to corrections due to higher-order Casimir operators of the gauge algebra

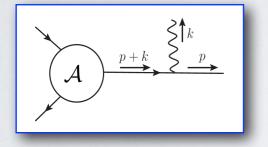
$$\Gamma_n\left(\frac{p_i}{\mu},\alpha_s(\mu^2)\right) = \Gamma_n^{\rm dip}\left(\frac{s_{ij}}{\mu^2},\alpha_s(\mu^2)\right) + \Delta_n\left(\rho_{ijkl},\alpha_s(\mu^2)\right), \qquad \rho_{ijkl} = \frac{p_i \cdot p_j \, p_k \cdot p_l}{p_i \cdot p_l \, p_j \cdot p_k} = \frac{s_{ij} s_{kl}}{s_{il} s_{jk}}.$$

Color operator notation

A powerful basis-independent notation uses colour operators `inserting' soft gluons

$$\mathcal{A}_{n+1}^{a \, b_1 \dots b_n} \bigg|_{\text{soft}} \propto \sum_{i=1}^n \left[\mathbf{T}_i^a \right]_{c_i}^{b_i} \mathcal{A}_n^{b_1 \dots c_i \dots b_n},$$

Soft gluon operators are generators of the algebra in the representation of the emitter



At leading power in k :

For different emitters :

$$g\mu^{\epsilon} \overline{u}_{s_i}(p_i) \gamma_{\alpha} \frac{\not p_i + \not k}{2p_i \cdot k} \left(T^c\right)_{c_i d_i} \widehat{\mathcal{A}}_{s_1 \dots s_n}^{c_1 \dots d_i \dots c_n} \left(\{p_j\}, k\right) \epsilon_{\lambda}^{* \alpha}(k) ,$$

$$g\mu^{\epsilon} \frac{\beta_{i} \cdot \epsilon_{\lambda}^{*}(k)}{\beta_{i} \cdot k} \left(T^{c}\right)_{c_{i}d_{i}} \left(\mathcal{A}_{n}\right)_{s_{1}...s_{n}}^{c_{1}...d_{i}...c_{n}} \left(\left\{p_{j}\right\}\right) \equiv g\mu^{\epsilon} \frac{\beta_{i} \cdot \epsilon_{\lambda}^{*}(k)}{\beta_{i} \cdot k} \mathbf{T}_{i} \mathcal{A}_{n}\left(\left\{p_{j}\right\}\right).$$

$$\mathbf{T}_{i}\Big|_{q, \text{out}} \to T^{a}_{cd}, \quad \mathbf{T}_{i}\Big|_{\bar{q}, \text{out}} \to -T^{a}_{dc}, \quad \mathbf{T}_{i}\Big|_{g, \text{out}} \to -\mathrm{i}f^{a}_{cd},$$

Colour operators obey identities inherited by the algebra and dictated by gauge invariance

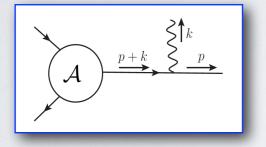
$$\left[\mathbf{T}_{i}^{a},\mathbf{T}_{i}^{b}\right] = \mathrm{i}f_{\ c}^{ab}\mathbf{T}_{i}^{c}, \qquad \mathbf{T}_{i}\cdot\mathbf{T}_{i} \equiv \mathbf{T}_{i}^{a}\mathbf{T}_{i}^{b}\delta_{ab} = C_{i}^{(2)}, \qquad \sum_{i=1}^{n}\mathbf{T}_{i} = 0,$$

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$$g\mu^{\epsilon} \frac{\beta_{i} \cdot \epsilon_{\lambda}^{*}(k)}{\beta_{i} \cdot k} \left(T^{c}\right)_{c_{i}d_{i}} \left(\mathcal{A}_{n}\right)_{s_{1}...s_{n}}^{c_{1}...d_{i}...c_{n}} \left(\left\{p_{j}\right\}\right) \equiv g\mu^{\epsilon} \frac{\beta_{i} \cdot \epsilon_{\lambda}^{*}(k)}{\beta_{i} \cdot k} \mathbf{T}_{i} \mathcal{A}_{n}\left(\left\{p_{j}\right\}\right).$$

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$$\begin{bmatrix} \mathbf{T}_i^a, \mathbf{T}_i^b \end{bmatrix} = \mathrm{i} f_{\ c}^{ab} \mathbf{T}_i^c, \qquad \mathbf{T}_i \cdot \mathbf{T}_i \equiv \mathbf{T}_i^a \mathbf{T}_i^b \,\delta_{ab} = C_i^{(2)}, \qquad \sum_{i=1}^n \mathbf{T}_i = 0, \qquad \text{the the set}$$

when acting on the amplitude

The dipole formula

Let's take a closer look at the structure of the infrared anomalous dimension matrix.

The dipole term :

$$\Gamma_n^{\rm dip}\left(\frac{s_{ij}}{\mu^2}, \alpha_s(\mu^2)\right) = \frac{1}{2}\,\widehat{\gamma}_K\left(\alpha_s(\mu^2)\right)\sum_{i=1}^n\sum_{j=i+1}^n\log\left(\frac{s_{ij}\,\mathrm{e}^{\mathrm{i}\pi\lambda_{ij}}}{\mu^2}\right)\mathbf{T}_i\cdot\mathbf{T}_j + \sum_{i=1}^n\gamma_i\left(\alpha_s(\mu^2)\right)\,,$$

The cusp anomalous dimension in the `Casimir scaling' limit:

$$\gamma_{K,r}(\alpha_s) = C_r^{(2)} \,\widehat{\gamma}_K(\alpha_s),$$

O.Almelid, C. Duhr, E. Gardi; J. Henn, B. Mistlberger.

$$F_{ijkl}(\{\rho\}) f_{abe} f_{cd}^{\ e} \mathbf{T}_i^a \mathbf{T}_j^a \mathbf{T}_k^c \mathbf{T}_l^d,$$

- The colour dipole is the natural structure arising at one loop from gluon exchange.
- The fact that it survives at two loops is a non-trivial consequence of symmetries.
- Field anomalous dimensions in color-uncorrelated terms govern collinear singularities.
- $\stackrel{\scriptstyle \bigvee}{\scriptstyle =}$ Unitarity phases contain crucial analytic information. For final-state pairs: $\lambda_{ij} = 1$.
- Final terms and the second sec
- Final The structure emerges from the constraints of scale invariance in the soft limit.

INFRARED VISIONS ON THE CELESTIAL SPHERE



On dipole correlations

Let us begin by disentangling collinear poles (which are colour-singlets) from soft poles (which are colour-correlated). We replace the running scale λ with the fixed scale μ in the logarithmic term, and perform the colour sum using colour conservation.

$$\Gamma_{n}^{\text{dipole}}\left(\frac{s_{ij}}{\lambda^{2}},\alpha_{s}(\lambda,\epsilon)\right) = \frac{1}{2}\widehat{\gamma}_{K}\left(\alpha_{s}(\lambda,\epsilon)\right)\sum_{i=1}^{n}\sum_{j=i+1}^{n}\ln\left(\frac{-s_{ij}+\mathrm{i}\eta}{\mu^{2}}\right)\mathbf{T}_{i}\cdot\mathbf{T}_{j}$$
$$-\sum_{i=1}^{n}\gamma_{i}\left(\alpha_{s}(\lambda,\epsilon)\right) - \frac{1}{4}\widehat{\gamma}_{K}\left(\alpha_{s}(\lambda,\epsilon)\right)\ln\left(\frac{\mu^{2}}{\lambda^{2}}\right)\sum_{i=1}^{n}C_{i}^{(2)}$$
$$\equiv\Gamma_{n}^{\text{corr.}}\left(\frac{s_{ij}}{\mu^{2}},\alpha_{s}(\lambda,\epsilon)\right) + \Gamma_{n}^{\text{singl.}}\left(\frac{\mu^{2}}{\lambda^{2}},\alpha_{s}(\lambda,\epsilon)\right),$$

At one loop, integrating the colour-correlated term yields single soft poles, while the singlet term yields single collinear and double soft-collinear poles

$$\alpha_s(\lambda,\epsilon) = \alpha_s(\mu) \left(\frac{\lambda^2}{\mu^2}\right)^{-\epsilon}, \qquad \int_0^{\mu^2} \frac{d\lambda^2}{\lambda^2} \alpha_s(\lambda,\epsilon) = -\frac{1}{\epsilon} \alpha_s(\mu), \quad \int_0^{\mu^2} \frac{d\lambda^2}{\lambda^2} \ln\left(\frac{\lambda^2}{\mu^2}\right) \alpha_s(\lambda,\epsilon) = -\frac{1}{\epsilon^2} \alpha_s(\mu), \quad (\epsilon < 0).$$

At h loops, multiple poles (up to order h+1) are generated by the β function. For conformal gauge theories the logarithm of the infrared factor has only single and double poles.

Celestial dipoles

Crucially, we now parametrise the light-cone momenta in celestial coordinates

$$p_i^{\mu} = \omega_i \left\{ 1 + z_i \bar{z}_i, \, z_i + \bar{z}_i, \, -i(z_i - \bar{z}_i), \, 1 - z_i \bar{z}_i \right\},$$

where the energy ω_i and the sphere coordinates z_i have simple transformation properties under the Lorentz group acting as SL(2, C):

$$\omega' = |cz+d|^2 \omega, \qquad z' = \frac{az+b}{cz+d},$$

Mandelstam invariants are distances on the sphere

$$s_{ij} = 2p_i \cdot p_j = 4\omega_i \omega_j |z_i - z_j|^2,$$

which unpacks the logarithms

$$\log\left(-s_{ij}+\mathrm{i}\eta\right) = \log\left(\left|z_i-z_j\right|^2\right) + \log\omega_i + \log\omega_j + 2\log 2 + \mathrm{i}\pi\,,$$

Energies give new singlet terms

$$\log(-s_{ij} + m) = \log(|z_i - z_j|) + \log \omega_i + \log \omega_j + 2\log 2 + m$$

$$\Gamma_n^{\text{dipole}}\left(\frac{s_{ij}}{\lambda^2}, \alpha_s(\lambda, \epsilon)\right) \equiv \widehat{\Gamma}_n^{\text{corr.}}\left(z_{ij}, \alpha_s(\lambda, \epsilon)\right) + \widehat{\Gamma}_n^{\text{singl.}}\left(\frac{\omega_i}{\lambda}, \alpha_s(\lambda, \epsilon)\right),$$

which take the form

$$\widehat{\Gamma}_{n}^{\text{singl.}}\left(\frac{\omega_{i}}{\lambda},\alpha_{s}(\lambda,\epsilon)\right) = -\sum_{i=1}^{n}\gamma_{i}\left(\alpha_{s}(\lambda,\epsilon)\right) - \frac{1}{4}\widehat{\gamma}_{K}\left(\alpha_{s}(\lambda,\epsilon)\right)\sum_{i=1}^{n}\ln\left(\frac{-4\omega_{i}^{2}+\mathrm{i}\eta}{\lambda^{2}}\right)C_{i}^{(2)},$$

Celestial dipoles

The colour-correlated term, responsible for all soft poles, is remarkably simple

$$\widehat{\Gamma}_{n}^{\text{corr.}}\left(z_{ij}, \alpha_{s}(\lambda, \epsilon)\right) = \frac{1}{2} \widehat{\gamma}_{K}\left(\alpha_{s}(\lambda, \epsilon)\right) \sum_{i=1}^{n} \sum_{j=i+1}^{n} \ln\left(|z_{ij}|^{2}\right) \mathbf{T}_{i} \cdot \mathbf{T}_{j}.$$

Scale and coupling dependence are completely factored from colour and kinematics, and equal for all dipoles. The scale integral can this be performed in full generality, yielding

$$\mathcal{Z}_{n}^{\text{corr.}}\left(z_{ij},\alpha_{s}(\mu),\epsilon\right) \equiv \exp\left[\int_{0}^{\mu} \frac{d\lambda}{\lambda} \widehat{\Gamma}_{n}^{\text{corr.}}\left(z_{ij},\alpha_{s}(\lambda,\epsilon)\right)\right]$$
$$= \exp\left[-K\left(\alpha_{s}(\mu),\epsilon\right)\sum_{i=1}^{n}\sum_{j=i+1}^{n}\ln\left(|z_{ij}|^{2}\right)\mathbf{T}_{i}\cdot\mathbf{T}_{j}\right],$$

The scale factor K is well-known in QCD from form-factor calculations, and gives the perturbative Regge trajectory in the high-energy limit of four-point amplitudes. It is

G. Korchemsky, I.A. Korchemskaya; V. Del Duca, C. Duhr, E. Gardi, LM, C. White; G. Falcioni, L. Vernazza, ...

$$K(\alpha_s(\mu),\epsilon) = -\frac{1}{2} \int_0^\mu \frac{d\lambda}{\lambda} \,\widehat{\gamma}_K(\alpha_s(\lambda,\epsilon)) \,.$$

The function K can be computed order by order in terms of the cusp and the β function

COLOUR ON THE CELESTIAL SPHERE



Hints of a celestial theory

The colour-correlated term in the anomalous dimension matrix is strongly reminiscent of conformal field theory results. One needs only go so far as Joe Polchinski's book to find

2.3 The expectation value of a product of exponential operators on the plane is

$$\left\langle \prod_{i=1}^{n} : e^{ik_i \cdot X(z_i,\bar{z}_i)} : \right\rangle = iC^X (2\pi)^D \delta^D \left(\sum_{i=1}^{n} k_i \right) \prod_{\substack{i,j=1\\i < j}}^{n} |z_{ij}|^{\alpha' k_i \cdot k_j} ,$$

with C^X a constant. This can be obtained as a limit of the expectation value (6.2.17) on the sphere, which we will obtain by several methods in chapter 6.

A correlator of vertex operators in a free-boson theory (such as the bosonic string) has the correct form, up to the substitution of momenta with colour matrices.

This was noticed by N. Kalyanapuram in 2011.11412, for the simple case of QED. He writes

Also Nande Pate and Strominger, 1705.00608

$$\ln\left(\mathcal{A}_{n,s=1}^{soft}|_{vir}\right) = -\frac{1}{8\pi^2\epsilon} \sum_{i\neq j} e_i e_j \ln|z_i - z_j|^2.$$

The result is formally reproduced by introducing vertex operators with electric charges

$$V_j(z_j, \overline{z}_j) :=: e^{ie_j \varphi(z_i, \overline{z}_j)}: \qquad \longrightarrow \qquad \langle V_1(z_1, \overline{z}_1) \cdots V_n(z_n, \overline{z}_n) \rangle = A_n^{soft}|_{vir, s=1}.$$

Lie-algebra-valued free bosons

It is natural to mimic the bosonic string, considering free bosons spanning the gauge algebra.

$$S(\phi) = \frac{1}{2\pi} \int d^2 z \, \partial_z \phi^a(z, \bar{z}) \, \partial_{\bar{z}} \phi_a(z, \bar{z}) \,,$$

The free bosons could be organised in a matrix field : gauge generators at different points must then be taken to commute

 $\Phi_r(z,\bar{z}) \equiv \phi_a(z,\bar{z}) T^a_{r,z} \,,$

The well-known results for free bosons in d=2 can be directly transcribed.

The equations of motions are:

$$\partial_z \, \partial_{\bar{z}} \, \phi^a(z, \bar{z}) \, = \, 0 \, ,$$

implying that the derivatives of the fields are (anti)holomorphic

A normal-ordered product can be defined, obeying the classical equation of motion

$$:\phi^a(z,\bar{z})\,\phi^b(w,\overline{w}):=\,\phi^a(z,\bar{z})\,\phi^b(w,\overline{w})+\frac{1}{2}\,\delta^{ab}\log|z-w|^2\,\,,$$

There is a traceless conserved energy-momentum tensor, and a conserved Noether current

$$T(z) = - : \partial_z \phi^a(z, \bar{z}) \, \partial_z \phi_a(z, \bar{z}):,$$

$$j^a(z) = \partial_z \phi^a(z, \bar{z}),$$

Matrix vertex operators

Guided by the QED example, we can tentatively define a matrix-valued vertex operator

$$V(z,\bar{z}) \equiv : e^{i\kappa \mathbf{T}_{z} \cdot \phi(z,\bar{z})} := : e^{i\kappa \Phi(z,\bar{z})} :,$$

A `single-copy' of the string vertex operator!

In colour space, this is a matrix in the representation of T_z , defined on the boundary sphere and acting on the bulk colour degrees of freedom. But is it a conformal primary field?

For conventional vertex operators (as for example for bosonic strings)

The same calculation yields

$$V_{\text{c.s.}}(z,\bar{z}) \equiv :e^{ik^{\mu}X_{\mu}(z,\bar{z})}: \longrightarrow h = \frac{1}{4}k^{\mu}k^{\nu}\eta_{\mu\nu} = \frac{k^2}{4},$$

$$V(z,\bar{z}) \equiv :e^{i\kappa \mathbf{T}_{z} \cdot \phi(z,\bar{z})} : \longrightarrow \qquad h = \frac{\kappa^{2}}{4} \mathbf{T}_{z} \cdot \mathbf{T}_{z} = \frac{\kappa^{2}}{4} C_{r}^{(2)},$$

Crucially, this is a positive real number and not a matrix. For consistency, two-point functions must evaluate to a power of the distance given by the conformal weight $\Delta = h + h$. Indeed

 $\langle V(z_1, \bar{z}_1) V(z_2, \bar{z}_2) \rangle \sim |z_{12}|^{-2\Delta},$ by colour conservation $\mathbf{T}_1 + \mathbf{T}_2 = 0$ Note analogies with other constructions. Vertex operator construction of Kac-Moody algebras: Reggeon fields for high-energy scattering: (Caron-Huot 2013) $U(z) = e^{ig_s T^a W^a(z)}.$ closely related

A conformal correlator

Our construction from the beginning targeted the n-point correlator

$$C_n(\{z_i\},\kappa) \equiv \left\langle \prod_{i=1}^n V(z_i,\bar{z}_i) \right\rangle.$$

The calculation is a textbook exercise: it can be done with oscillators, after expanding the free fields in modes on the sphere, or computing the path integral (Polchinski). The result is

$$\mathcal{C}_n\left(\{z_i\},\kappa\right) = C(N_c) \exp\left[\frac{\kappa^2}{2} \sum_{i=1}^n \sum_{j=i+1}^n \ln\left(|z_{ij}|^2\right) \mathbf{T}_i \cdot \mathbf{T}_j\right],$$

reproducing the structure of the gauge theory infrared operator. Note that

$$\sum_{i=1}^{n} \mathbf{T}_{i} = 0,$$

- Final Stress Field on the second seco
- Field normalisation K maps to the integral K, carrying scale and regulator dependence.
- In a path integral evaluation on a curved surface (say, a finite sphere with radius R) the correlator acquires a scale-dependent `Weyl' factor, which in this setting maps to an (undetermined) colour-singlet collinear contribution.

$$\mathcal{W}_n\Big(\{z_i\},\kappa\Big) = \exp\left[-\frac{1}{2}\sum_{i=1}^n C_i^{(2)}g(z_i,\bar{z}_i)\right],\,$$

A tree-level soft theorem

 $\mathbf{T}_i = 0,$

Real emission of a soft massless gauge boson from a fixed angle hard amplitude factorises in any non-abelian theory in the form

$$\langle c | \otimes \langle \lambda | \mathcal{A}_{g, f_1 \dots f_n}(k, p_1, \dots, p_n) \rangle_{\text{soft}} = \epsilon_{\lambda}(k) \cdot J^c(k) | \mathcal{A}_{f_1 \dots f_n}(p_1, \dots, p_n) \rangle$$

The tree-level soft-gluon current has the classic eikonal form and is gauge-invariant

$$\mathbf{J}^{\mu}(k) = g \sum_{i=1}^{n} \mathbf{T}_{i} \frac{\beta_{i}^{\mu}}{\beta_{i} \cdot k}, \qquad k \cdot \mathbf{J}^{\mu}(k) = g$$

The tree-level soft theorem is reproduced by the Ward identity for the Noether current associated with invariance under field translations in the Lie algebra. Using the conformal operator product expansion one finds

A. Strominger, T. He, P. Mitra, A. Nande, M. Pate, W. Fan, A. Fotopoulos, T.R. Taylor, ...

$$\left\langle \partial_z \phi^a(z, \bar{z}) \prod_{i=1}^n V(z_i, \bar{z}_i) \right\rangle \simeq -\frac{\mathrm{i}}{2} \sum_{i=1}^n \frac{\mathbf{T}_i^a}{z - z_i} \mathcal{C}_n\left(\{z_i\}, \kappa\right).$$

where the poles as $z \rightarrow z_i$ are collinear poles, since the celestial theory is energy-independent.

Collinear limits

The operator product expansion governs the collinear limit on the sphere. One can transcribe the textbook result substituting colour operators for momenta.

$$: e^{i\kappa \mathbf{T}_1 \cdot \phi(z_1, \bar{z}_1)} :: e^{i\kappa \mathbf{T}_2 \cdot \phi(z_2, \bar{z}_2)} : \sim |z_{12}|^{\kappa^2 \mathbf{T}_1 \cdot \mathbf{T}_2} : e^{i\kappa (\mathbf{T}_1 + \mathbf{T}_2) \cdot \phi(z, \bar{z})} :,$$

Note that the exact collinear limit is outside the validity of the original factorisation. But it can be approached: on the gauge theory side, one defines a splitting anomalous dimension

$$\Gamma_{\mathrm{Sp.}}(p_1, p_2) \equiv \left. \Gamma_n(p_1, p_2, \dots, p_n) - \Gamma_{n-1}(p, p_3, \dots, p_n) \right|_{\mathbf{T}_p \to \mathbf{T}_1 + \mathbf{T}_2}.$$

(Becher-Neubert 2009)

The OPE encodes collinear factorisation: the n-point correlator reduces to (n-1)-points, with the `merged' point carrying the sum of the colours of (only!) the two collinear particles. The calculation of the splitting function is then the same as in the gauge theory, but requires reinstating the energy dependence, which is not encoded by the conformal correlator.

$$\Gamma_{\mathrm{Sp.}}(p_1, p_2) = \frac{1}{2} \widehat{\gamma}_K(\alpha_s) \left[\ln \left(\frac{-s_{12} + \mathrm{i}\eta}{\mu^2} \right) \mathbf{T}_1 \cdot \mathbf{T}_2 - \ln x \ \mathbf{T}_1 \cdot \left(\mathbf{T}_1 + \mathbf{T}_2 \right) - \ln(1-x) \ \mathbf{T}_2 \cdot \left(\mathbf{T}_1 + \mathbf{T}_2 \right) \right],$$

MANY QUESTIONS



Many Questions

 $\stackrel{\scriptstyle{\lor}}{=}$ The choice of the gauge coupling.

Our construction lends support to the idea the the cusp anomalous dimension should be taken as the definition of the strong coupling in the infrared. How far can one take this definition?

Scale and regulator dependence.

It is remarkable, and necessary, that infrared singularities be hidden in the matching condition between the gauge theory and the conformal theory. How can one make this correspondence more precise?

Beyond the free theory.

The celestial conformal theory certainly has corrections involving structure constants (as confirmed by the structure of Δ). The deformed theory is still conformal. What drives the deformation?

Constraints from vast field theory data.

Soft and collinear factorisation kernels are known to three loops, and in the massive case to two loops. In most cases their remarkable simplicity is only partly explained. How can we harness these data to constrain the celestial theory?

The exploration has just begun!

