

Conformal Blocks and Light-Ray Operators in Celestial CFT

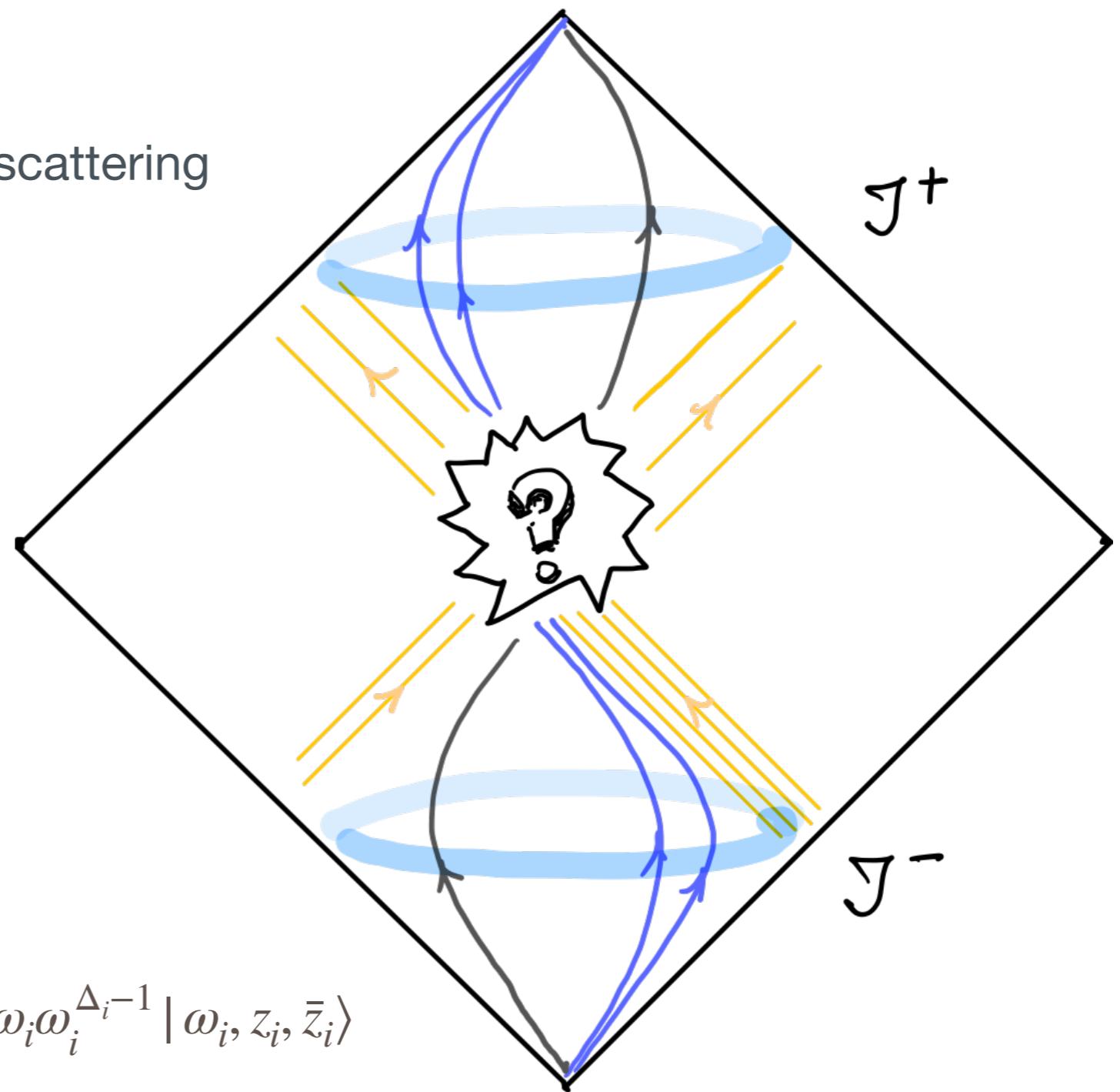
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Celestial Amplitudes and Flat Space Holography, Corfu, 2021

Motivation

- Celestial amplitudes: new toolkit for scattering in asymptotically flat spacetimes
- Study scattering in basis of boost eigenstates
- Map from momentum space to celestial sphere for massless external states:



$$|p_i\rangle \rightarrow |\omega_i, z_i, \bar{z}_i\rangle \rightarrow |\Delta_i, z_i, \bar{z}_i\rangle = \int_0^\infty d\omega_i \omega_i^{\Delta_i-1} |\omega_i, z_i, \bar{z}_i\rangle$$

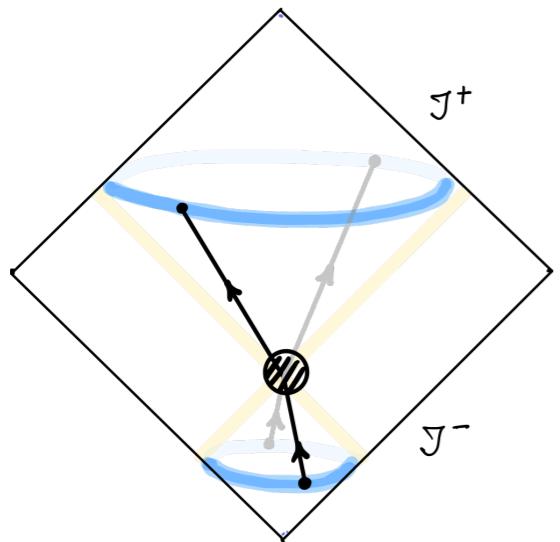
Celestial holography



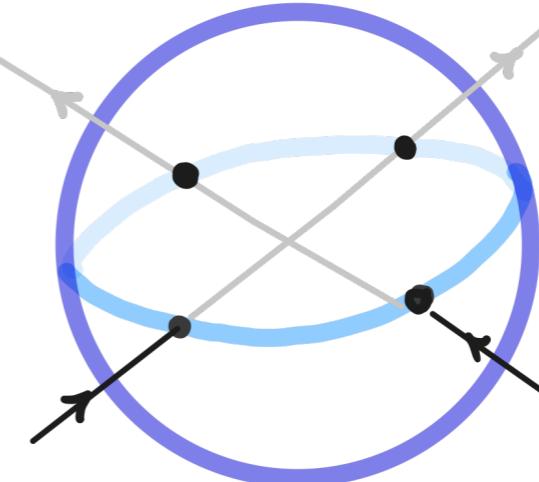
Outline

- Conformal blocks for scalar scattering in CCFT
 - massive scalar exchanges of positive integral dimension
 - spinning massive exchanges: light-ray transforms of massive scalars on the principal series
- Comments on light-transforms and shadows
- Open problems

Massless scalar 4-point scattering



Momentum space



Celestial sphere

$$\mathbb{A}(p_i) = \mathcal{M}(s, t) \delta^{(4)}\left(\sum_{i=1}^4 p_i\right)$$

$$s = -(p_1 + p_2)^2 \equiv \omega^2$$

$$t = -(p_1 + p_3)^2 \equiv -z\omega^2$$



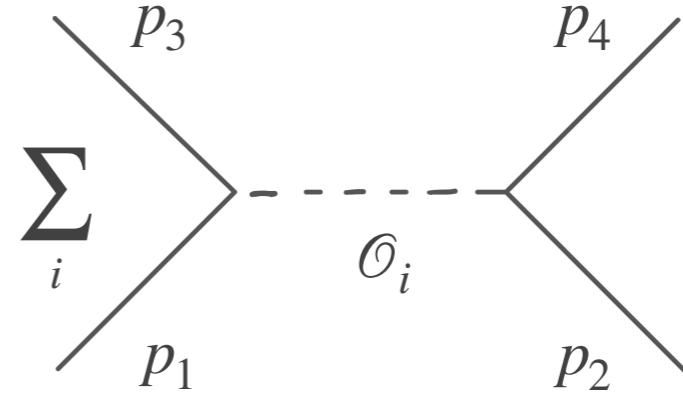
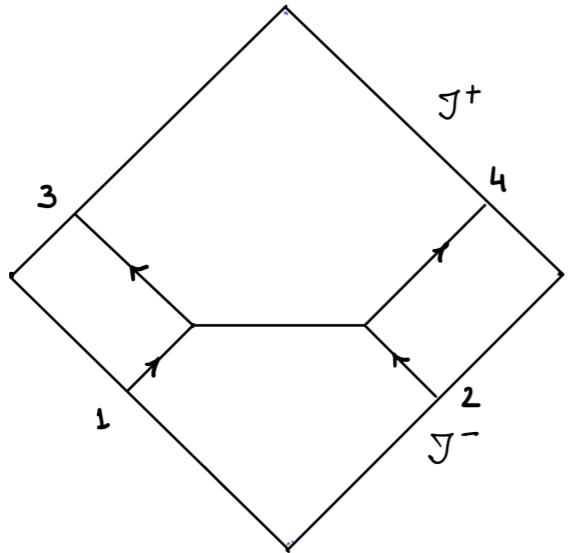
$$\widetilde{\mathcal{A}}(z_i, \bar{z}_i; \beta) = K(z_i, \bar{z}_i) X(z, \beta) \int_0^\infty d\omega \omega^{\beta-1} \mathcal{M}(\omega^2, -z\omega^2)$$

$$\beta = \sum_{i=1}^4 \Delta_i - 4, \quad z = -\frac{t}{s} = \frac{z_{13}z_{24}}{z_{12}z_{34}} \in [0,1]$$

$$K(z_i, \bar{z}_i) = \prod_{i < j} z_{ij}^{\frac{h}{3} - h_i - h_j} \bar{z}_{ij}^{\frac{\bar{h}}{3} - \bar{h}_i - \bar{h}_j}$$

$$X(z, \beta) = \delta(z - \bar{z}) |z(1-z)|^{\frac{1}{6}(\beta+4)}$$

Tree-level example



$$\begin{aligned}
 \mathcal{M}(s, t) = -g^2 \frac{1}{t - m^2} \quad &\rightarrow \quad \widetilde{\mathcal{A}}(z_i, \bar{z}_i; \beta) = g^2 K(z_i, \bar{z}_i) X(z, \beta) \int_0^\infty d\omega \omega^{\beta-1} \frac{1}{z\omega^2 + m^2} \\
 &= I_{13-24}(z_i, \bar{z}_i) \underbrace{N_{gm}(\beta) |z|^2 |1-z|^{h_{13}-h_{24}} \delta(z-\bar{z})}_{f_t(z, \bar{z})}
 \end{aligned}$$

Goal: Decompose $f_t(z, \bar{z})$ into conformal blocks.

Roadmap: Treat z, \bar{z} as real independent variables and expand $f_t(z, \bar{z})$ on a complete set of orthogonal solutions to the two particle conformal Casimir over the Lorentzian square, $z, \bar{z} \in [0, 1]$.

Conformal partial waves

- 1-3 conformal Casimir equation $(\mathcal{D}_z + \mathcal{D}_{\bar{z}}) \Psi_{h,\bar{h}}(z, \bar{z}) = [h(h-1) + \bar{h}(\bar{h}-1)] \Psi_{h,\bar{h}}(z, \bar{z})$
 $\mathcal{D}_z = z^2(1-z)\frac{\partial^2}{\partial z^2} - (1-h_{13}+h_{24})z^2\frac{\partial}{\partial z} + h_{13}h_{24}z$
admits solutions of the form $\Psi_{h,\bar{h}}(z, \bar{z}) = \Psi_h(z)\Psi_{\bar{h}}(\bar{z})$.

- Conformal partial waves $\Psi_h(z) = \frac{1}{2} (Q(h)k_h(z) + Q(1-h)k_{1-h}(z))$, $h = \frac{1}{2} + \alpha$, $\alpha \in i\mathbb{R}$,
- where $k_{h,\bar{h}}(z, \bar{z}) = k_h(z)k_{\bar{h}}(\bar{z})$, $k_h(z) = z^h {}_2F_1(h-h_{12}, h+h_{34}; 2h; z)$ are $SL(2, \mathbb{C})$ blocks.

- $Q(h)$ fixed by **orthogonality** $\langle \Psi_h, \Psi_{h'} \rangle = \int_0^1 \frac{dz}{\mu(z)} \Psi_h(z) \Psi_{h'}(z) = \frac{N(h)}{2} [\delta(h+h') + \delta(h-h')]$
- **Completeness** $\int_{\frac{1}{2}-i\infty}^{\frac{1}{2}+i\infty} \frac{dh}{N(h)} \Psi_h(z) \Psi_h(z') = \mu(z) \delta(z-z')$, $\mu(z) = z^2(1-z)^{h_{12}-h_{34}}$

Conformal blocks for celestial scattering

$$f_t(z, \bar{z}) = N_{gm}(\beta) |z|^2 |1-z|^{h_{13}-h_{24}} \delta(z - \bar{z}) = \int_{\frac{1}{2}-i\infty}^{\frac{1}{2}+i\infty} \frac{dh}{N(h)} \int_{\frac{1}{2}-i\infty}^{\frac{1}{2}+i\infty} \frac{d\bar{h}}{N(\bar{h})} g(h, \bar{h}) \Psi_{h, \bar{h}}(z, \bar{z})$$

- Upon deforming the contour

$$f_t(z, \bar{z}) = n(\beta) \left[\sum_{n=1}^{\infty} D^{nn} C_{13n} C_{24n} \cos \pi \left(\frac{n}{2} + h_{13} \right) \cos \pi \left(\frac{n}{2} + h_{24} \right) k_{\frac{1+n}{2}}(z) k_{\frac{1+n}{2}}(\bar{z}) \right.$$

$$\left. + \frac{1}{2} \int_{-i\infty}^{i\infty} d\alpha C_{13\alpha}^L C_{24\alpha}^L D^{L, \alpha\alpha} k_{\frac{1}{2}+\alpha}(z) k_{\frac{1}{2}-\alpha}(\bar{z}) \right].$$

- Massive scalar exchanges $C_{ijn} = \frac{g}{m^4} \left(\frac{m}{2} \right)^{2h_i+2h_j} B \left(\frac{1+n}{2} + h_{ij}, \frac{1+n}{2} - h_{ij} \right)$, $D^{nn} = \frac{nm^2}{2\pi}$.

Conformal blocks for celestial scattering

$$f_t(z, \bar{z}) = N_{gm}(\beta) |z|^2 |1-z|^{h_{13}-h_{24}} \delta(z - \bar{z}) = \int_{\frac{1}{2}-i\infty}^{\frac{1}{2}+i\infty} \frac{dh}{N(h)} \int_{\frac{1}{2}-i\infty}^{\frac{1}{2}+i\infty} \frac{d\bar{h}}{N(\bar{h})} g(h, \bar{h}) \Psi_{h, \bar{h}}(z, \bar{z})$$

$$\langle f_t(z, \bar{z}), \Psi_{h, \bar{h}}(z, \bar{z}) \rangle$$

- Upon deforming the contour

$$f_t(z, \bar{z}) = n(\beta) \left[\sum_{n=1}^{\infty} D^{nn} C_{13n} C_{24n} \cos \pi \left(\frac{n}{2} + h_{13} \right) \cos \pi \left(\frac{n}{2} + h_{24} \right) k_{\frac{1+n}{2}}(z) k_{\frac{1+n}{2}}(\bar{z}) \right.$$

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Conformal blocks for celestial scattering

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$\langle f_t(z, \bar{z}), \Psi_{h, \bar{h}}(z, \bar{z}) \rangle$

- Upon deforming the contour

$$f_t(z, \bar{z}) = n(\beta) \left[\sum_{n=1}^{\infty} D^{nn} C_{13n} C_{24n} \cos \pi \left(\frac{n}{2} + h_{13} \right) \cos \pi \left(\frac{n}{2} + h_{24} \right) k_{\frac{1+n}{2}}(z) k_{\frac{1+n}{2}}(\bar{z}) \right]$$

$$+ \frac{1}{2} \int_{-i\infty}^{i\infty} da C_{13\alpha}^L C_{24\alpha}^L D^{L, \alpha\alpha} k_{\frac{1}{2}+\alpha}(z) k_{\frac{1}{2}-\alpha}(\bar{z}) \right].$$

- Massive scalar exchanges $C_{ijn} = \frac{g}{m^4} \left(\frac{m}{2} \right)^{2h_i+2h_j} B \left(\frac{1+n}{2} + h_{ij}, \frac{1+n}{2} - h_{ij} \right)$, $D^{nn} = \frac{nm^2}{2\pi}$.
- C_{ijn} matches 3-point of 2 massless and 1 massive scalars
- D^{nn} massive 2-point

Conformal blocks for celestial scattering

$$f_t(z, \bar{z}) = N_{gm}(\beta) |z|^2 |1-z|^{h_{13}-h_{24}} \delta(z - \bar{z}) = \int_{\frac{1}{2}-i\infty}^{\frac{1}{2}+i\infty} \frac{dh}{N(h)} \int_{\frac{1}{2}-i\infty}^{\frac{1}{2}+i\infty} \frac{d\bar{h}}{N(\bar{h})} g(h, \bar{h}) \Psi_{h, \bar{h}}(z, \bar{z})$$

$\langle f_t(z, \bar{z}), \Psi_{h, \bar{h}}(z, \bar{z}) \rangle$

- Upon deforming the contour

$$f_t(z, \bar{z}) = n(\beta) \left[\sum_{n=1}^{\infty} D^{nn} C_{13n} C_{24n} \cos \pi \left(\frac{n}{2} + h_{13} \right) \cos \pi \left(\frac{n}{2} + h_{24} \right) k_{\frac{1+n}{2}}(z) k_{\frac{1+n}{2}}(\bar{z}) \right.$$

$$\left. + \frac{1}{2} \int_{-i\infty}^{i\infty} da C_{13\alpha}^L C_{24\alpha}^L D^{L,\alpha\alpha} k_{\frac{1}{2}+\alpha}(z) k_{\frac{1}{2}-\alpha}(\bar{z}) \right].$$

- Light-ray exchanges

$$C_{ij\alpha}^L = -\pi i \frac{g}{m^4} \left(\frac{m}{2} \right)^{2h_i+2h_j} \frac{1}{\alpha},$$

matches 3-point of 2 massless and 1 light-transformed massive scalar

$$D^{L,\alpha\alpha} = -\frac{m^2 \alpha^2}{\pi^2 i}.$$

light-transformed massive 2-point

Light-transformed conformal primary wavefunctions

- Light transforms relate the bulk conformal primary wavefunctions $\phi_\Delta(z, \bar{z}) = \frac{i^\Delta \Gamma(\Delta)}{(-q \cdot X)^\Delta}$ of $\Delta = 1 + i\lambda, J = 0$ to conformal primary wavefunctions

$$\Phi(z, \bar{z}; X) \propto \int_{-\infty}^{\infty} \frac{dz'}{(z' - z)^{2-\Delta}} \frac{1}{(-q(z', \bar{z}) \cdot X)^\Delta} \propto \frac{(\partial_z q \cdot X)^{1-\Delta}}{-q \cdot X}, \quad \Delta_L = 1, J_L = -i\lambda.$$

- Two, three point functions of massless particles are singular in conformal primary basis, eg. tree-level gluons

$$\widetilde{\mathcal{A}}_3 \equiv \langle \mathcal{O}_1^-(z_1, \bar{z}_1) \mathcal{O}_2^-(z_2, \bar{z}_2) \mathcal{O}_3^+(z_3, \bar{z}_3) \rangle \propto \delta(\lambda_1 + \lambda_2 + \lambda_3) z_{12}^{1-i(\lambda_1+\lambda_2)} z_{23}^{i\lambda_1-1} z_{13}^{i\lambda_2-1} \delta(\bar{z}_{13}) \delta(\bar{z}_{23}).$$

[Pasterski, Shao, Strominger '18]

- Light-transform leads to finite 3-point, eg. ambidextrous light transform of \mathcal{O}^- and \mathcal{O}^+ .

[Sharma '21]

New basis for celestial amplitudes?

- Shadows also turn singular 3-point functions into non-singular ones (prescription less clear).
- Evidence that a “good” basis for CCFT will involve light-transforms and/or shadows.
- Study symmetry generators and associated constraints in the light/shadow bases.
- Massless momenta $P = \hat{q}e^{\partial_\Delta}$ transform to
$$\widetilde{P} = \frac{(\Delta - 1)^2}{2} \left(\partial_z \partial_{\bar{z}} \hat{q} + \frac{1}{\Delta - 1} \partial_z \hat{q} \partial_{\bar{z}} + \frac{1}{\Delta - 1} \partial_{\bar{z}} \hat{q} \partial_z + \frac{\hat{q} \partial_z \partial_{\bar{z}}}{(\Delta - 1)^2} \right) e^{-\partial_\Delta} \quad \text{upon shadow.}$$
- Can be used to derive constraints on celestial amplitudes involving shadows!

[work in progress w. Laurent Freidel]

Open questions

- Principle to dictate choice of basis?
- Transform amplitudes to this basis, figure out how symmetries act: Poincare, soft symmetries, etc.
- 4-point functions, conformal blocks in the new basis \Rightarrow simplifications?
[Fan, Fotopoulos, Stieberger, Taylor, Zhu '21]
- Use constraints to implement a bootstrap program \Rightarrow non-perturbative aspects of gravity in AFS?