

The Gravity Dual of the Berkooz-Douglas matrix model

V. Filev

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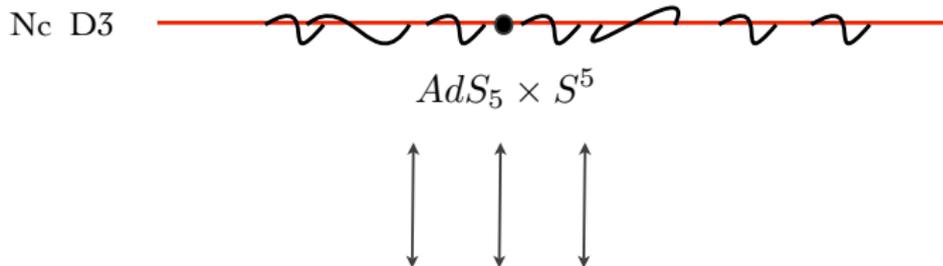
related to work with Y. Asano, D. O'Connor and S. Kovacik

Corfu September 2019

- 1 AdS/CFT correspondence
 - Pure Super Yang-Mills
 - Adding flavours
- 2 D0/D4 system
 - Lower dimensional correspondence
 - Berkooz-Douglas matrix model
- 3 Back-reacted D0/D4 background
 - Uplift to 11D supergravity
 - BPS equations
 - Solution

AdS/CFT correspondence

Type IIB String Theory on



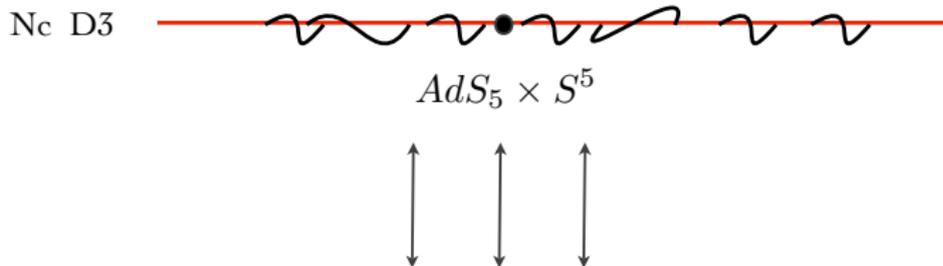
$\mathcal{N} = 4$ $SU(N_c)$ SUSY YM

- Gubser-Klebanov-Polyakov-Witten formula:

$$\langle e^{\int d^d x \phi_0(x) \mathcal{O}(x)} \rangle_{\text{CFT}} = \mathcal{Z}_{\text{string}}[\phi_0(x)]$$

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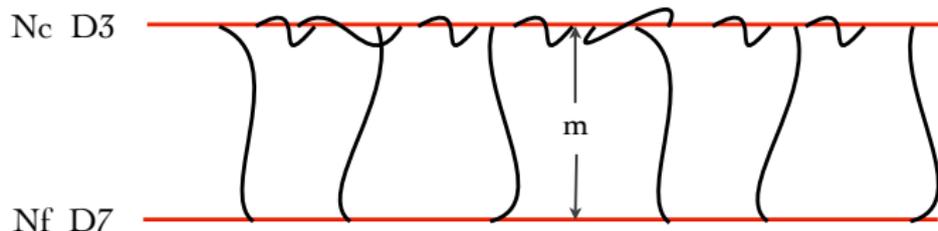
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Adding flavours D3/D7 Karch & Katz

Generalizing the correspondence

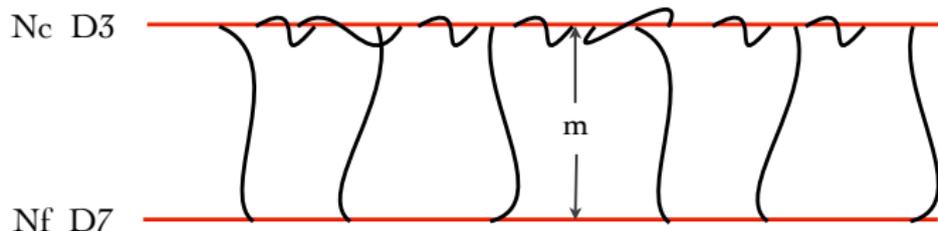


	0	1	2	3	4	5	6	7	8	9
D3	-	-	-	-	·	·	·	·	·	·
D7	-	-	-	-	-	-	-	-	·	·

- Adding N_f massive $\mathcal{N} = 2$ Hypermultiplets:

$$m_q \int d^2\theta \tilde{Q} Q \rightarrow \text{SYM} \quad \text{with} \quad m_q = m/2\pi\alpha'$$

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String spectrum

3-3 strings	pure $\mathcal{N}=4$ SYM adjoint of $SU(N_c)$
3-7 strings	Q_i fundamental chiral field
7-3 strings	\tilde{Q}^i anti-fundamental chiral field
7-7 strings	gauge field on the D7 brane frozen by infinite volume

- The probe is described by a Dirac-Born-Infeld action
$$S \propto \int d^7\xi e^{-\Phi} \sqrt{||G_{ab} - 2\pi\alpha' \mathcal{F}_{ab}||}$$
- The profile of the D-brane encodes the fundamental condensate of theory. The semi-classical fluctuations correspond to meson-like excitations.
- The D-brane gauge field can describe: external electromagnetic field, chemical potential, electric current etc.
- Numerous applications: thermal and quantum phase transitions, chiral symmetry breaking, magnetic catalysis etc.

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Lower dimensional correspondence

- The **D3/D7** system is T-dual to the **D0/D4** system, share many common properties (meson melting transition, meson spectra)
- The dual theory of the **D0/D4** set-up is a flavoured version of the **BFSS** matrix model - the Berkooz-Douglas (**BD**) matrix model.
- The **BD** matrix model is **1D** quantum mechanics and is super renormalisable, avoiding the fine tuning problem.
- Aspects of the **BD** model in the probe approximation are studied in [arXiv:1512.02536, 1605.05597, 1612.09281] by Y. Asano, V. Filev, S. Kovacik and D. O'Connor.

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Berkooz-Douglas matrix model

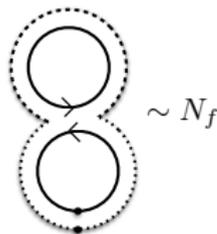
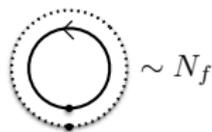
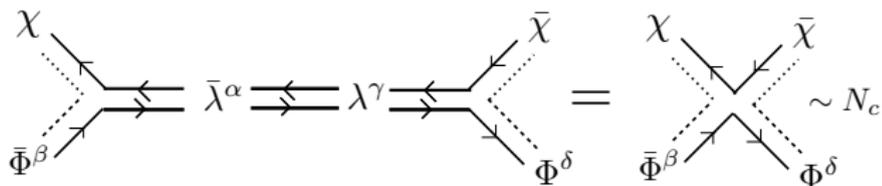
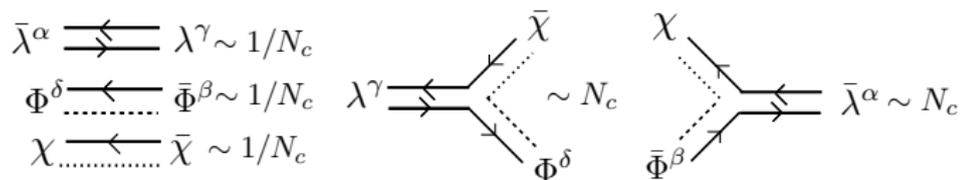
- Original motivation - M_5 brane density [hep-th/9610236](https://arxiv.org/abs/hep-th/9610236) (Berkooz & Douglas).
- Reducing the [D5/D9](https://arxiv.org/abs/hep-th/0112081) system (Van Raamsdonk, [hep-th/0112081](https://arxiv.org/abs/hep-th/0112081)):

$$\mathcal{L} = \frac{1}{g^2} \text{Tr} \left(\frac{1}{2} D_0 X^a D_0 X^a + \frac{i}{2} \lambda^{\dagger\rho} D_0 \lambda_\rho + \frac{1}{2} D_0 \bar{X}^{\rho\dot{\rho}} D_0 X_{\rho\dot{\rho}} + \frac{i}{2} \theta^{\dagger\dot{\rho}} D_0 \theta_{\dot{\rho}} \right) + \frac{1}{g^2} \text{tr} \left(D_0 \bar{\Phi}^\rho D_0 \Phi_\rho + i \chi^\dagger D_0 \chi \right) + \mathcal{L}_{\text{int}}$$

where:

$$\begin{aligned} \mathcal{L}_{\text{int}} = & \frac{1}{g^2} \text{Tr} \left(\frac{1}{4} [X^a, X^b][X^a, X^b] + \frac{1}{2} [X^a, \bar{X}^{\rho\dot{\rho}}][X^a, X_{\rho\dot{\rho}}] - \frac{1}{4} [\bar{X}^{\alpha\dot{\alpha}}, X_{\beta\dot{\beta}}][\bar{X}^{\beta\dot{\beta}}, X_{\alpha\dot{\alpha}}] \right) \\ & - \frac{1}{g^2} \text{tr} \left(\bar{\Phi}^\rho (X^a - m^a)(X^a - m^a) \Phi_\rho \right) \\ & + \frac{1}{g^2} \text{tr} \left(\bar{\Phi}^\alpha [\bar{X}^{\beta\dot{\alpha}}, X_{\alpha\dot{\alpha}}] \Phi_\beta + \frac{1}{2} \bar{\Phi}^\alpha \Phi_\beta \bar{\Phi}^\beta \Phi_\alpha - \bar{\Phi}^\alpha \Phi_\alpha \bar{\Phi}^\beta \Phi_\beta \right) \\ & + \frac{1}{g^2} \text{Tr} \left(\frac{1}{2} \bar{\lambda}^\rho \gamma^a [X^a, \lambda_\rho] + \frac{1}{2} \bar{\theta}^{\dot{\alpha}} \gamma^a [X^a, \theta_{\dot{\alpha}}] - \sqrt{2} i \varepsilon_{\alpha\beta} \bar{\theta}^{\dot{\alpha}} [X_{\beta\dot{\alpha}}, \lambda_\alpha] \right) \\ & + \frac{1}{g^2} \text{tr} \left(\bar{\chi} \gamma^a (X^a - m^a) \chi + \sqrt{2} i \varepsilon_{\alpha\beta} \bar{\chi} \lambda_\alpha \Phi_\beta - \sqrt{2} i \varepsilon_{\alpha\beta} \bar{\Phi}^\alpha \bar{\lambda}_\beta \chi \right) \end{aligned}$$

Quenched versus dynamical



Uplift to 11D supergravity

- Both the **D0** and **D4**-branes can be uplifted to **11D**.
- The **D0**-brane is a **KK** - brane (uplifts to **KK** momentum).
- The **D4**-brane uplifts to a **$\mathcal{M}5$** -brane.
- Therefore the **D0/D4** system uplifts to a **$\mathcal{M}5$** -brane background with quantised momentum along the **\mathcal{M}** -theory circle.
- A magnetic analogue of this system, the **D2/D6** system has been already studied in the literature by **Cherkiz** and **Hashimoto** ([hep-th/0210105](#)). The uplifted background is a **$\mathcal{M}2$** -membrane background with a **Taub-NUT** geometry in the transverse space.
- We will consider the most general ansatz consistent with the symmetries and use **SUSY** to restrict it.

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The ansatz

- We consider the ansatz:

$$ds_{11}^2 = -K_1(u, v) dt^2 + K_3(u, v)(dx_{11} + A_0(u, v) dt)^2 + K_2(u, v)(du^2 + u^2 d\Omega_3^2) + K_4(u, v)(dv^2 + v^2 d\Omega_4^2),$$

$$\mathcal{F}_{(4)} = F'(v) v^4 \sin^3 \psi \sin \tilde{\alpha} \cos \tilde{\alpha} d\psi \wedge d\tilde{\alpha} \wedge d\tilde{\beta} \wedge d\tilde{\gamma},$$

$$d\Omega_3^2 = d\alpha^2 + \sin^2 \alpha d\beta^2 + \cos^2 \alpha d\gamma^2,$$

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- Charge conservation fixes $F'(v)$:

$$\int \mathcal{F}_{(4)} = F'(v) v^4 = \frac{8}{3} \pi^2 v^4 F'(v) = -Q_5$$

- Which results in:

$$F(v) = 1 + \frac{Q_5}{8\pi^2 v^3} \equiv 1 + \frac{v_5^3}{v^3}$$

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Gravitino variation

- We demand that the gravitino variation vanish:

$$\delta\psi_\mu = \nabla_\mu \varepsilon - \frac{1}{6} \frac{F'(v)}{K_4(u, v)^2} \Gamma_\mu \bar{\Gamma}^{78910} \varepsilon \quad \text{for } \mu \in S^4,$$

$$\delta\psi_\mu = \nabla_\mu \varepsilon + \frac{1}{12} \frac{F'(v)}{K_4(u, v)^2} \Gamma_\mu \bar{\Gamma}^{78910} \varepsilon \quad \text{for } \mu \notin S^4.$$

- We use the conventions / projections:

$$\bar{\Gamma}^{012345678910} = \delta_1$$

$$\bar{\Gamma}^{012345} \varepsilon = \delta_2 \varepsilon$$

$$\bar{\Gamma}_{01} \varepsilon = \delta_3 \varepsilon$$

- where $\bar{\Gamma}^i$ are flat and we have:

$$\delta_1^2 = \delta_2^2 = \delta_3^2 = 1$$

- Clearly the background preserves **1/4** of the original **SUSY** consistent with the description of **SUSY** brane intersection.

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BPS equations

- This boils down to solving:

$$0 = \frac{\partial_u K_1}{K_1} - \delta_3 \left(\frac{K_3}{K_1} \right)^{1/2} \partial_u A_0$$

$$0 = \frac{\partial_v K_1}{K_1} - \delta_3 \left(\frac{K_3}{K_1} \right)^{1/2} \partial_v A_0 + \delta_1 \delta_2 \frac{1}{3} \frac{F'(v)}{K_4^{3/2}}$$

$$0 = \frac{\partial_u K_1}{K_1} + \frac{\partial_u K_3}{K_3}$$

$$F'(v) = -\frac{3}{2} \delta_1 \delta_2 K_4^{3/2} \left(\frac{\partial_v K_1}{K_1} + \frac{\partial_v K_3}{K_3} \right)$$

$$F'(v) = -3 \delta_1 \delta_2 K_4^{3/2} \frac{\partial_v K_2}{K_2}$$

$$F'(v) = \delta_1 \delta_2 \frac{3}{2} K_4^{1/2} \partial_v K_4$$

$$\partial_u K_2 = 0$$

$$\partial_u K_4 = 0$$

- With the choice $\delta_1 = \delta_2$ and $\delta_3 = 1$ we arrive at:

$$K_1 = \left(1 + \frac{v^3}{v^3}\right)^{-1/3} H(u, v)^{-1}$$

$$K_2 = \left(1 + \frac{v^3}{v^3}\right)^{-1/3}$$

$$K_3 = \left(1 + \frac{v^3}{v^3}\right)^{-1/3} H(u, v)$$

$$K_4 = \left(1 + \frac{v^3}{v^3}\right)^{2/3}$$

$$A_0(u, v) = H(u, v)^{-1} - 1$$

$$F(v) = 1 + \frac{v^3}{v^3}$$

The metric

- The resulting metric is given by:

$$ds_{11}^2 = \left(1 + \frac{v_5^3}{v^3}\right)^{-1/3} \left(-\frac{dt^2}{H(u, v)} + H(u, v) (dx_{11} + (H(u, v)^{-1} - 1) dt)^2 + du^2 + u^2 d\Omega_3^2\right) + \left(1 + \frac{v_5^3}{v^3}\right)^{2/3} (dv^2 + v^2 d\Omega_4^2) .$$

- And the reduced 10D metric is:

$$ds_{10}^2 = -H(u, v)^{-1/2} \left(1 + \frac{v_4^3}{v^3}\right)^{-1/2} dt^2 + H(u, v)^{1/2} \left[\frac{(du^2 + u^2 d\Omega_3^2)}{\left(1 + \frac{v_4^3}{v^3}\right)^{1/2}} + \left(1 + \frac{v_4^3}{v^3}\right)^{1/2} (dv^2 + v^2 d\Omega_4^2) \right]$$

$$e^\Phi = \left(1 + \frac{v_4^3}{v^3}\right)^{-1/4} H(u, v)^{3/4},$$

$$C_0 = (H(u, v)^{-1} - 1) dt, \quad F_4 = -3 v_4^3 \omega_{S^4}$$

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$$ds_{10}^2 = -H(u, v)^{-1/2} \left(1 + \frac{v^3}{v^3}\right)^{-1/2} dt^2 + H(u, v)^{1/2} \left[\frac{(du^2 + u^2 d\Omega_3^2)}{\left(1 + \frac{v^3}{v^3}\right)^{1/2}} + \left(1 + \frac{v^3}{v^3}\right)^{1/2} (dv^2 + v^2 d\Omega_4^2) \right]$$

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$$C_0 = (H(u, v)^{-1} - 1) dt, \quad F_4 = -3 v_4^3 \omega_{S^4}$$

Harmonic equation

- $H(u, v)$ is not fixed by SUSY. We define the angular momentum along $\xi = \partial/\partial x_{11}$:

$$J_{x_{11}} \propto \int_{\partial\Sigma} \star(\nabla_\mu \xi_\nu dx^\mu \wedge dx^\nu) = \int_\Sigma d \star(\nabla_\mu \xi_\nu dx^\mu \wedge dx^\nu) .$$

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- The resulting harmonic equation can be written as:

$$\partial_v^2 H(u, v) + \frac{4}{v} \partial_v H(u, v) + \left(1 + \frac{v^3}{v^3}\right) \left(\partial_u^2 H(u, v) + \frac{3}{u} \partial_u H(u, v)\right) = 0$$

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Perturbative solution at infinity

- It is natural to consider the expansion:

$$H(u, v) = \sum_{n=0}^{\infty} \left(\frac{v_5^3}{v^3} \right)^n H_n(u, v)$$

- and Fourier transform each component:

$$H(u, v) = 1 + \frac{r_0^7}{(2\pi)^4} \int d^4 p e^{i\vec{p}\cdot\vec{u}} h(p, v) = 1 + \frac{r_0^7}{4\pi^2} \int_0^{\infty} dp p^2 \frac{J_1(pu)}{u} h(p, v)$$

- $h_n(p, v)$ satisfy the recursive relation:

$$\partial_v^2 h_n(p, v) + \frac{4-6n}{v} \partial_v h_n(p, v) + \left(\frac{9n(n-1)}{v^2} - p^2 \right) h_n(p, v) = p^2 h_{n-1}(p, v)$$

- with general solution:

$$h_n(p, v) = A_n e^{-pv} (pv)^{3n} \frac{1+pv}{v^3} + B_n e^{pv} (pv)^{3n} \frac{1-pv}{v^3},$$

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- We could compare with a result from the probe brane calculation. To leading order in N_f they should agree (we have $v_5^3 \sim N_f/N_c \lambda$).
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General solution

- To obtain the general solution one needs to use information in the IR. In the 11D uplift for small v the metric is:

$$ds_{11}^2 = \frac{z^2}{L^2} \left(\frac{dt^2}{H} + H(dx_{11} + (H^{-1} - 1)dt)^2 + du^2 + u^2 d\Omega_3^2 \right) + \frac{L^2 dz^2}{z^2} + \frac{L^2}{4} d\Omega_4^2$$

- Where $z = 2\sqrt{v_5 v}$ and $L = 2v_5$.
- This is an $AdS_7 \times S^4$ space-time with a spherical wave.
- The independent solutions are given in terms of Bessel K_3 and I_3 functions. Only the K_3 function is regular at the origin, however it is the I_3 function that corresponds to the perturbative solution at infinity.
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