

Where we are on B-decay Discrepancies

Diego Guadagnoli
CNRS, LAPTh Annecy

Overall message

The TH picture has evolved while, remarkably staying coherent – in spite of all the constraints

*Based on 1903.10434, with
J. Aebischer, W. Altmannshofer, M. Reboud, P. Stangl and D. M. Straub*

B-discrepancies

2019

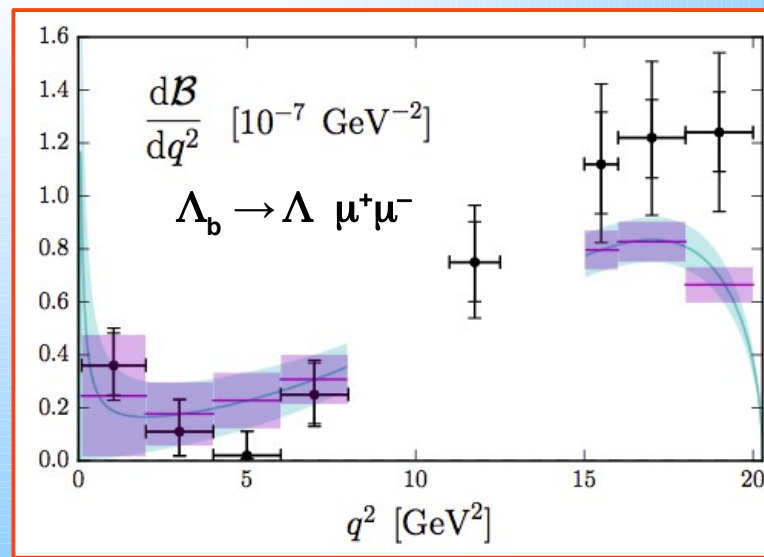
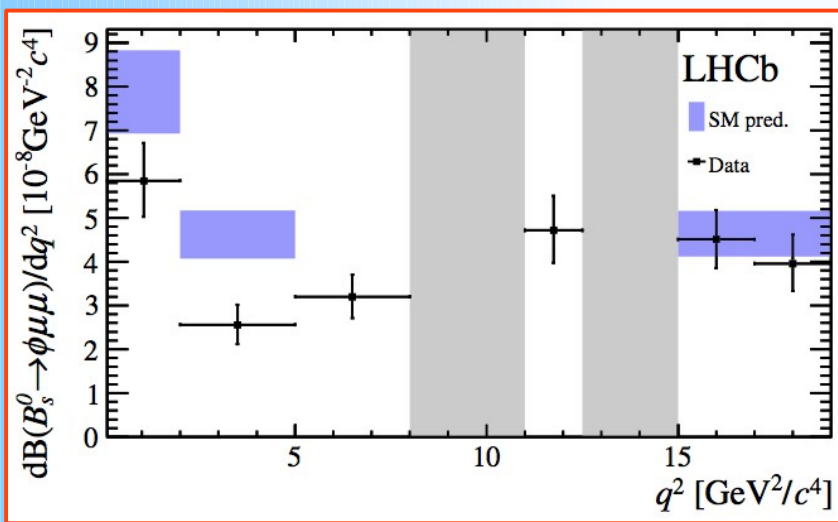
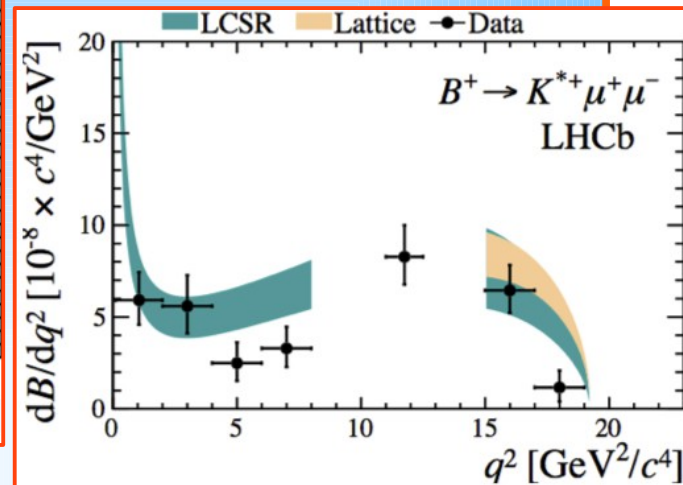
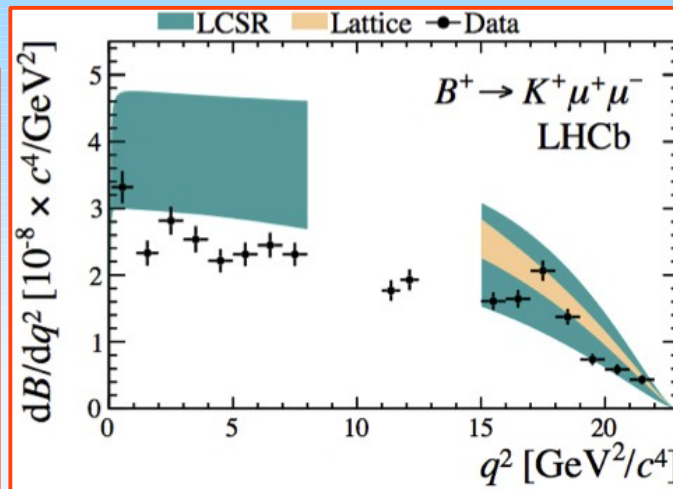
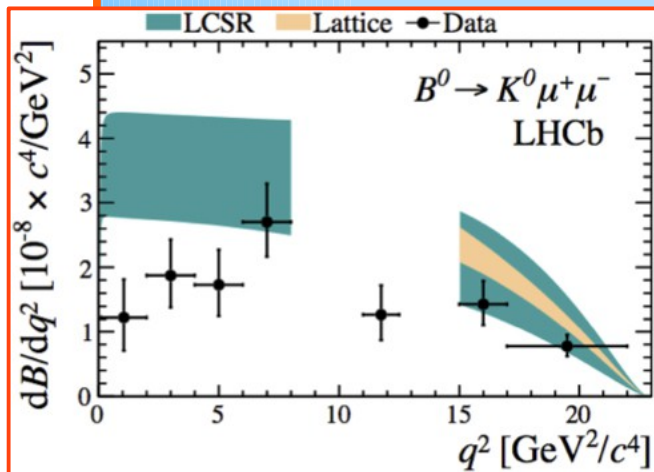
Data overview

4 groups of interesting datasets, w/ different challenges

① *$b \rightarrow s \mu\mu$ BR data < SM*

Challenge: $B \rightarrow$ light meson f.f.'s

$b \rightarrow s \mu\mu$ BR data < SM



Data overview

4 groups of interesting datasets, w/ different challenges

- 1 $b \rightarrow s \mu\mu$ BR data $<$ SM
Challenge: $B \rightarrow$ light meson f.f.'s
- 2 $B \rightarrow K^* \mu\mu$ angular data
Challenge: charm loops

Data overview

4 groups of interesting datasets, w/ different challenges

- 1 $b \rightarrow s \mu\mu$ BR data $<$ SM
Challenge: $B \rightarrow$ light meson f.f.'s
- 2 $B \rightarrow K^* \mu\mu$ angular data
Challenge: charm loops
- 3 $b \rightarrow s \mu\mu$ / $b \rightarrow s ee$ ratios
Challenge: (mostly) stats

$b \rightarrow s \mu\mu / b \rightarrow s ee$ ratios

$$R_K(q_{\min}^2, q_{\max}^2) \equiv \frac{\Gamma(B^+ \rightarrow K^+ \mu\mu)}{\Gamma(B^+ \rightarrow K^+ ee)} \Big|_{[q_{\min}^2, q_{\max}^2]}$$

A probe of Lepton-Universality Violation, by construction

- $R_K(1\text{GeV}^2, 6\text{GeV}^2) = 0.846_{-0.054}^{+0.060} \pm 0.016$
(2.5σ effect)
- $R_{K^*0}(1.1\text{GeV}^2, 6.0\text{GeV}^2) = 0.685_{-0.069}^{+0.113} \pm 0.047$
($\sim 2.4\sigma$ effect)
- $R_{K^*0}(0.045\text{GeV}^2, 1.1\text{GeV}^2) = 0.660_{-0.070}^{+0.110} \pm 0.024$
($\sim 2.2\sigma$ effect)

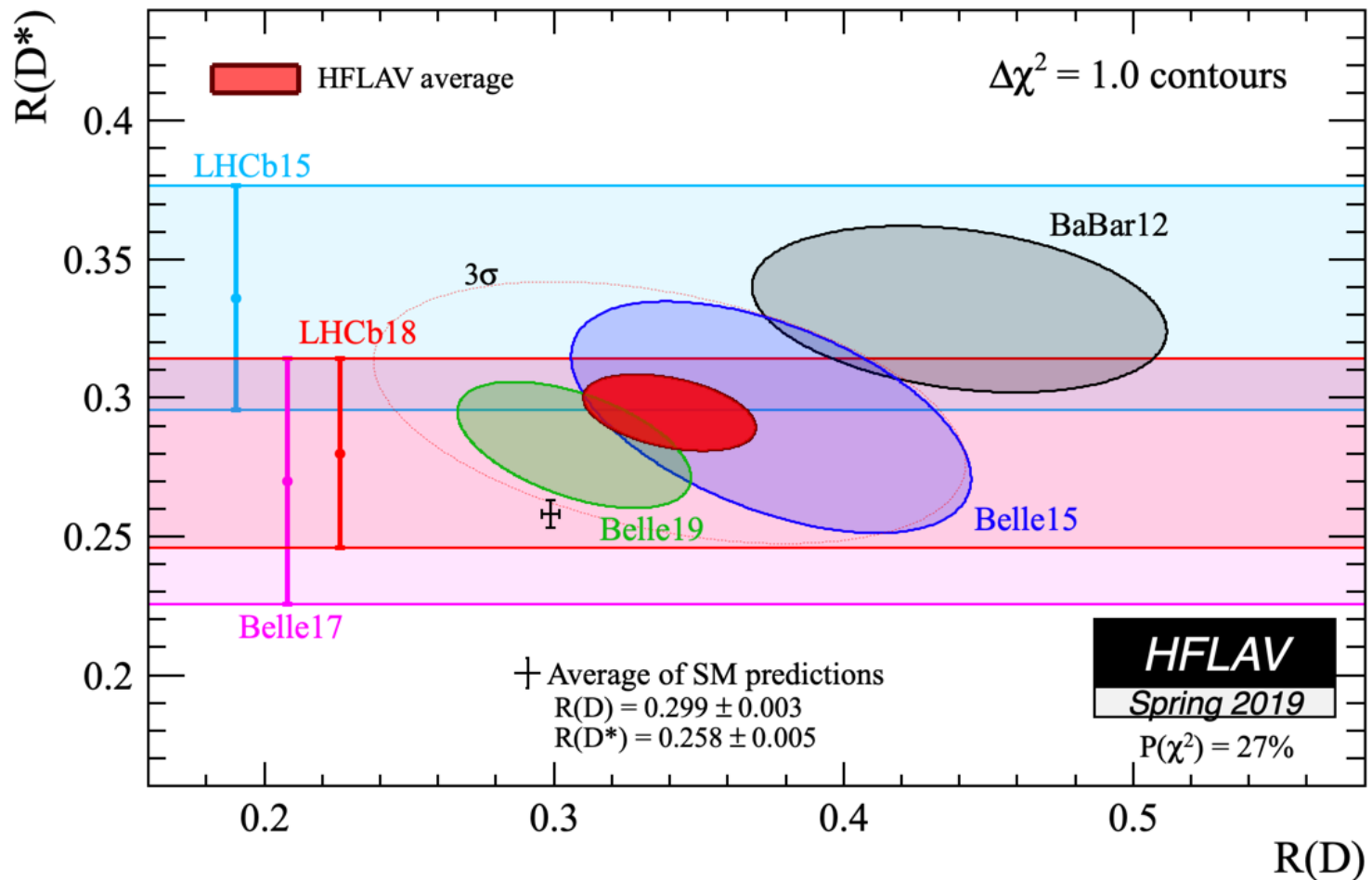
Data overview

4 groups of interesting datasets, w/ different challenges

- 1 $b \rightarrow s \mu\mu$ BR data $<$ SM
Challenge: $B \rightarrow$ light meson f.f.'s
- 2 $B \rightarrow K^* \mu\mu$ angular data
Challenge: charm loops
- 3 $b \rightarrow s \mu\mu$ / $b \rightarrow s ee$ ratios
Challenge: (mostly) stats
- 4 $b \rightarrow c \tau\nu$ / $b \rightarrow c \ell\nu$ ratios
Challenge: stats + syst

$b \rightarrow c \tau \nu$ / $b \rightarrow c \ell \nu$ ratios

$$R_D^{(*)} \equiv \frac{\Gamma(B \rightarrow D^{(*)} \tau \nu)}{\Gamma(B \rightarrow D^{(*)} \ell \nu)}$$



Basic TH considerations

- ① $(b \rightarrow s \mu\mu \text{ BR data} < \text{SM})$
- +
- ② $(B \rightarrow K^* \mu\mu \text{ angular data})$



*may be alleviated by
more conservative TH
assumptions*

Basic TH considerations

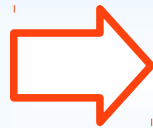
- ① $(b \rightarrow s \mu\mu \text{ BR data} < \text{SM})$
- +
- ② $(B \rightarrow K^* \mu\mu \text{ angular data})$



*may be alleviated by
more conservative TH
assumptions*

But

- ① + ② + ③
- $(R_{K^{(*)}})$



*Explicable (quantitatively)
w/ two semi-leptonic operators*



*substantial improvement
w.r.t. SM alone*

Basic TH considerations

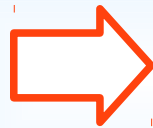
- ① $(b \rightarrow s \mu\mu \text{ BR data} < \text{SM})$
- +
- ② $(B \rightarrow K^* \mu\mu \text{ angular data})$



*may be alleviated by
more conservative TH
assumptions*

But

- ① + ② + ③
- $(R_{K^{(*)}})$



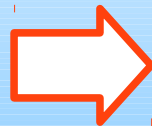
*Explicable (quantitatively)
w/ two semi-leptonic operators*



*substantial improvement
w.r.t. SM alone*

And

- ① + ② + ③ + ④
- $(R_{D^{(*)}})$



*Explicable (quantitatively)
w/ single-mediator
simplified models*

Data updates

Including all data as of Moriond 19 crucial to get full picture

(a) R_K update from LHCb Run1 + 1/3 of Run2

$$R_K \cong 0.85 (1 \pm 7\%) \quad \Rightarrow \quad 2.5\sigma$$

Data updates

Including all data as of Moriond 19 crucial to get full picture

(a) R_K update from LHCb Run1 + 1/3 of Run2

$$R_K \cong 0.85 (1 \pm 7\%) \quad \Rightarrow \quad 2.5\sigma$$

(b) R_{K^*} update from Belle

$$R_{K^*} \cong 0.90 (1 \pm 30\%) \quad \Rightarrow \quad (\text{compatible w/ LHCb's } R_{K^*})$$

Data updates

Including all data as of Moriond 19 crucial to get full picture

(a) R_K update from LHCb Run1 + 1/3 of Run2

$$R_K \cong 0.85 (1 \pm 7\%) \quad \Rightarrow \quad 2.5\sigma$$

(b) R_{K^*} update from Belle

$$R_{K^*} \cong 0.90 (1 \pm 30\%) \quad \Rightarrow \quad (\text{compatible w/ LHCb's } R_{K^*})$$

(c) $B_s \rightarrow \mu\mu$ from ATLAS

Data updates

Including all data as of Moriond 19 crucial to get full picture

(a) R_K update from LHCb Run1 + 1/3 of Run2

$$R_K \cong 0.85 (1 \pm 7\%) \quad \Rightarrow \quad 2.5\sigma$$

(b) R_{K^*} update from Belle

$$R_{K^*} \cong 0.90 (1 \pm 30\%) \quad \Rightarrow \quad (\text{compatible w/ LHCb's } R_{K^*})$$

(c) $B_s \rightarrow \mu\mu$ from ATLAS

(d) $R_{D^{(*)}}$ from Belle (semil. tagging) \Rightarrow (SM-like, 1.2σ)

Data updates

Including all data as of Moriond 19 crucial to get full picture

(a) R_K update from LHCb Run1 + 1/3 of Run2

$$R_K \cong 0.85 (1 \pm 7\%) \quad \Rightarrow \quad 2.5\sigma$$

(b) R_{K^*} update from Belle

$$R_{K^*} \cong 0.90 (1 \pm 30\%) \quad \Rightarrow \quad (\text{compatible w/ LHCb's } R_{K^*})$$

(c) $B_s \rightarrow \mu\mu$ from ATLAS

(d) $R_{D^{(*)}}$ from Belle (semil. tagging) \Rightarrow (SM-like, 1.2σ)

and more ($\Lambda_b \rightarrow \Lambda \mu\mu$, $b \rightarrow s \gamma$ & $b \rightarrow s g$, ...)

EW-scale

Effective-Theory picture

$b \rightarrow s$ EFT picture

- *One starts from the following Hamiltonian*

$$H(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4\pi} \left[\bar{b}_L \gamma^\lambda s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_\lambda \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_\lambda \gamma_5 \mu \right) \right]$$

$b \rightarrow s$ EFT picture

- One starts from the following Hamiltonian

$$H(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4\pi} \left[\bar{b}_L \gamma^\lambda s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_\lambda \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_\lambda \gamma_5 \mu \right) \right]$$

About equal size & opposite sign
in the SM (at the m_b scale)

$(V - A) \times (V - A)$ interaction

b → **s** EFT picture

- One starts from the following Hamiltonian

$$H(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4\pi} \left[\bar{b}_L \gamma^\lambda s_L \cdot \boxed{C_9^{(\mu)}} \bar{\mu} \gamma_\lambda \mu + \boxed{C_{10}^{(\mu)}} \bar{\mu} \gamma_\lambda \gamma_5 \mu \right]$$

About equal size & opposite sign
in the SM (at the m_b scale)

(V - A) x (V - A) interaction

- The best-performing BSM scenarios to explain the data involve

$$O_9 \propto \bar{b}_L \gamma^\lambda s_L \cdot \bar{\mu} \gamma_\lambda \mu \quad O_{10} \propto \bar{b}_L \gamma^\lambda s_L \cdot \bar{\mu} \gamma_\lambda \gamma_5 \mu$$

b → s EFT picture

- One starts from the following Hamiltonian

$$H(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4\pi} \left[\bar{b}_L \gamma^\lambda s_L \cdot \boxed{C_9^{(\mu)}} \bar{\mu} \gamma_\lambda \mu + \boxed{C_{10}^{(\mu)}} \bar{\mu} \gamma_\lambda \gamma_5 \mu \right]$$

About equal size & opposite sign
in the SM (at the m_b scale)

$(V - A) \times (V - A)$ interaction

- The best-performing BSM scenarios to explain the data involve

$$O_9 \propto \bar{b}_L \gamma^\lambda s_L \cdot \bar{\mu} \gamma_\lambda \mu \quad O_{10} \propto \bar{b}_L \gamma^\lambda s_L \cdot \bar{\mu} \gamma_\lambda \gamma_5 \mu$$

- Specifically, either O_9 alone,
- or $O_9 - O_{10}$ \Rightarrow again, $(V - A) \times (V - A)$
well-suited to UV-complete models

1-Wilson-coeff. picture

Coeff.	best fit	1σ	2σ	pull
$C_9^{bs\mu\mu}$	-0.97	[-1.12, -0.81]	[-1.27, -0.65]	5.9σ
$C_9^{\prime bs\mu\mu}$	+0.14	[-0.03, +0.32]	[-0.20, +0.51]	0.8σ
$C_{10}^{bs\mu\mu}$	+0.75	[+0.62, +0.89]	[+0.48, +1.03]	5.7σ
$C_{10}^{\prime bs\mu\mu}$	-0.24	[-0.36, -0.12]	[-0.49, +0.00]	2.0σ
$C_9^{bs\mu\mu} = C_{10}^{bs\mu\mu}$	+0.20	[+0.06, +0.36]	[-0.09, +0.52]	1.4σ
$C_9^{bs\mu\mu} = -C_{10}^{bs\mu\mu}$	-0.53	[-0.61, -0.45]	[-0.69, -0.37]	6.6σ

1-Wilson-coeff. picture

Coeff.	best fit	1σ	2σ	pull
$C_9^{bs\mu\mu}$	-0.97	[-1.12, -0.81]	[-1.27, -0.65]	5.9 σ
$C_9'^{bs\mu\mu}$	+0.14	[-0.03, +0.32]	[-0.20, +0.51]	0.8 σ
$C_{10}^{bs\mu\mu}$	+0.75	[+0.62, +0.89]	[+0.48, +1.03]	5.7 σ
$C_{10}'^{bs\mu\mu}$	-0.24	[-0.36, -0.12]	[-0.49, +0.00]	2.0 σ
$C_9^{bs\mu\mu} = C_{10}^{bs\mu\mu}$	+0.20	[+0.06, +0.36]	[-0.09, +0.52]	1.4 σ
$C_9^{bs\mu\mu} = -C_{10}^{bs\mu\mu}$	-0.53	[-0.61, -0.45]	[-0.69, -0.37]	6.6 σ

- Two scenarios stand out: C_9 alone or $C_9 = -C_{10}$ ($\mu\mu$ -channel only)

1-Wilson-coeff. picture

Coeff.	best fit	1σ	2σ	pull
$C_9^{bs\mu\mu}$	-0.97	[-1.12, -0.81]	[-1.27, -0.65]	5.9 σ
$C_9^{\prime bs\mu\mu}$	+0.14	[-0.03, +0.32]	[-0.20, +0.51]	0.8 σ
$C_{10}^{bs\mu\mu}$	+0.75	[+0.62, +0.89]	[+0.48, +1.03]	5.7 σ
$C_{10}^{\prime bs\mu\mu}$	-0.24	[-0.36, -0.12]	[-0.49, +0.00]	2.0 σ
$C_9^{bs\mu\mu} = C_{10}^{bs\mu\mu}$	+0.20	[+0.06, +0.36]	[-0.09, +0.52]	1.4 σ
$C_9^{bs\mu\mu} = -C_{10}^{bs\mu\mu}$	-0.53	[-0.61, -0.45]	[-0.69, -0.37]	6.6 σ

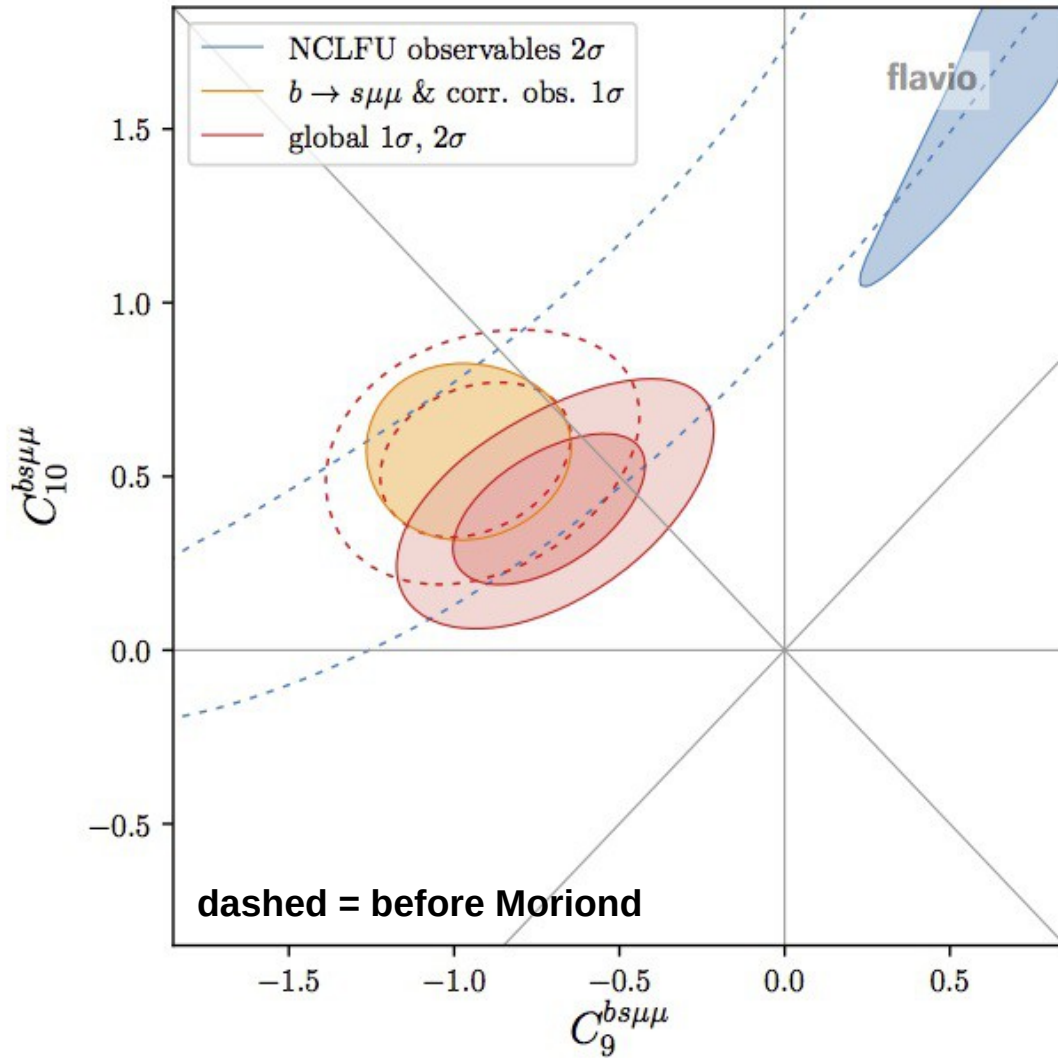
- Two scenarios stand out: C_9 alone or $C_9 = -C_{10}$ ($\mu\mu$ -channel only)
- $C_9 = -C_{10}$ now better than C_9 alone

1-Wilson-coeff. picture

Coeff.	best fit	1σ	2σ	pull
$C_9^{bs\mu\mu}$	-0.97	[-1.12, -0.81]	[-1.27, -0.65]	5.9 σ
$C_9'^{bs\mu\mu}$	+0.14	[-0.03, +0.32]	[-0.20, +0.51]	0.8 σ
$C_{10}^{bs\mu\mu}$	+0.75	[+0.62, +0.89]	[+0.48, +1.03]	5.7 σ
$C_{10}'^{bs\mu\mu}$	-0.24	[-0.36, -0.12]	[-0.49, +0.00]	2.0 σ
$C_9^{bs\mu\mu} = C_{10}^{bs\mu\mu}$	+0.20	[+0.06, +0.36]	[-0.09, +0.52]	1.4 σ
$C_9^{bs\mu\mu} = -C_{10}^{bs\mu\mu}$	-0.53	[-0.61, -0.45]	[-0.69, -0.37]	6.6 σ

- Two scenarios stand out: C_9 alone or $C_9 = -C_{10}$ ($\mu\mu$ -channel only)
- $C_9 = -C_{10}$ now better than C_9 alone
- C_{10} alone also ok, but $B \rightarrow K^* \mu\mu$ unresolved

C_9 vs. $C_9 = -C_{10}$

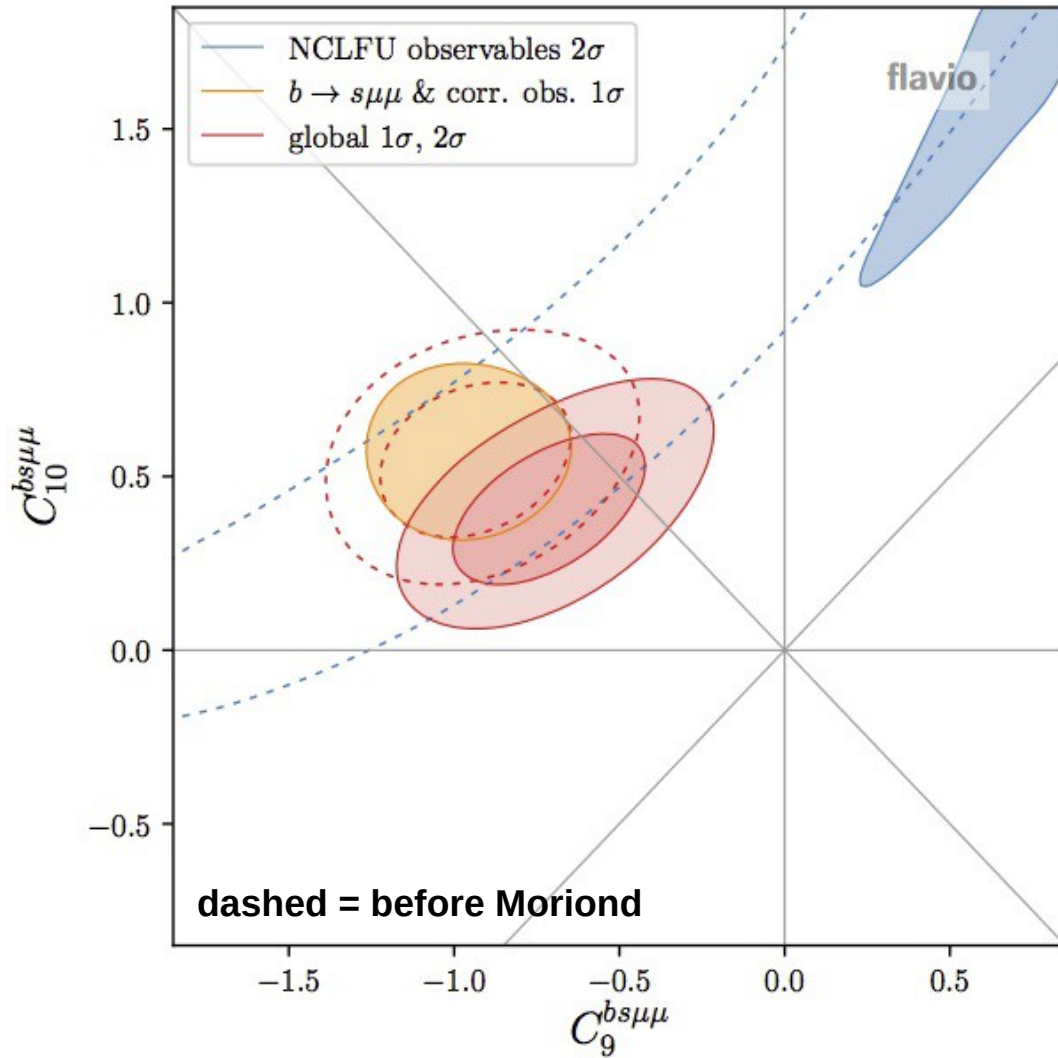


Main point

- **Before Moriond**

$R_{K^{(*)}}$ & $b \rightarrow s\mu\mu$ in nearly perfect agreement

$$C_9 \text{ vs. } C_9 = -C_{10}$$



Main point

- **Before Moriond**

$R_{K^{(*)}}$ & $b \rightarrow s\mu\mu$ in nearly perfect agreement

- **After Moriond**

some tension in C_9 dir.



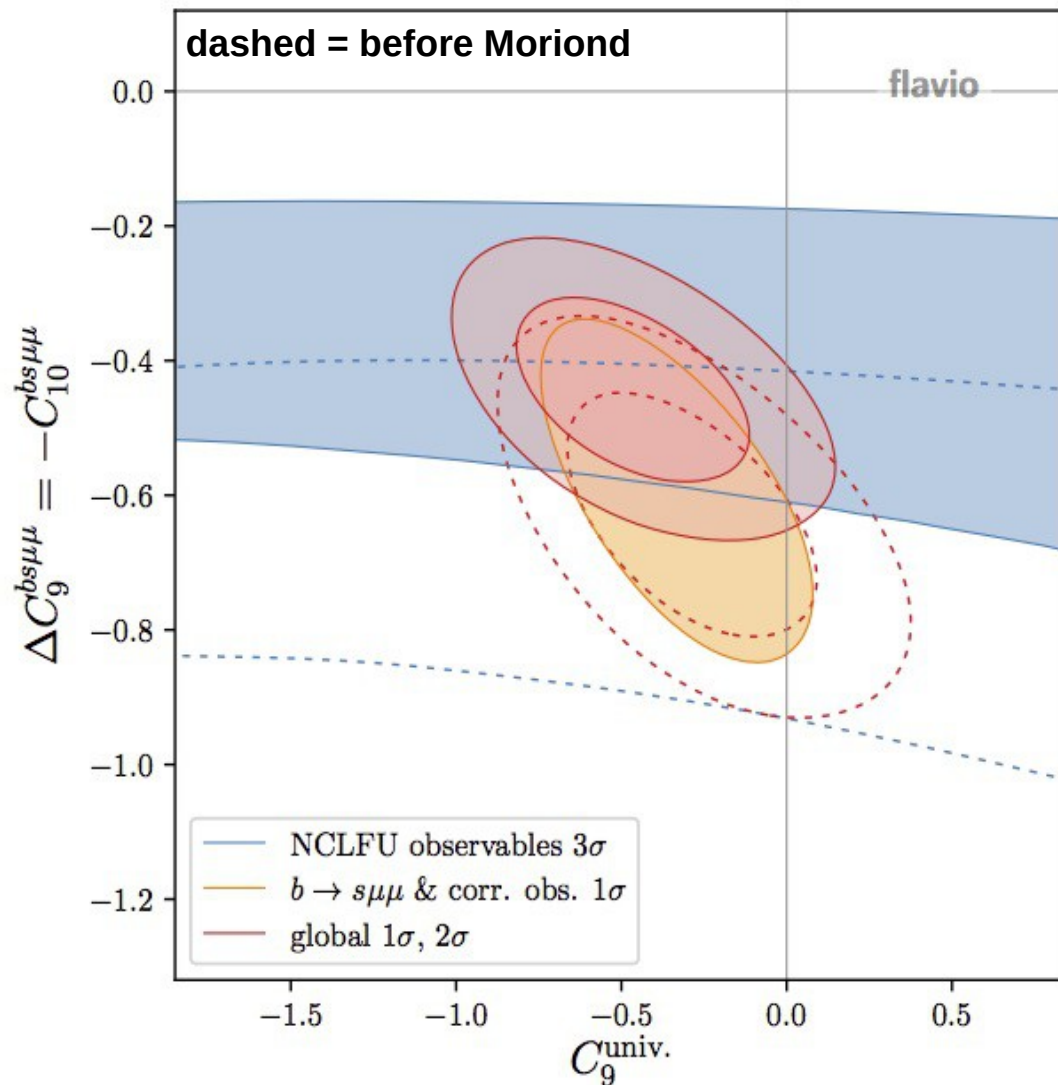
but see next for a UV interpretation

Univ. vs. non-univ. Wilson coeffs.

- Note: a $C_9^{univ.}$ component would shift $b \rightarrow s$ $\mu\mu$ data but not $R_{K^{(*)}}$

Univ. vs. non-univ. Wilson coeffs.

- Note: a $C_9^{univ.}$ component would shift $b \rightarrow s \mu\mu$ data but not $R_{K^{(*)}}$



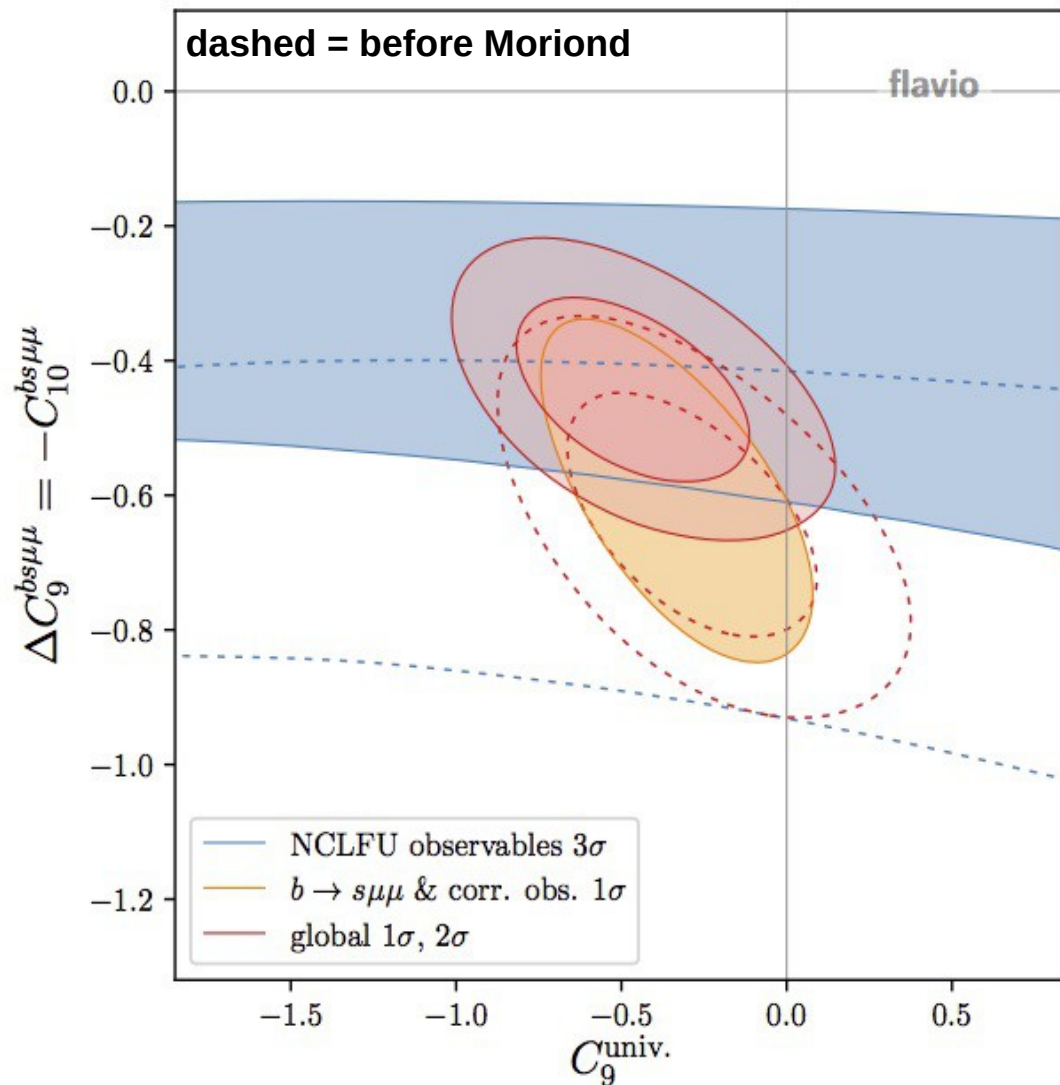
Notes

y-axis: μ -specific shift in
 $C_9 = -C_{10}$

x-axis: additional, lepton-univ.
 shift in C_9 only

Univ. vs. non-univ. Wilson coeffs.

- Note: a $C_9^{univ.}$ component would shift $b \rightarrow s \mu\mu$ data but not $R_{K^{(*)}}$



Notes

y-axis: μ -specific shift in
 $C_9 = -C_{10}$

x-axis: additional, lepton-univ.
 shift in C_9 only

After Moriond

Data tend to prefer $C_9^{univ.} \neq 0$

How to justify $C_9 = -C_{10}$ or $C_9^{univ.}$

above the EW scale?

The SMEFT picture

SMEFT 101

- *If NP is at a scale $\Lambda \gg M_{EW}$, with nothing new in between*



Effects below Λ are described by ops. constructed with SM fields, and invariant under the full SM group: $SU(3)_c \times SU(2)_L \times U(1)_Y$

This defines the SMEFT

SMEFT 101

- *If NP is at a scale $\Lambda \gg M_{EW}$, with nothing new in between*



Effects below Λ are described by ops. constructed with SM fields, and invariant under the full SM group: $SU(3)_c \times SU(2)_L \times U(1)_Y$

This defines the SMEFT

- *Non-redundant op. basis for SMEFT discussed in [B. Grzadkowski et al., JHEP 2010]*

SMEFT 101

- If NP is at a scale $\Lambda \gg M_{EW}$, with nothing new in between



Effects below Λ are described by ops. constructed with SM fields, and invariant under the full SM group: $SU(3)_c \times SU(2)_L \times U(1)_Y$

This defines the SMEFT

- Non-redundant op. basis for SMEFT discussed in [B. Grzadkowski et al., JHEP 2010]

- Such approach allows to address model-independently the question

What operators, above the EW scale, can generate contributions to $C_9^{(\mu)} = -C_{10}^{(\mu)}$ or C_9^{univ} ?

SMEFT picture

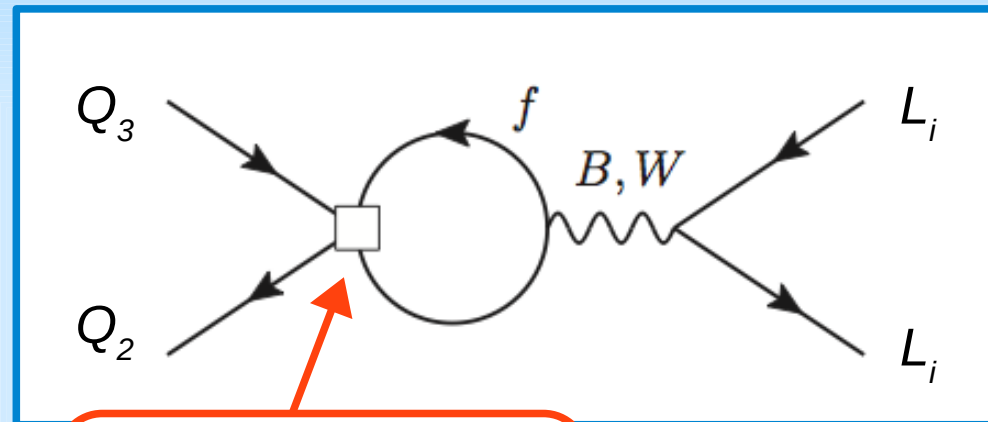
- 1 Contributions to muonic $C_9 = -C_{10}$ may come from SMEFT ops. directly matching onto $O_{9,10}$

$$[O_{LQ}^{(1)}]_{2223} = \bar{L}_2 \gamma^\lambda L_2 \cdot \bar{Q}_2 \gamma_\lambda Q_3$$

$$[O_{LQ}^{(3)}]_{2223} = \bar{L}_2 \gamma^\lambda \sigma^a L_2 \cdot \bar{Q}_2 \gamma_\lambda \sigma^a Q_3$$

SMEFT picture

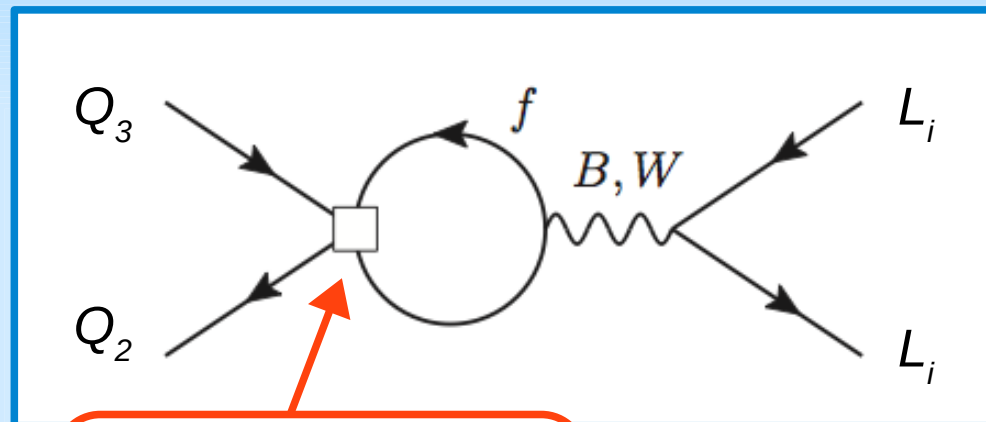
- ② Contributions to $C_9^{univ.}$ can come from RGE effects:



any suitable SMEFT
4-fermion op. here

SMEFT picture

- ② Contributions to $C_9^{univ.}$ can come from RGE effects:



any suitable SMEFT
4-fermion op. here

- Case $f = \tau$ especially interesting



Connection with “semi-tauonic” ops.

$$[O_{LQ}^{(1)}]_{3323}$$

$$[O_{LQ}^{(3)}]_{3323}$$

responsible for $b \rightarrow c \tau \nu$

Semi-tauonic ops. in short

- $[O_{LQ}^{(3)}]_{3323} \supset \bar{\tau} \gamma_L^\lambda \nu \cdot \bar{c} \gamma_{\lambda L} b$
 \Rightarrow *can explain $R_{D^{(*)}}$*

Semi-tauonic ops. in short

- $[O_{LQ}^{(3)}]_{3323} \supset \bar{\tau} \gamma_L^\lambda \nu \cdot \bar{c} \gamma_{\lambda L} b$

⇒ *can explain $R_{D^{(*)}}$*

⇩ *also induces $C_9^{univ.}$ w/ the right sign
to potentially accommodate $b \rightarrow s \mu\mu$*

[Crivellin-Greub-Müller-Saturnino]

Semi-tauonic ops. in short

- $[O_{LQ}^{(3)}]_{3323} \supset \bar{\tau} \gamma_L^\lambda \nu \cdot \bar{c} \gamma_{\lambda L} b$

⇒ can explain $R_{D^{(*)}}$

⇩ also induces $C_9^{univ.}$ w/ the right sign
to potentially accommodate $b \rightarrow s \mu \mu$
[Crivellin-Greub-Müller-Saturnino]

- $[O_{LQ}^{(1) \text{ or } (3)}]_{2223} \supset \bar{\mu} \gamma_L^\lambda \mu \cdot \bar{s} \gamma_{\lambda L} b$

⇒ can explain $R_{K^{(*)}}$

Semi-tauonic ops. in short

- $[O_{LQ}^{(3)}]_{3323} \supset \bar{\tau} \gamma_L^\lambda \nu \cdot \bar{c} \gamma_{\lambda L} b$

⇒ can explain $R_{D^{(*)}}$

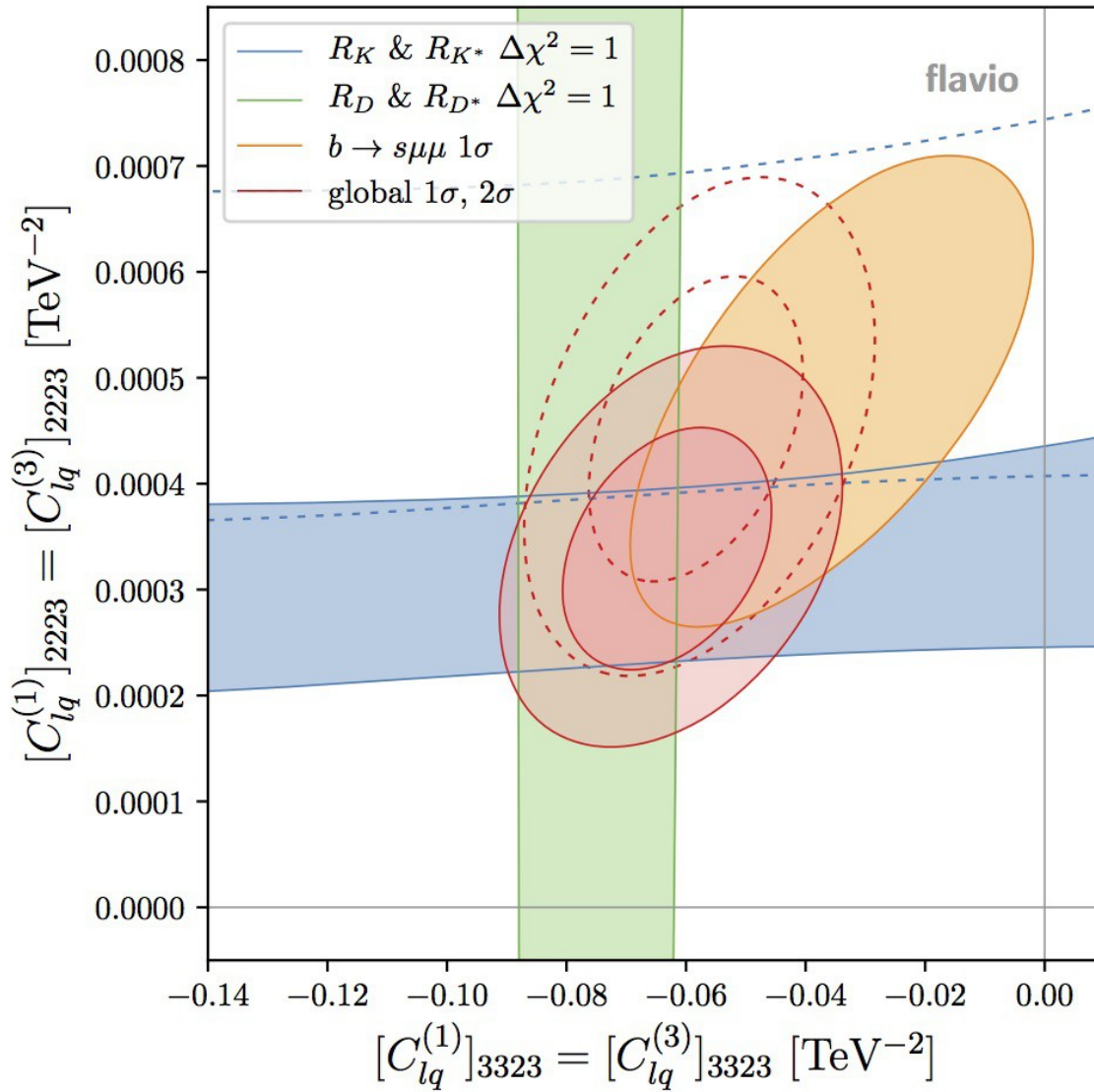
↳ also induces $C_9^{univ.}$ w/ the right sign
to potentially accommodate $b \rightarrow s \mu \mu$
[Crivellin-Greub-Müller-Saturnino]

- $[O_{LQ}^{(1) \text{ or } (3)}]_{2223} \supset \bar{\mu} \gamma_L^\lambda \mu \cdot \bar{s} \gamma_{\lambda L} b$

⇒ can explain $R_{K^{(*)}}$

- *Caveat:* one must have $[C_{LQ}^{(1)}]_{3323} \simeq [C_{LQ}^{(3)}]_{3323}$
to avoid the $B \rightarrow K^{(*)} \nu \nu$ constraint
[Buras-Girrbach-Niehoff-Straub]

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \quad \text{vs.} \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223}$$



Before Moriond (dashed)

$R_{K^{(*)}}$ (blue) and $b \rightarrow s \mu\mu$ (orange) were in perfect agreement

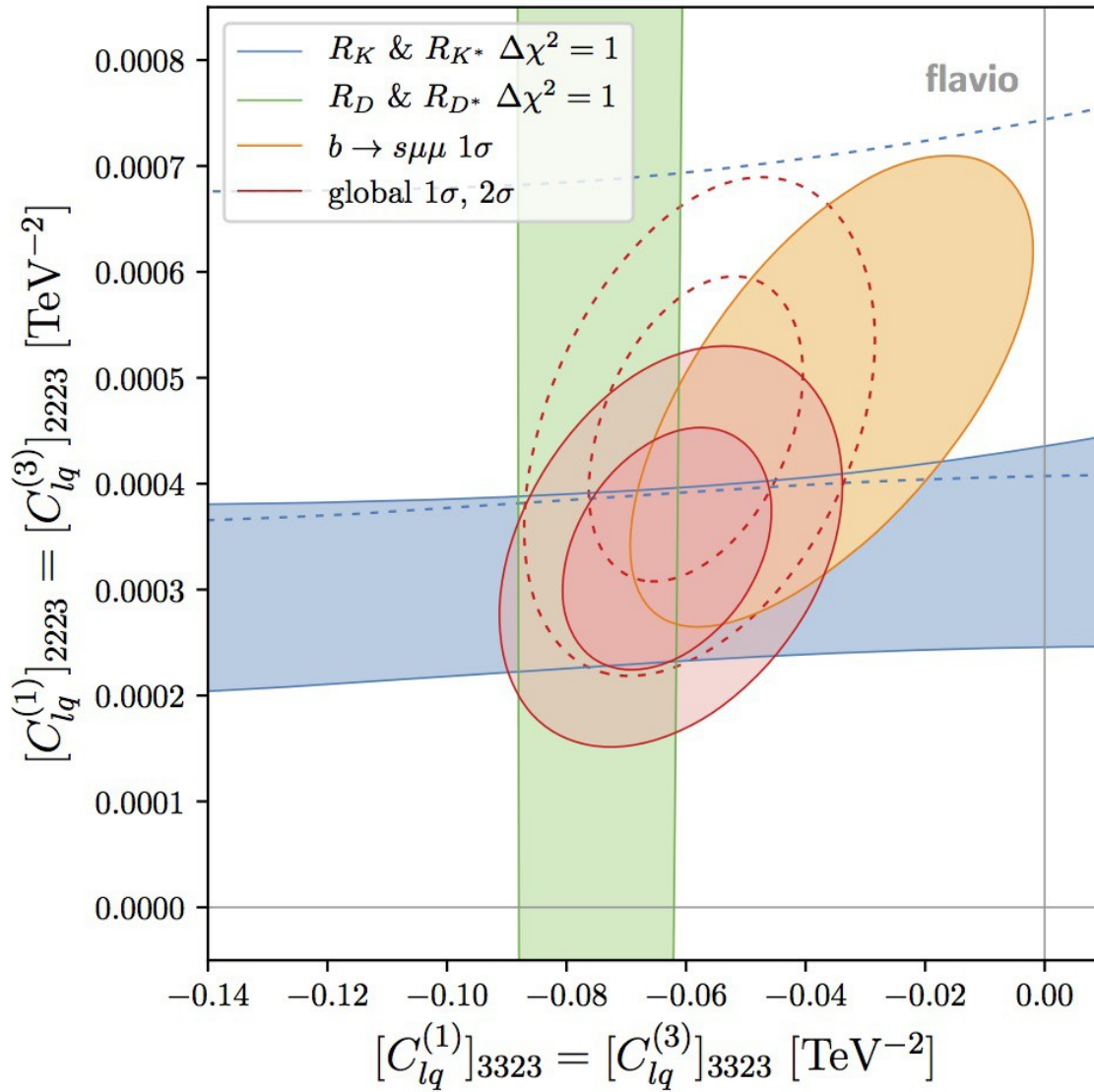


in a region close to 0 in the x-axis



$R_{D^{(*)}}$ not explained

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \quad \text{vs.} \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223}$$



Before Moriond (dashed)

$R_{K^{(*)}}$ (blue) and $b \rightarrow s \mu\mu$ (orange) were in perfect agreement



in a region close to 0 in the x-axis



$R_{D^{(*)}}$ not explained

After Moriond

$R_{K^{(*)}}$ and $b \rightarrow s \mu\mu$ intersect

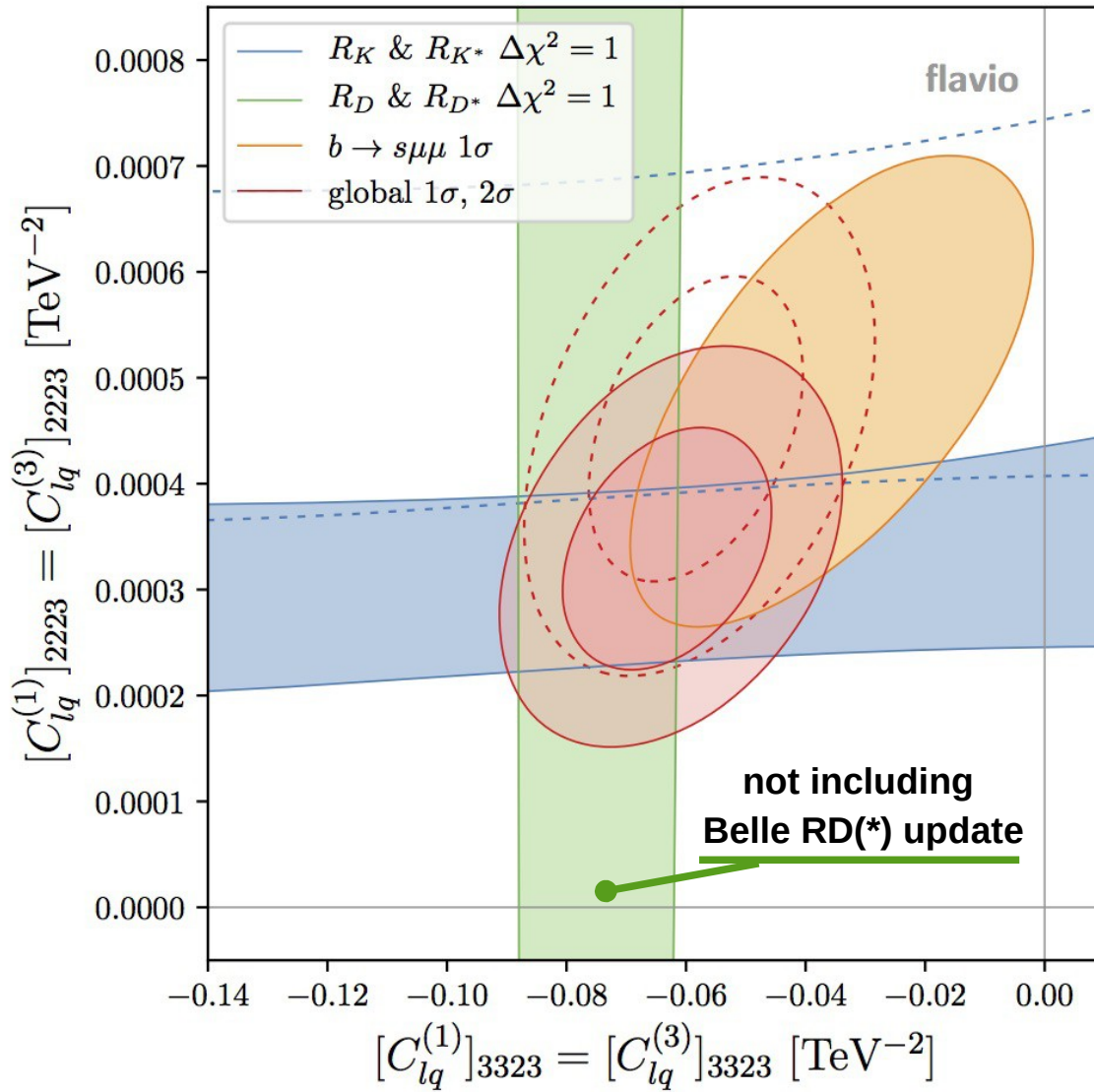


in a region with x-axis values well below 0



Quantitative agreement with the $R_{D^{(*)}}$ constraint (green)

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \quad \text{vs.} \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223}$$



Before Moriond (dashed)

$R_{K^{(*)}}$ (blue) and $b \rightarrow s\mu\mu$ (orange) were in perfect agreement



in a region close to 0 in the x-axis



$R_{D^{(*)}}$ not explained

After Moriond

$R_{K^{(*)}}$ and $b \rightarrow s\mu\mu$ intersect

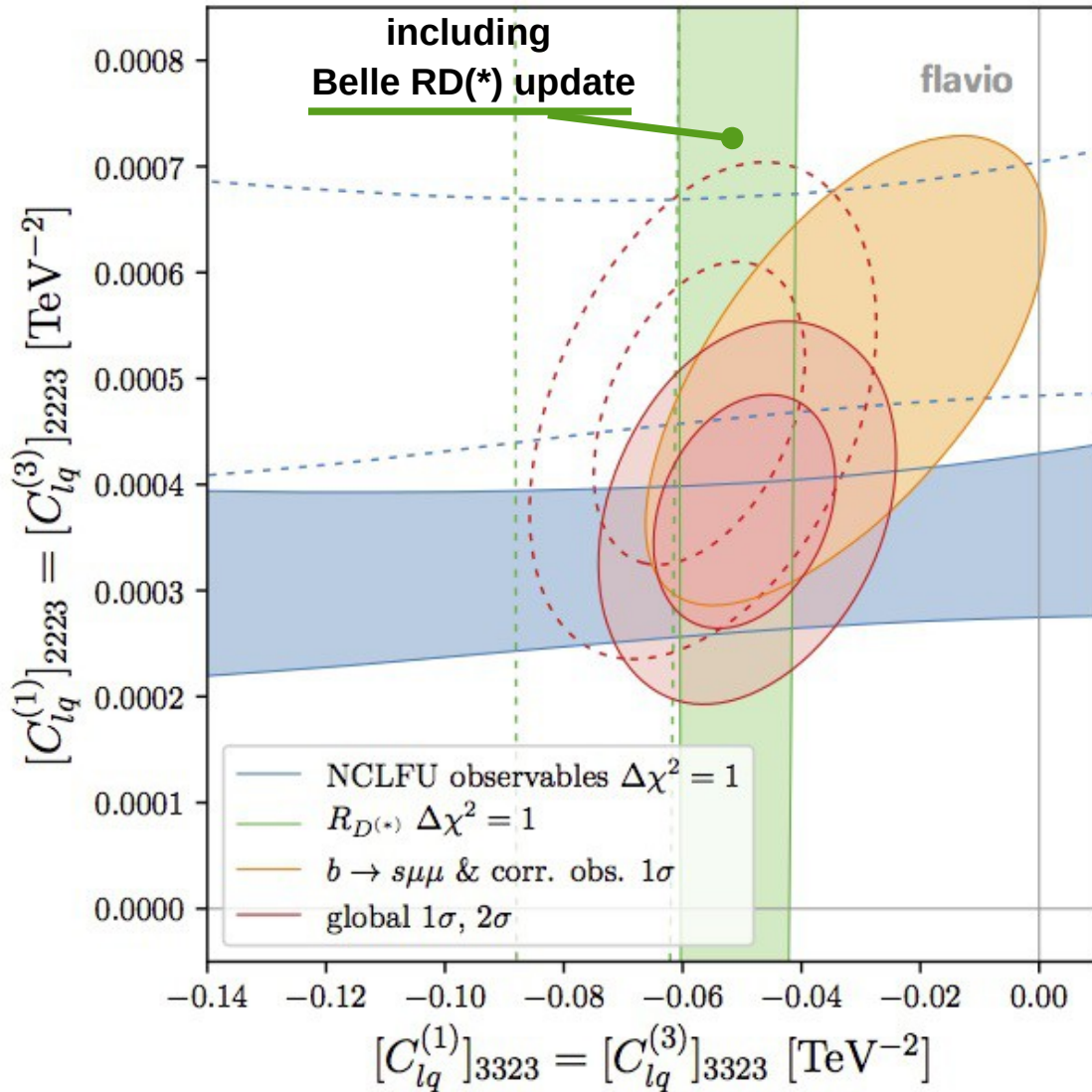


in a region with x-axis values well below 0



Quantitative agreement with the $R_{D^{(*)}}$ constraint (green)

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \quad \text{vs.} \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223}$$



Before Moriond (dashed)

$R_{K^{(*)}}$ (blue) and $b \rightarrow s\mu\mu$ (orange) were in perfect agreement



in a region close to 0 in the x-axis



$R_{D^{(*)}}$ not explained

After Moriond

$R_{K^{(*)}}$ and $b \rightarrow s\mu\mu$ intersect



in a region with x-axis values well below 0



Quantitative agreement with the $R_{D^{(*)}}$ constraint (green)

Beyond EFTs:

The picture within “simplified” models

The U_1 leptoquark

- $U_1 \sim (\mathbf{3}, \mathbf{1})_{2/3}$ is the only single mediator known to yield

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \neq 0 \quad \&\& \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223} \neq 0$$

[Alonso-Grinstein-Martin-Camalich, Calibbi-Crivellin-Ota, 2015]

The U_1 leptoquark

- $U_1 \sim (\mathbf{3}, \mathbf{1})_{2/3}$ is the only single mediator known to yield

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \neq 0 \quad \&\& \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223} \neq 0$$

[Alonso-Grinstein-Martin-Camalich, Calibbi-Crivellin-Ota, 2015]

- Define the couplings:

$$\mathcal{L}_{U_1} \supset g_{lq}^{ji} \bar{Q}^i \gamma^\mu L^j U_\mu + \text{h.c.}$$

The U_1 leptoquark

- $U_1 \sim (\mathbf{3}, \mathbf{1})_{2/3}$ is the only single mediator known to yield

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \neq 0 \quad \&\& \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223} \neq 0$$

[Alonso-Grinstein-Martin-Camalich, Calibbi-Crivellin-Ota, 2015]

- Define the couplings:

$$\mathcal{L}_{U_1} \supset g_{lq}^{ji} \bar{Q}^i \gamma^\mu L^j U_\mu + \text{h.c.}$$



$$\delta R_{K(*)} |_{\text{in } \mu \text{ channel}} \propto g_{lq}^{22} \quad \& \quad g_{lq}^{23}$$

$$\delta R_{D(*)} |_{\text{in } \tau \text{ channel}} \propto g_{lq}^{32} \quad \& \quad g_{lq}^{33}$$

The U_1 leptoquark

- $U_1 \sim (\mathbf{3}, \mathbf{1})_{2/3}$ is the only single mediator known to yield

$$[C_{LQ}^{(1)}]_{3323} = [C_{LQ}^{(3)}]_{3323} \neq 0 \quad \&\& \quad [C_{LQ}^{(1)}]_{2223} = [C_{LQ}^{(3)}]_{2223} \neq 0$$

[Alonso-Grinstein-Martin-Camalich, Calibbi-Crivellin-Ota, 2015]

- Define the couplings:

$$\mathcal{L}_{U_1} \supset g_{lq}^{ji} \bar{Q}^i \gamma^\mu L^j U_\mu + \text{h.c.}$$

$$\delta R_{K(*)} |_{\text{in } \mu \text{ channel}} \propto g_{lq}^{22} \quad \& \quad g_{lq}^{23}$$

$$\delta R_{D(*)} |_{\text{in } \tau \text{ channel}} \propto g_{lq}^{32} \quad \& \quad g_{lq}^{33}$$

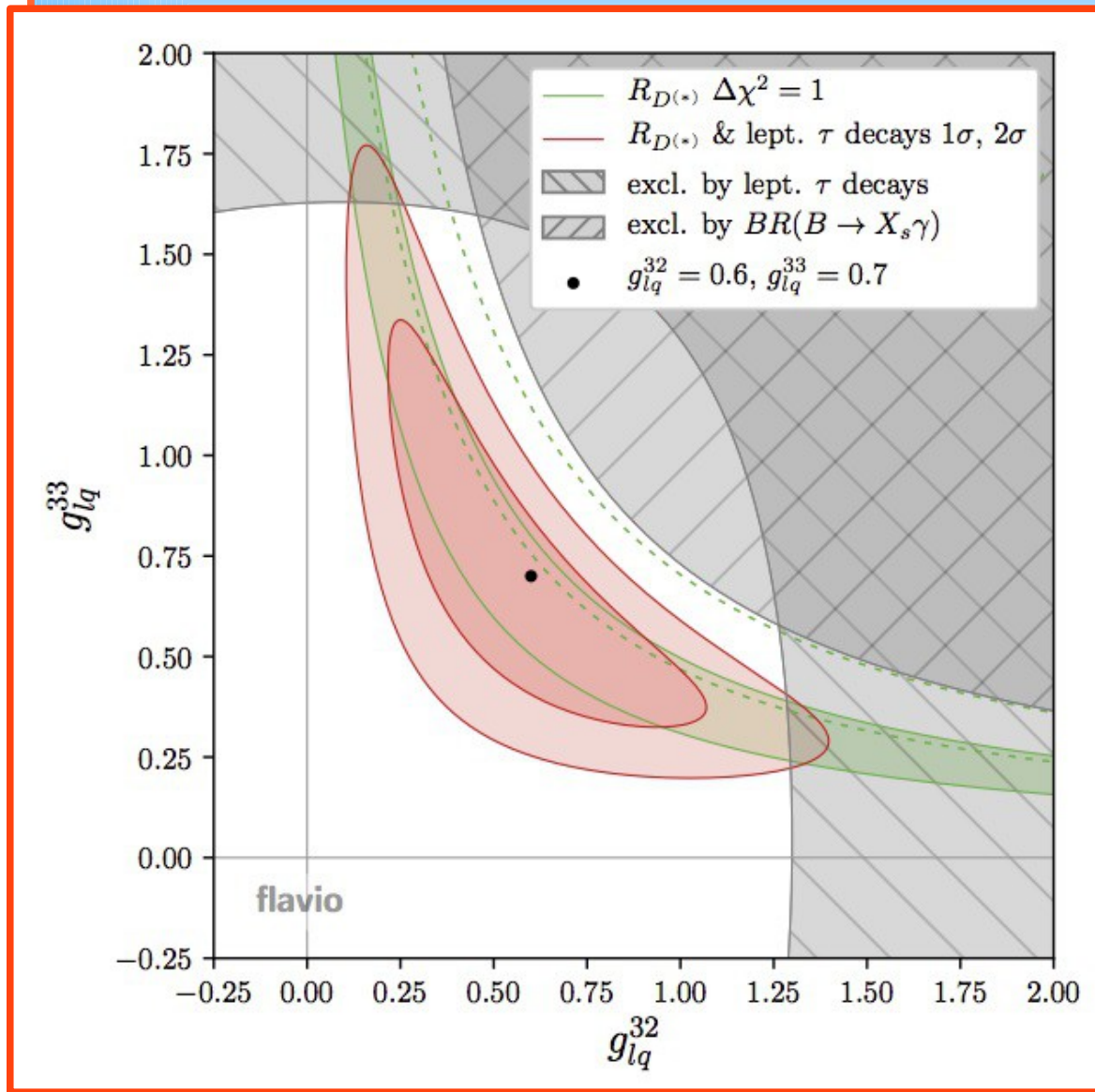


these couplings also famously constrained by

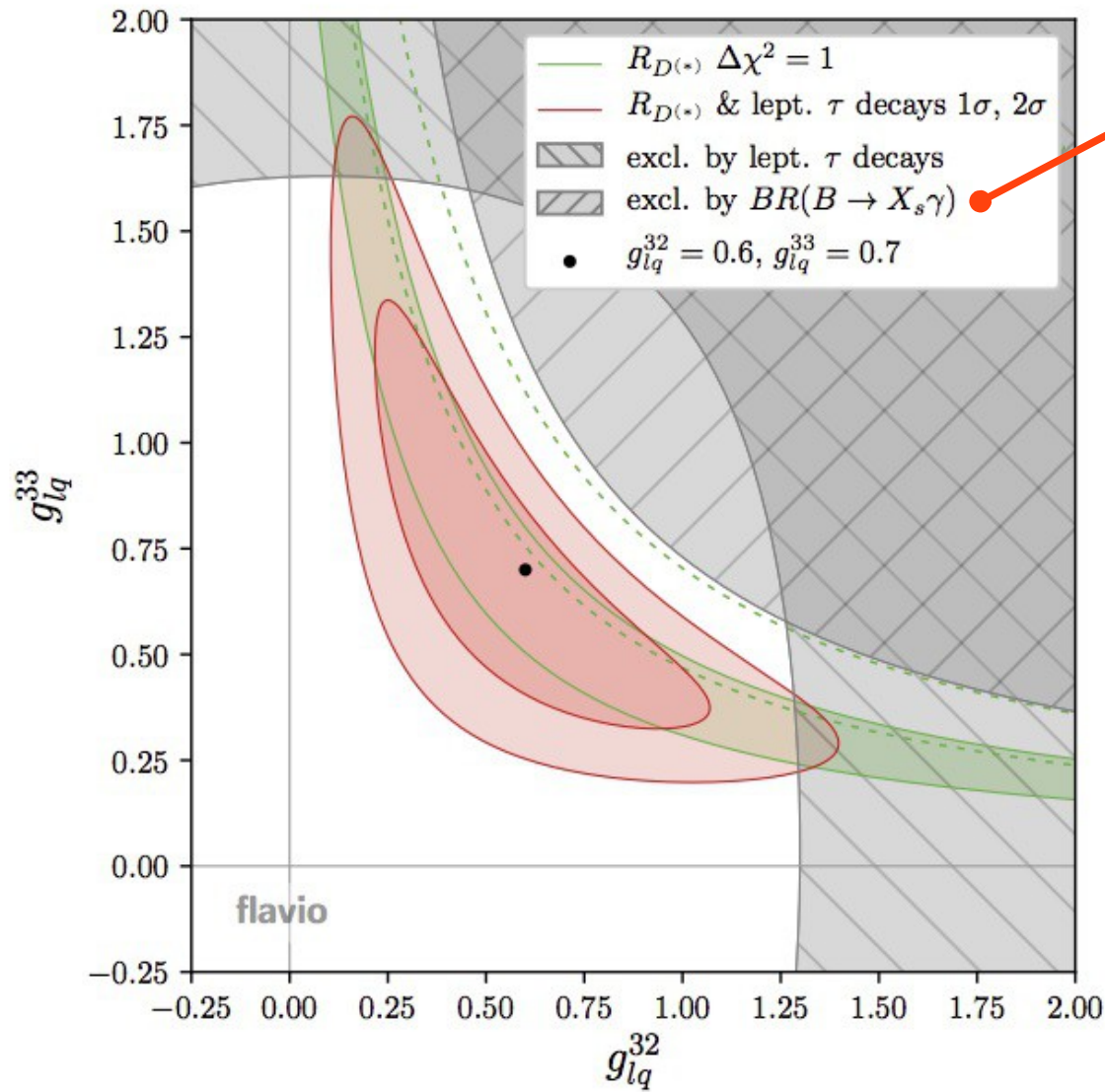
$$\tau \rightarrow \ell \nu \nu \quad [\text{Feruglio-Paradisi-Pattori}]$$

(hence far from obvious that an $R_{D(*)}$ description achievable)

U₁ LQ: g_{lq}^{32} vs. g_{lq}^{33}



U₁ LQ: g_{lq}^{32} vs. g_{lq}^{33}

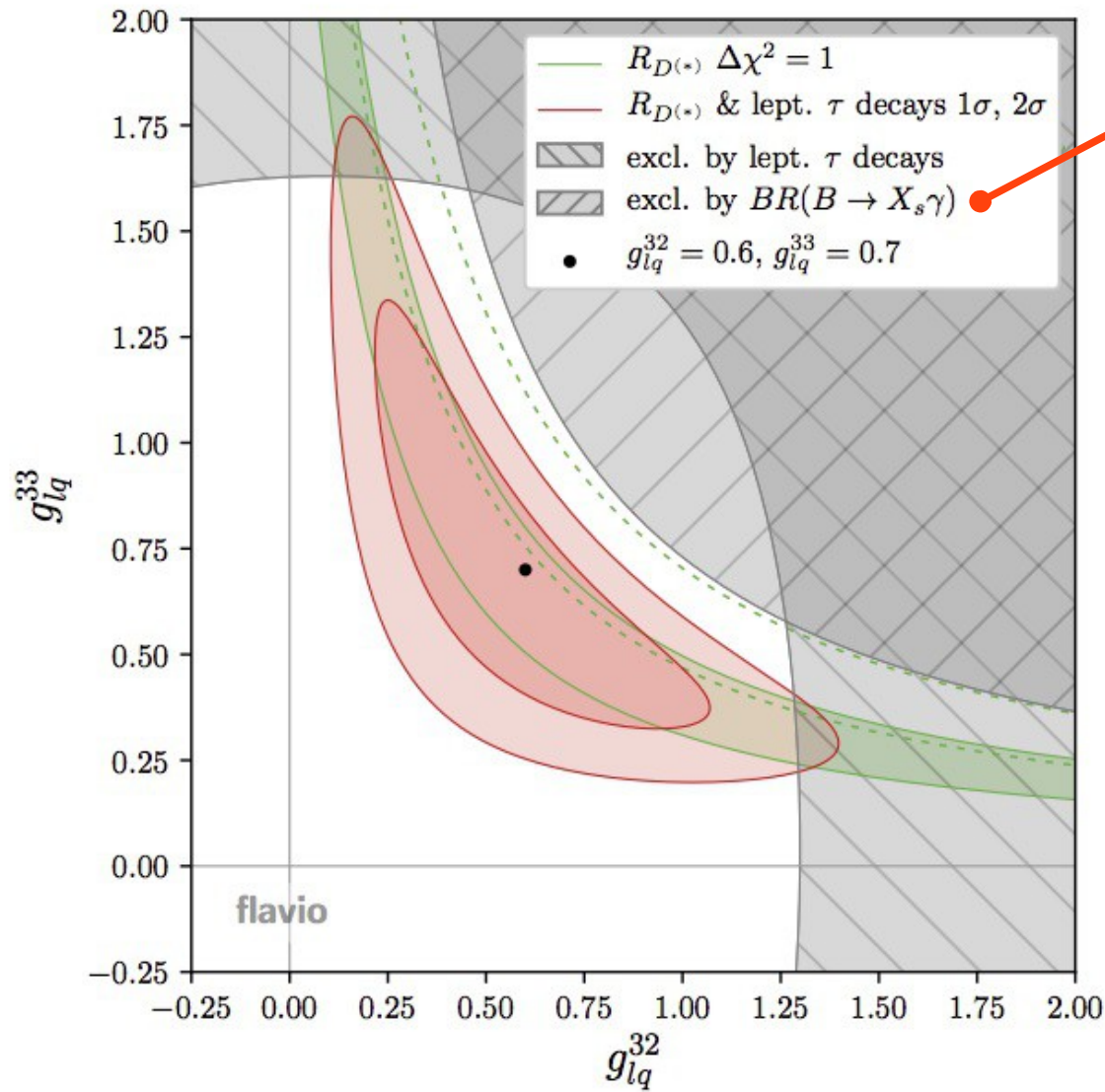


Model-dependent constraint

See discussion in

[Cornella-Fuentes-Isidori, 2019;
Calibbi-Crivellin-Li, 2018;
Bordone *et al.*, 2018]

U₁ LQ: g_{lq}^{32} vs. g_{lq}^{33}

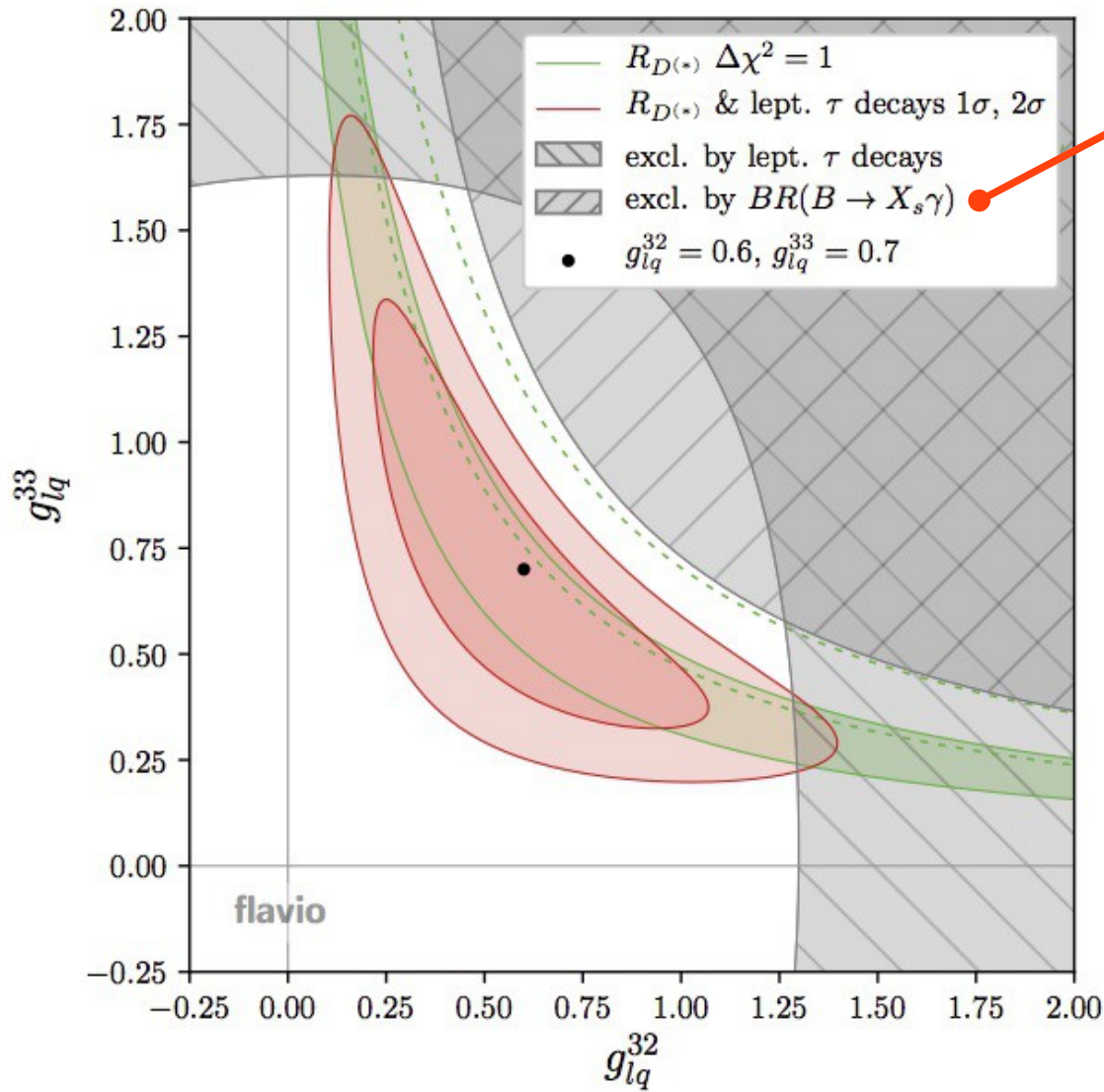


Model-dependent constraint

See discussion in
 [Cornella-Fuentes-Isidori, 2019;
 Calibbi-Crivellin-Li, 2018;
 Bordone et al., 2018]

- $R_{D^{(*)}}$ and $\tau \rightarrow \ell \nu \nu$ select a non-trivial region

U₁ LQ: g_{lq}^{32} vs. g_{lq}^{33}

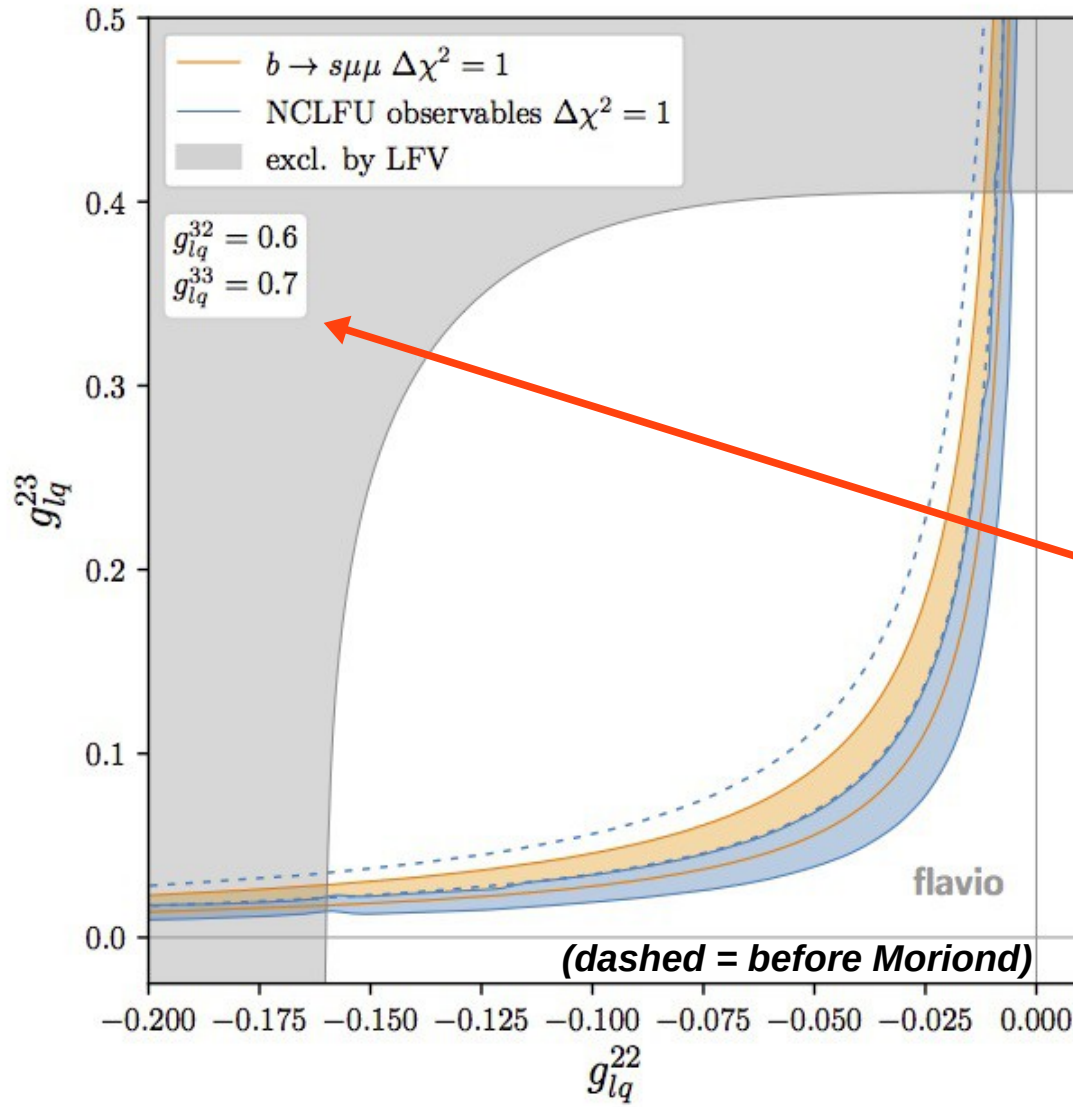


Model-dependent constraint

See discussion in
 [Cornella-Fuentes-Isidori, 2019;
 Calibbi-Crivellin-Li, 2018;
 Bordone et al., 2018]

- $R_{D^{(*)}}$ and $\tau \rightarrow \ell \nu \nu$ select a non-trivial region
- We pick a benchmark point, then constrain the other two couplings

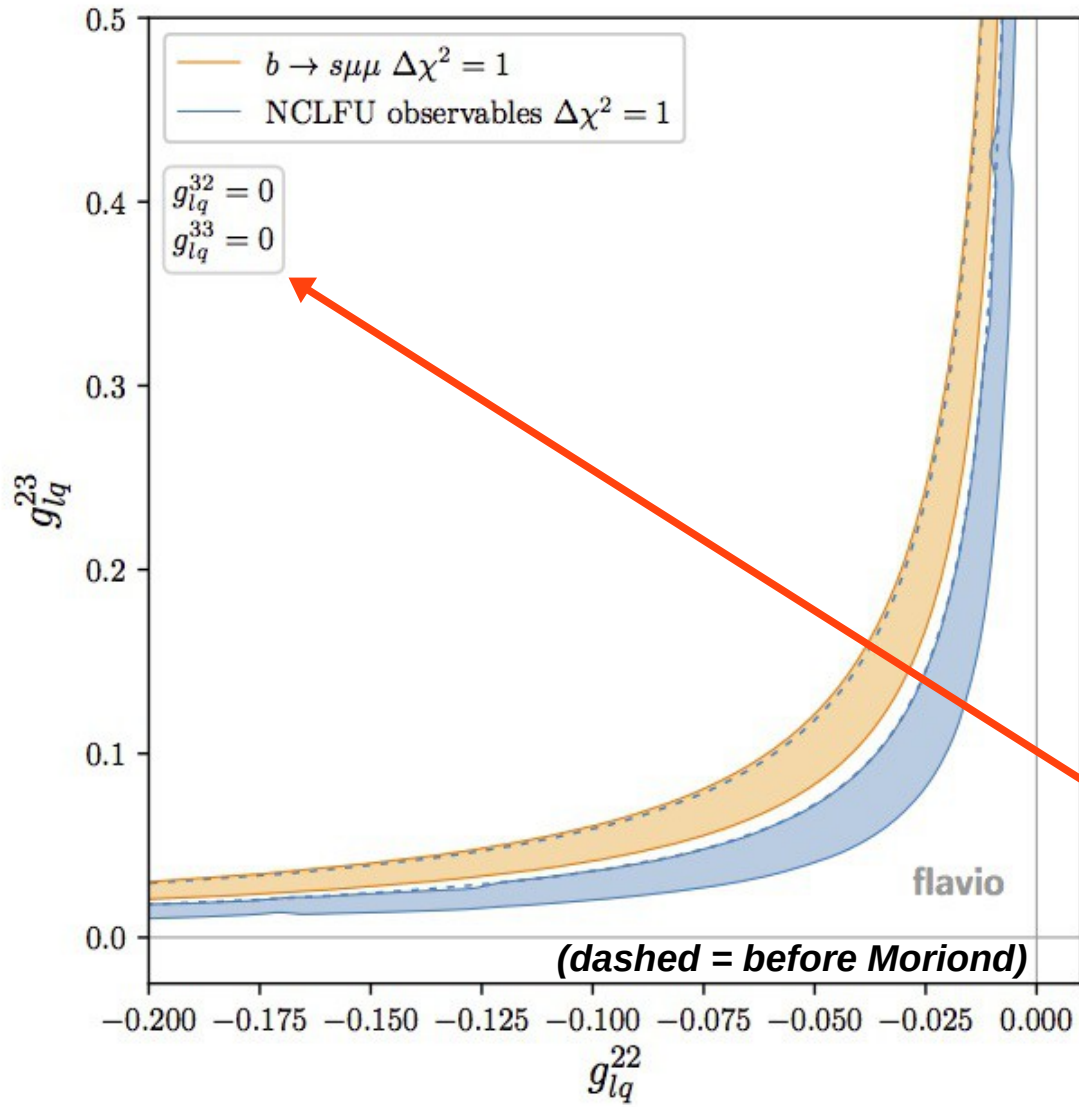
U_1 LQ: g_{lq}^{22} vs. g_{lq}^{23}



The plane of muonic couplings shows that the picture is fully consistent

- The $R_{K^{(*)}}$ and $b \rightarrow s \mu \mu$ regions now nicely overlap
- Tauonic couplings fixed to the benchmark value explaining $R_{D^{(*)}}$

U_1 LQ: g_{lq}^{22} vs. g_{lq}^{23}



The plane of muonic couplings shows that the picture is fully consistent

- The $R_{K^{(*)}}$ and $b \rightarrow s\mu\mu$ regions now nicely overlap
- Tauonic couplings fixed to the benchmark value explaining $R_{D^{(*)}}$
- If they were set to zero, the picture would work much less

Conclusions

- *Semi-lept. B-decay data still display preference for new effects in 4-f ops. w/ LH quarks*

Conclusions

- *Semi-lept. B-decay data still display preference for new effects in 4-f ops. w/ LH quarks*
- *Solution with muonic $C_9 = -C_{10}$ now favoured over pure C_9*
- *Even better description obtained with additional $C_9^{univ.}$
Allows to connect $b \rightarrow s$ with $b \rightarrow c$ discrepancies*

Conclusions

- *Semi-lept. B-decay data still display preference for new effects in 4-f ops. w/ LH quarks*
- *Solution with muonic $C_9 = -C_{10}$ now favoured over pure C_9*
- *Even better description obtained with additional $C_9^{univ.}$
Allows to connect $b \rightarrow s$ with $b \rightarrow c$ discrepancies*
- *One gets a coherent picture all the way from the WET, to the SMEFT, to simplified models*

Conclusions

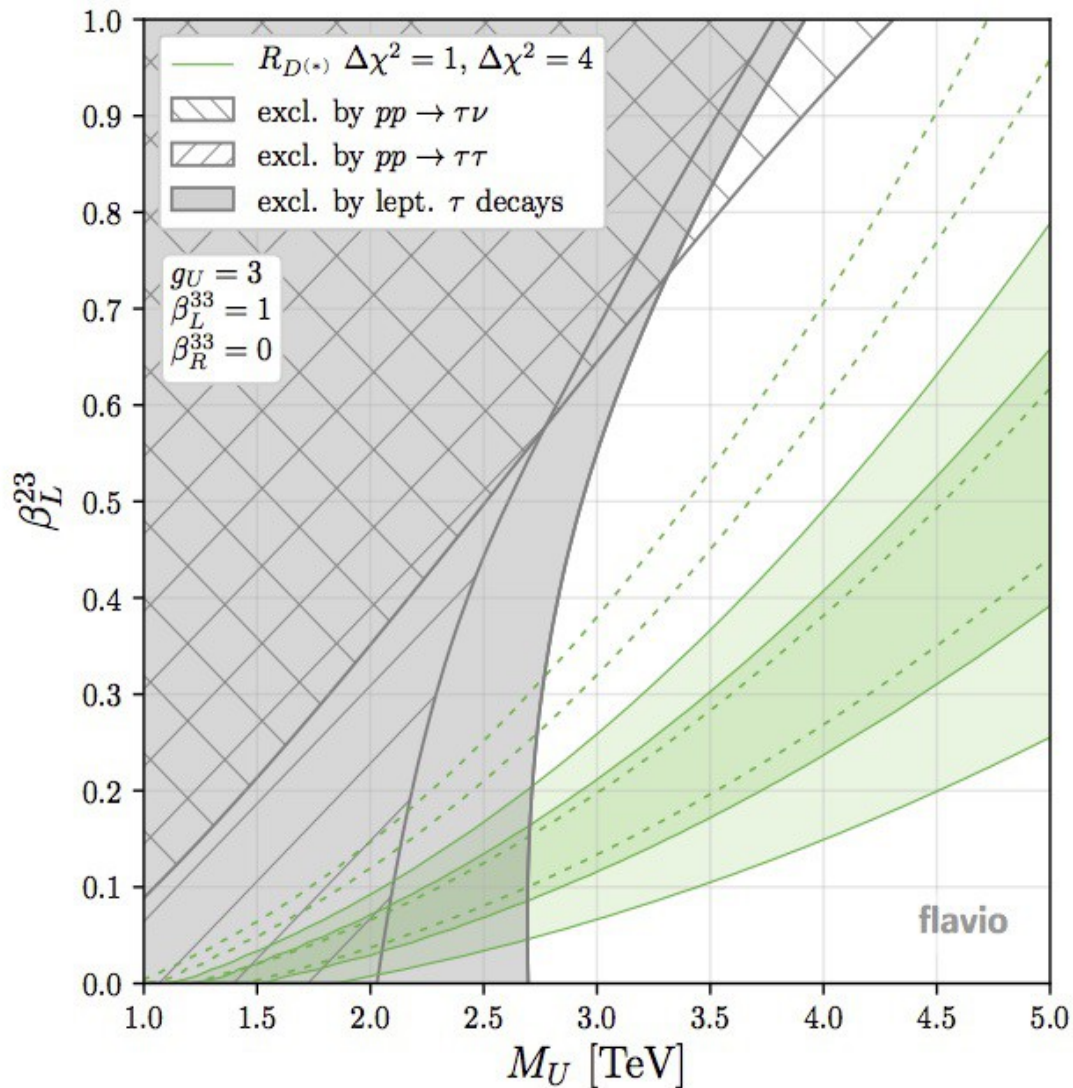
- *Semi-lept. B-decay data still display preference for new effects in 4-f ops. w/ LH quarks*
- *Solution with muonic $C_9 = -C_{10}$ now favoured over pure C_9*
- *Even better description obtained with additional $C_9^{univ.}$
Allows to connect $b \rightarrow s$ with $b \rightarrow c$ discrepancies*
- *One gets a coherent picture all the way from the WET, to the SMEFT, to simplified models*
- *Time for more tests
More LUV observables will clarify the situation soon*

Spare

U_1 LQ: direct constraints

Aren't such tauonic couplings also constrained by direct searches?

E.g. $pp \rightarrow \tau\tau$ or $\tau\nu$



For the sake of comparison, we normalized coupling values to those used in

[Baker-Fuentes-Isidori-König]

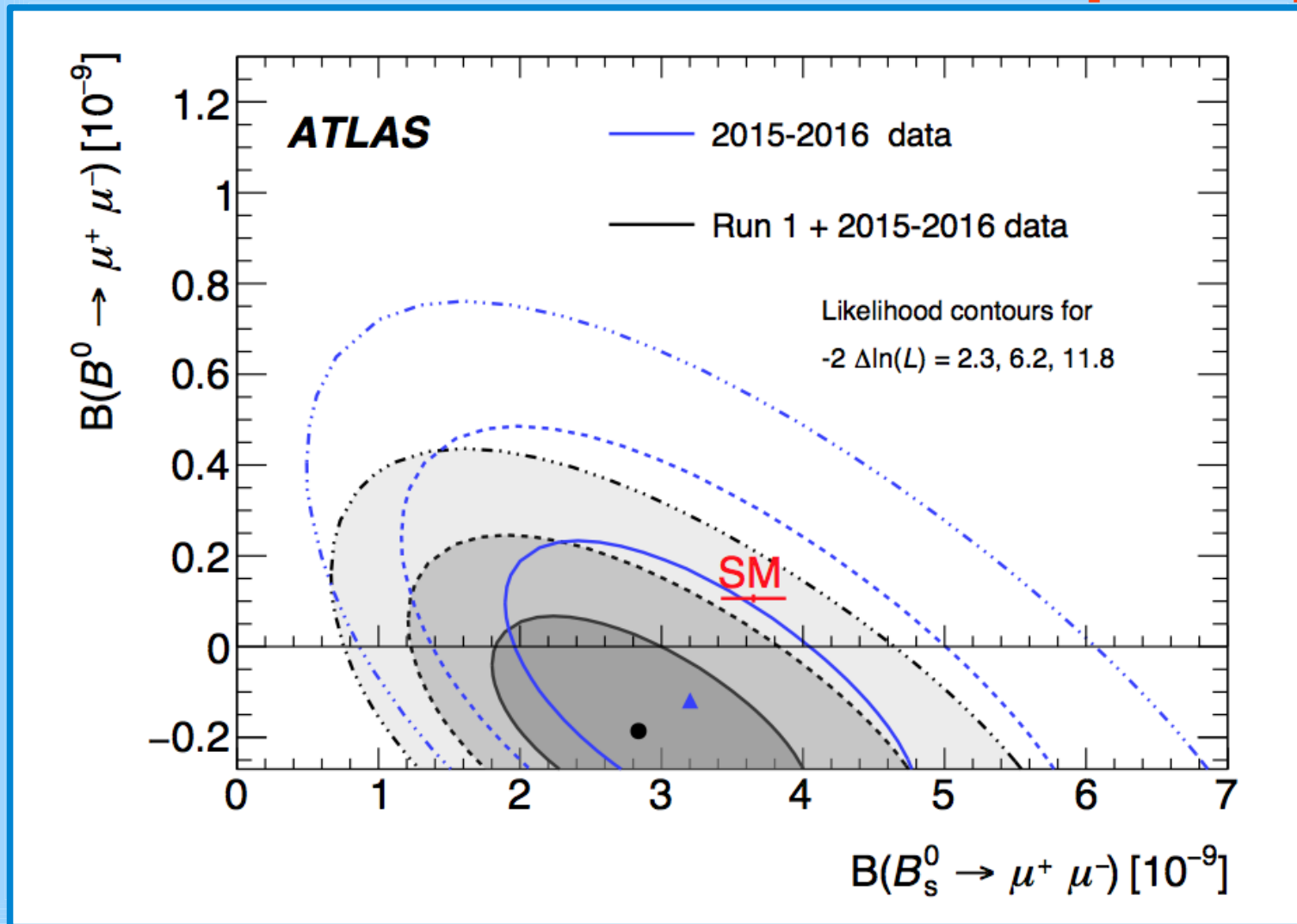
With $g_{lq}^{3i} \neq 0$ (as required by $R_{D^{(*)}}$)
LUV constraints are stronger than direct ones

Note also the large g_U value used here (as said, for comparison).

All constraints scale down with lower g_U values.

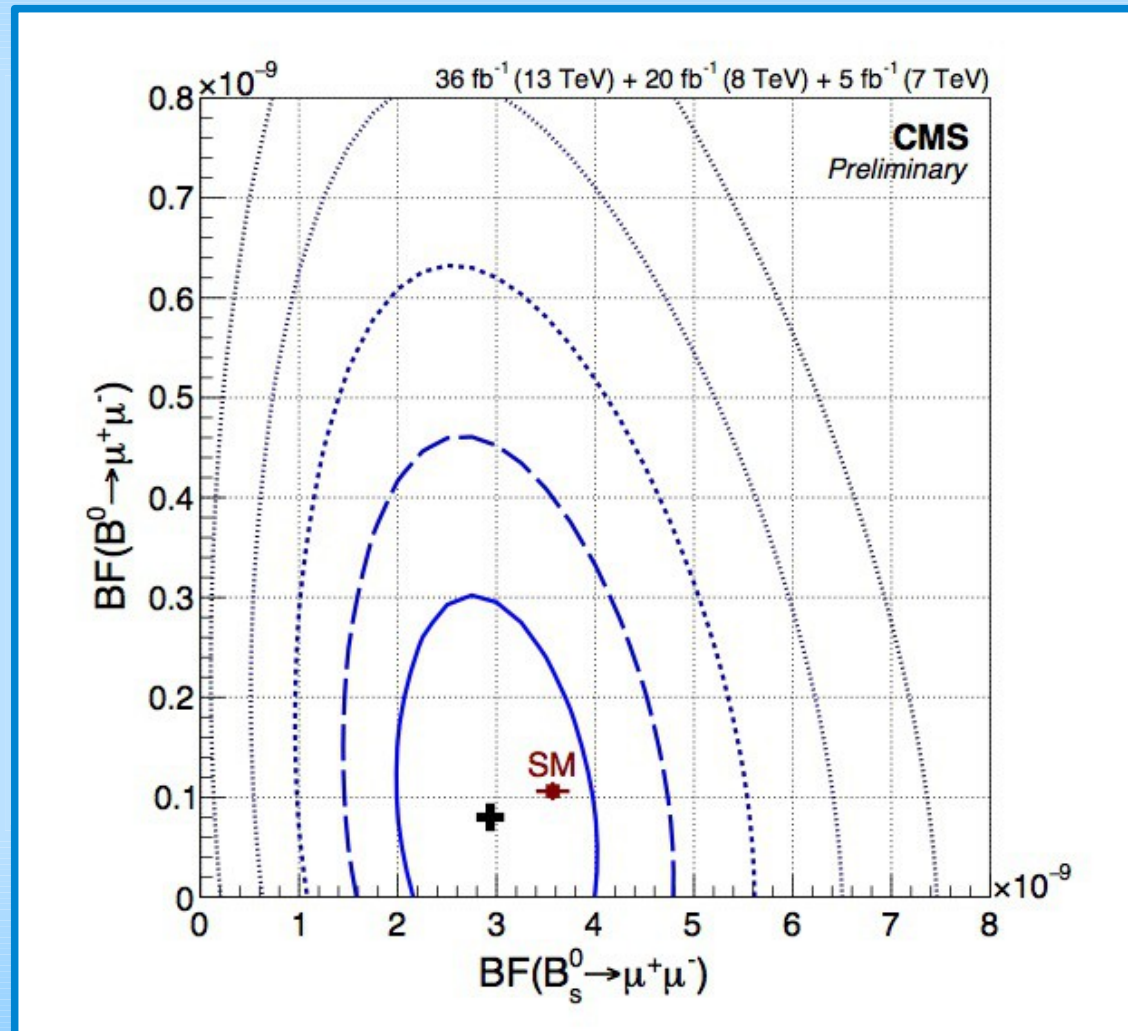
$B_s \rightarrow \mu\mu$ update

[1812.03017]



New update from CMS

$$BR(B_s \rightarrow \mu\mu) = (2.9_{-0.6}^{+0.7} \pm 0.2) \times 10^{-9}$$



CMS PAS BPH-16-004 (2019/08/04)